## Gamma-rays at ELI-NP-GBS

#### GENERATION AND MONITORING

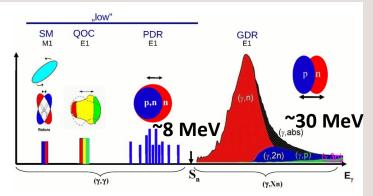


LIA-COSMA WORKSHOP - BUCHAREST (ROMANIA) - K. DUPRAZ

#### Gamma-ray requirements

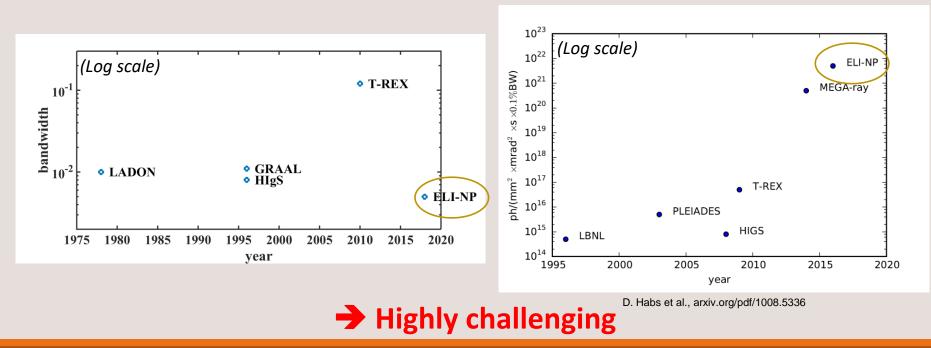
#### Gamma-beam specifications:

- Energies  $\gamma$  (E<sub>v</sub>) : 0.2 19.5 MeV
- Bandwidth (Δ́E/E) : <0.5%
- Spectral density (TASD) : >5000 γ/(s.eV)
- Linear polarization: >95%

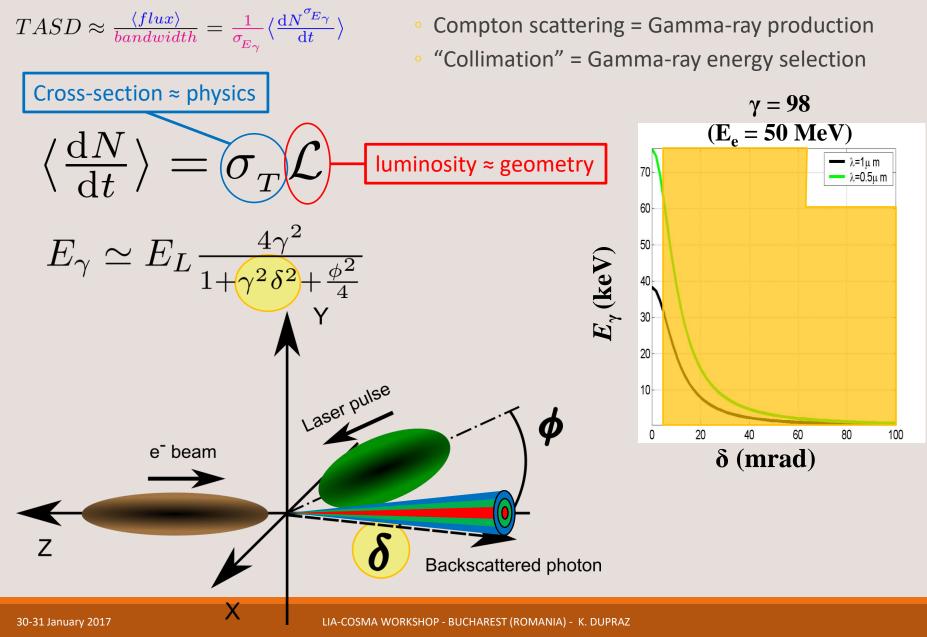


AIP Conference Proceedings 1462, 177 (2012)

