

The SoLid experiment

Short baseline Oscillation search with Lithium-6 Detector

SoLid



Leonidas N. Kalousis (VUB)

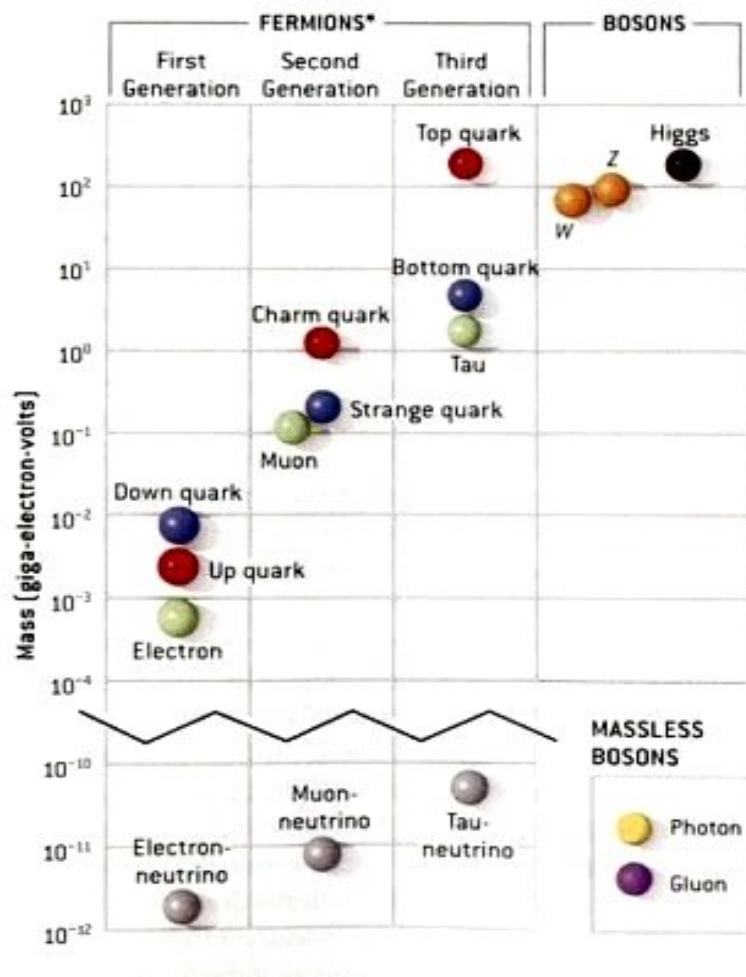
for the SoLid collaboration

March 2, 2017

Introduction

- The recent emergence of the *Reactor Antineutrino Anomaly* has revived the interest in very short-baseline experiments probing the disappearance of ν_e and $\bar{\nu}_e$
 - Source experiments using high-intensity neutrino and antineutrino emitters coupled with large-scale detectors
 - Very short-baseline reactor experiments
- SoLid is a reactor project that aims to resolve the anomaly employing a novel detector design
 - $\sim 6.0 - 9.0$ m from the BR2 research reactor core in SCK•CEN (Belgium)
 - Volume segmentation and robust neutron identification capabilities
 - Synergy with reactor monitoring, nuclear non-proliferation efforts
- SoLid (phase I) is currently under construction
 - Scan the allowed parameter region within a year of data taking

Neutrinos

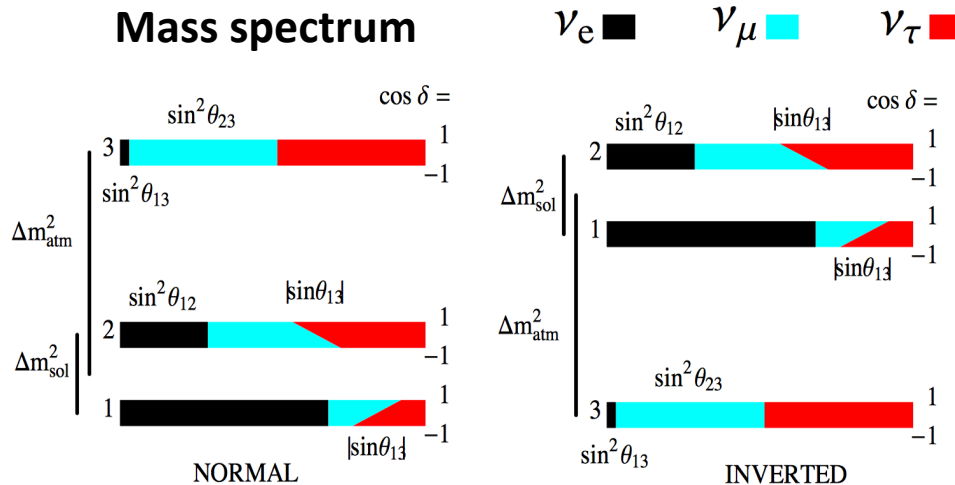


- Originally, incorporated as massless particles in the Standard Model (SM) of particle physics
 - Left-handed helicity states only
- Neutrino oscillations first discovered in 1998 by Super-Kamiokande
 - Now confirmed by several experiments
 - Solar and atmospheric oscillations

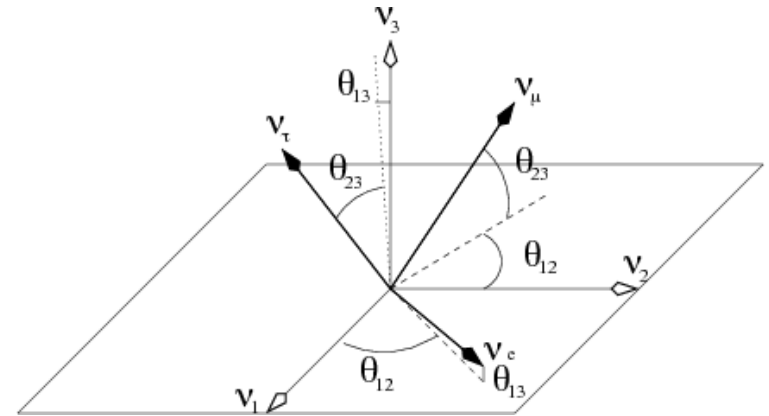
Open questions

- *What are their masses ?*
- *Are they Majorana particles ?*
- *Is there CP violation in the ν sector ?*
- *Do sterile neutrinos exist ?*

Neutrino masses and mixing



O. Mena and S. Parke Phys. Rev. D 69, 117301 (2004)



Neutrino mixing through a rather simple schematic

- Lepton mixing matrix ($c_{ij}=\cos\theta_{ij}$, $s_{ij}=\sin\theta_{ij}$) :

$$|\nu_\alpha\rangle = \sum_{i=1}^3 \mathcal{U}_{\alpha i}^* |\nu_i\rangle$$

$$\mathcal{U} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\text{Atmospheric}} \times \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix}}_{\text{1-3}} \times \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{Solar}} \times \underbrace{\text{diag}\{e^{i\alpha_1}, e^{i\alpha_2}, 1\}}_{\text{0}\nu 2\beta}$$

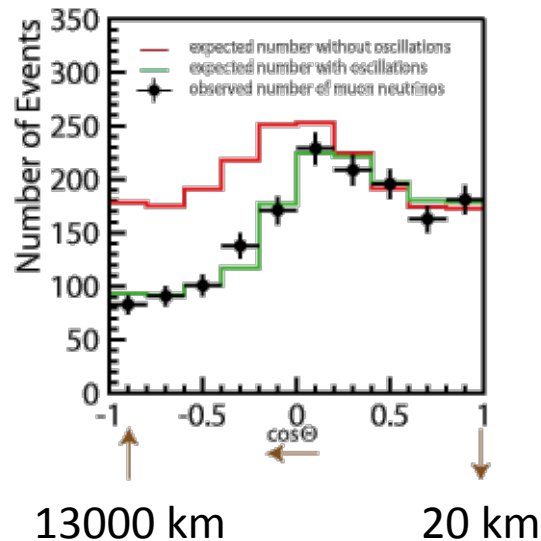
Lepton flavor violation through neutrino oscillations

$$P_{\nu_\alpha \rightarrow \nu_\beta} = \sum_{i,j} U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^* e^{-i \frac{\Delta m_{ij}^2 L}{2E}}$$

- Oscillation patterns driven by the squared-mass differences, Δm_{ij}^2
- The formula depends on the neutrino energy (E) and distance (L)

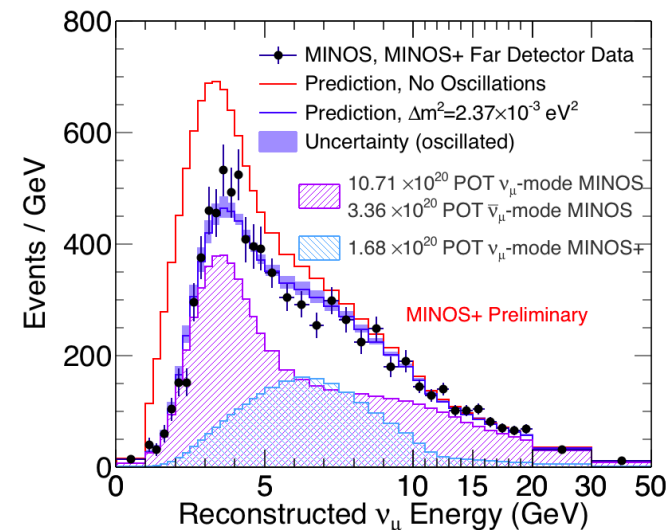
Super-Kamiokande,

<http://www-sk.icrr.u-tokyo.ac.jp/sk/sk/atmos-e.html>



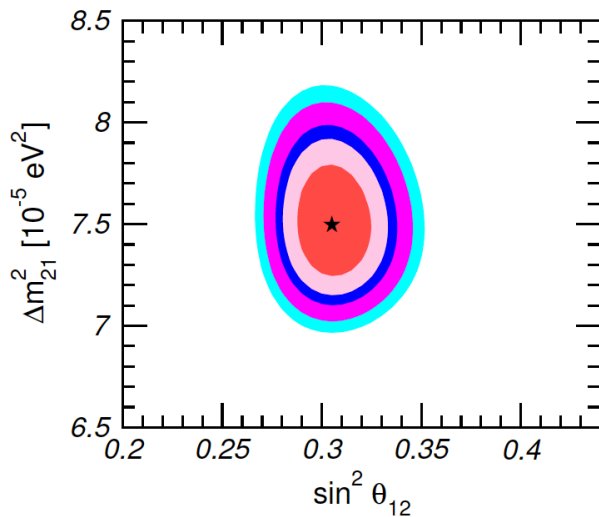
MINOS and MINOS+,

<http://www-nucl.fnl.gov/PublicInfo/forscientists.html>

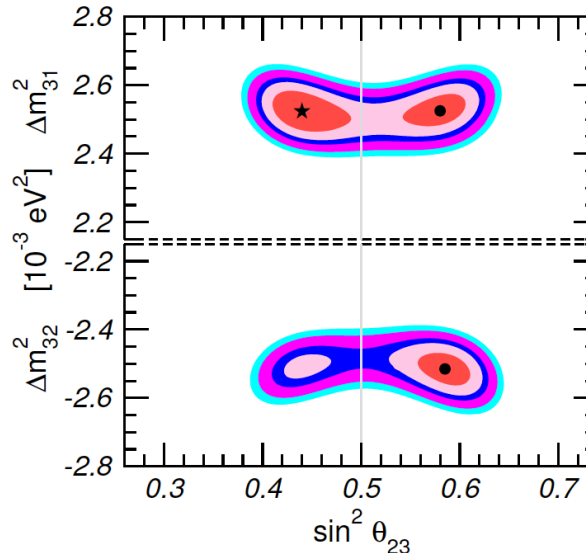


Current picture

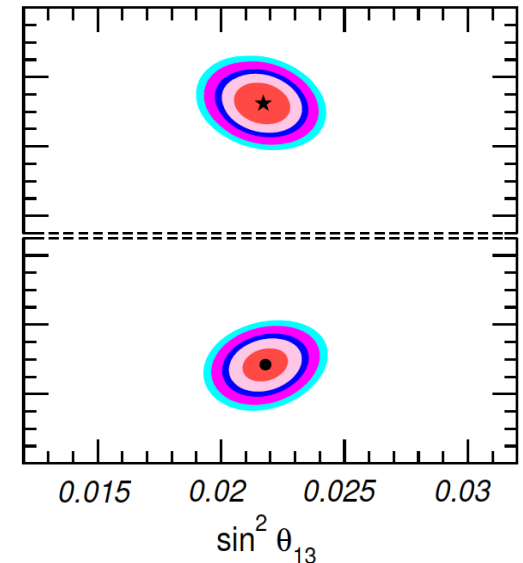
Solar sector



Atmospheric sector



1-3 sector



Solar exp. and KamLAND

$$\Delta m_{12}^2 \approx 7.6 \cdot 10^{-5} \text{ eV}^2$$

$$\sin^2(2\theta_{12}) \approx 0.85$$

(large mixing angle)

Super-K, MINOS, T2K *et al.*

$$\Delta m_{23}^2 \approx 2.4 \cdot 10^{-3} \text{ eV}^2$$

$$\sin^2(2\theta_{23}) \approx 1.0$$

(almost maximal mixing)

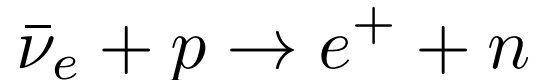
Daya Bay, Double Chooz,

RENO, T2K and Nova

$$\sin^2(2\theta_{13}) \approx 0.1$$

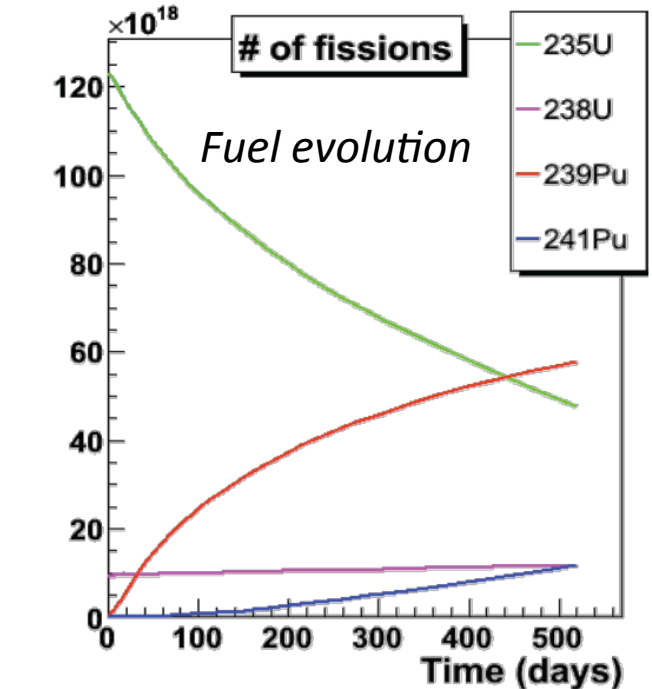
Reactor antineutrinos

- Reactors are copious sources of $\bar{\nu}_e$
 - Beta decays of fission fragments
 - Low energy antineutrinos; isotropic flux
 - An 1 GW_{th} power reactor emits $2 \cdot 10^{20} \bar{\nu}_e/\text{sec}$
- The most common detection channel is inverse beta decay (IBD):



- Number of events detected:

$$n = \frac{1}{4\pi R^2} \frac{P_{th}}{\langle E_f \rangle} N_p \sigma_f \epsilon$$



Cross-section per fission

$$\sigma_f = \int_0^\infty S(E_\nu) \sigma_{V-A}(E_\nu) dE_\nu$$

Reactor spectrum re-evaluation

- New reactor antineutrino spectra pioneered by Saclay
 - Work stimulated by Double Chooz, Phys. Rev. C **83**, 054615 (2011)
- Conversion with “*true*” distribution reproducing >90% of ILL data and five effective branches to the remaining 10%
 - 3% net increase wrt old spectrum for ^{235}U , ^{239}Pu and ^{241}Pu
 - Off equilibrium effects increase neutrino yield
 - Decrease of neutron life-time

A ~6 % increase (confirmed by P. Huber)

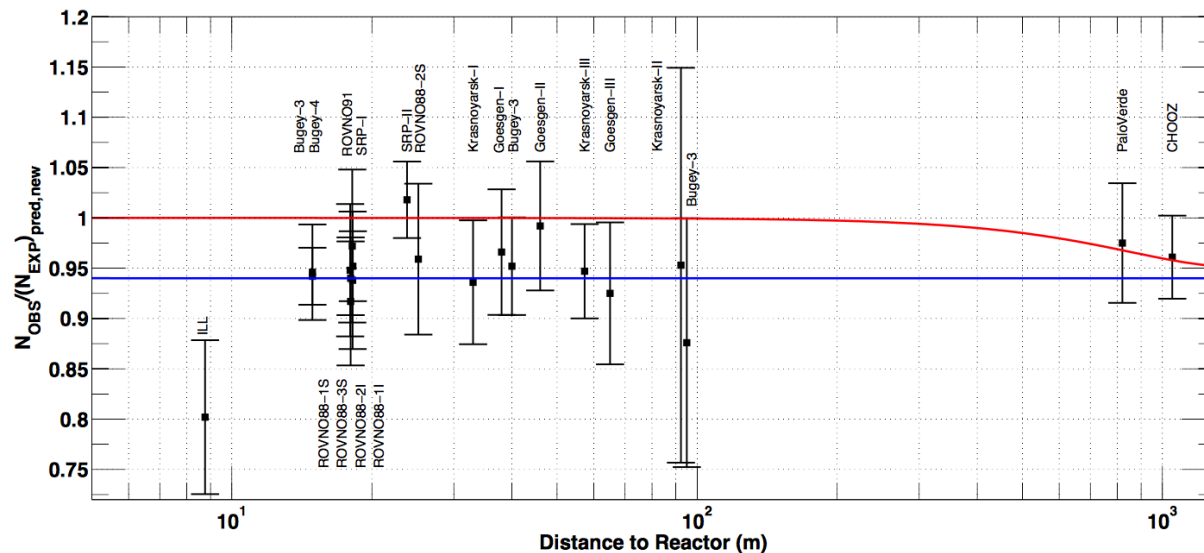
$$\sigma_f^{pred} = \sum_k f_k \sigma_{f,k}^{pred}$$

	old [3]	new
$\sigma_{f,^{235}\text{U}}^{pred}$	$6.39 \pm 1.9\%$	$6.61 \pm 2.11\%$
$\sigma_{f,^{239}\text{Pu}}^{pred}$	$4.19 \pm 2.4\%$	$4.34 \pm 2.45\%$
$\sigma_{f,^{238}\text{U}}^{pred}$	$9.21 \pm 10\%$	$10.10 \pm 8.15\%$
$\sigma_{f,^{241}\text{Pu}}^{pred}$	$5.73 \pm 2.1\%$	$5.97 \pm 2.15\%$
σ_f^{pred}	$5.824 \pm 2.7\%$	$6.102 \pm 2.7\%$

Phys. Rev. D **83** 073006 (2011)

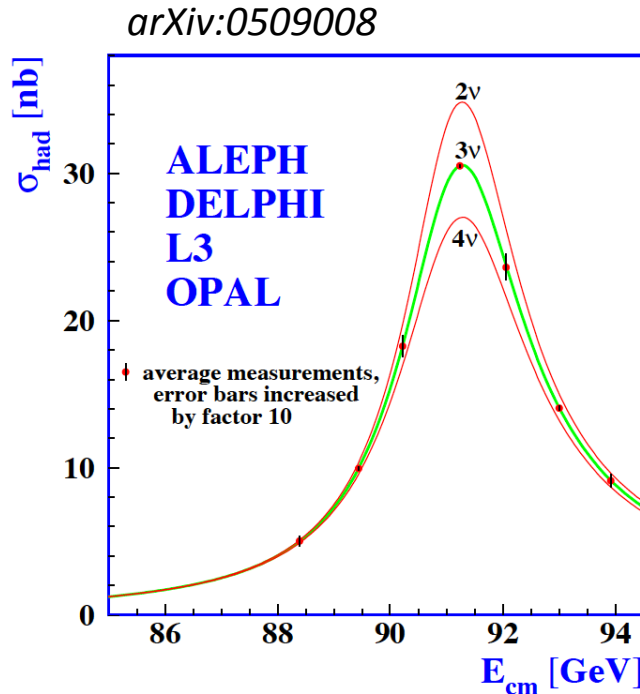
Reactor Antineutrino Anomaly

Mention et al., Phys. Rev. D **83** 073006 (2011)



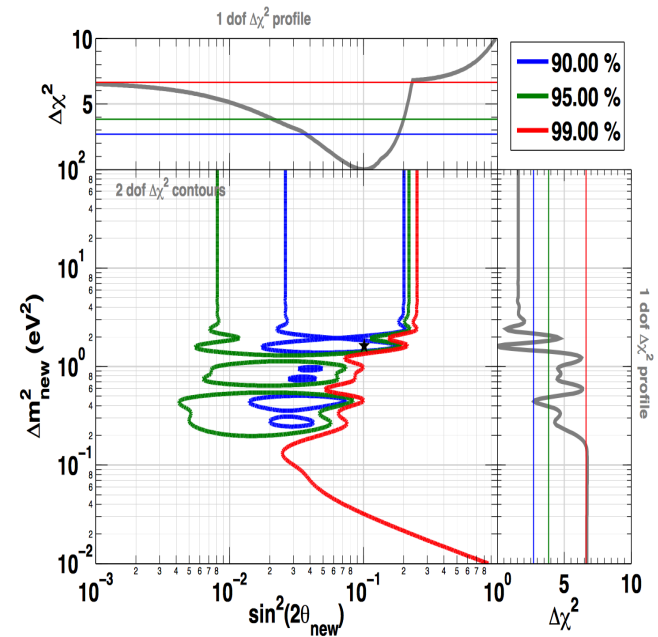
- All previous experiments short baselines ($<100\text{m}$) shifted with respect to re-evaluated spectra
 - Updated observed/predicted averaged event ratio: $R = 0.938 \pm 0.023$ (2.7σ)
- Possible explanations:
 - Wrong estimation of antineutrino spectrum
 - A possible hint for new physics ...

