

### Measuring Distances to the Galactic Supernova Remnants Using Red Clump Stars Reporter : Susu Shan

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SN 1987A, 30 years later Réunion Island, 22February 2017 E-mail: <u>shansusu@nao.cas.cn</u>

### Outline

# BACKGROUND Method Results

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### BACKGROUND

Reliable distance determination to Galactic Supernova Remnants (SNRs) is important to constrain its physical parameters like the age, the size and the explosion energy.

➢Only about half of Galactic SNRs have distance measurements

Measure the distances of Galactic SNRs by taking advantage of the extinction (Av)-distance (d) relation.

## Red clump stars(RCs) as a standard candle



 An obvious concentration region in the color-magnitude diagrams.
 The absolute magnitude of RCs varies fairly little.
 D(pc)=10<sup>[(m-M-A(D)+5)/5]</sup>



Build the extinction (Av)-distance(d) relation in each direction of Av-known SNRs

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### Method

We take SNR G40.5-0.5 as an example
➢ Data: 2MASS All-Sky Point Sources Catalog
➢ Centering SNR G40.5-0.5, extract stars in the 0.5 deg<sup>2</sup> area around the SNR



Color-magnitude diagram for the stars within 0.5 deg<sup>2</sup> of G40.5-0.5, the red dot and lines show the fitted location of the RC peak and its extent with  $1\sigma$ .



Overlap the extinction of the SNR on the extinctiondistance relation we build by RCs.

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### Result

- ➢Obtain distances of 22 SNRs and upper/lower limits for 34 SNRs in the first quarter of Galaxy.
- Among them, 3 SNRs' distances are given first time.
- Compared with previous distance-known SNRs, we find that most of our distances are consistent with previous results.
- ➢ In the future, we will extend our work to the whole Galaxy. We can use the distances to re-calibrate the Sigma-diameter relation and estimates SNRs' other distance-dependent parameters.

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Part of the result			
Sourse Name	A <sub>V</sub> (mag)	D <sub>known</sub> kpc	D <sub>(ourwork)</sub> kpc
G54.1+0.3	7.7±0.5	6.2	6.4
G63.7+1.1	$16\pm5$	3.8-6	> 4
G65.8-0.5	$2.4 \pm 0.4$	-	2.5
G66.0+0.0	$2.0\pm0.2$	-	2-3
G67.7+0.9	$1.9 \pm 0.2$	-	2
G82.2+5.3	$1.9 \pm 1.1$	2.0	1.8

### Thanks for your attention!

Sec. 5





红团簇星是小质量恒星演化到早期氦核燃烧阶段。