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Discovery of Titanium-K Lines in the Northeastern Jet of Cas A

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Outline

Introduction

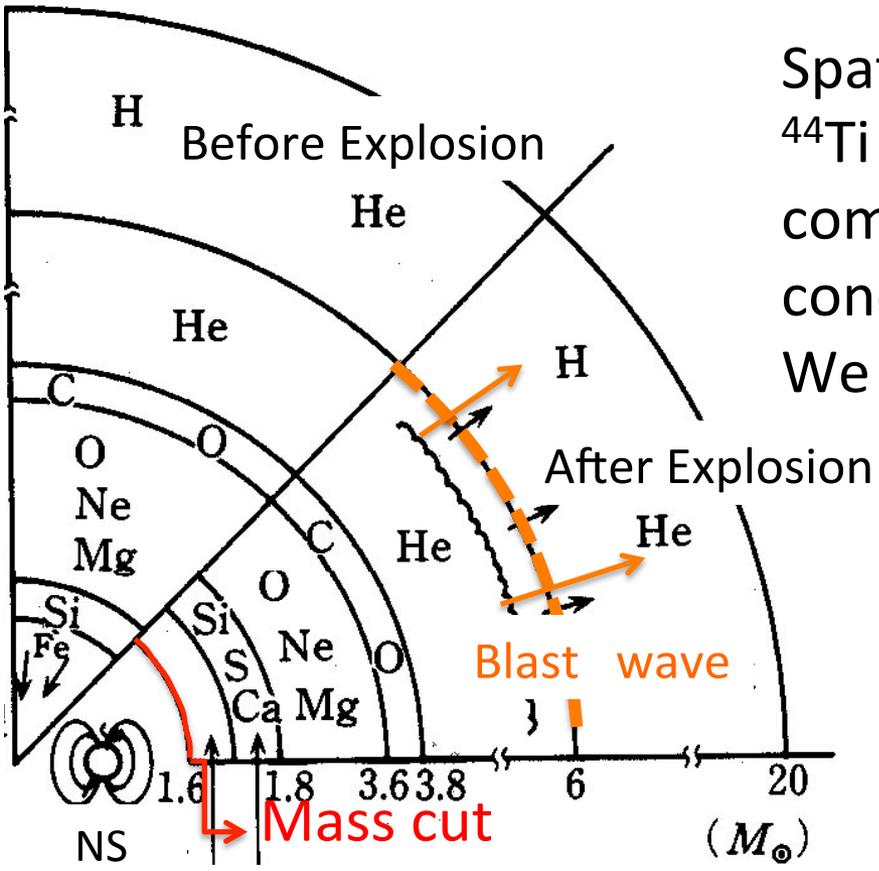
- ^{44}Ti in explosive nucleosynthesis
- Cassiopeia A: Previous results

Cas A jet with Chandra

- Spectral and image analysis
- Estimation of Ti mass in jet
- Comparison to nucleosynthesis model

Summary

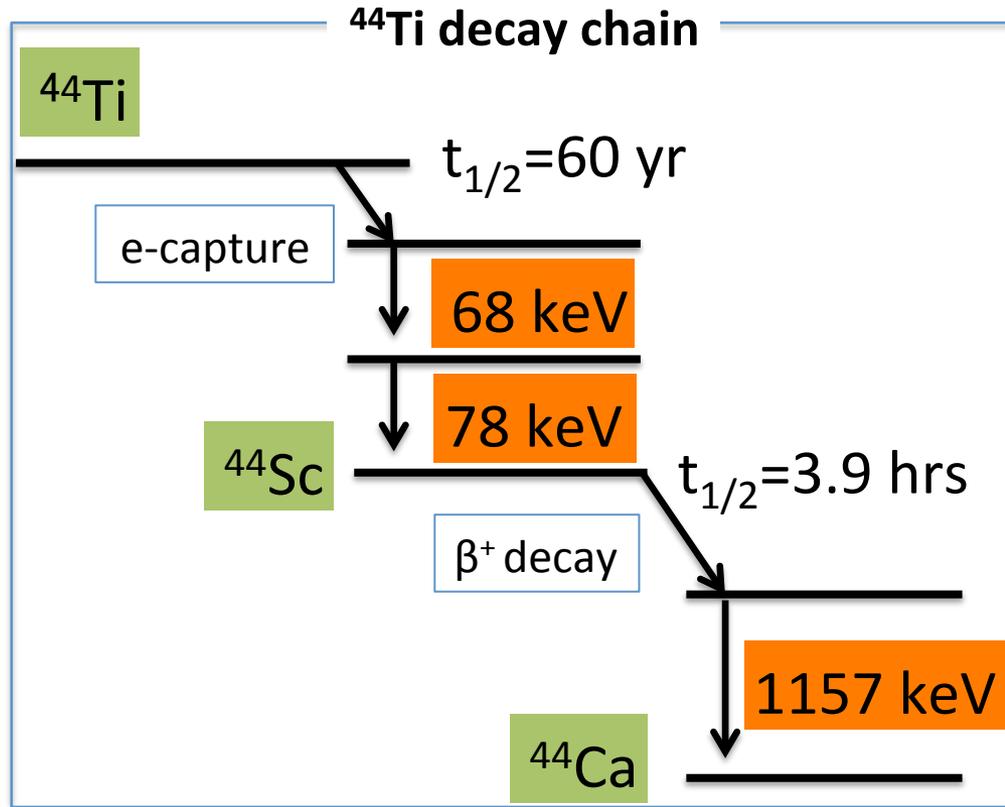
^{44}Ti in Explosive Nucleosynthesis



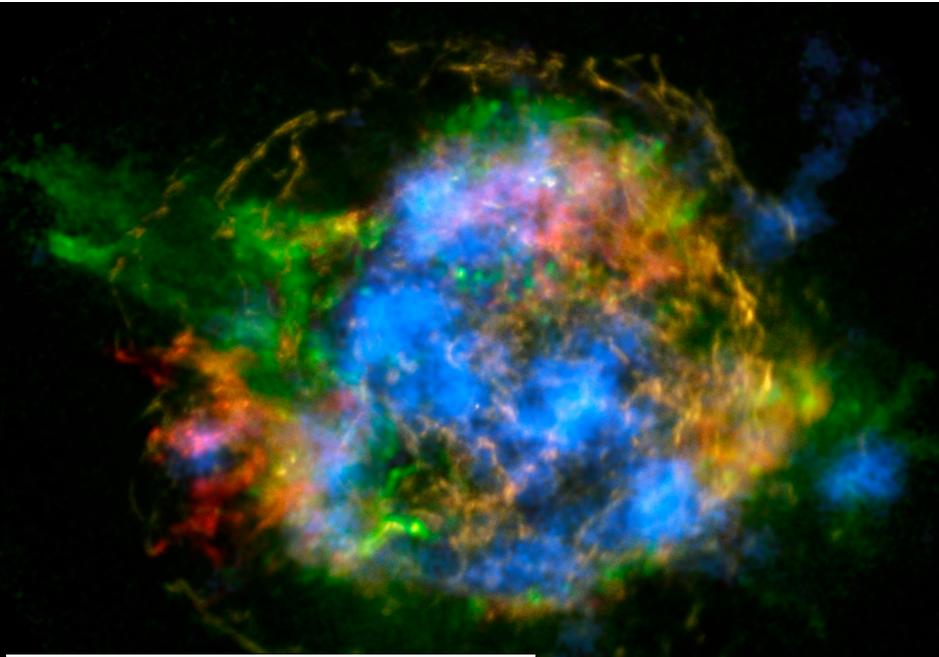
Spatial distribution and the amount of ^{44}Ti produced at mass cut (between compact object and ejecta) reflect the condition of explosion. We can learn about explosion mechanism.