#### Bandwidth + energy tunability -> Compton scattering



## Production of the gamma-beam

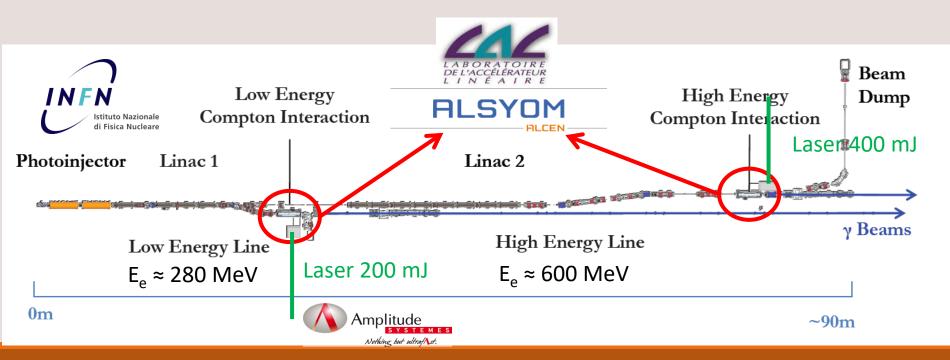


## ELI-NP gamma beam source (accelerator)

Requirements:

- 2 interaction points (Low/High Energies)
- Accelerator hall size fixed
- → compact accelerator design (hybrid band S and C = SPARC, Frascati)

= State of the Art, low development risk



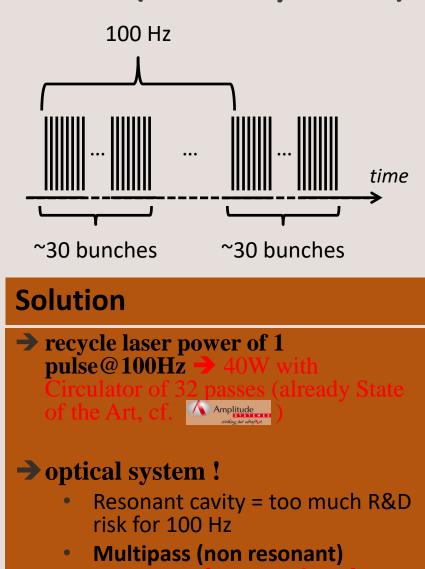
## ELI-NP gamma beam source (laser system)

#### LASER REQUIRED FOR COMPTON INTERACTION POINT

- Accelerator characteristics constrained by electrons beam quality  $\rightarrow$ 
  - Laser average power > 1 kW
  - green light,
  - high quality,
  - pulsed mode > 100Hz @few ps

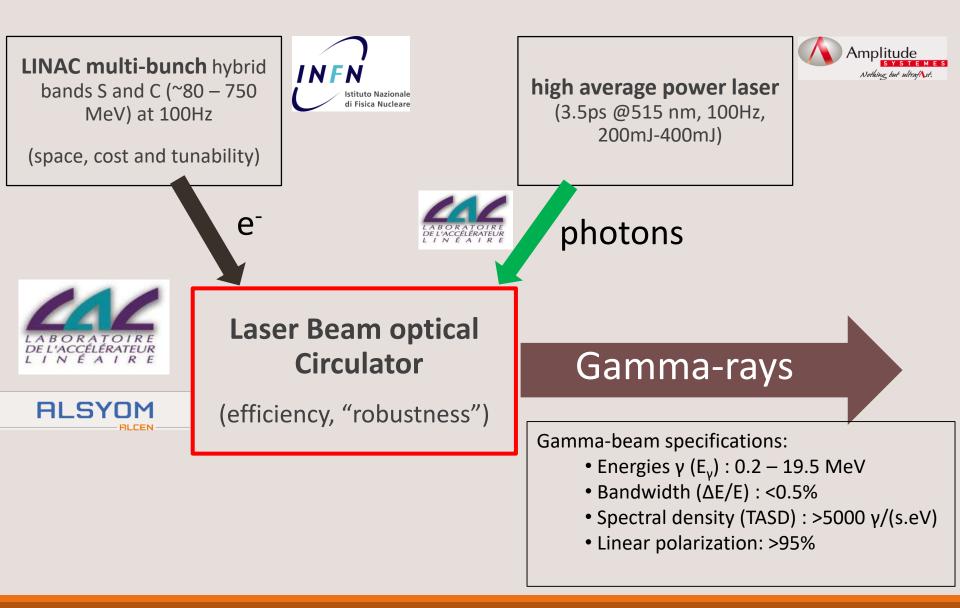
#### Better than state of the art !