Forth neutrino hypothesis



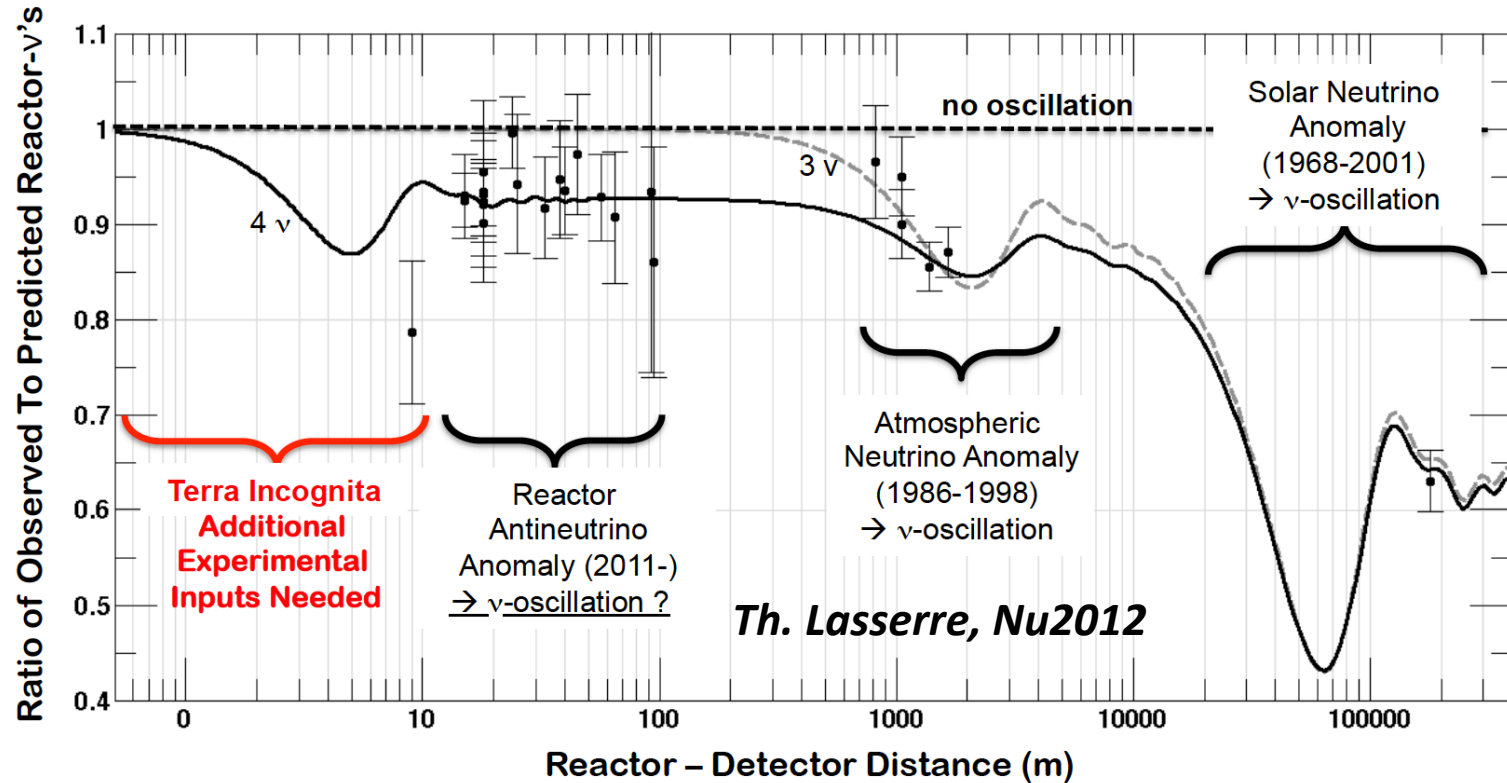
- LEP has constrain the number of (active) neutrinos that couple the Z boson
 - Open possibilities for very heavy or sterile neutrinos
- Sterile neutrinos invoked to explain the LSND excess

$$P_{\nu_e \rightarrow \nu_e} = 1 - \sin^2(2\theta_{ee}) \sin^2\left(\frac{\Delta m_{41}^2 L}{E}\right)$$



Best fit: $\sin^2(2\vartheta) \approx 0.1$
and $\Delta m^2 \approx 1.5 \text{ eV}^2$

Terra incognita

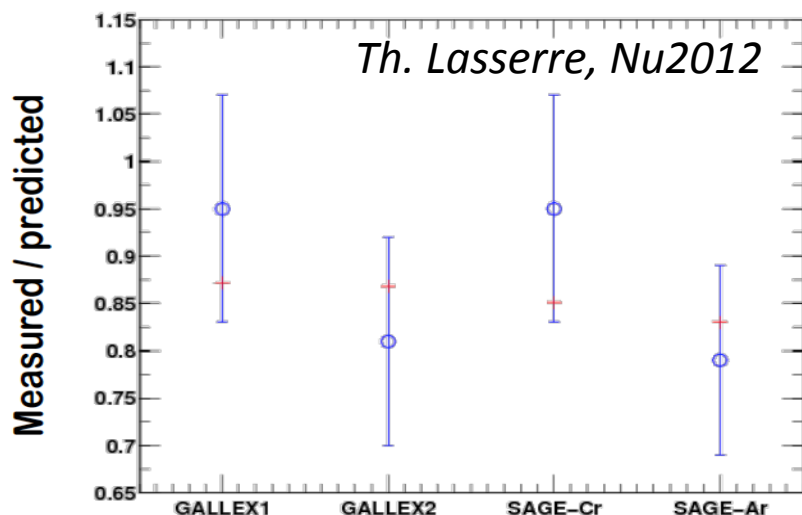


- Oscillations due to an (additional) sterile neutrino are not excluded from other data sets
 - Hints from other experiments in the same channel
 - Further input is needed ! Very short-baseline reactor experiment

The Gallium anomaly and the T2K ν_e disappearance result

Gallex and SAGE

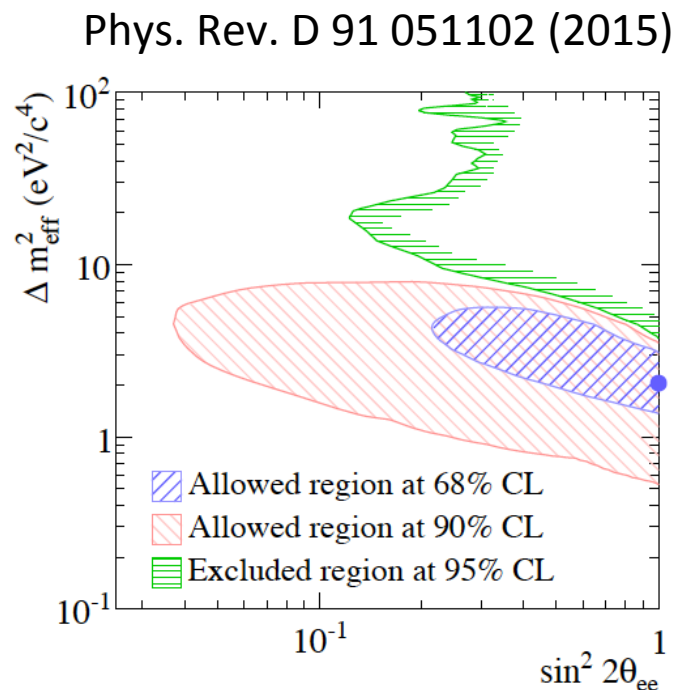
- Four calibration runs with intense MiC neutrino sources
 - ^{51}Cr source, 750 keV ν_e emitter
 - ^{37}Ar source, 810 keV ν_e emitter



- Low counted rates in all runs

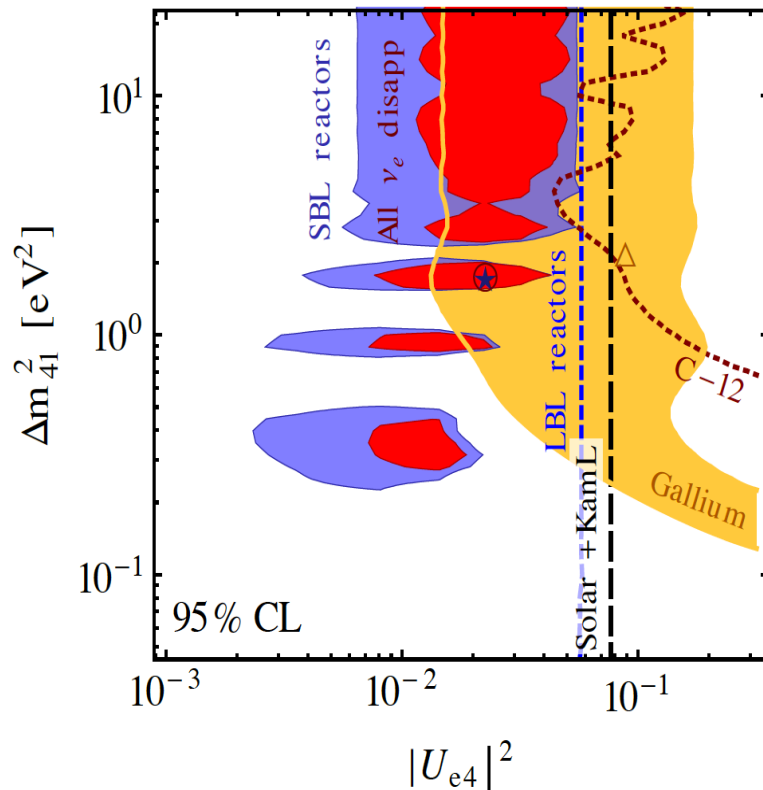
Tokai to Kamioka (T2K)

- Analysis using the beam ν_e contamination at near detector



Global ν_e disappearance analysis

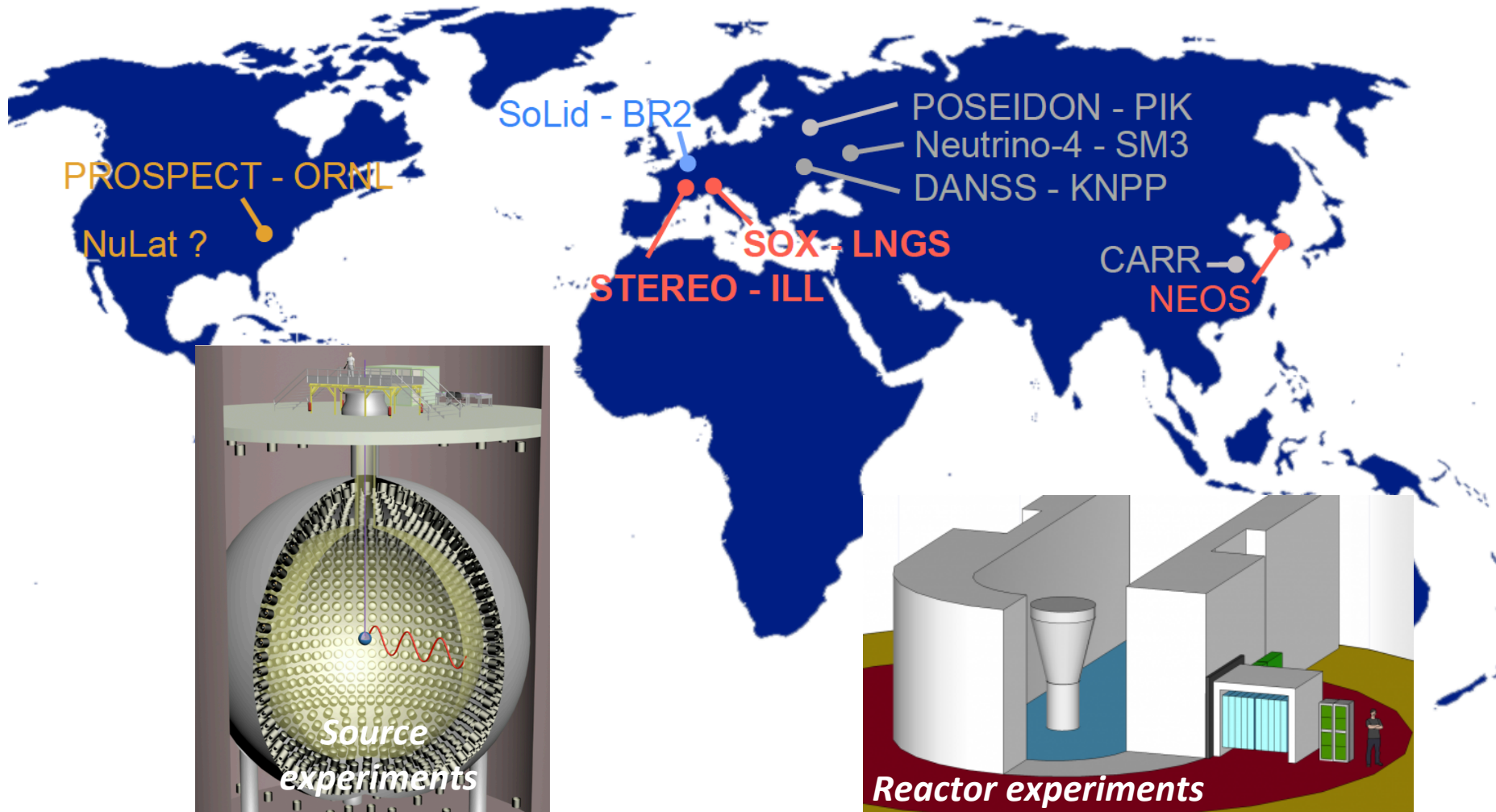
$$\sin^2 2\theta_{ee} \equiv 4|U_{e4}|^2(1 - |U_{e4}|^2)$$



JHEP 1305, 050 (2013)

- Reactor and Gallium anomalies appear to be quite compatible with each other
- Constraints from:
 - LSND and KARMEN ^{12}C data
 - Medium baseline reactor exp. (Chooz, Palo Verde, etc ...)
 - Solar experiments and KamLAND
- Tension between appearance and disappearance experiments
 - No ν_μ disappearance
 - *Recent results from Daya Bay, NEOS and IceCube*

New experimental tests needed



Also: Tritium decay experiments, ie., KATRIN

The SoLid collaboration



Universiteit
Antwerpen



Vrije
Universiteit
Brussel

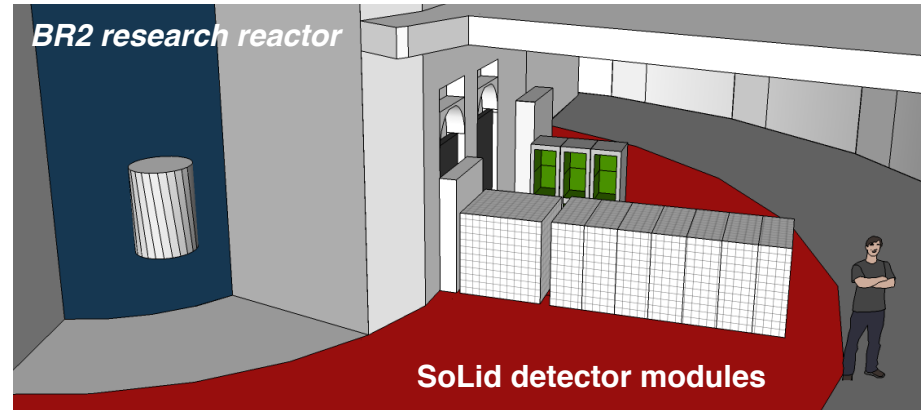
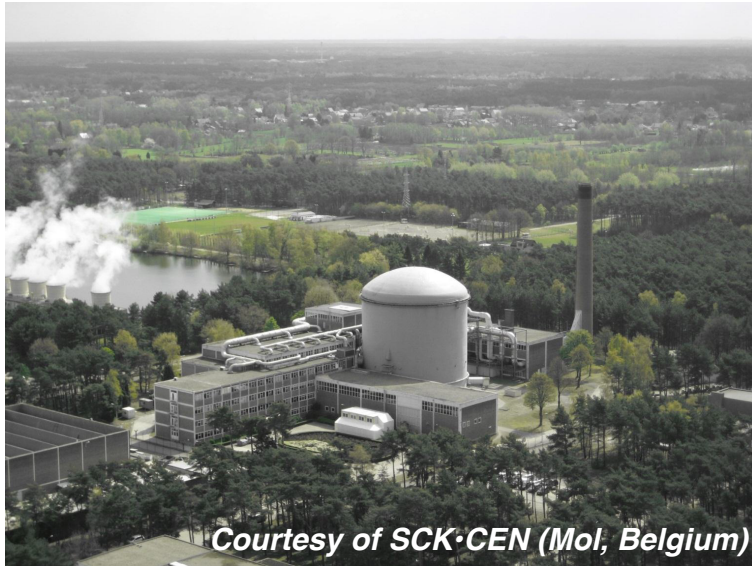


University of
BRISTOL

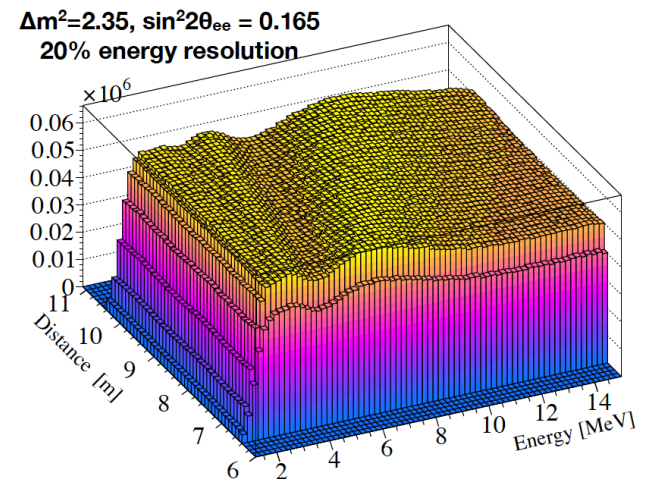


VirginiaTech
Invent the Future

Experimental layout



- Detector modules installed at a distance of $\sim 6.0 - 9.0$ m from the BR2 reactor
 - Precise reactor antineutrino oscillometry
- ^{235}U flux measurement
 - Improve reactor flux prediction
 - Demonstrate reactor monitoring



Challenges

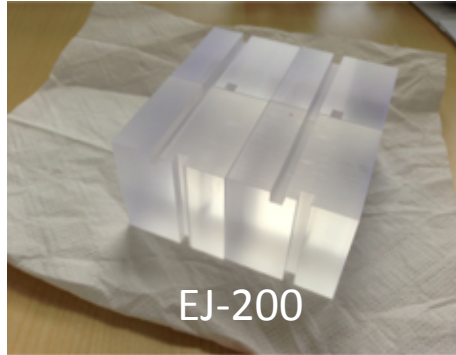
- **Small oscillation effect (10%)**
 - Large statistics, good understanding of systematics
- **Requires compact reactor core ($d < 1$ m)**
 - A few meters oscillation length
- **Cover a large baseline range (6.0 - 9.0 m)**
 - Good vertex and energy resolution
- **Control of background is the key**
 - Close proximity to a nuclear reactor
 - Low overburden (almost on surface)

BR2 reactor at SCK•CEN

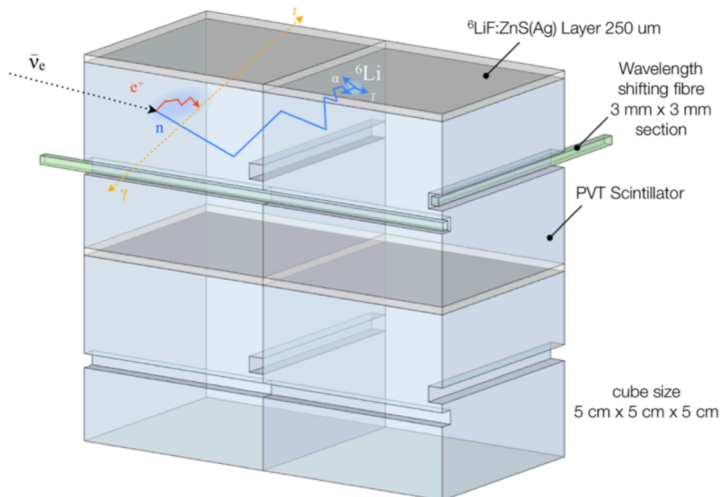


- Research nuclear reactor
 - Highly enriched in ^{235}U
- High operating thermal power
 - Typical values between 40 and 80 MW
- Compact antineutrino source
 - 50 cm effective core diameter
- 150 days per year duty cycle
 - Reactor off running data for background understanding and subtraction
- Low reactor correlated background rate (compared with other sites)
- Large available space covering baselines of 5.5 to 12 m
- *Good collaboration with SCK•CEN*