Inner composition at before and after SNe
[Kumagai, Nomoto, 1991]

^{44}Ti decays into ^{44}Sc by e-capture (half-life 60yr) and ^{44}Sc decays into ^{44}Ca by beta decay (half-life 3.9hrs)



X-ray Observation of Cassiopeia A



Chandra (Fe, Si/Mg)
NuSTAR (^{44}Ti (68 keV))

[NASA/JPL-Caltech/
CXC/SAO]

- Age: ~ 330 yr [Fesen 2006]
- Distance: 3.4 kpc [Reed+ 1995]
- Progenitor mass: $\sim 16 M_{\odot}$ [Lee+ 2014]
- Initially synthesized ^{44}Ti mass:
 $1.37 \pm 0.19 \times 10^{-4} M_{\odot}$ [Siegert+ 2015]

Cassiopeia A (Cas A)

Detections of 1157 keV, 68 keV, 78 keV lines (^{44}Ti decay).

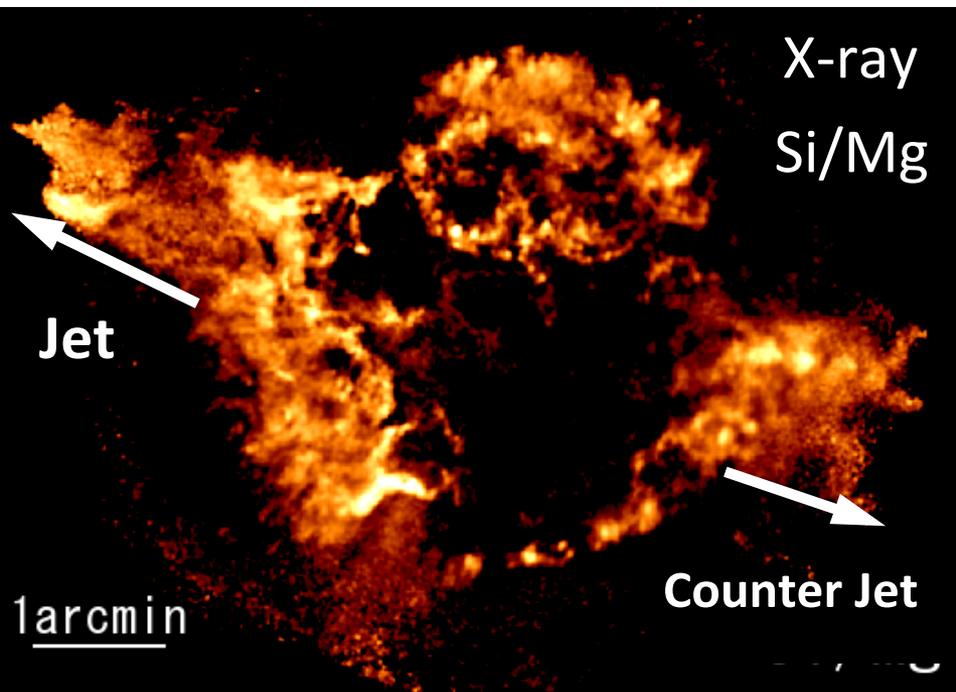
^{44}Ti 's spatial distribution differs from that of Fe K lines measured with the Chandra X-ray Observatory.

No any Ti-K line detections so far.

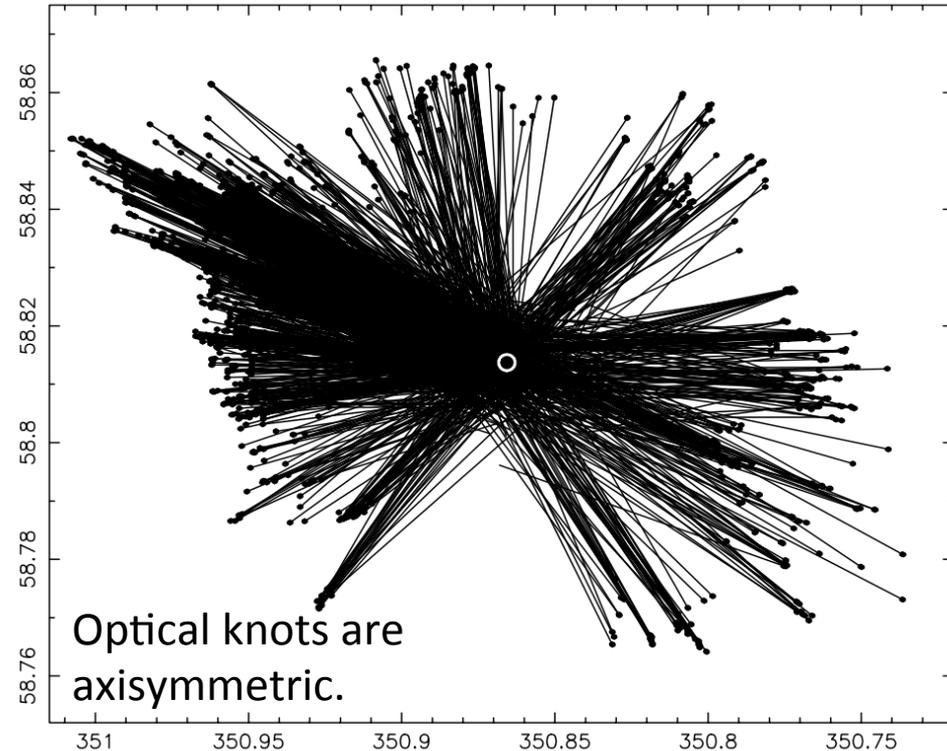
Stable ^{48}Ti is mainly produced decay from ^{48}Cr . ^{48}Cr mass can be enhanced by “alpha-rich freeze out”, which produces ^{44}Ti .

Modest amount of ^{44}Ti and Fe in northeastern jet.