= unreachable even with R&D



Laser beam circulator

## Overview design



## Optical system: laser beam circulator

 $\Gamma_i D_{MPS}$ 

z

 $M_2$ 

 $2\omega_M$ 

фм

#### Requirements:

- Constant crossing angle φ (small bandwidth)
- Unique laser-electron beam interaction point

L A B O R A T O I R E DE L'ACCÉLÉRATEUR

mplitude

Nothing but ultrafist

➔ 1 Mirror-Pair System (MPS) per pass = Optical plan switching

- No optical aberration
  - → 2 high-grade quality parabolic mirrors

x

3D

ParisTech

 $(M_2)$ 

boratoire Charles Fa

<u>"Dragon-Shape"</u>

e

 $(M_1)$ 

**2D** 

Conceller Conceller

e

 $M_1$ 

### **Recirculator constraints**

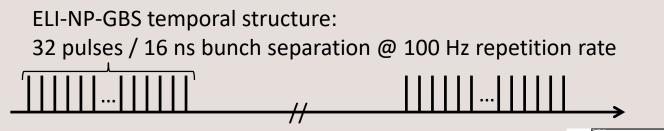
- Mirror surface quality
- Frozen geometry (parabolic mirrors distance)
  => Tight alignment (few μm, μrad) with 7 degrees of freedom
- 30 cm  $(M_2)$ D = 2.4 m' $\Gamma_i D_{MPS}$  $2\omega_M$ 3 cm  $\phi_M$

- MPS parallelism (< 3 µrad)
- Synchronization (few 100fs)

## Online Monitoring

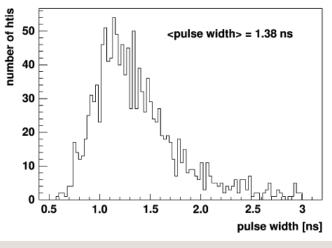
A. Martens - Nuclear Photonics (Monterey, USA)

## **Diamond sensors**

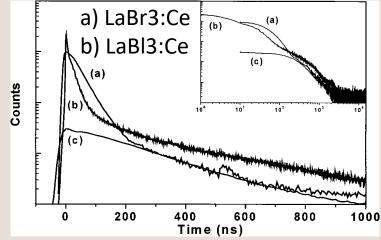


Solution with scintillating material not convenient

- → Slow component polution
- ➔ Difficult to insert in ELI-NP-GBS accelerator:
- (relatively) large volume occupied by the calorimeter
- high-vacuum implementation



Kölbl et al., IEEE Trans. Nucl. Science 6 (2004)