Detector concept

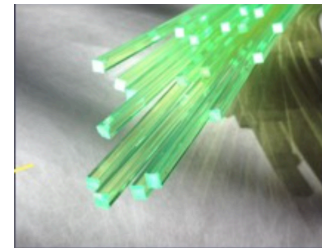


- $5 \times 5 \times 5 \text{ cm}^3$ PVT cubes
- Non-flammable scintillator

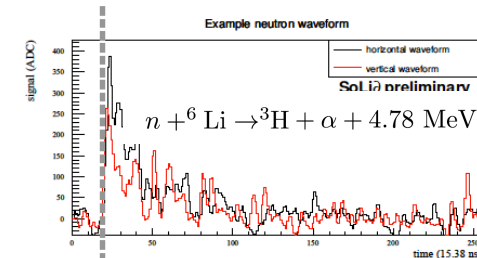
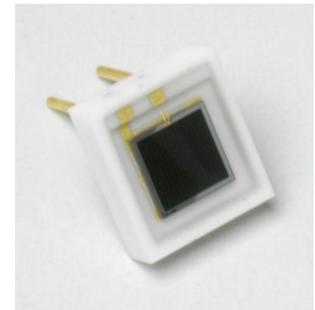


- Adjacent planes of cubes

- Cubes are optically separated (wrapped in Tyvek)
- $^6\text{LiF:ZnS(Ag)}$ for neutron identification
- Light collected through optical fibers and silicon photomultipliers (SiPMs require low-voltage)

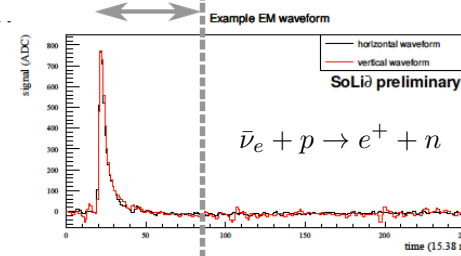


Squared BCF-91A fiber



Delayed
signal

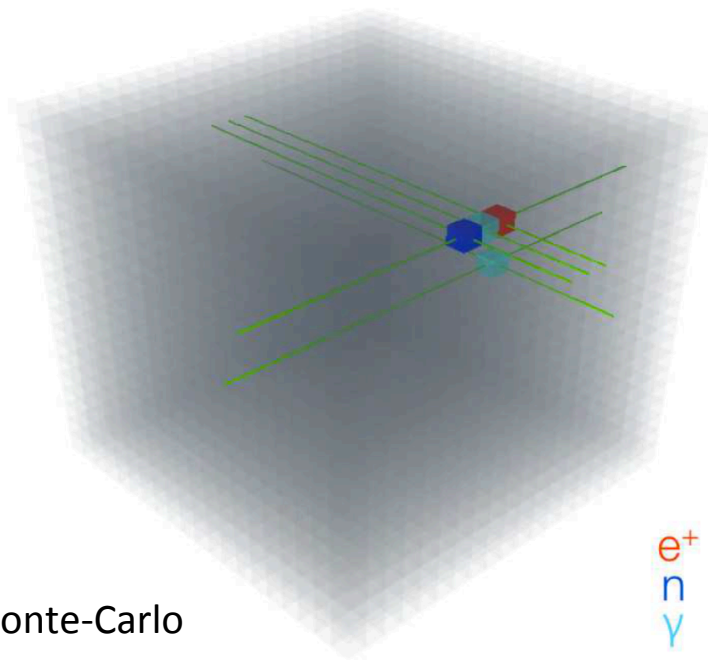
$\Delta t < 250 \mu\text{s}$



Prompt
signal

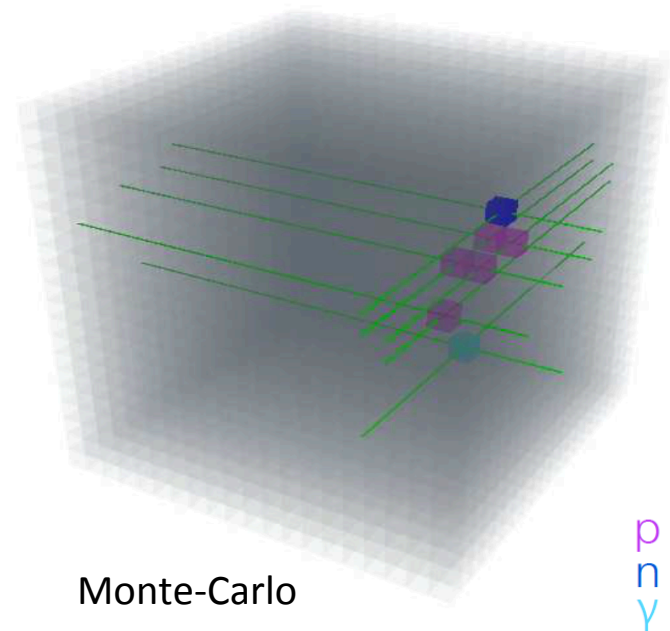
Event topology in SoLid

Inverse beta decay event



Monte-Carlo

Fast neutron event

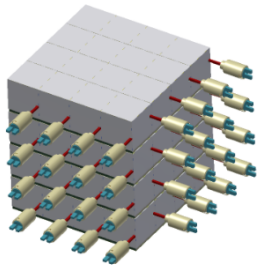


Monte-Carlo

- High granularity allows for signal localization and thus enhances significantly background rejection
- Fast neutron rejection possible through event topology

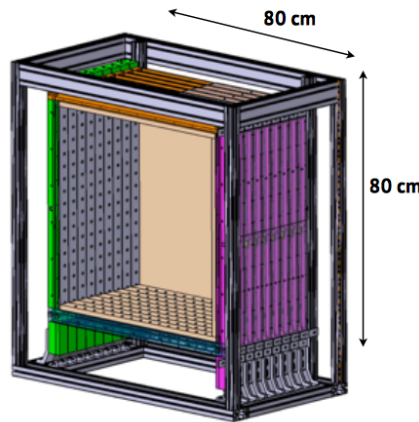
Detector development

NEMENIX, 2013



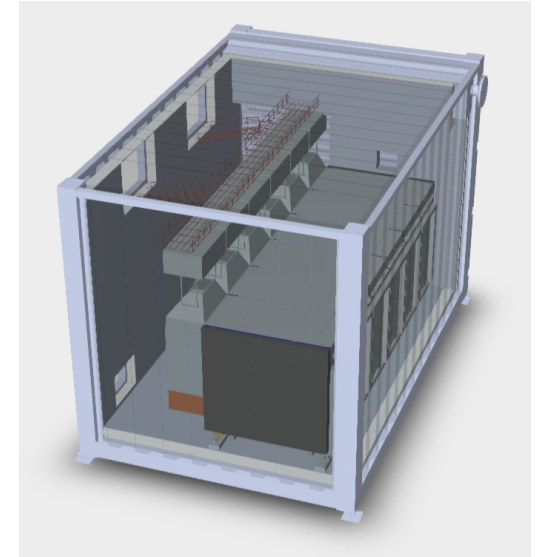
*64 cubes totally
8 kg active mass*

SM1 prototype,
2014 -2015



*9 planes of 16×16 cubes
288 kg active mass*

Phase I detector, 2017



*60 planes of 16×16 cubes
2.0 tons active mass*

Proof of principle

- Validate neutron identification
- Demonstrate prompt-delayed signal selection
- Background measurement

First large scale prototype

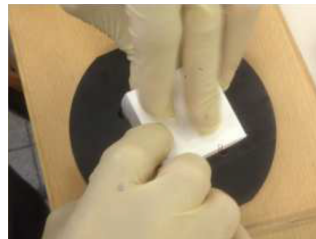
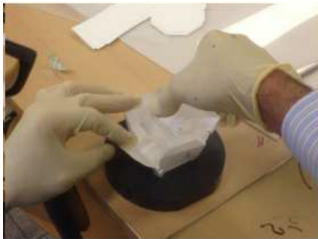
- Demonstrate scalability and test production schedule
- Probe background rejection
- Analysis tools, physics results

Real scale system

- Improved design
- Implement neutron trigger
- Perform high precision measurements

SM1 prototype

- 2304 cubes machined and assembled
 - Wrapped with Tyvek and carefully weighted
 - Number of protons determined with better than 1 % accuracy



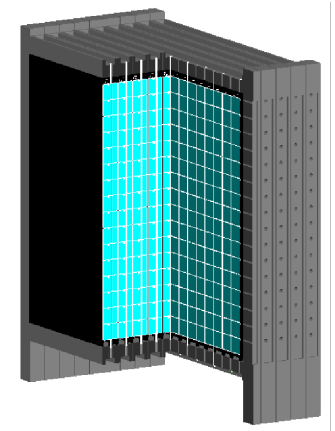
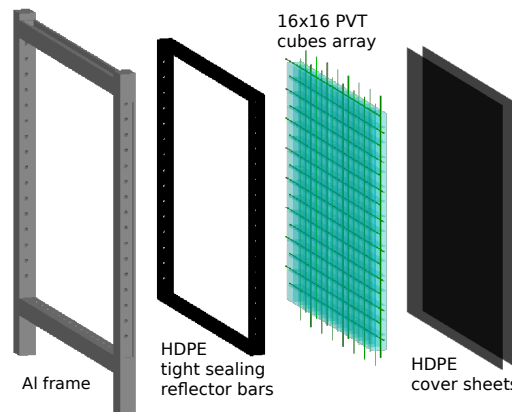
Plane under construction



Assembled plane



- 16×16 PVT cubes grouped together to form a single plane
 - Mechanical support with aluminum frame
 - HDPE to reduce neutron dissipation



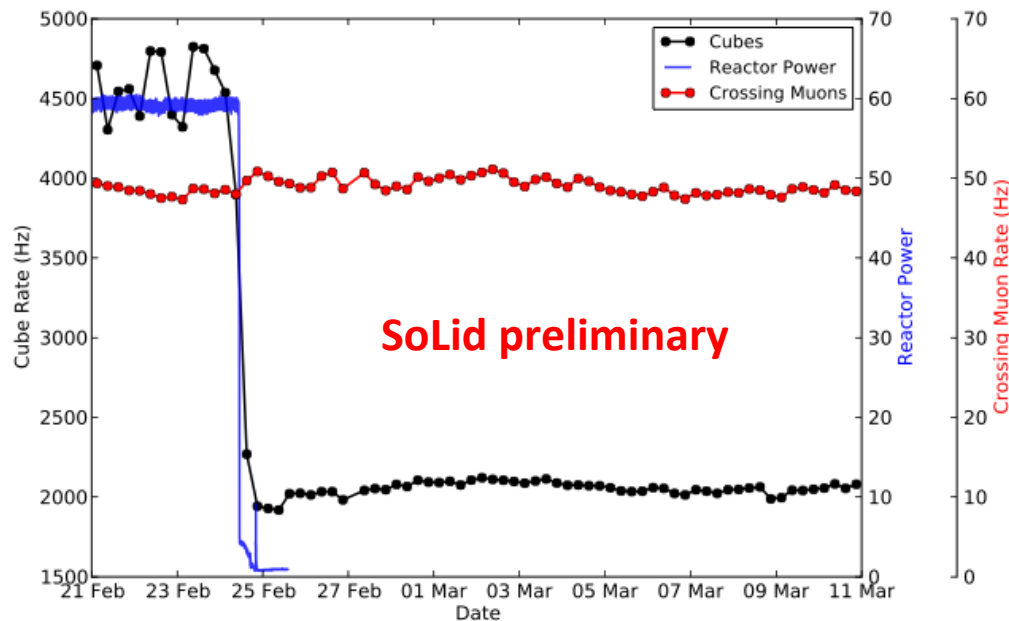
9 planes totally, 288 kg
288 readout channels
80 × 80 × 45 cm

Deployment at BR2

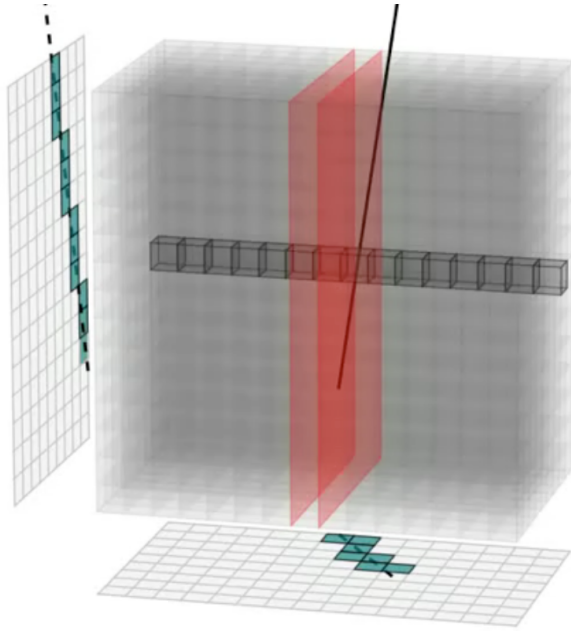


Data taking

- SM1 run at 12/14 - 03/15
 - Detector commissioning
 - 3 - 4 days reactor on
 - ~ 1 month reactor off
- Detector calibration
 - ^{60}Co and AmBe (04/15)
 - ^{252}Cf in situ (08/15)

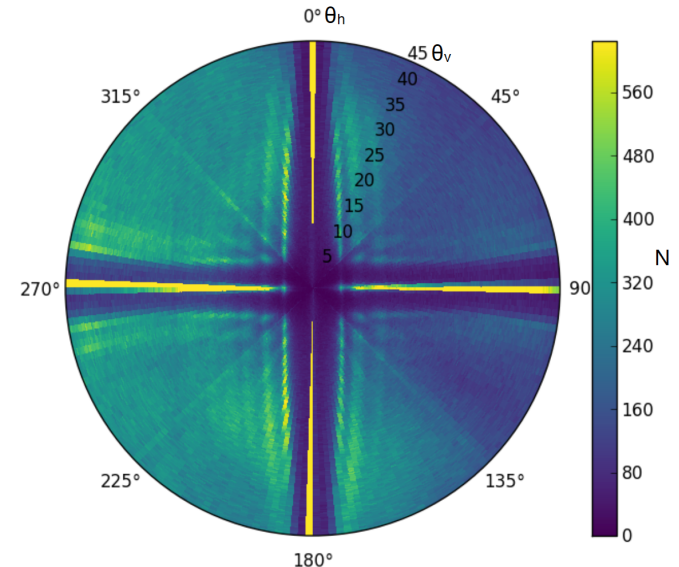
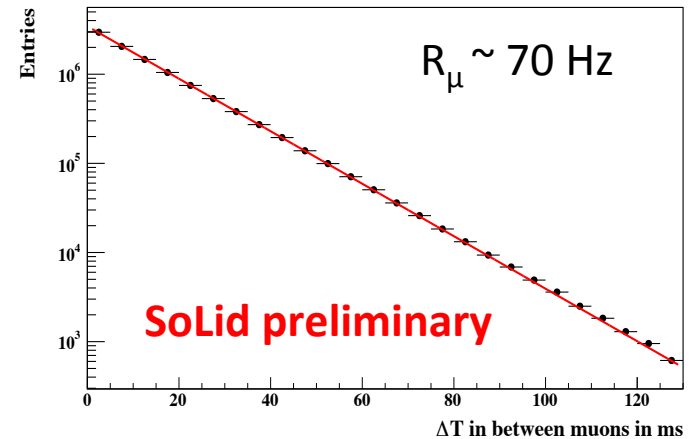


Cosmic muons



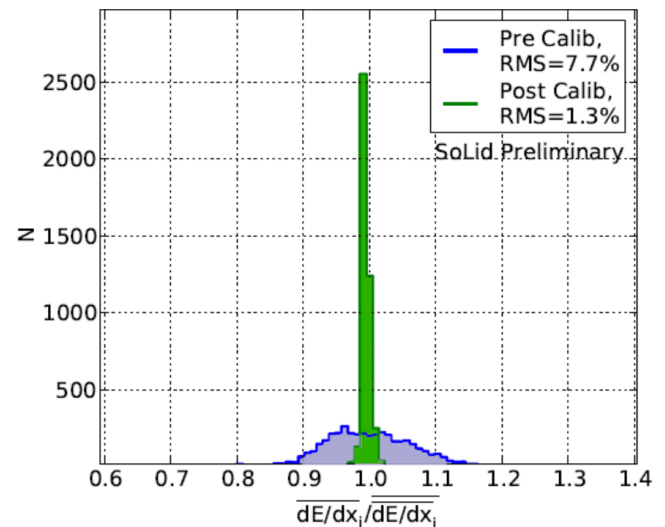
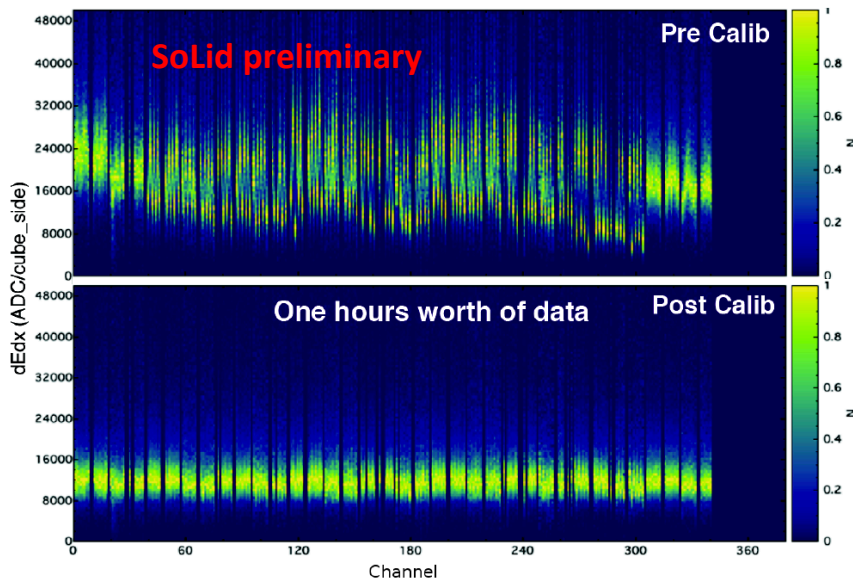
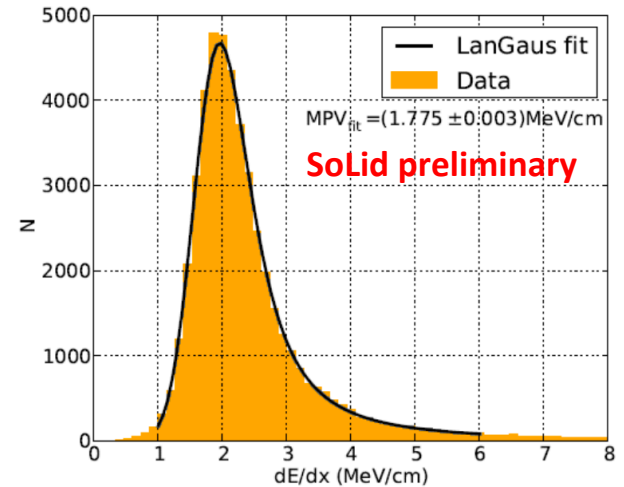
Crossing muon event

- Excellent muon tracking due to detector segmentation
- Detector calibration and stability monitoring using cosmic muons
- Provides handle on muon correlated background rejection