Jet in Cassiopeia A



Cas A has a **Si-rich** jet extending northeastward and counter jet extending southwestward



The speed of Jet: [Fesen+ 2006]
10000~14000 km/s

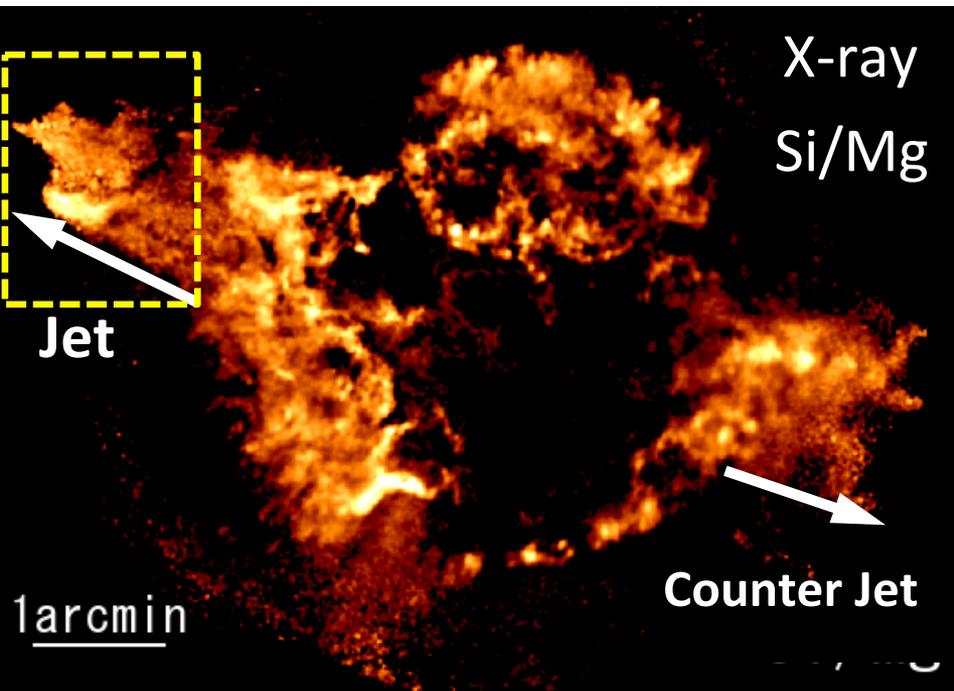
Non-Jet ejecta speed: [Rutherford+ 2013]
1000~3000 km/s

Jet's kinetic energy: [Fesen+ 2015]
 $\sim 1 \times 10^{50}$ erg

Jet area has $\sim 10\%$ flux compared with other ejecta area.

“Jet” origin remain unsettled whether axisymmetric circumstellar interaction [Blondin, 1996] or bipolar explosion?

Jet in Cassiopeia A



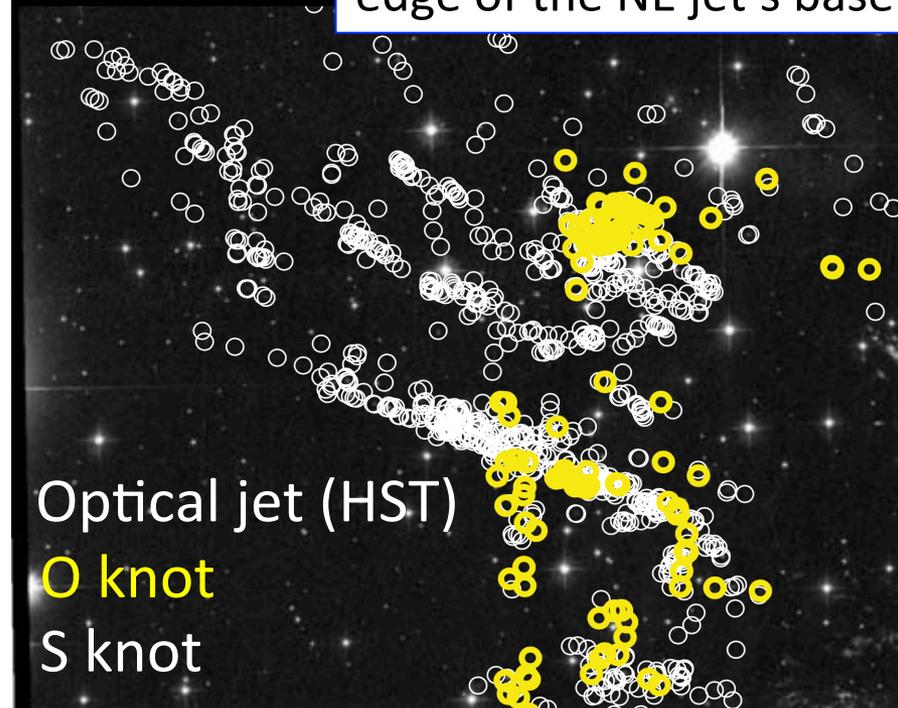
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[Hammell and
Fesen 2008]



**S jet penetrated
surrounding O-layer?**

Synchrotron component is
overlaid in counter jet direction.
Focus only NE jet for analysis.