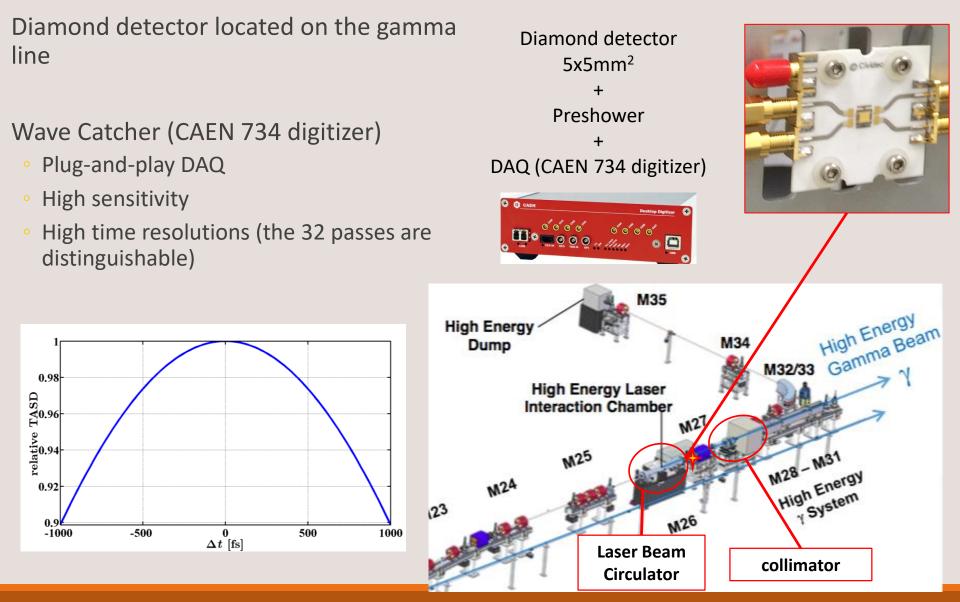


Van Loef et. Al, Appl. Phys. Let. 10 (2001)

A compact (easy to integrate), fast (few ns pulse width), radiation hard sensor is needed

- ➔ Diamond sensors
  - O(1ns) pulse width
    - Diamond sensors are naturally radiation hard (much more than scintillating materials)

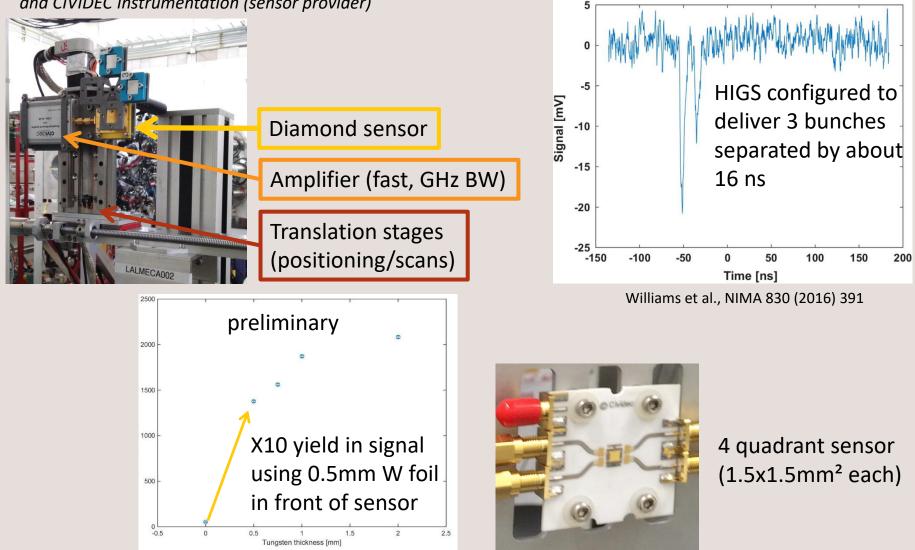
## **Online Synchronization**



A. Martens - Nuclear Photonics (Monterey, USA)

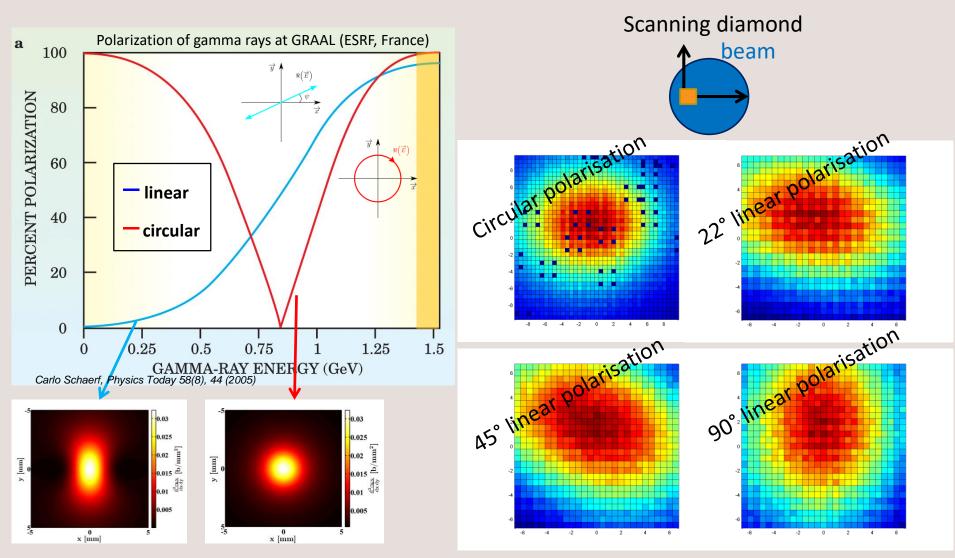
#### First experiments @ HIGS & newSUBARU