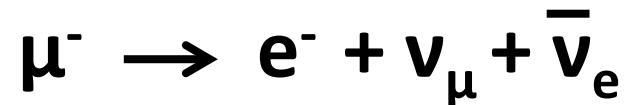
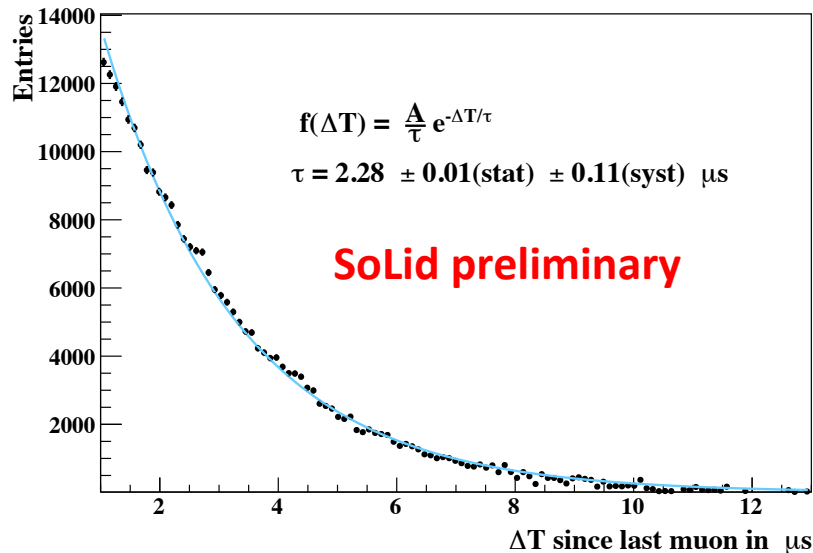
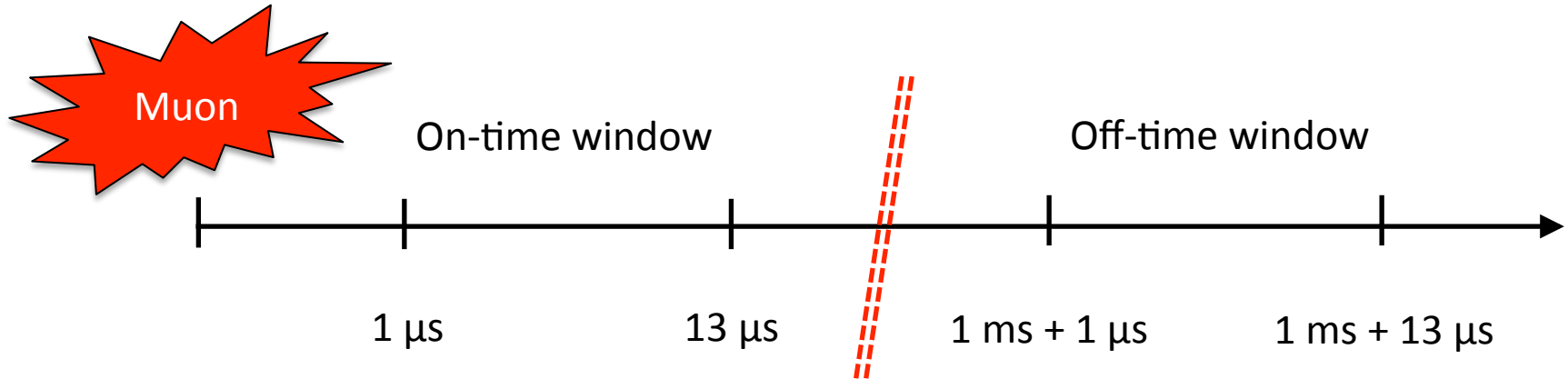


Calibration with cosmics

- In-situ energy calibration using dE/dx
 - Channels inter-calibration
 - Cube response equalization
- Light yield measured: 25 PA/cube
- SiPM gain measured with dark rate
 - No need for light injectors

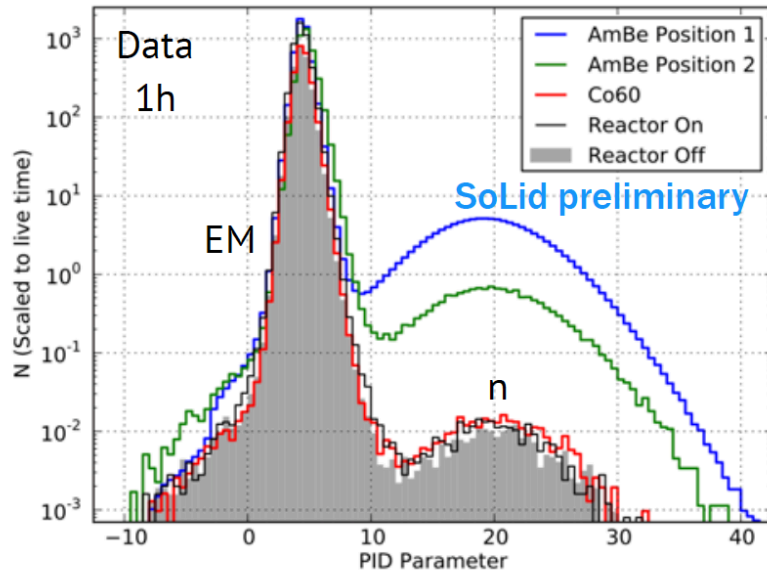


Muon daughters: michel electrons



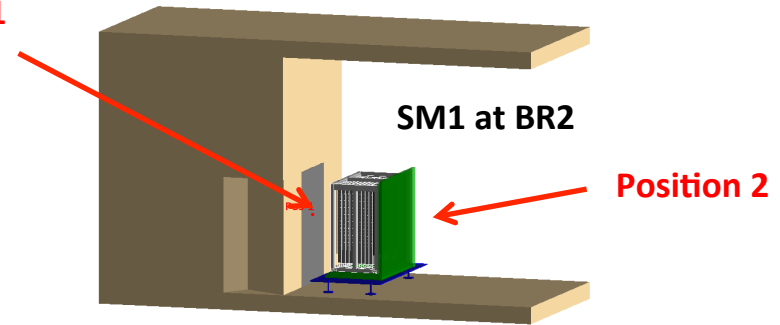
- Michel decay probes the tagging of prompt-delayed coincidences
 - DAQ is well behaving
- Large sample of michel electrons that can be used for calibration
 - Higher energy range than IBD

Neutron identification

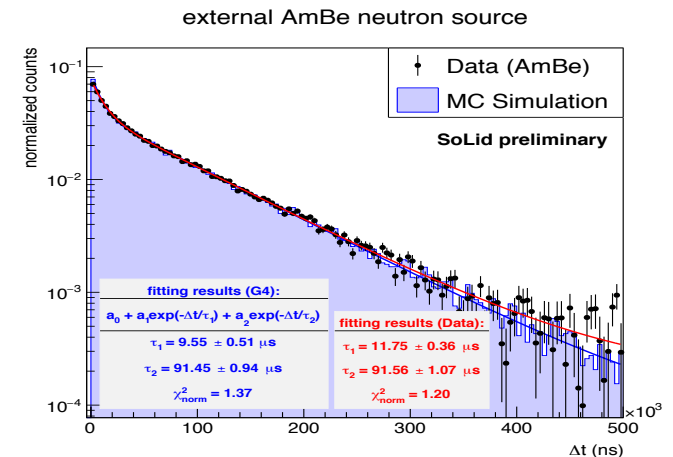


- Excellent neutron identification
 - Tail-to-total algorithm using information from both channels
- Can distinguish a neutron in millions of signals !

Position 1

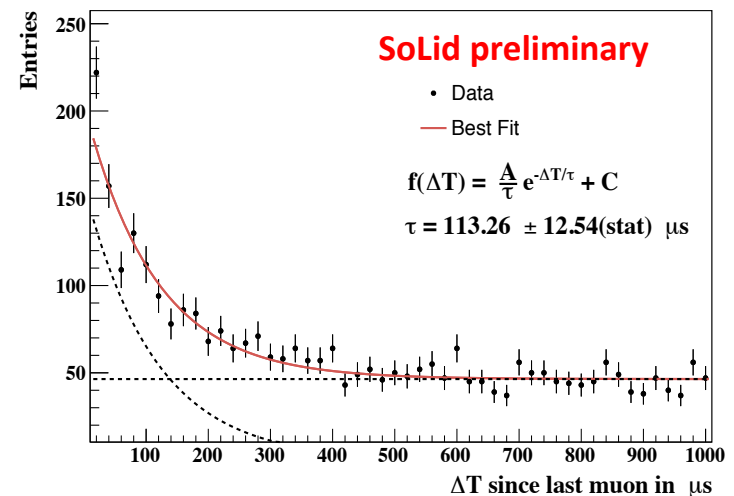
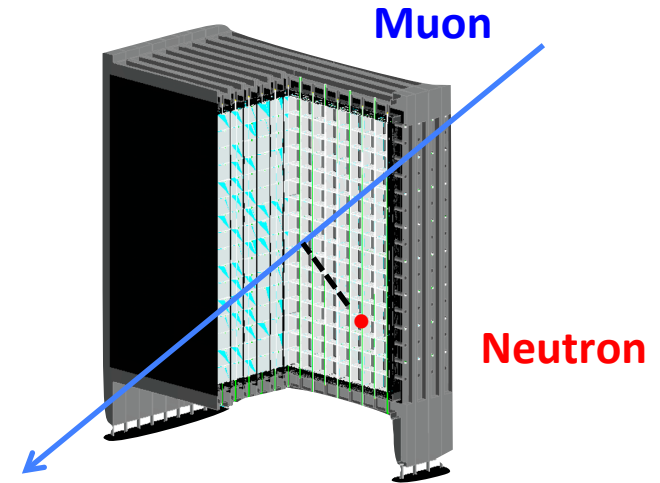


- Calibration runs with AmBe source
 - AmBe is a fast neutron emitter
- Study prompt-delayed coincidence
 - Time-space correlation, neutron efficiency, etc...



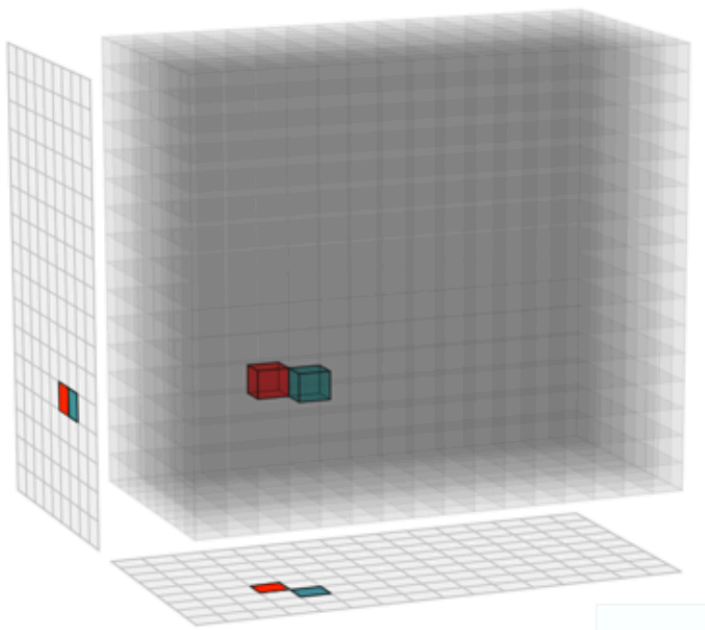
Muon daughters: spallation neutrons

- Muon induced (spallation) neutrons traced in SM1 data
 - Similar selection as with michels
 - Capture time in good agreement with AmBe data
- Control sample that can serve different purposes:
 - Detector stability versus time
 - Neutron identification studies
 - Tune neutron selection

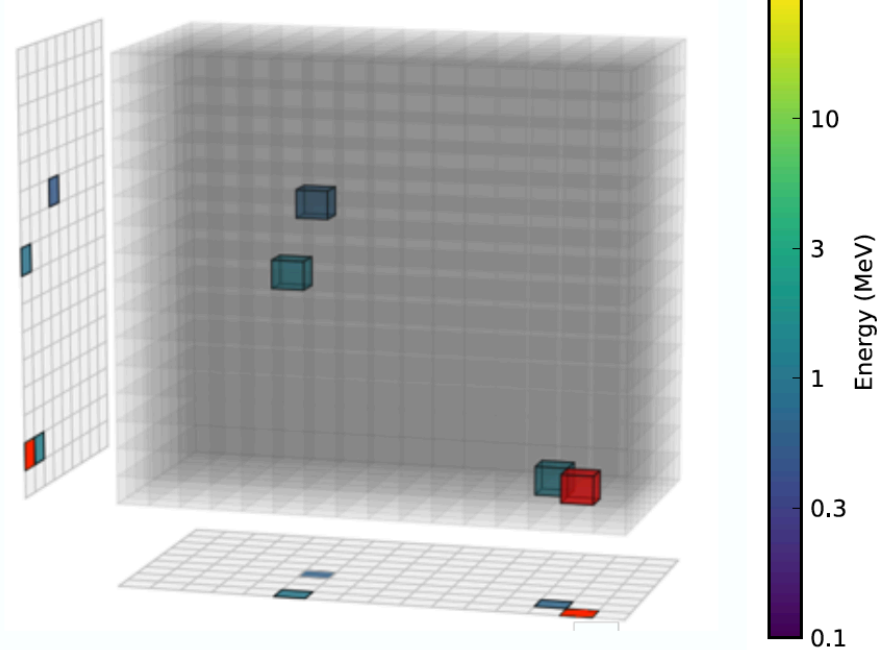


IBD candidates

SM1 data



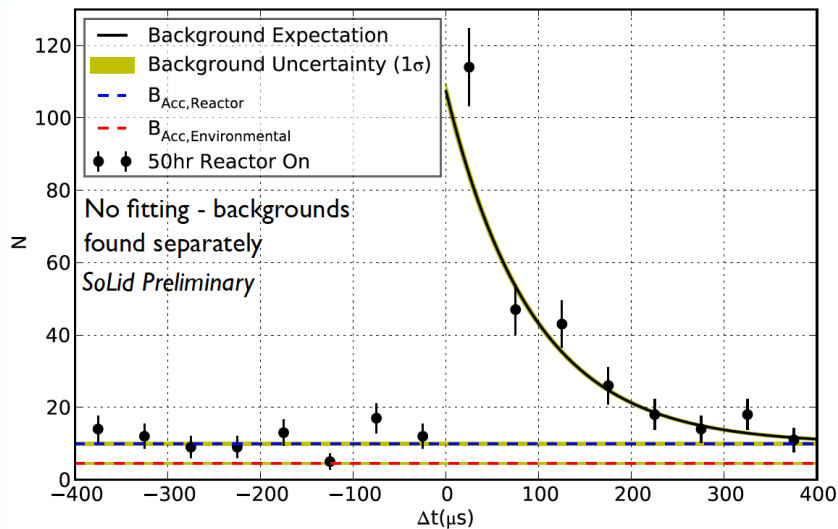
SM1 data



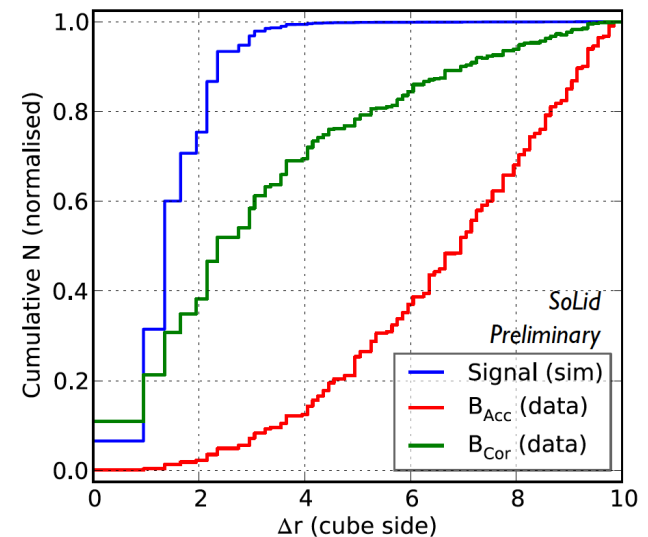
EM signal (prompt) shown in green
Neutrons (delayed) signal are shown in red

IBD selection

- The selection of IBD candidates depends on the time (Δt) and space (Δr) correlation of prompt-delayed coincidences



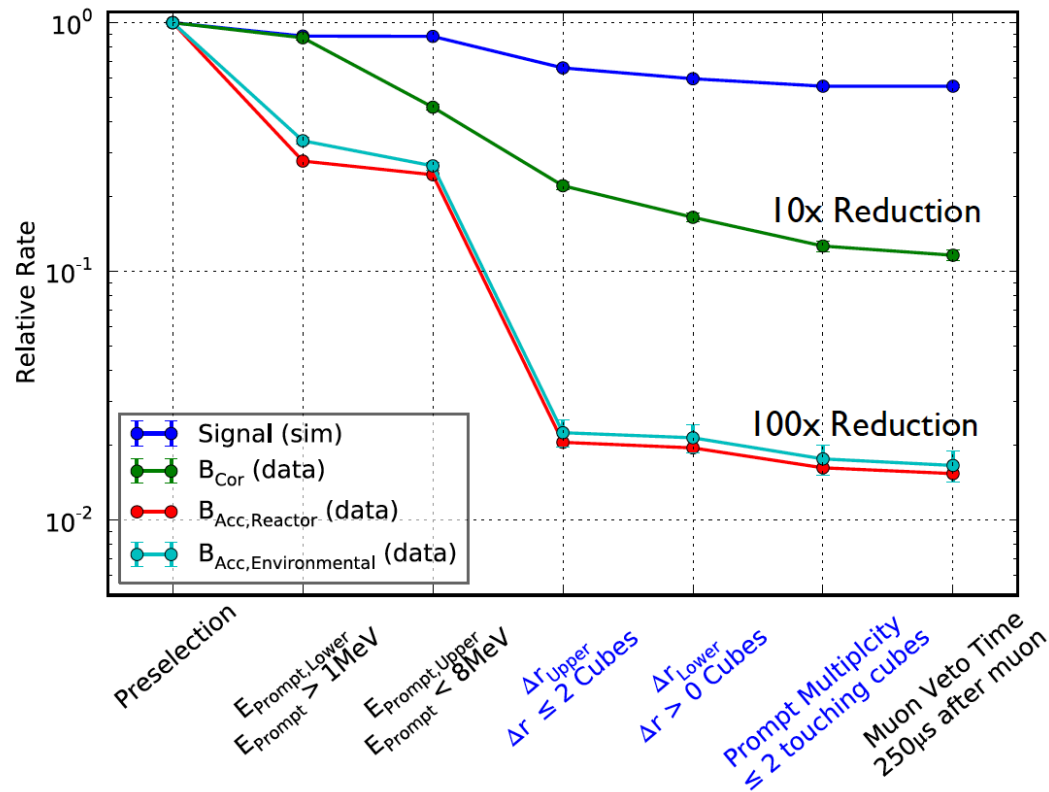
Accidental background has a flat contribution as it was expected



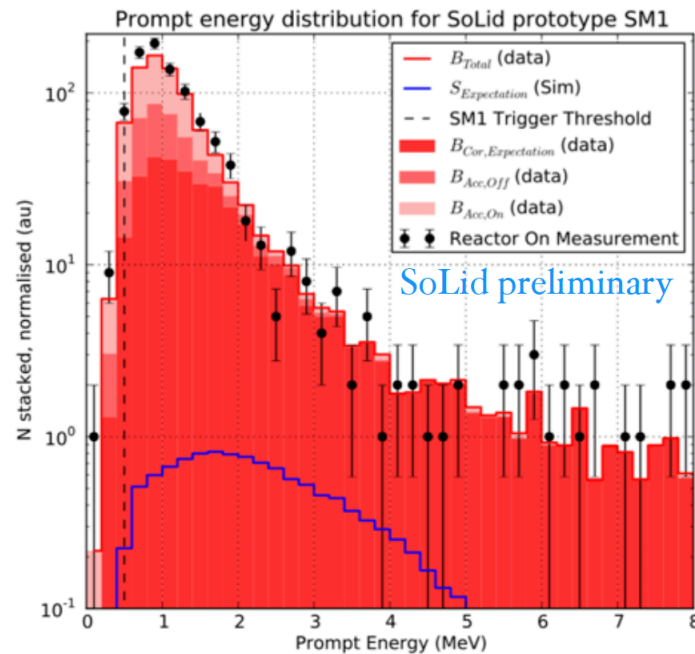
Large background reduction using a Δr cut (volume segmentation)

Background reduction

- Very good signal-background discrimination has been achieved with SM1



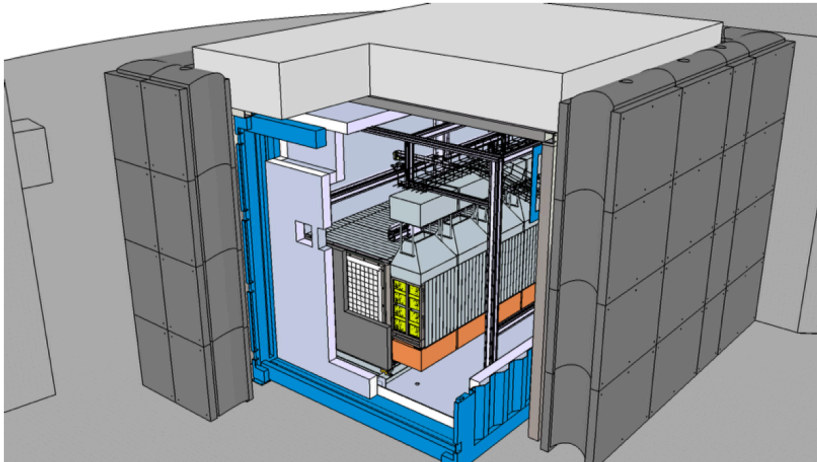
Signal and background



- Correlated background estimated using reactor off data and accidentals using off-time windows
- Good agreement between reactor on data and expectation; Validation of background estimation methods