Outline

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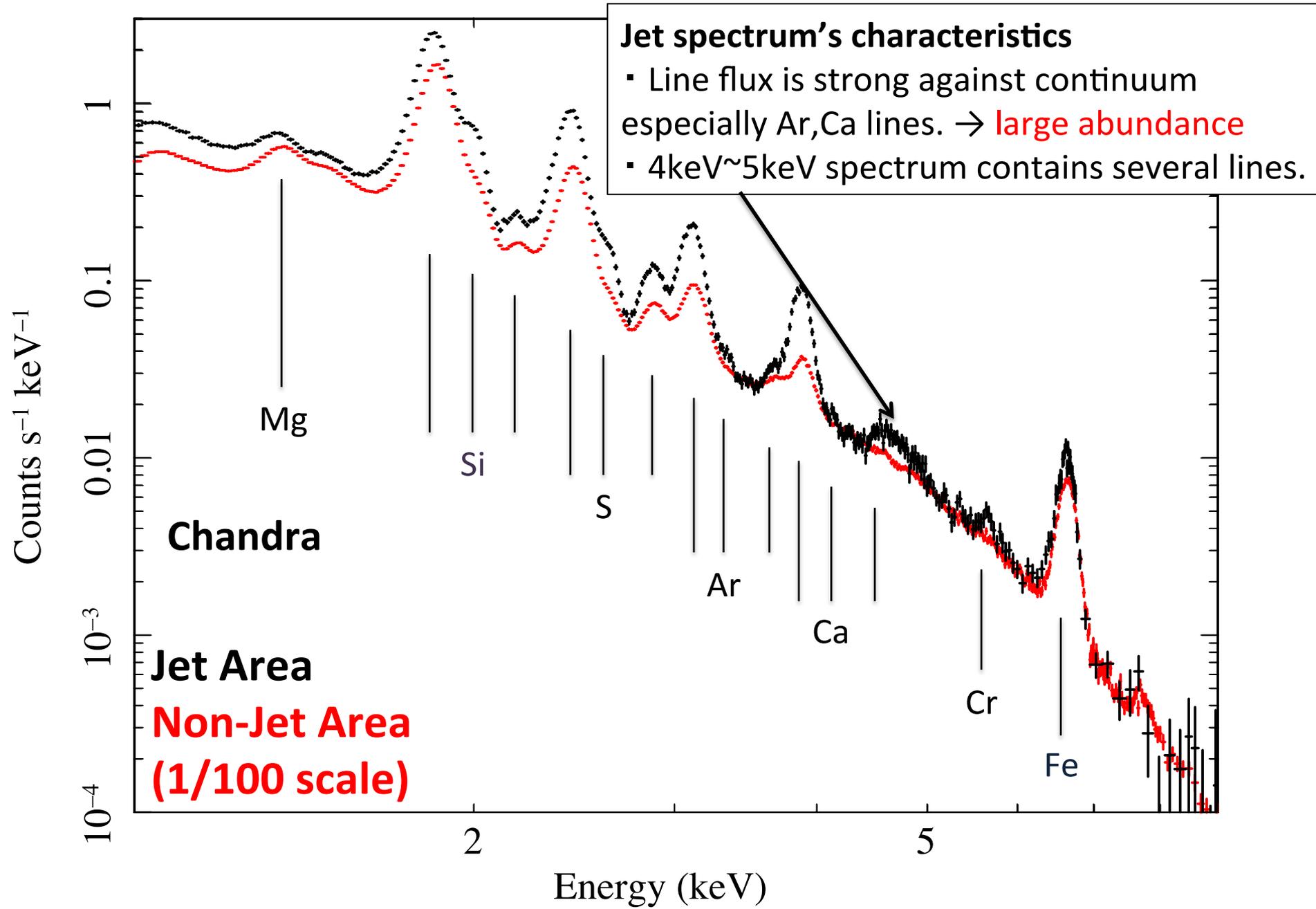
- ^{44}Ti in explosive nucleosynthesis
- Cassiopeia A: Previous results

Cas A jet with Chandra

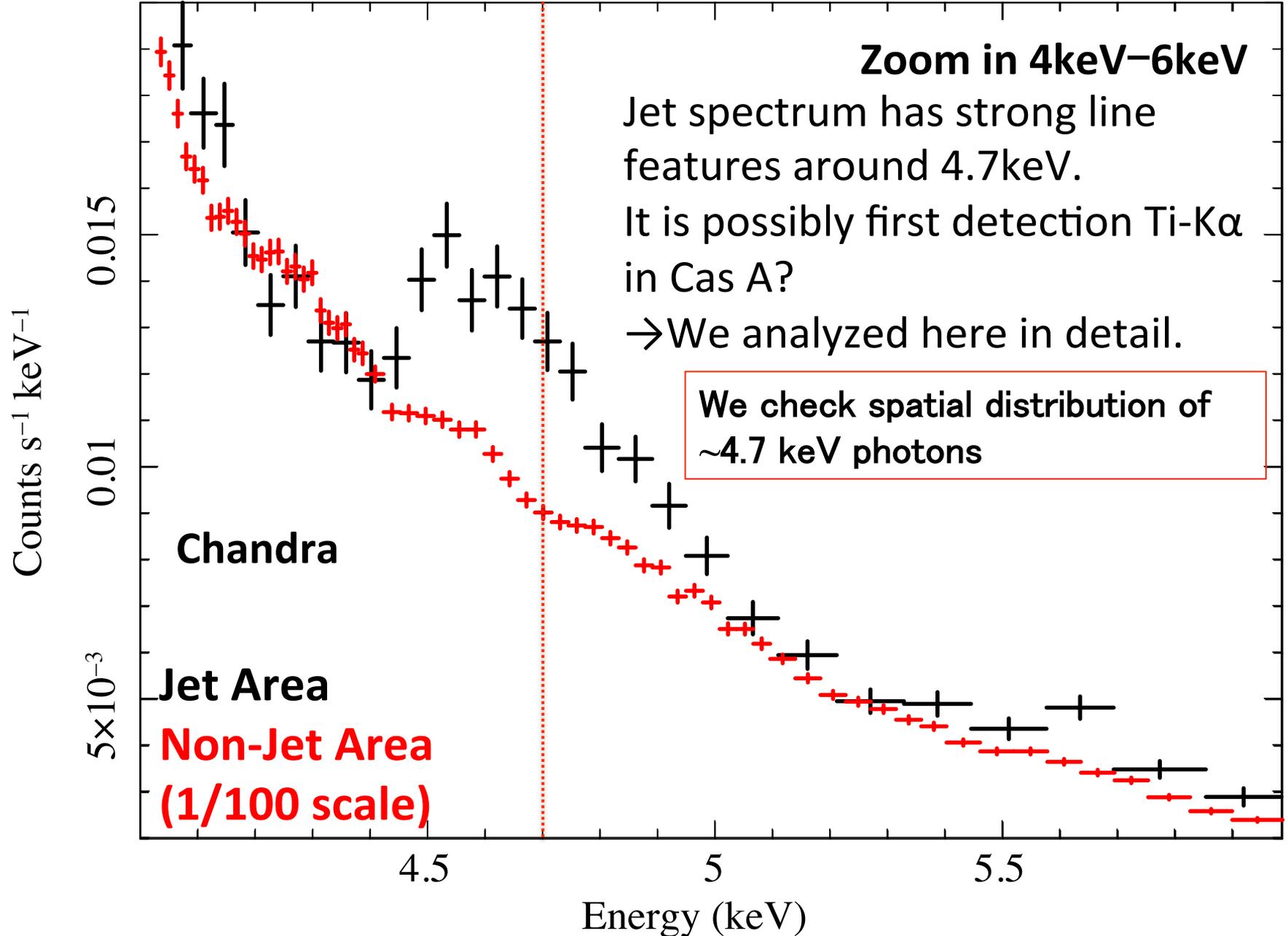
- Spectral and image analysis for Ti line
- Estimation of Ti mass in jet
- Comparison to nucleosynthesis model