Phase I detector

Construction phase has started



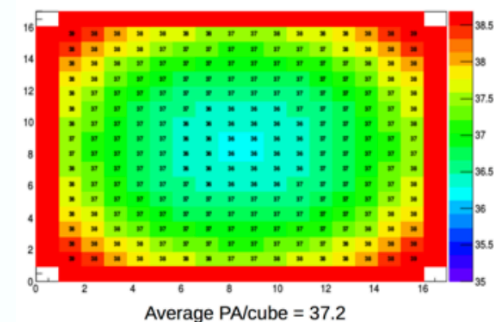
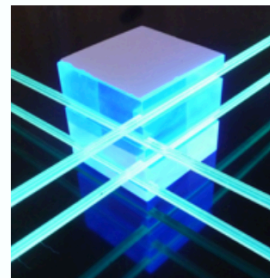
- 6 modules of 10 planes, 16×16 cubes
- 60 planes totally, 15360 cubes
- Temperature controlled system
- 3200 readout channels, max 0.5 TB per day

Light yield and uniformity

- Four fiber readout
 - 37 PA/MeV, 66% increase with respect to SM1
 - 7% variation of light yield across the detector

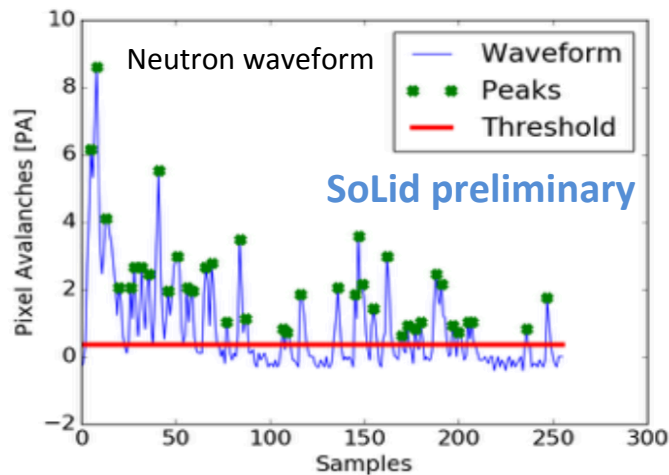
Neutron detection efficiency

- Additional LiF:ZnS sheets
 - Increases ^6Li capture efficiency from 51% to 66%
 - Reduces capture time from 105 μs to 66 μs
- New screens with improved transparency



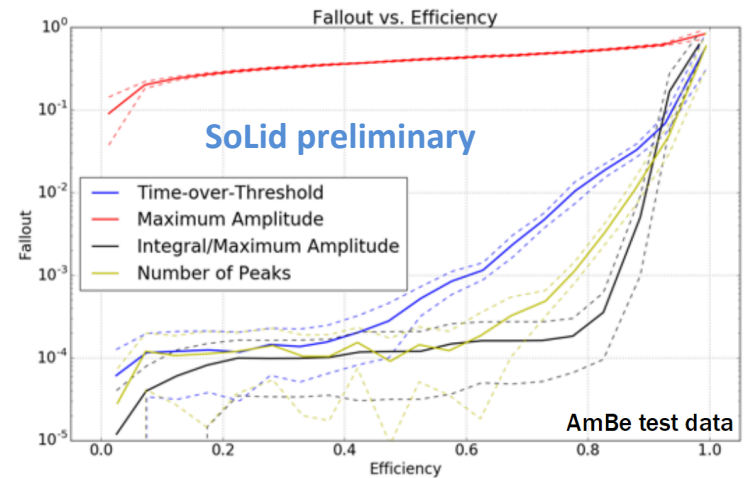
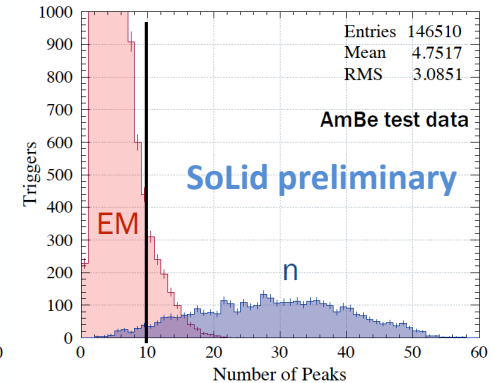
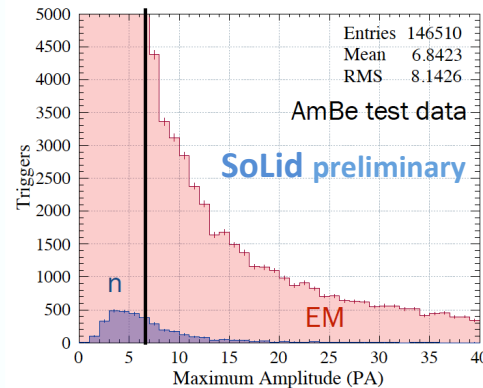
Neutron trigger

- SM1 had a rather low neutron detection efficiency of $\sim 5\%$, due to high trigger threshold (6.5 PA)



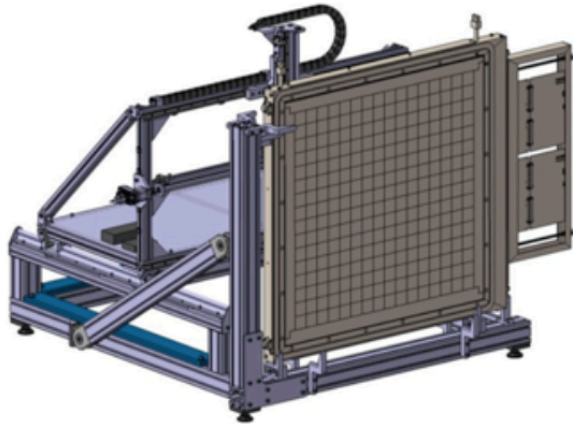
- Phase I detector is designed to have a neutron selection at the trigger level (implemented at the firmware level)
 - Buffer time $\pm 500 \mu\text{s}$ and ± 2 planes around a neutron event
 - Zero suppression threshold at 1.5 PA
 - Reduces dramatically the amount of data
 - Retains high IBD efficiency
- We expect n detection efficiency of $\sim 70\%$

Neutron (n) and electromagnetic (EM) signals



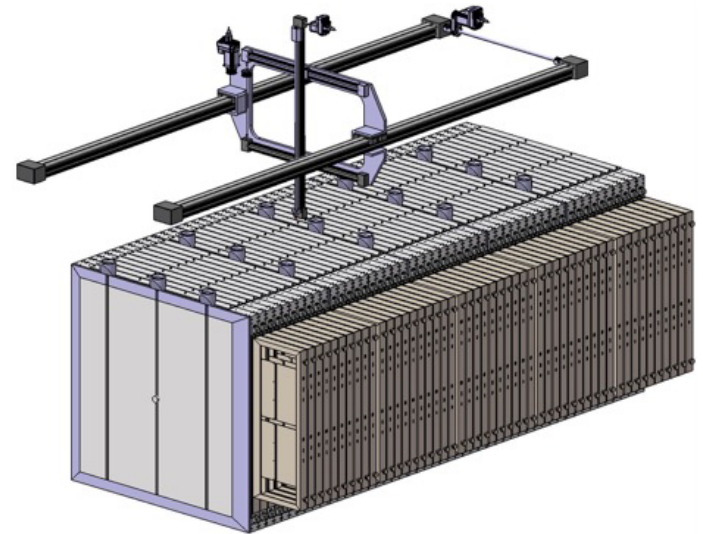
Detector calibration

Off-site calibration system (CALIPSO)



- Plane characterization and commissioning
- Cube to cube equalization
- Neutron and EM signals benchmark

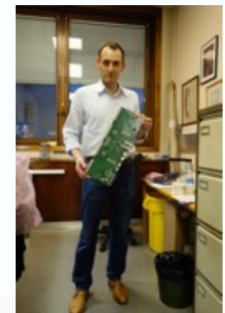
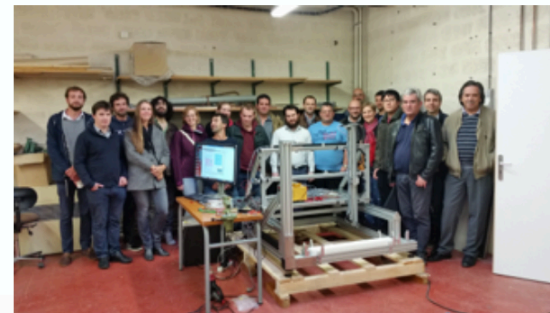
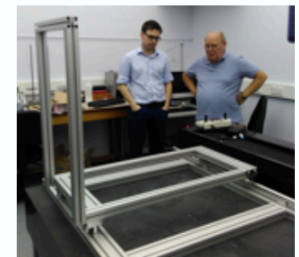
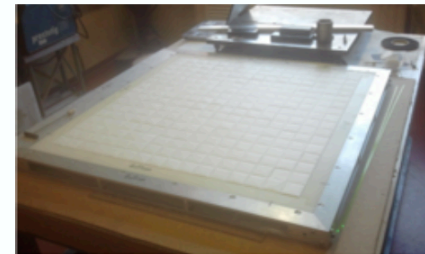
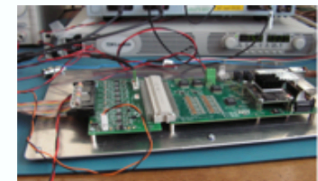
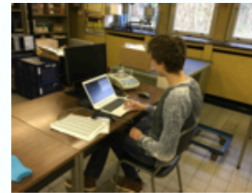
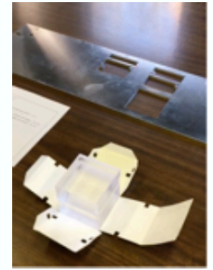
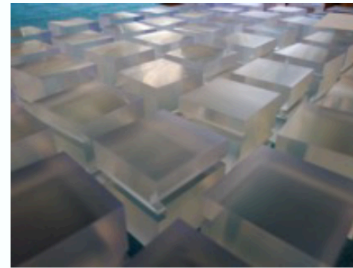
In-situ calibration system (CROSS)



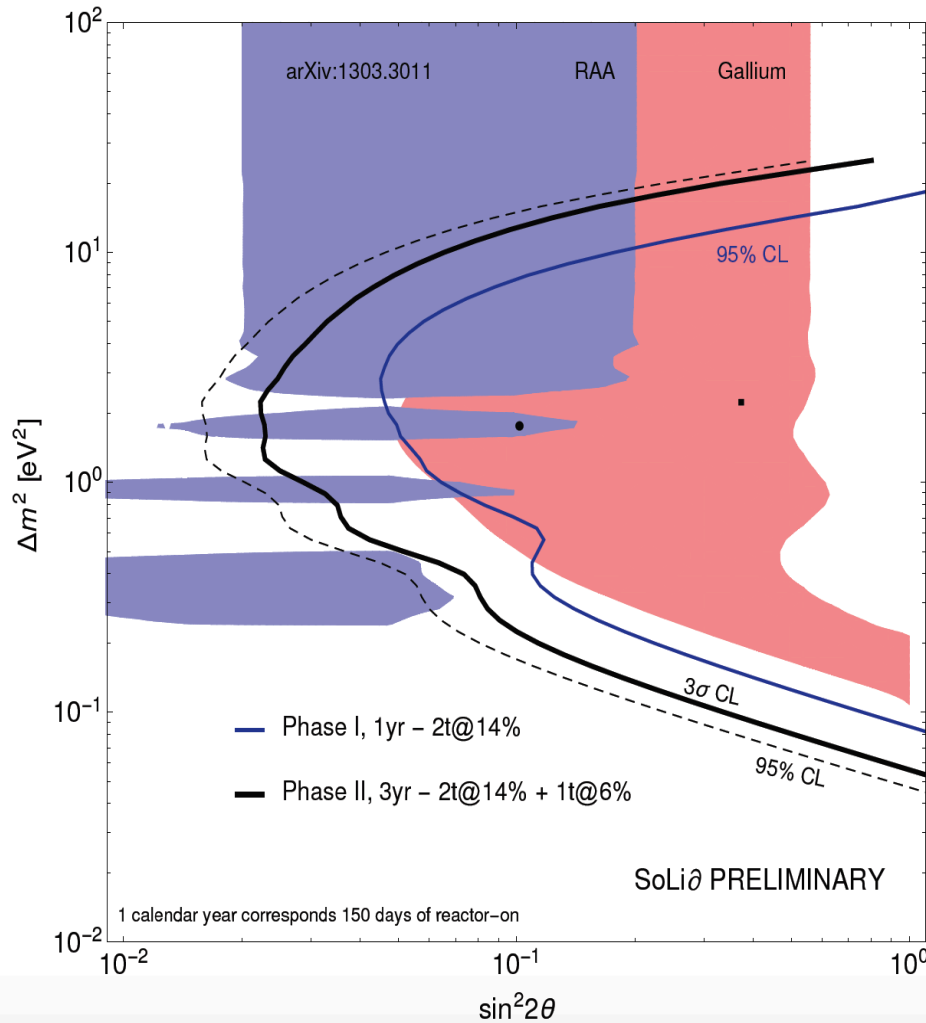
- In-situ radioactive sources deployment
- Precise energy scale and neutron detection efficiency determination

Detector construction

- SoLid detector is under construction !
- Cube production and frame assembly is well underway
- Staged production of electronics
- Trigger firmware implemented and DAQ currently under development
- Container has been built and delivered



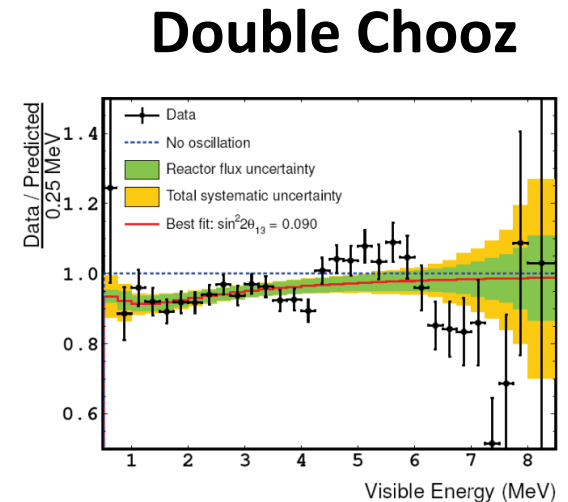
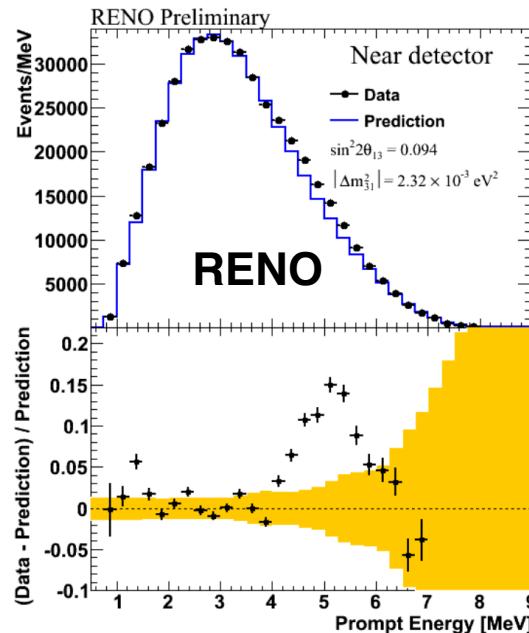
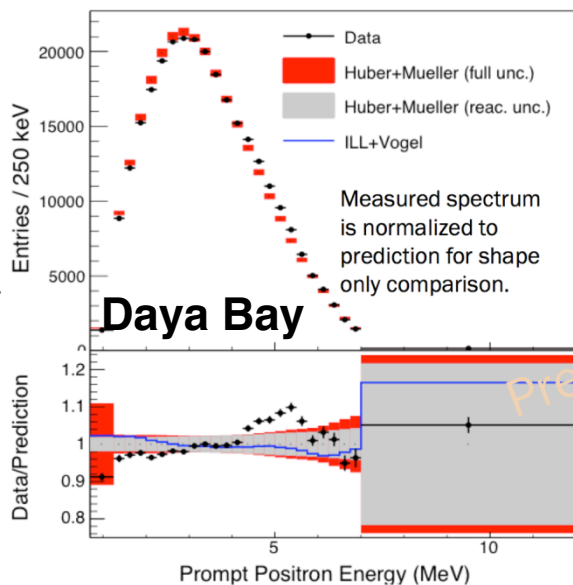
Sensitivity to sterile neutrino



- Two-dimensional fit in energy (E) and distance (L)
 - Good control on detector systematics
 - Energy resolution is crucial
- *Best fit covered within the first year of data taking*

Other physics goals

- Precise measurement of the reactor $\bar{\nu}_e$ spectrum
 - Recent interest after the 5 MeV bump observation



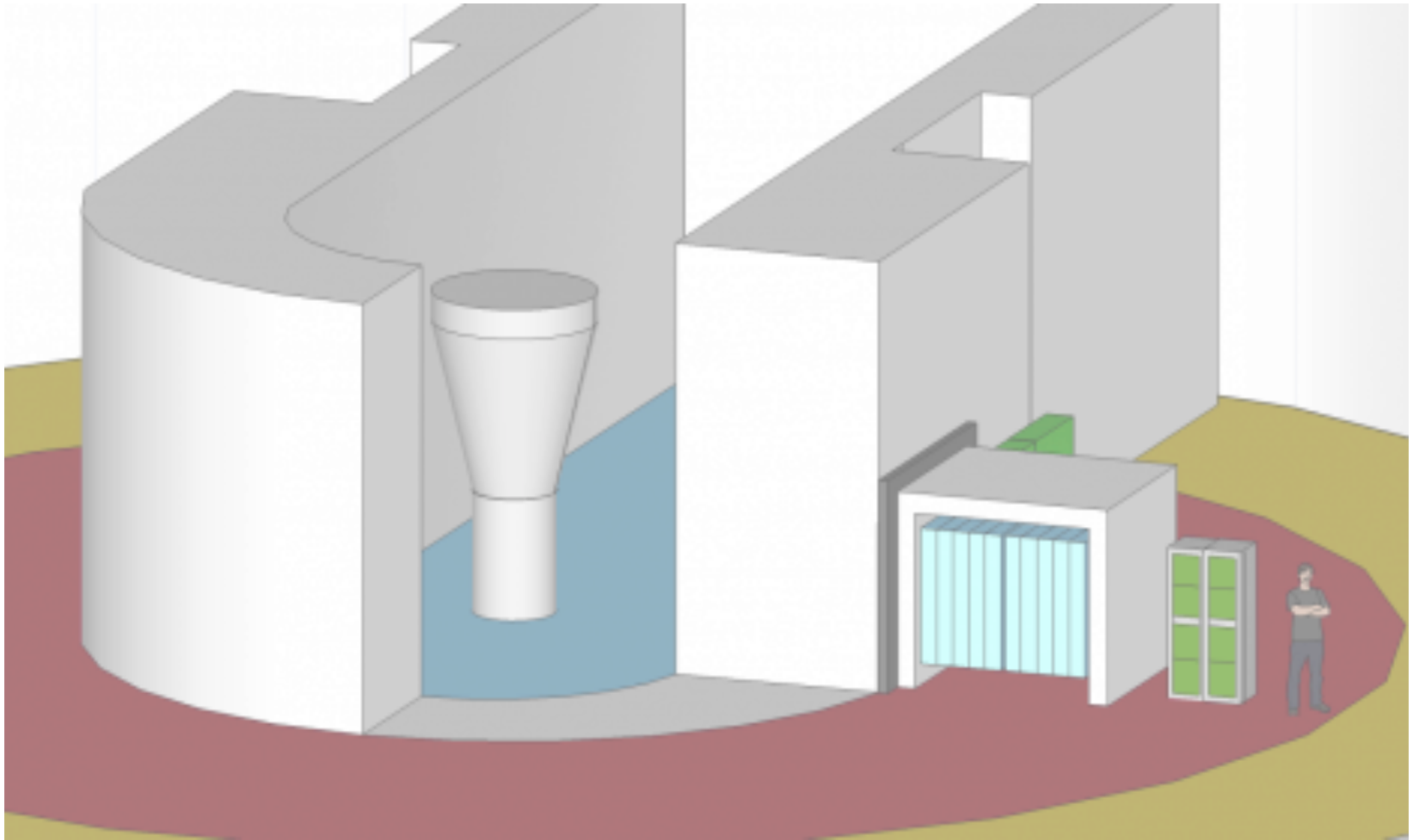
- Synergy with reactor monitoring and nuclear weapons non-proliferation efforts
- For instance: Huber *et al.*, PRL **113**, 042503 (2014)

Ending themes

- The SoLid experiment will make a very sensitive search for $\bar{\nu}_e$ disappearance using a novel detector design
- SM1 operation has been very successful
 - Excellent EM/neutron identification
 - Low background at BR2 has been confirmed
 - Precise calibration with muons (cube equalization $\sim 1.5\%$)
- Entered construction phase for the 2.0 ton detector
 - Funded by FWO, Hercules (BE), ERC (EU) and ANR (FR) grants
 - Increased light yield; more optical fibers
 - Real-time neutron trigger (data reduction)
- Detector commissioned and deployed in BR2 by summer 2017
 - Stay tuned for a high-precision measurement !

Thank you for your attention !

Leonidas N. Kalousis
leonidas.kalousis@vub.ac.be



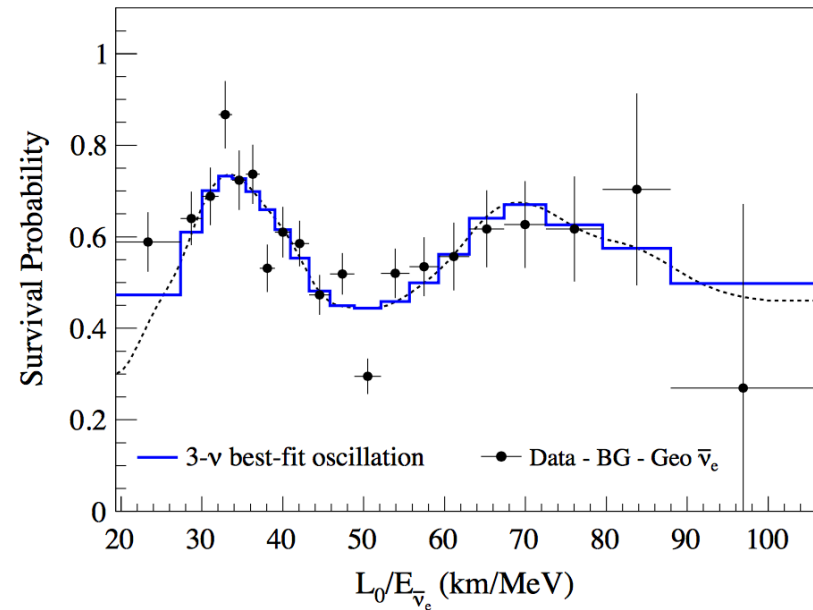
SPARES

Oscillation patterns



KamLAND detector is surrounded by several reactors at different distances

KamLAND, *arXiv:1303.4667*



- Signature of neutrino oscillations seen by many experiments
 - Clear oscillatory patterns that cannot be reproduced by other possible mechanisms (neutrino decay, etc ...)

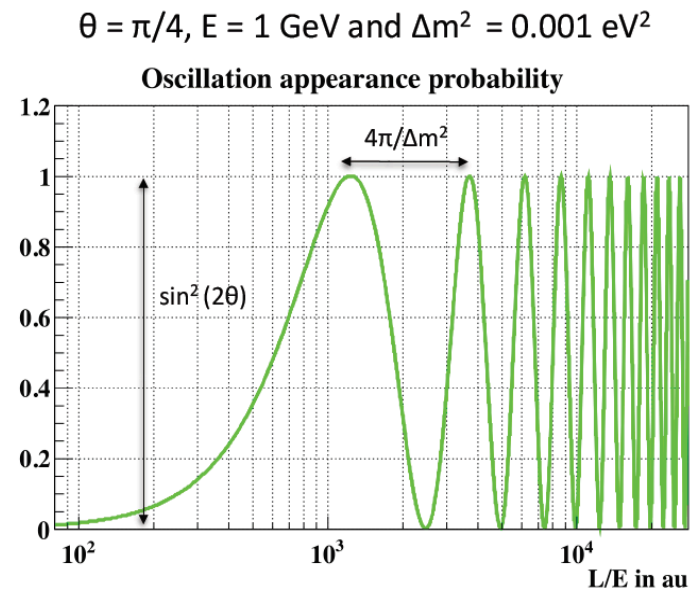
Two flavour approximation

- Three flavour are highly suppressed since $|\Delta m^2_{31}| \ll \Delta m^2_{21}$ and $\cos^2(2\theta_{13}) \approx 1.0$
- Dominant oscillations are well described by effective two-flavour oscillations.
- One mixing angle no complex phase.

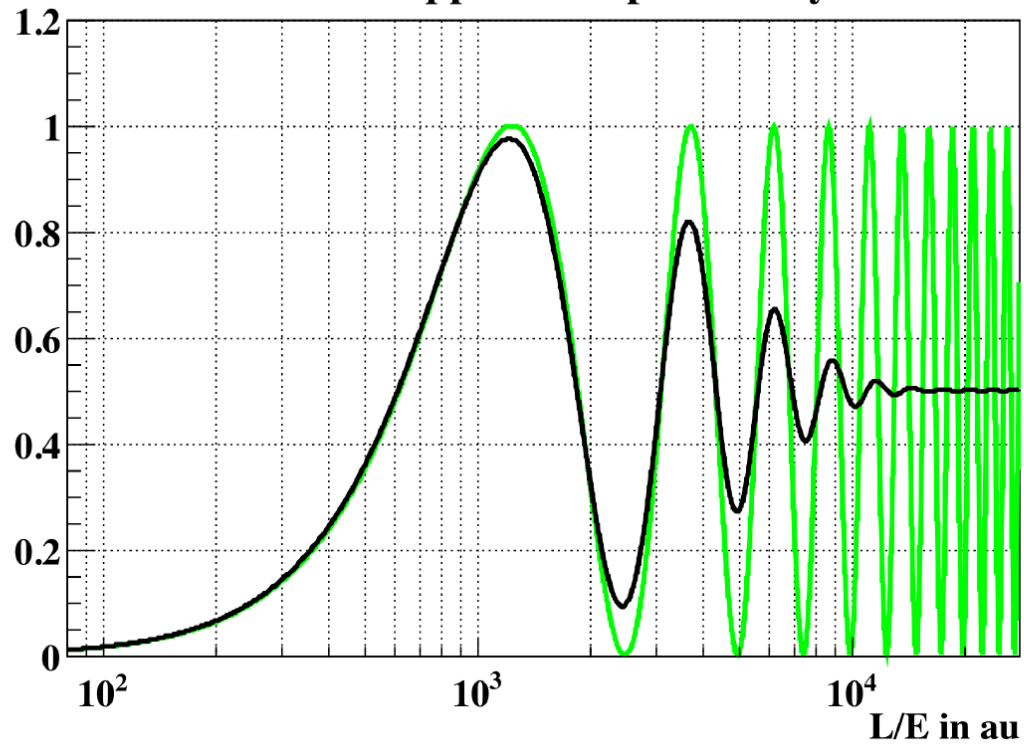
$$\mathcal{U} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$$

$$p(\nu_\alpha \rightarrow \nu_\beta) = \sin^2 2\theta \sin^2 \frac{\Delta m^2 L}{4E}$$

Appearance probability
Typical oscillatory behaviour



Oscillation appearance probability



$\langle E \rangle = 1 \text{ GeV}$ and $\Delta E = 0.1 \text{ GeV}$

Find

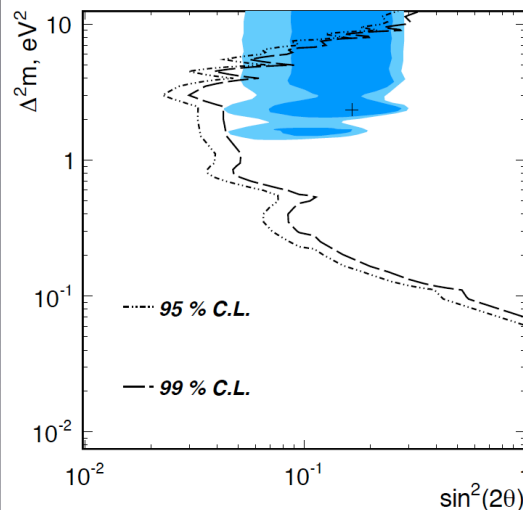
challenge

Previous

Effect of core size

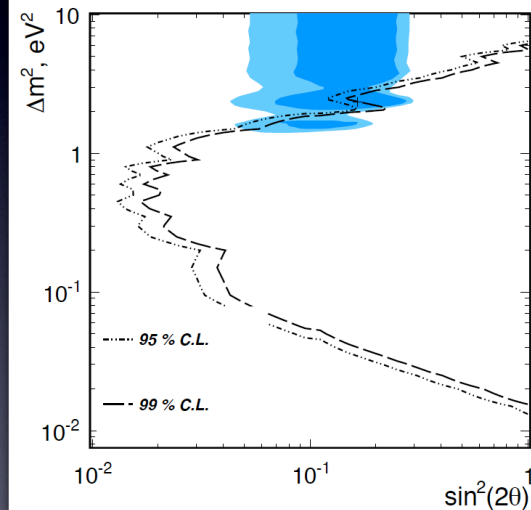
ILL core 0.4 x 0.8 m @ 6.5 m

MARS-a oscillation sensitivity: all stat=195600 eve (340 eve/day, 2 y)



Power reactor core 3 m x 3.5 m @ 12 m

MARS-a oscillation sensitivity: all stat=195600 eve (340 eve/day, 2 y)



Search for ν_s with ${}^3\text{H}$ β dec

Find

KATRIN

Previous

- Source: ${}^3_1\text{H} \rightarrow {}^3_2\text{He} + e^- + \bar{\nu}_e$

- β spectrum shape depends on:

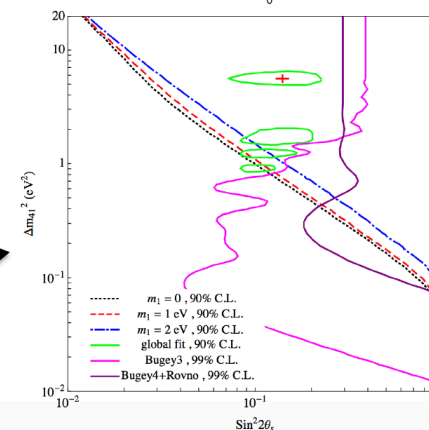
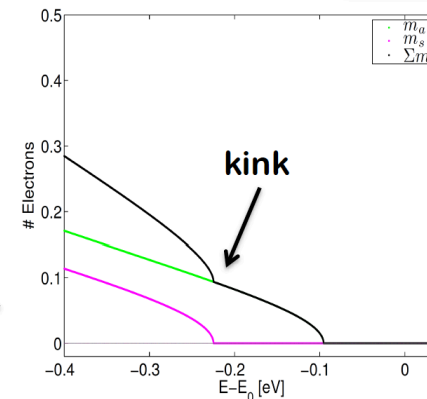
$$\langle m_\beta \rangle = \sqrt{\sum_{i=1,2,3,\dots} |U_{ei}|^2 m_i^2}$$

- Hypothetical 4th ν contribution

$$\langle m_\beta \rangle_4 = |U_{e4}| \sqrt{\Delta m_{41}^2}$$

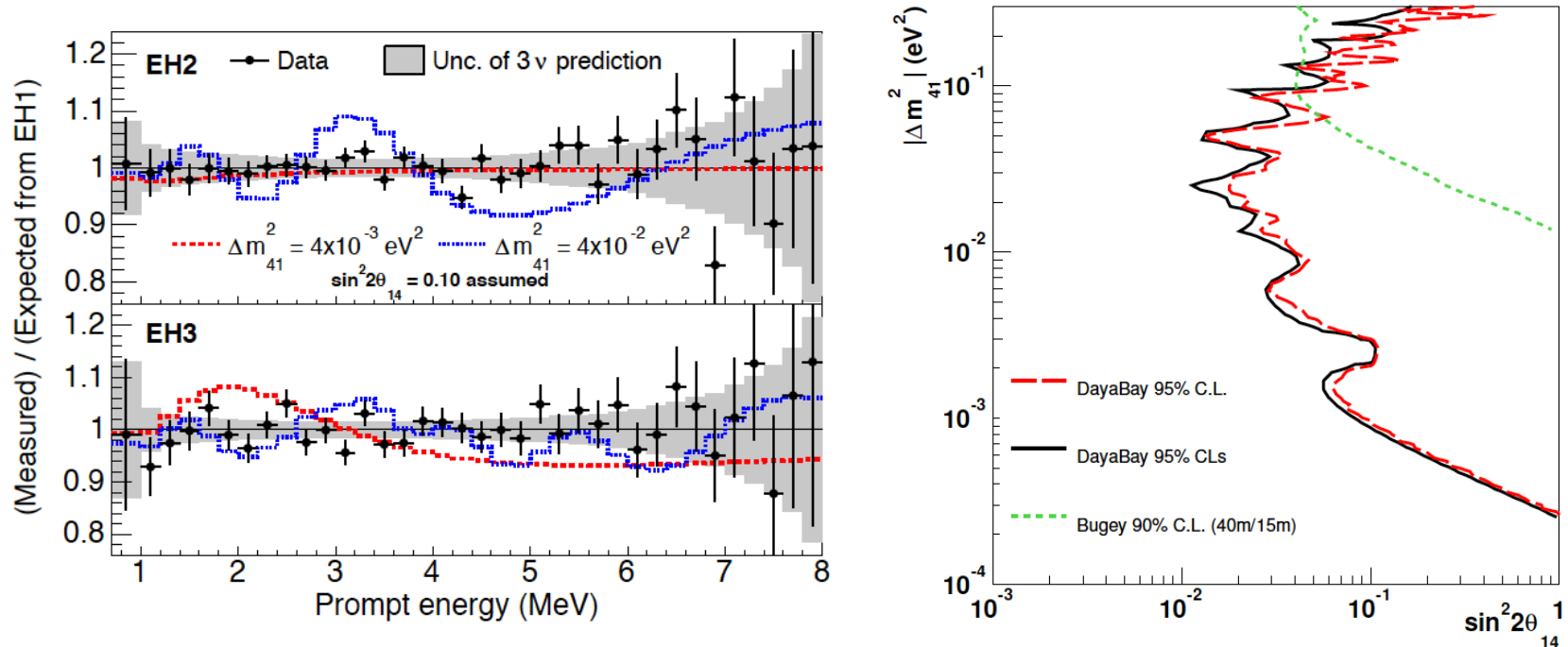
→ Search for a kink few eV below end point

- KATRIN** –as designed – can test the ν_e disappearance anomalies (sensitivity to be assessed with syst.)



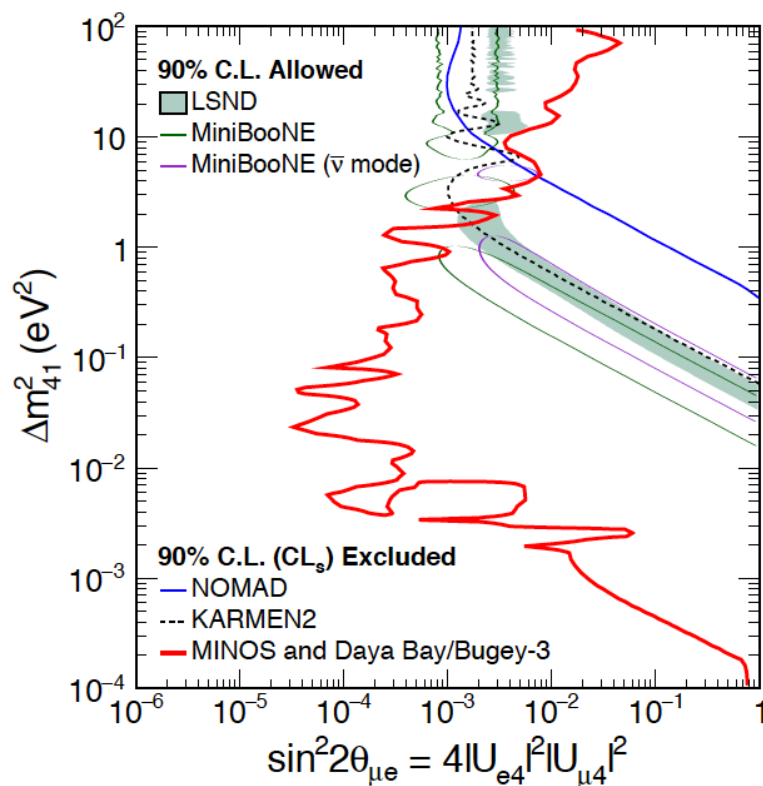
	Tech	Reactor	P [MW]	L (m)	M (tonnes)	Starting dates
Nucifer	LS+Gd	OSIRIS	70	7	0.8	ended 2015
POSEIDON	LS+Gd	PIK	100	5-8	~3	not funded
STEREO	LS+Gd	ILL	57	8.8-11.2	1.75	Aug 2016
Neutrino-4	LS+Gd	SM3	100	6-12	1.5	2014 - ended ?
PROSPECT	LS + Gd/ ⁶ Li	ORNL HFIR	85	7-18	1 & 10	awaits funding
SoLid	PVT + ⁶ LiF:ZnS	SCK•CEN BR2	45-80	5.5-11	1.44/2.88	2016
DANSS	PS + Gd	KNPP	3000	9.7-12.2	0.9	2016 ?
NEOS	PS + Gd/ ⁶ Li	Hanaro/ Younggwa	30-2800	6-?	~1	2015 at PWR
CeSOX	LS	-	N/A	5-14	20	Dec 2016

Sterile neutrino search in Daya Bay



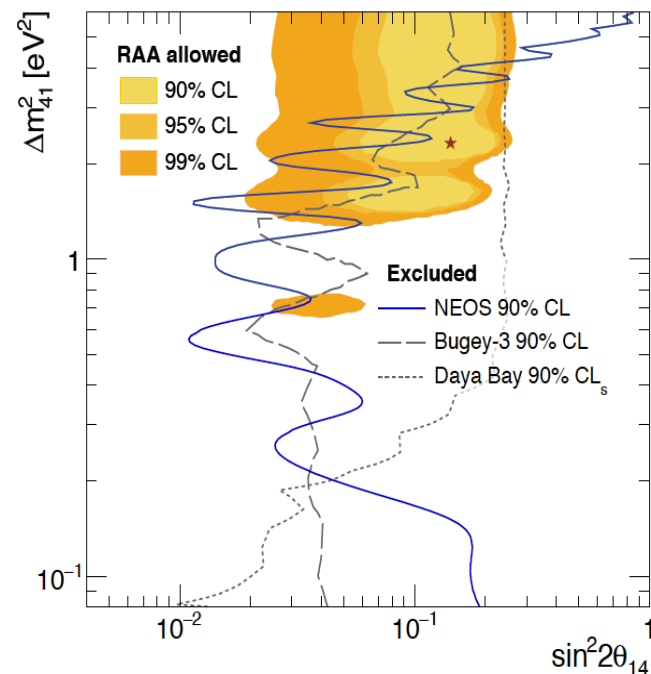
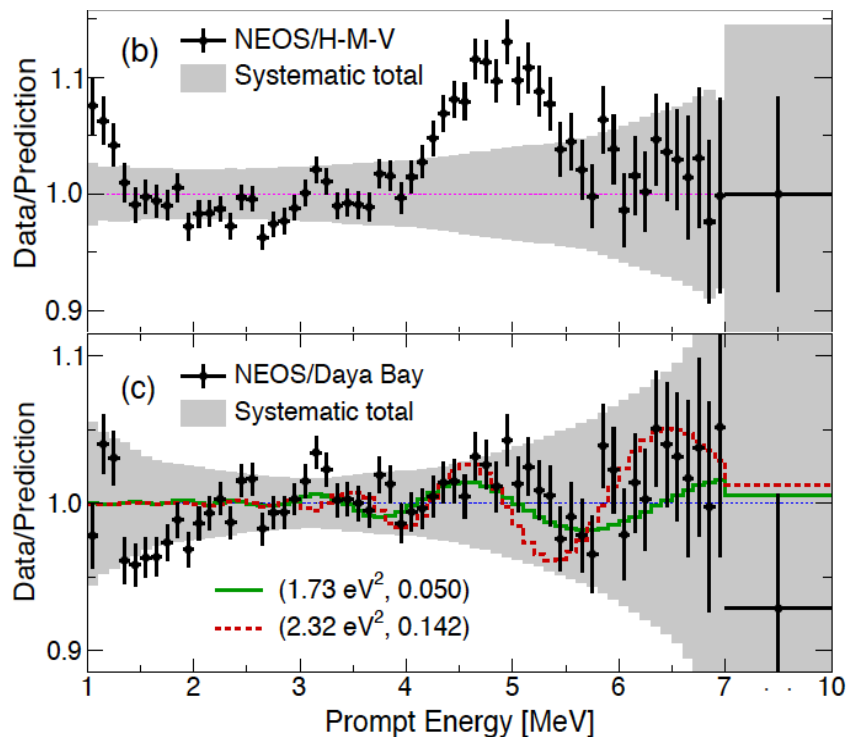
- Compromised sensitivity on the *interesting region* due to the large distance and reactor size
 - Still, important constraints provided by DB
 - Similar analyses put forward by Double Chooz and RENO