Summary

Comparison jet to non-jet area spectrum



Comparison jet to non-jet area spectrum

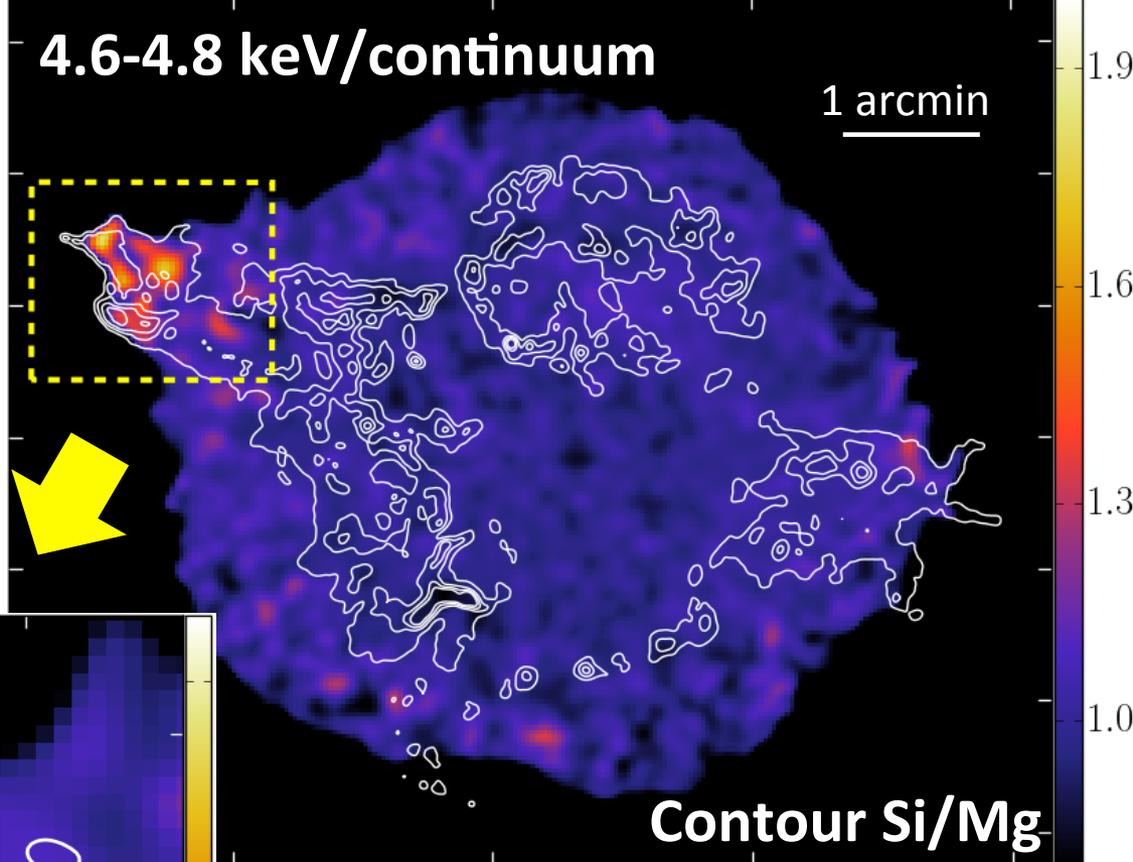


Spatial distribution:

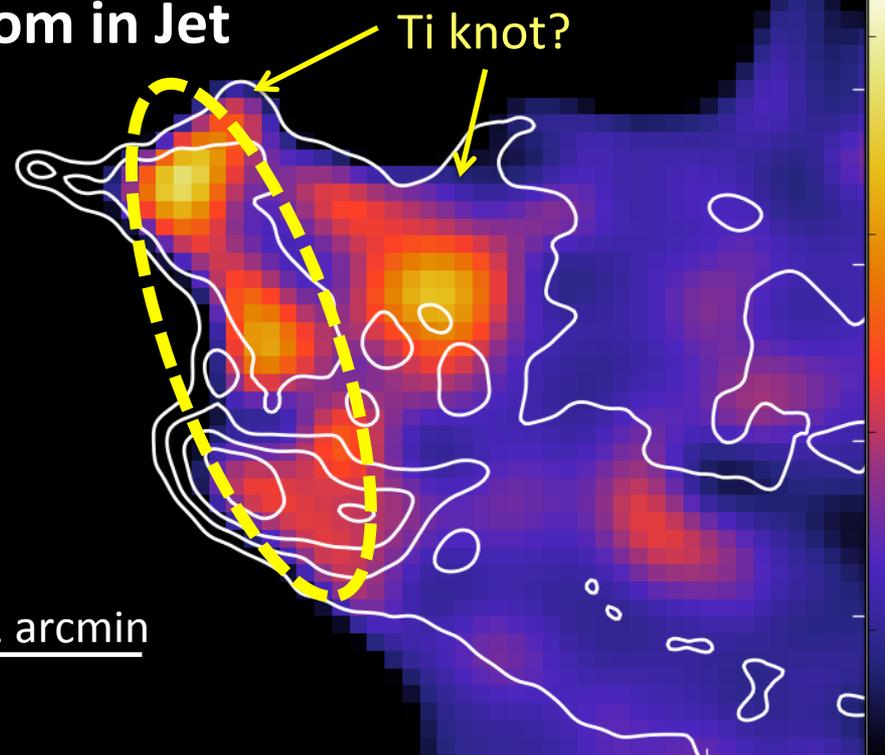
~4.7 keV line

- ~4.7 keV line map shows
- Detected only tip of Jet
- different distribution from Si, Fe
- Knot-like distribution

4.6-4.8 keV/continuum



Zoom in Jet



- “knots” are seen in jet
→ Spectral analysis for sub-regions.
- Ti-K α is blended with Ca line
→ We chose the elliptical region (yellow) where Ca line is weak and Ti line is strong.