Bougey-3, Daya Bay and MINOS



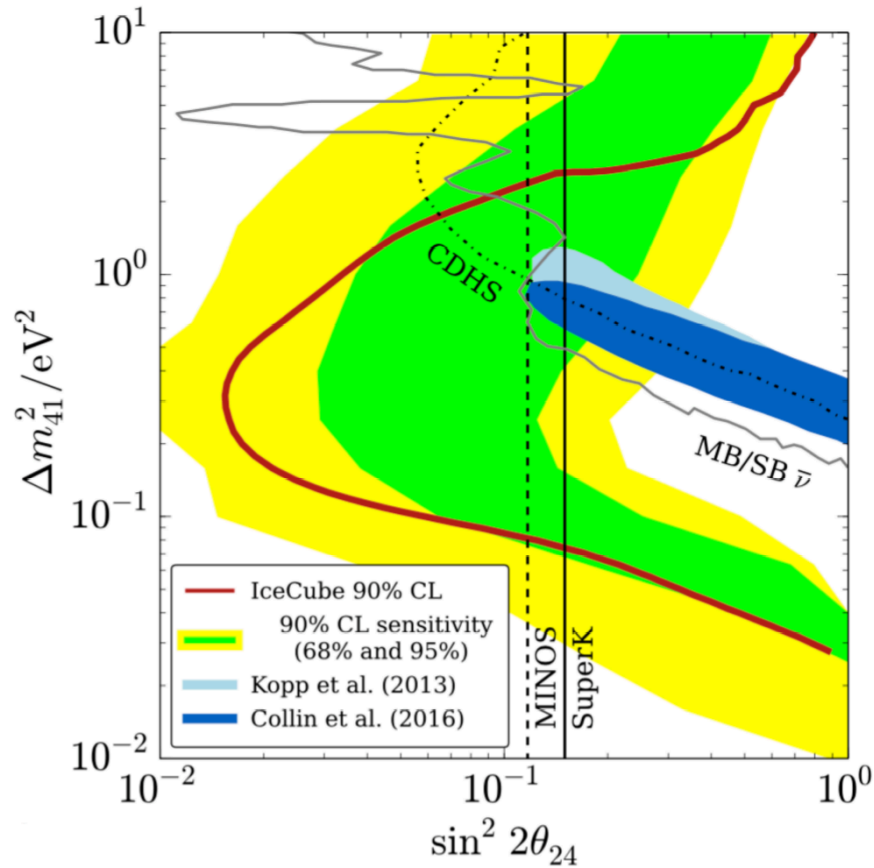
- Combination of Bougey-3, Daya Bay and MINOS
- Strong constraints imposed on the appearance channel
 - Exclusion contour covers part of the LSND/KARMEN allowed region
 - Complementary information to ICARUS/OPERA

The NEOS experiment in Korea



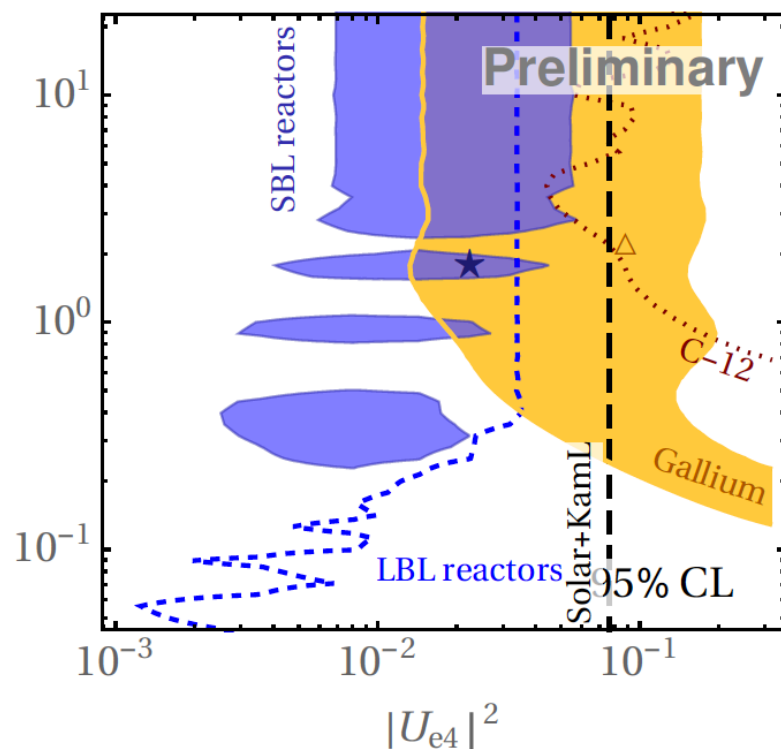
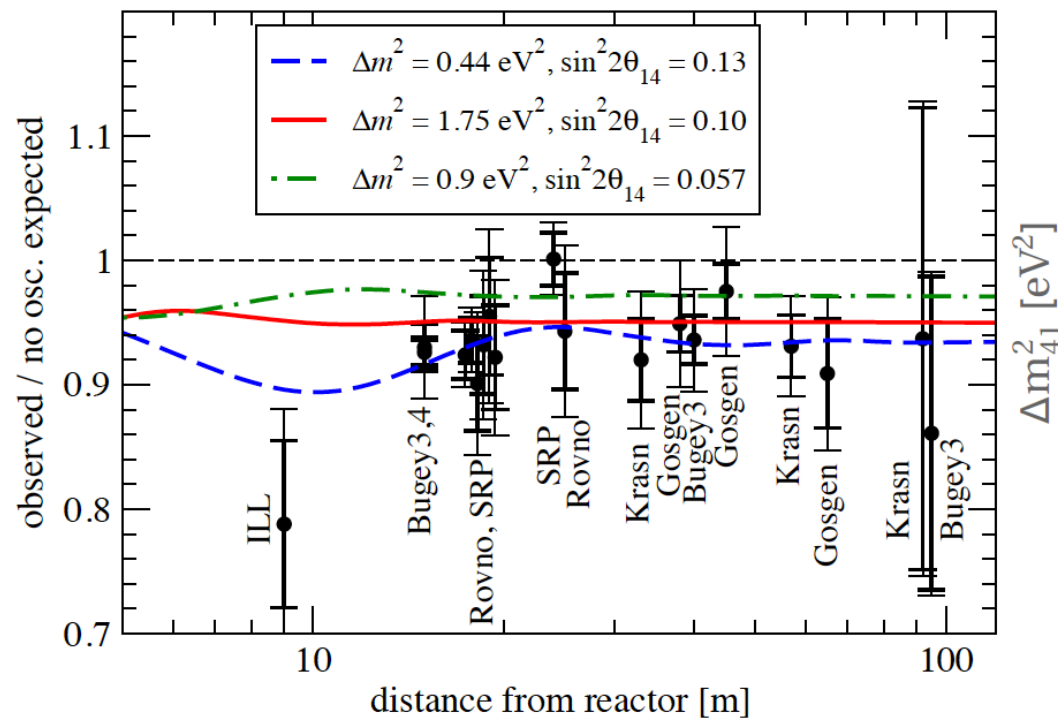
- Experiment in a power reactor
 - Confirmation of the so-called $\sim 5 \text{ MeV}$ bump seen by Daya Bay, Double Chooz and RENO
 - Lower sensitivity on the large Δm^2

IceCube



- Matter-induced oscillations produced by the *potential* existence of sterile neutrinos

$\bar{\nu}_e$ disappearance in the 3+1 scenario



	$\sin^2 2\theta_{14}$	Δm_{41}^2 [eV ²]	$\chi^2_{\min}/\text{dof}$ (GOF)	$\Delta\chi^2_{\text{no osc}}/\text{dof}$ (CL)
SBL rates only	0.13	0.44	11.5/17 (83%)	11.4/2 (99.7%)
SBL incl. Bugey3 spect.	0.10	1.75	58.3/74 (91%)	9.0/2 (98.9%)
SBL + Gallium	0.11	1.80	64.0/78 (87%)	14.0/2 (99.9%)
global ν_e disapp.	0.09	1.78	403.3/427 (79%)	12.6/2 (99.8%)

The Three LArTPC SBN Program

A Proposal for a Three Detector
Short-Baseline Neutrino Oscillation Program
in the Fermilab Booster Neutrino Beam

Submitted FNAL PAC January 2015

arXiv:1503.01520

$$\frac{\langle L_\nu \rangle}{\langle E_\nu \rangle} \sim \frac{600 \text{ m}}{700 \text{ MeV}} \sim \mathcal{O}(1 \text{ km/GeV})$$

Detector

Distance from
BNB Target

Active LAr
Mass

SBND

110 m

112 ton

MicroBooNE

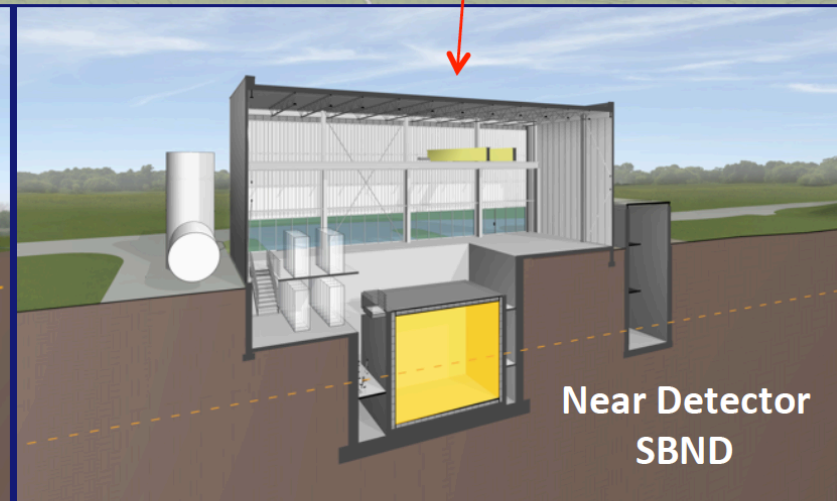
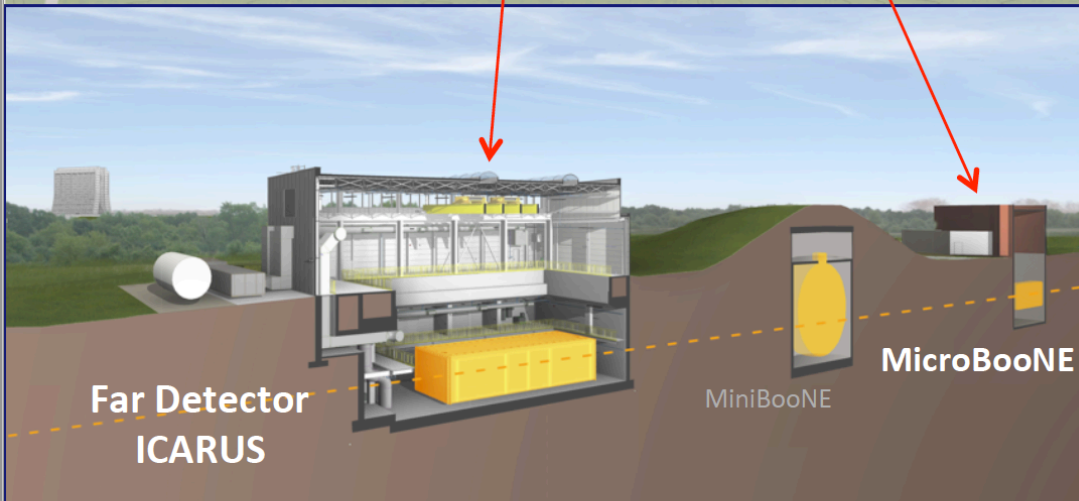
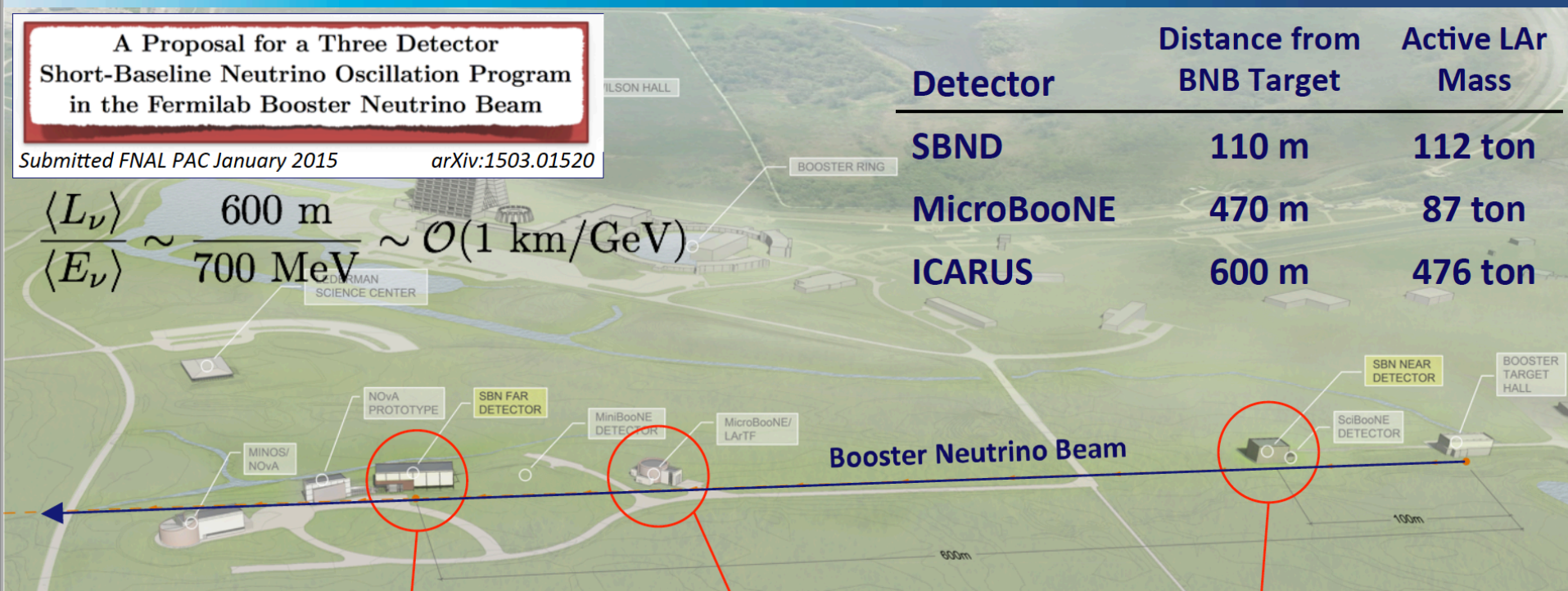
470 m

87 ton

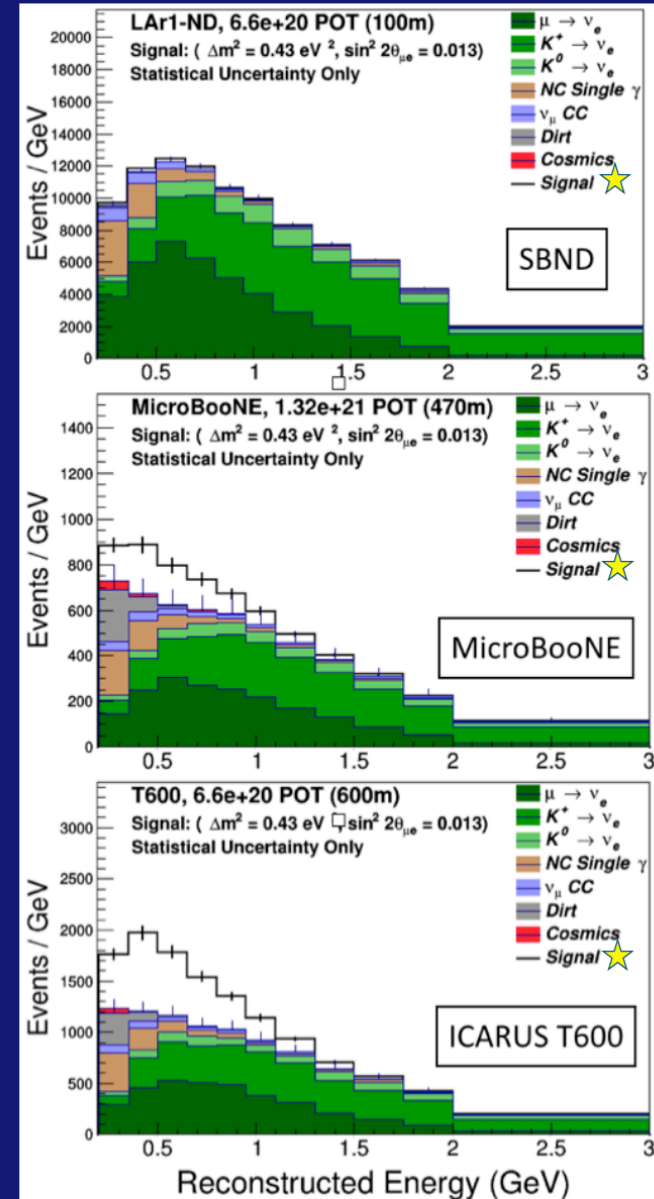
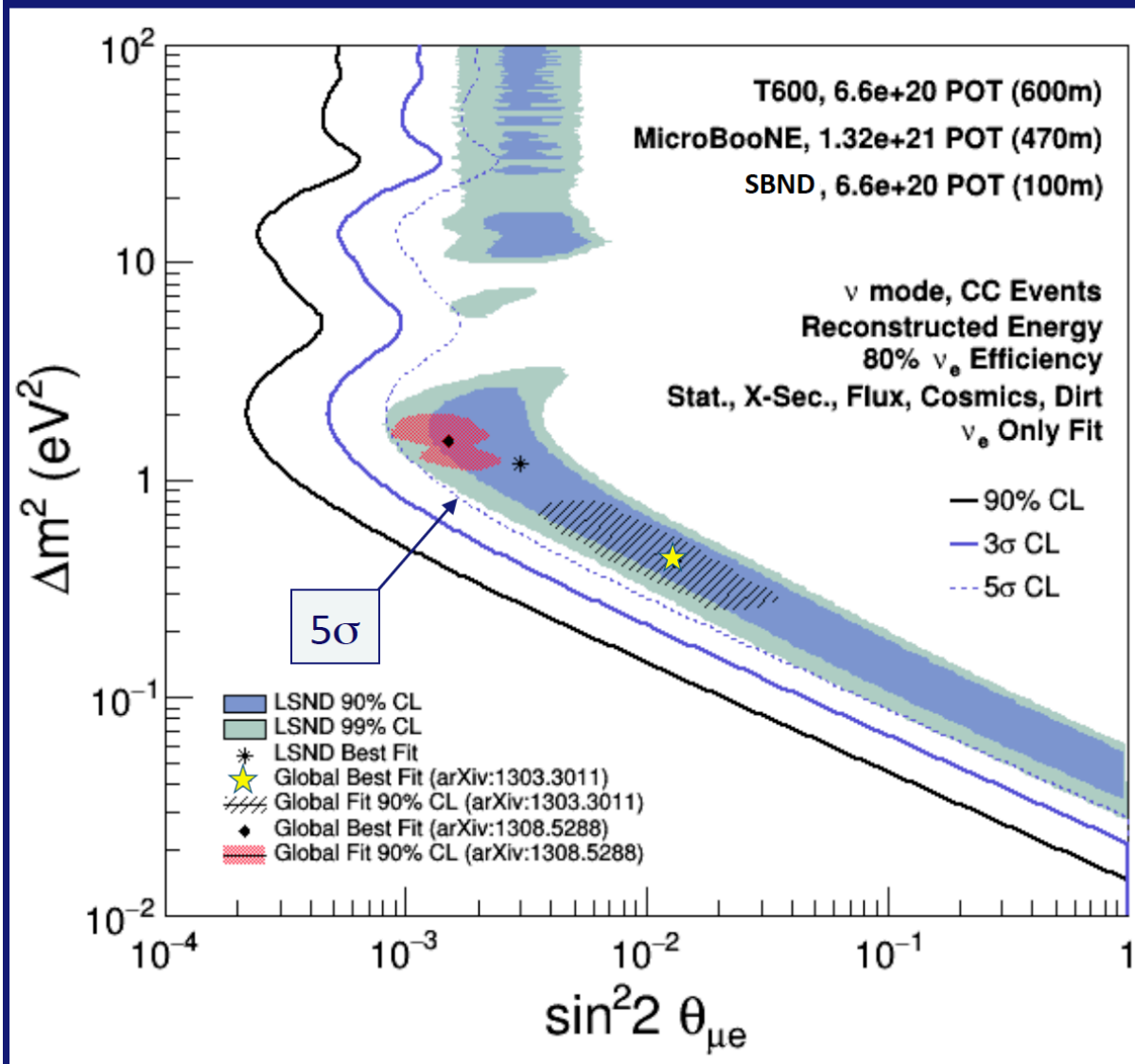
ICARUS