Spectral Analysis 1: Ti line

The feature in 4.4-5.0 keV:
a peak at 4.75 keV

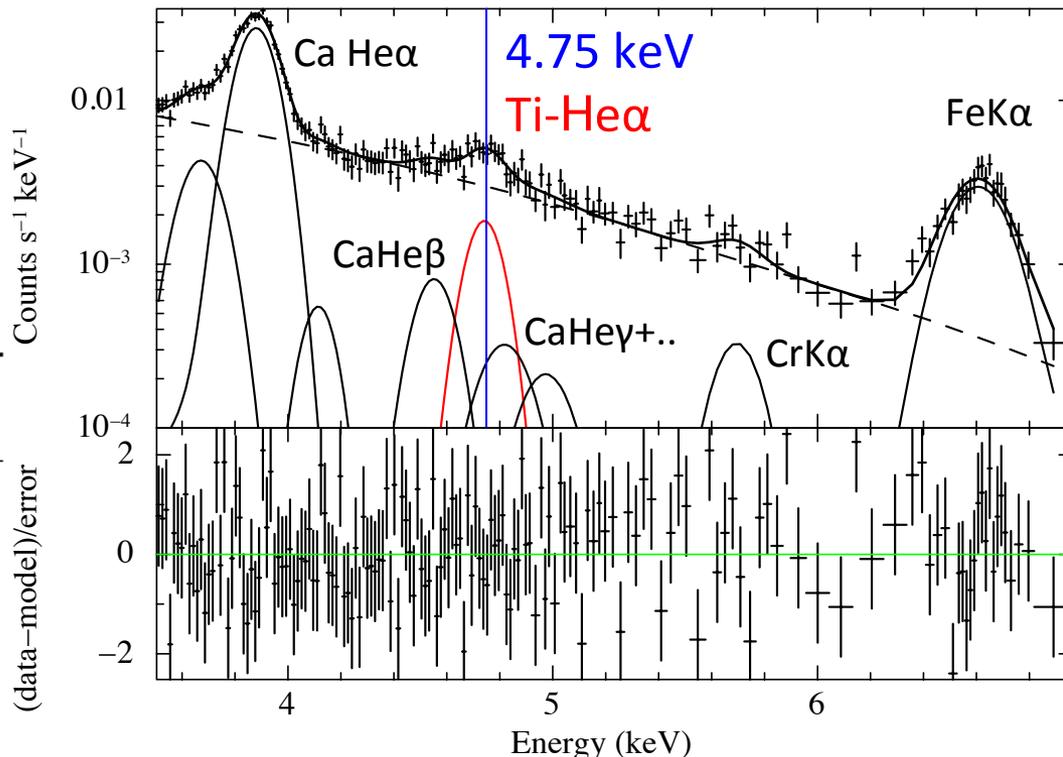
→ This cannot be explained
by broadening of Ca lines

Ion	E_{line} (keV)	transition	Emissivity ($\text{ph cm}^3 \text{s}^{-1}$)
Ca XIX	4.584	n:3→1 l:1→0	5.874e-19
Ti XXI	4.708	n:2→1 s:1→0	4.280e-20
Ti XXI	4.756	n:2→1 l:1→0	1.184e-19
Ca XIX	4.822	n:4→1 l:1→0	1.875e-19
Ca XX	4.862	n:3→1 l:1→0	8.401e-20
Ca XX	4.864	n:3→1 l:1→0	1.670e-19

The range of 4.5-4.8 keV line list ($kT_e = 3 \text{ keV}$) [ATOMDB]

Plasma with electron temperature
of 3 keV without Ti lines cannot
explain this component.

**Detected Ti-K line is
statistically significant.**



Jet spectrum model: BREMSS + gauss

Add a gaussian model as Ti K-line

$$\chi^2 / \text{d.o.f} = 273.6 / 228 \rightarrow 233.9 / 226$$

$$\Delta\chi^2 \sim 40 \rightarrow 5.6 \sigma$$

$$\text{Line energy } 4.75_{-0.04}^{+0.02} \text{ keV}$$

→ Ti Heα: 4.71-4.76 keV is consistent.

Discussion 1: Ti's isotope

The analysis by using Non-equilibrium ionization model
(vvpshock model in XSPEC)

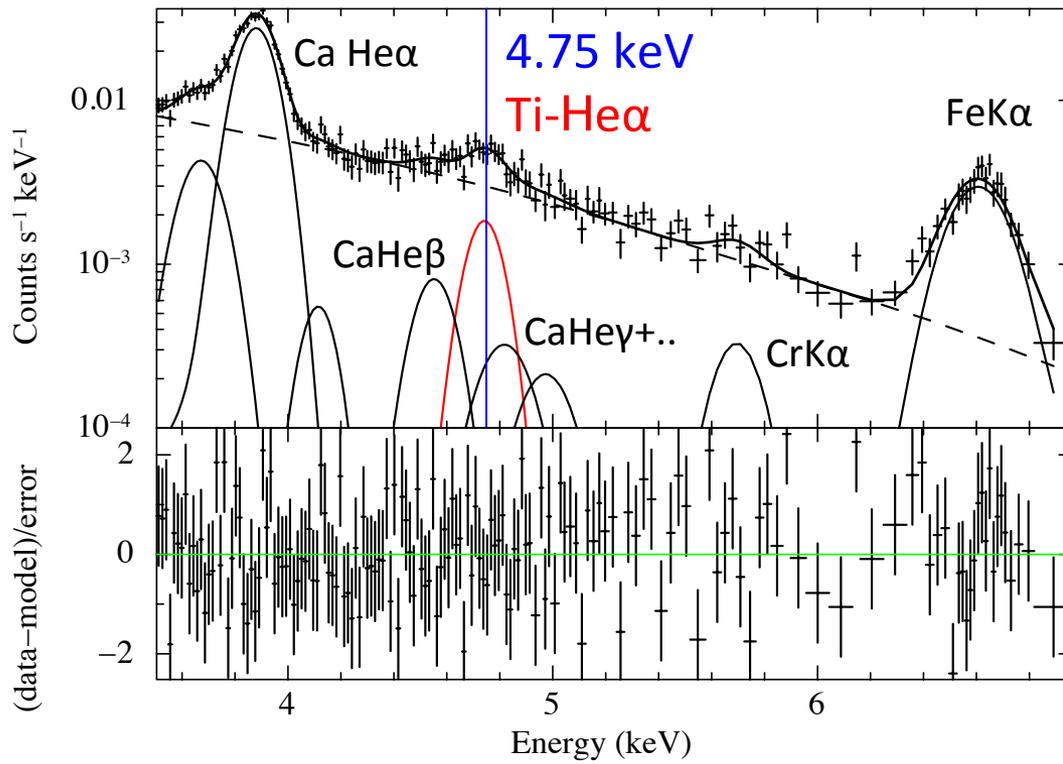
The amount of Ti in jet:
 $0.6-3 \times 10^{-5} M_{\odot}$

If this Ti is ^{44}Ti ,
68/78 keV line flux from jet is
 $>4 \times 10^{-5} \text{ photons/s/cm}^2$
at present.

→ larger than observed with
NuSTAR.

(**$1.5 \pm 0.3 \times 10^{-5} \text{ ph/cm}^2/\text{s}$**
for Cas A total)[Grefenstette+, 2014]

Ti in jet traced by
Ti-K is not ^{44}Ti .



Jet spectrum model: BREMSS + gauss

Stable Ti mass in nucleosynthesis model
[Thielemann et al. 1995]

$^{46}\text{Ti}: 3 \times 10^{-6} M_{\odot}$, $^{47}\text{Ti}: 5 \times 10^{-6} M_{\odot}$
 $^{48}\text{Ti}: 1 \times 10^{-4} M_{\odot}$, $^{49}\text{Ti}: 4 \times 10^{-6} M_{\odot}$

Detection of ^{48}Ti K α line

Discussion 2: Comparison to ^{48}Ti mass ^{48}Ti mass $\sim 1 \times 10^{-5} M_{\odot}$

Assuming that Ti mass in jet has isotropic distribution in Cas A

- ^{48}Ti mass in jet: $\sim 1 \times 10^{-5} M_{\odot}$
- Solid angle Ω of jet: ~ 0.1 str

→ **Total ^{48}Ti mass: $\sim 1 \times 10^{-3} M_{\odot}$**

Nucleosynthesis results [Thielemann+ 1995]

→ **Synthesized ^{48}Ti : $\sim 1 \times 10^{-4} M_{\odot}$**

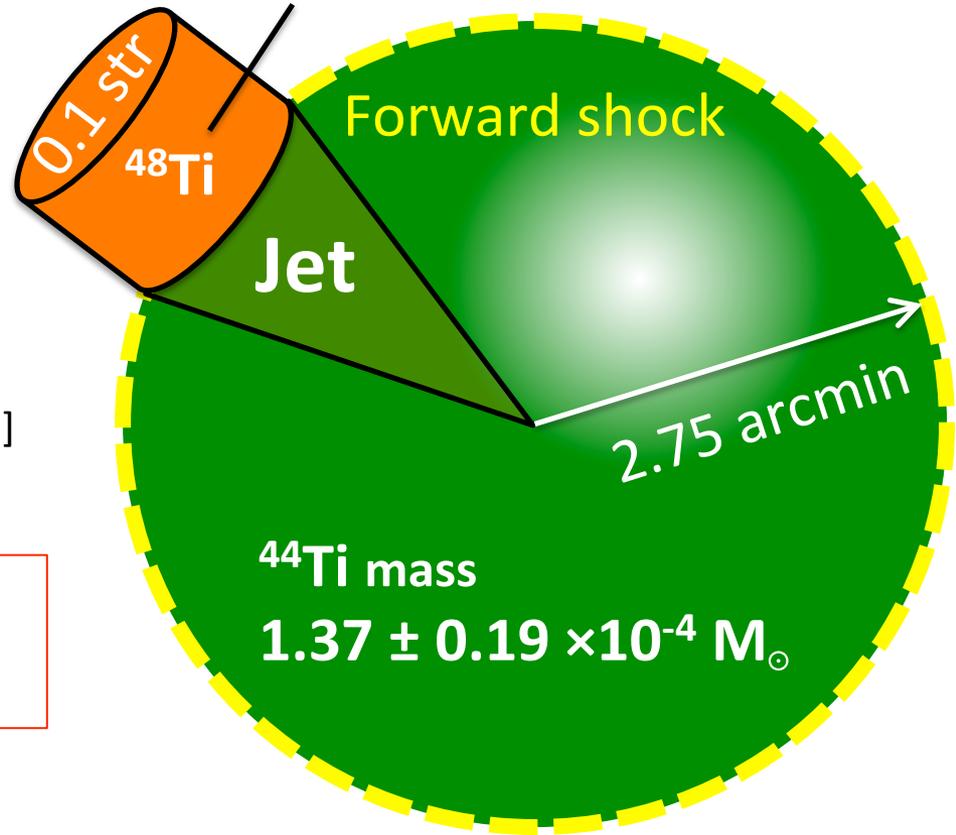
^{48}Ti mass in jet is enhanced in comparison with other regions

Synthesized ^{44}Ti is also larger than spherical model.

Model: $0.2 - 2 \times 10^{-4} M_{\odot}$

Observation: $1.37 \pm 0.19 \times 10^{-4} M_{\odot}$

→ The effect of aspherical explosion?



Cas A's configuration

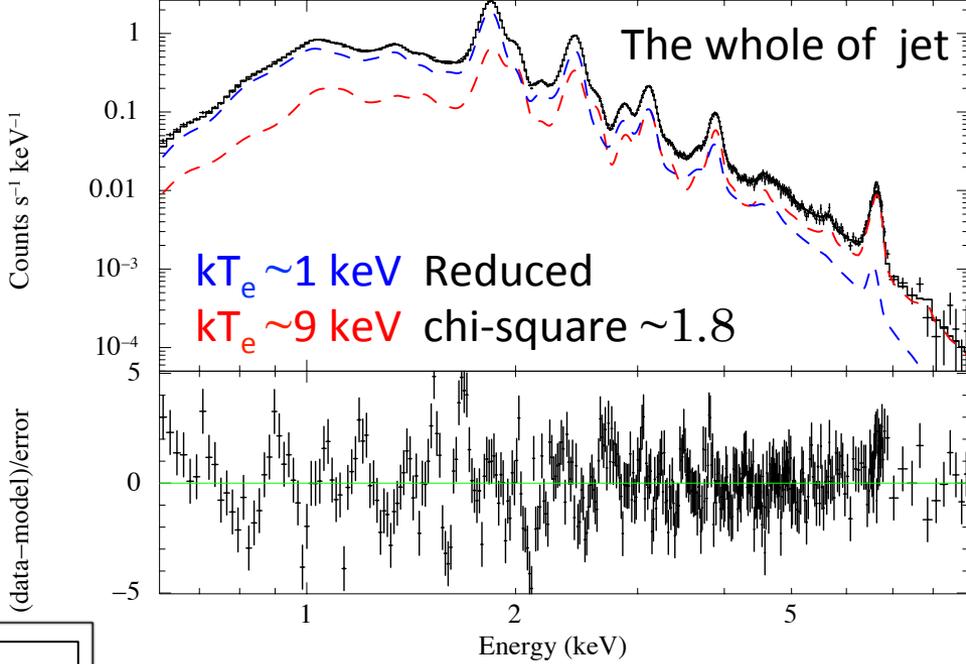
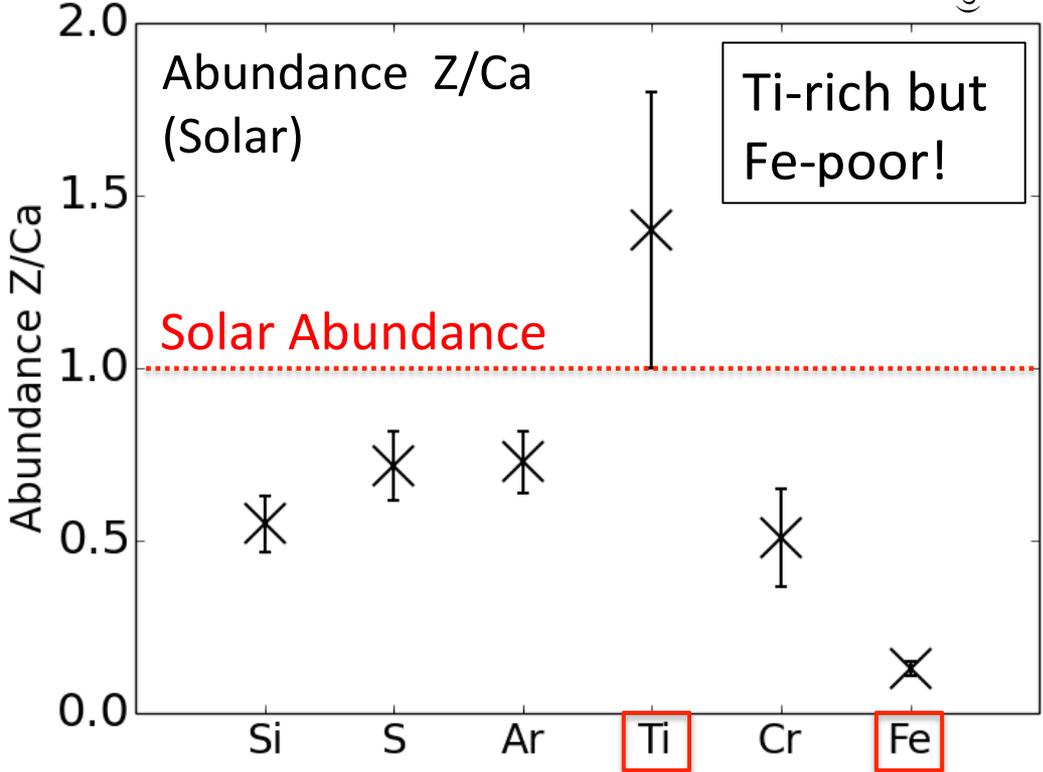
How about abundances of other elements?