600 m

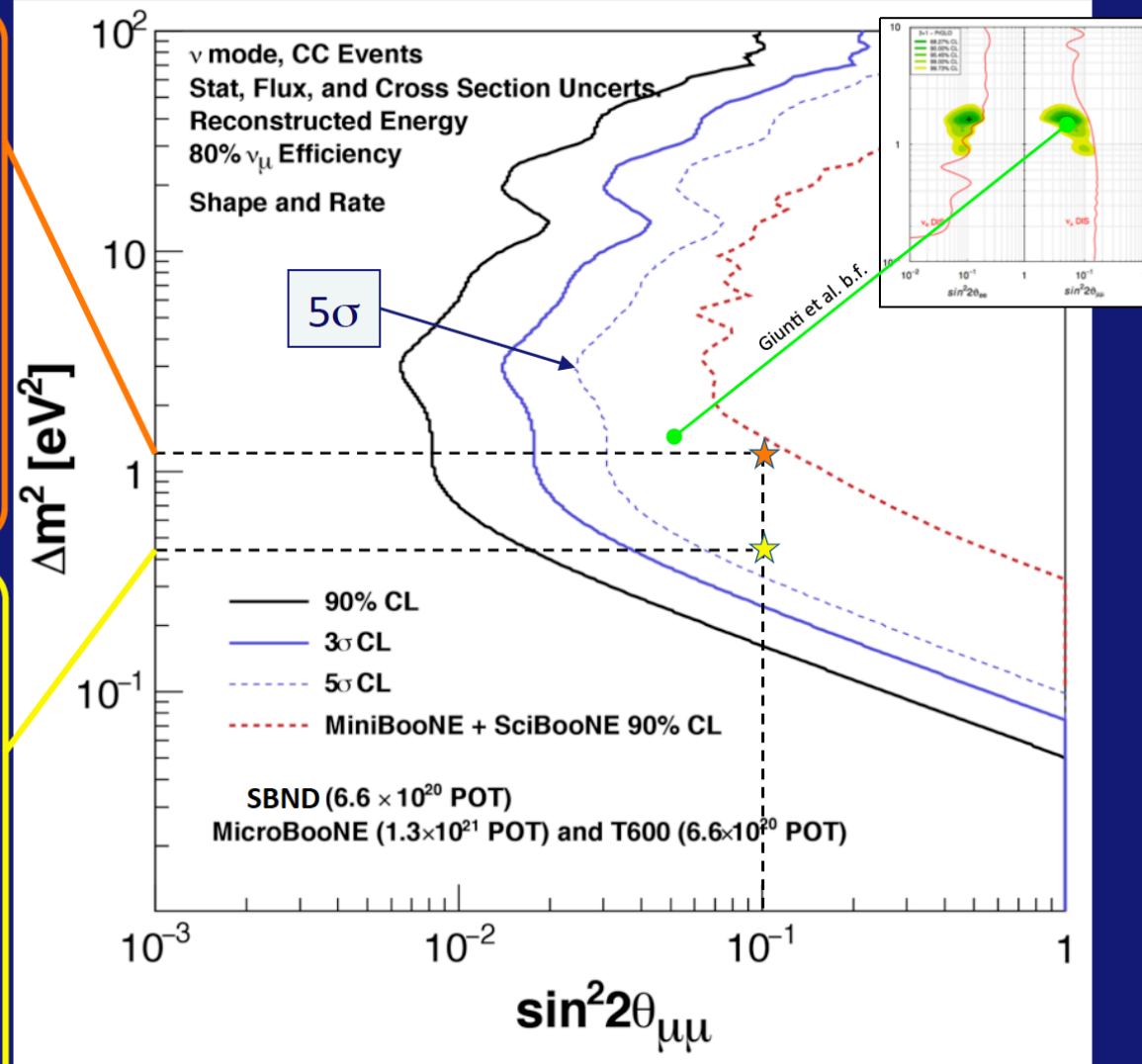
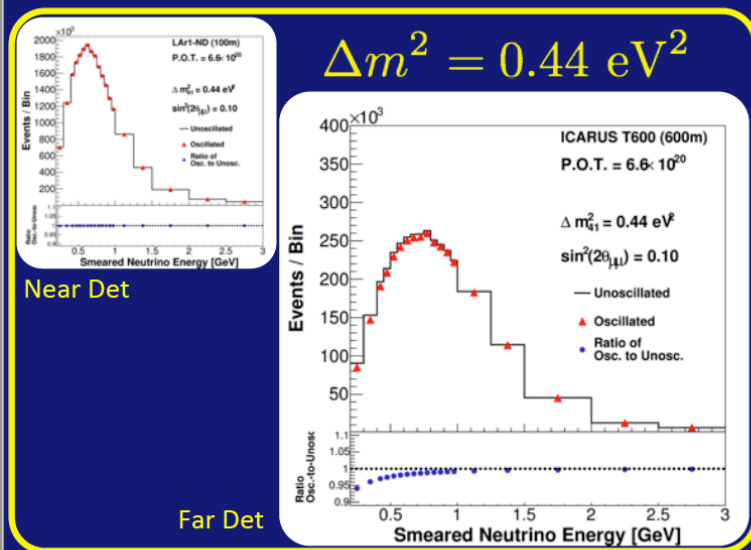
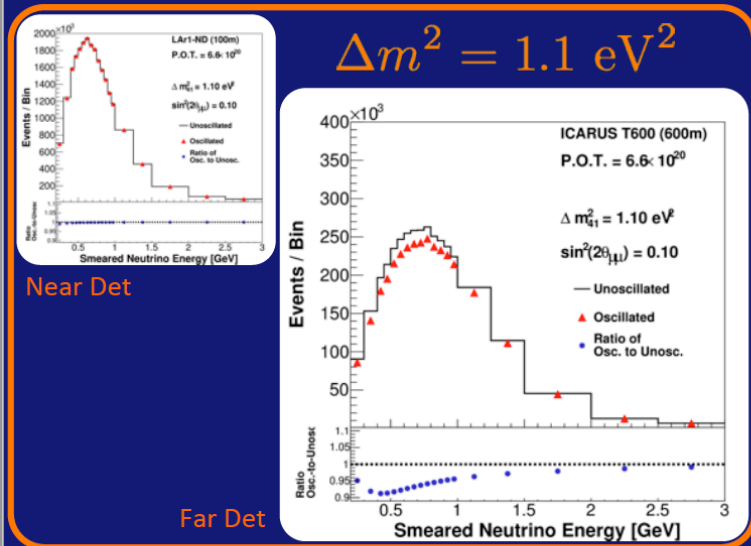
476 ton



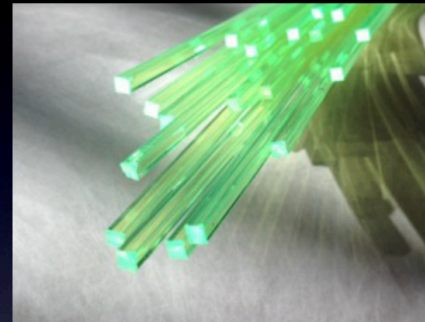
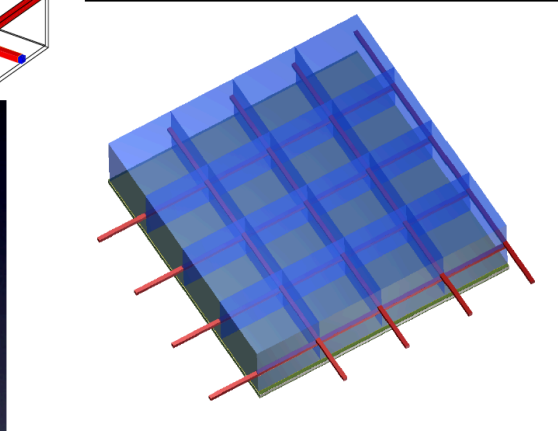
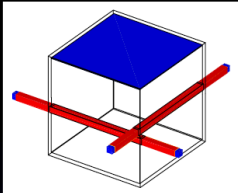
SBN $\nu_\mu \rightarrow \nu_e$ Oscillation Sensitivity



SBN $\nu_\mu \rightarrow \nu_x$ Oscillation Sensitivity

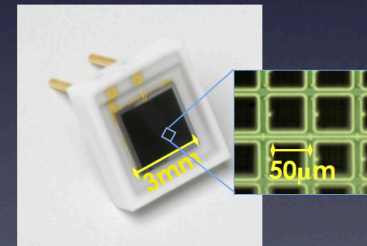


Read out



Squared BCF-91A fibre

- Both scintillator light collected by wavelength shifting fibre
- MPPC read out at both end of fibre

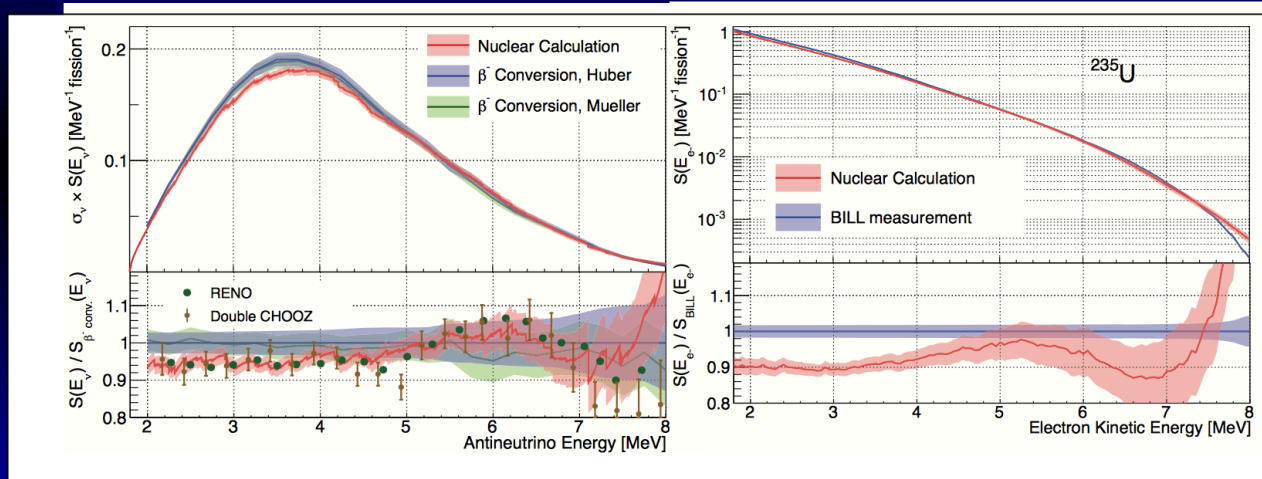


MPPC 3 mm x 3 mm
50 µm pixel pitch
60-65% active area
Pixel RC cst~13 ns
PDE ~ 30-40%

Explanations?

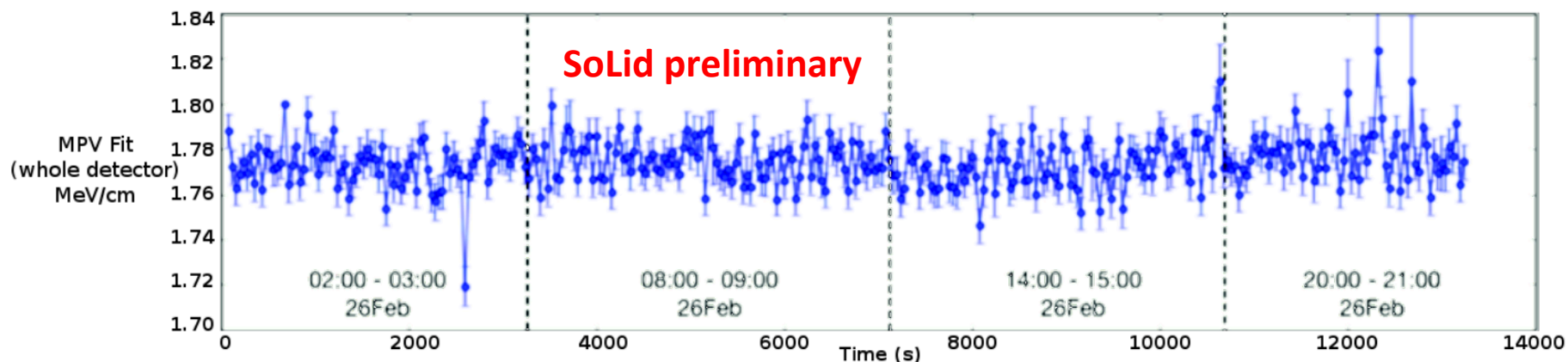
Direct summation of latest ENSDF database,
assuming allowed beta-spectrum shape

Dwyer and Langford, 2014



This direct summation, as all other direct summations,
does not agree with the Schreckenbach total
beta-spectrum.

Calibration stability



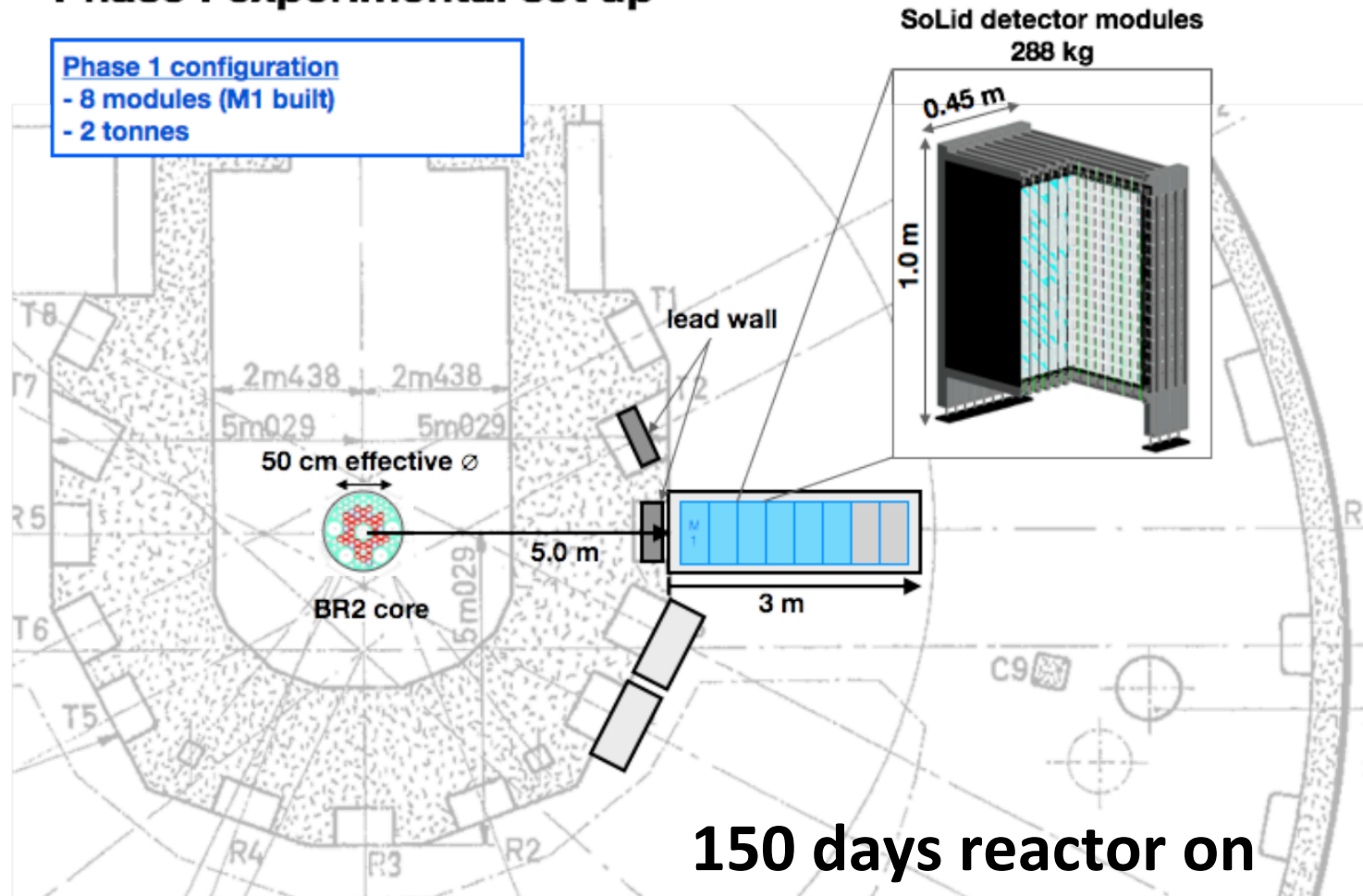
- Very good stability over time
 - A few % deviations in the energy scale only
 - Temperature well-controlled in BR2

Future plans: 2016-2017

Phase I experimental set up

Phase 1 configuration

- 8 modules (M1 built)
- 2 tonnes



150 days reactor on

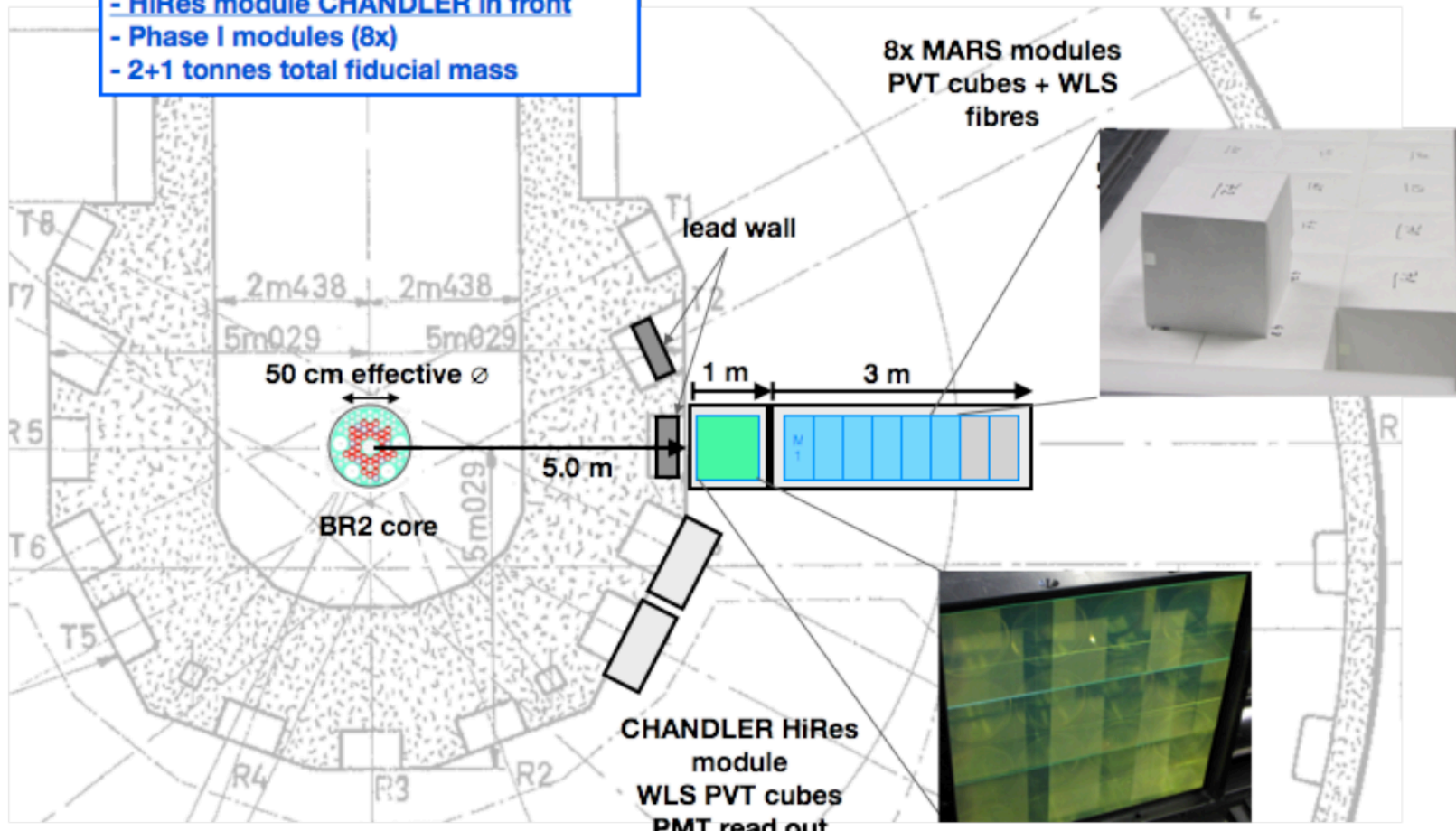
Future plans: 2017-2020

Phase II experimental set up

Configuration:

- HiRes module CHANDLER in front
- Phase I modules (8x)
- 2+1 tonnes total fiducial mass

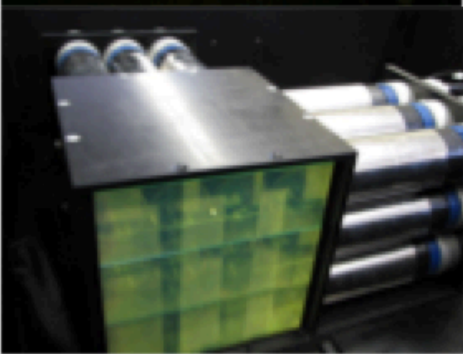
450 days reactor on



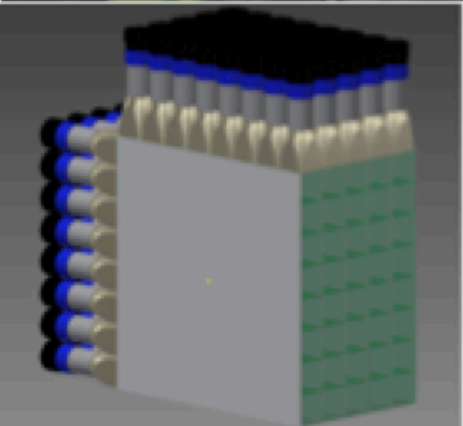
CHANDLER R&D Effort



Cube String Studies have been used to study light production, light collection, light attenuation, energy resolution and wavelength shifter concentration.



MicroCHANDLER is a $3 \times 3 \times 3$ prototype which we are using to test our full electronics chain, develop the data acquisition system, study neutron capture identification and measure background rates.



MiniCHANDLER is a **fully funded** systems test ($8 \times 8 \times 5$) which is currently under construction and will be deployed at a commercial nuclear power plant. It will be operational winter 2016.

J. Link, Aspen 2016