Spectral Analysis 2:

Jet Abundance

Using two component Non-equilibrium ionization model, and estimate abundance pattern in jet.

Plot the elements abundance to Ca. (the unit is solar abundance.)



Jet Spectrum with result of fitting by using NEI model (example)

Jet composed of incomplete Si-burning “ash”?

Incomplete Si-burning ($5 \times 10^9 \text{ K} > T > 4 \times 10^9 \text{ K}$)
→ Rich intermediate mass element. (Ar, Ca, Ti...)

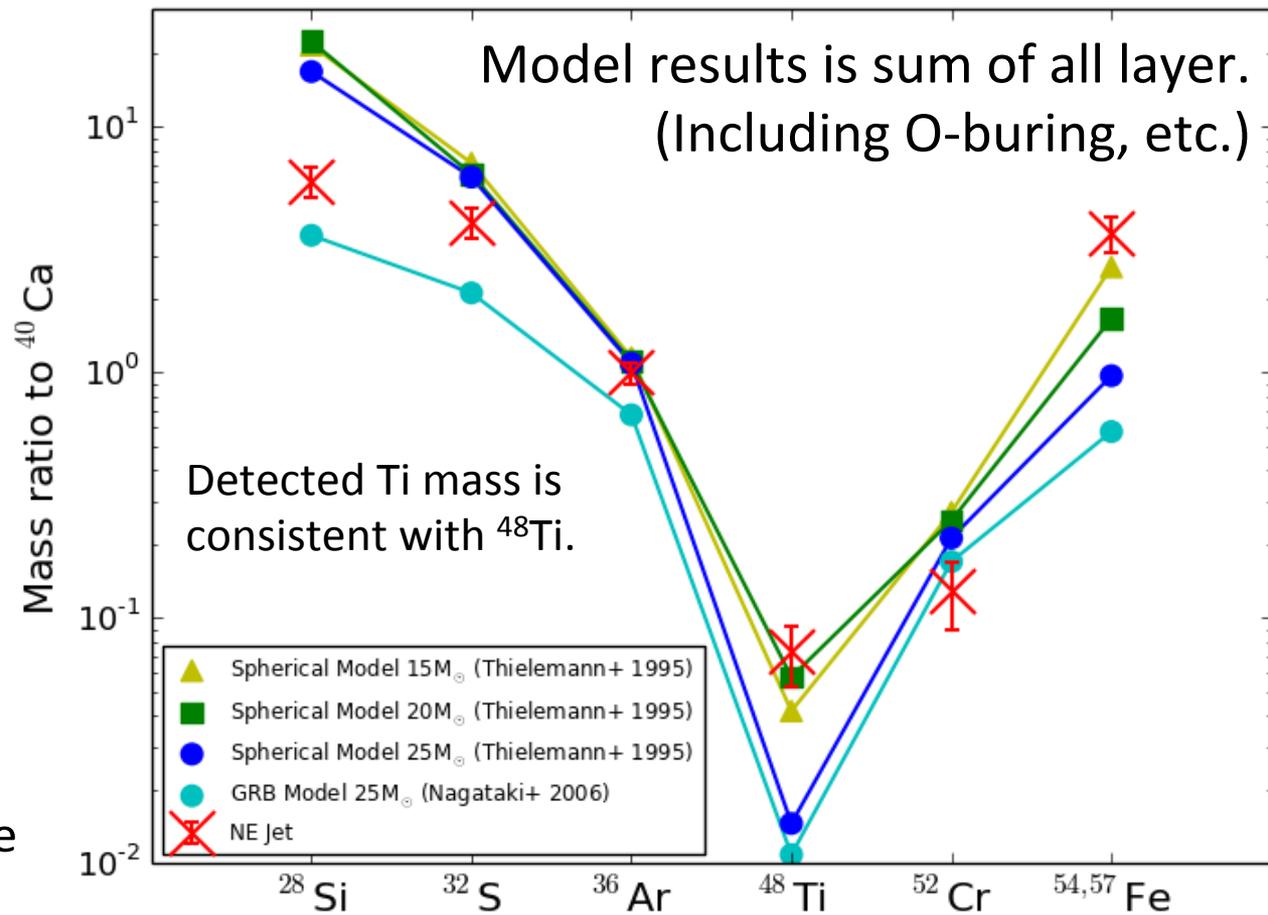
Discussion 3: Jet Abundance

We chose isotope which is mainly synthesized by incomplete Si-burning for comparison:

^{28}Si , ^{32}S , ^{36}Ar , ^{40}Ca , ^{48}Ti ,
 ^{52}Cr , ^{54}Fe , ^{57}Fe
(including daughter nuclei)

We used 3 spherical nucleosynthesis model whose difference is progenitor mass and GRB model.

Calculate mass ratio to ^{40}Ca .



Comparison between observation and model.

Jet abundance is consistent with incomplete Si-burning elements (spherical model is used for now).

SUMMARY

- We have discovered **Ti-K α feature** at 4.7 keV in the NE jet of Cassiopeia A.
- Estimated Ti mass is **$0.6 \sim 3 \times 10^{-5} M_{\odot}$** in Cas A jet.
- If this titanium is ^{44}Ti , it is inconsistent with 68 keV observation. We conclude that Ti-K α is mainly produced by **stable Ti**.
- Ti in jet traced by Ti-K is ^{48}Ti from the viewpoint of nucleosynthesis.
- The analysis by using Non-equilibrium ionization model lead jet abundance, and it suggests jet is composed of **incomplete Si-burning elements**.