Thanks to HIGS/TUNL colleagues and CIVIDEC Instrumentation (sensor provider)



#### A. Martens - Nuclear Photonics (Monterey, USA)

#### Gamma-beam imaging



Gamma ray beam profile distribution **→** Gamma ray beam polarisation

(for 100% polarized beam)

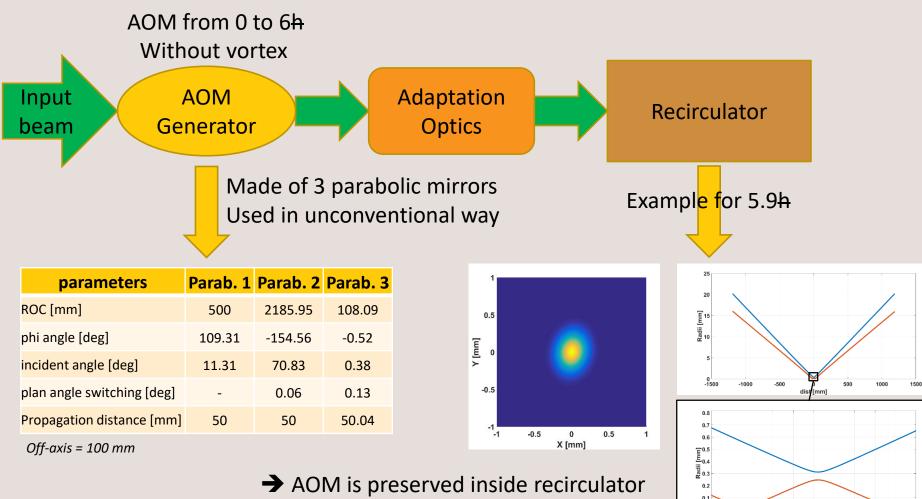
# Prototype and Upgrade

#### Prototype design



30-31 January 2017

#### Possible upgrade: Angular Orbital Momentum



➔ Down to 15% of TASD @6h

25

5

-5 0 dist[mn 10 15 20

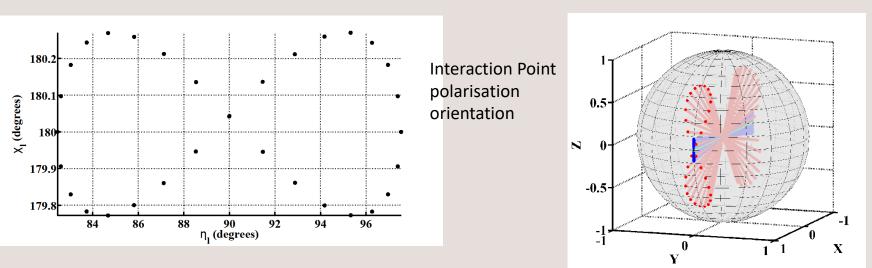
-25 -20 -15 -10

### Summary

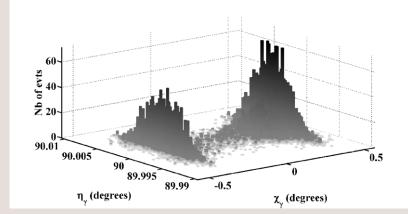
- Highly challenging machine to produced:
  - TASD: 5000 γ/(s.eV)
  - Bandwidth: < 0.5%
  - Degree of Polarization: > 95%
  - Polarization preserved
  - AOM preserved
- Online monitoring with diamond detector for:
  - Electrons-Laser beams synchronization
  - Gamma-rays flux
  - Polarization

## Thank you for your attention

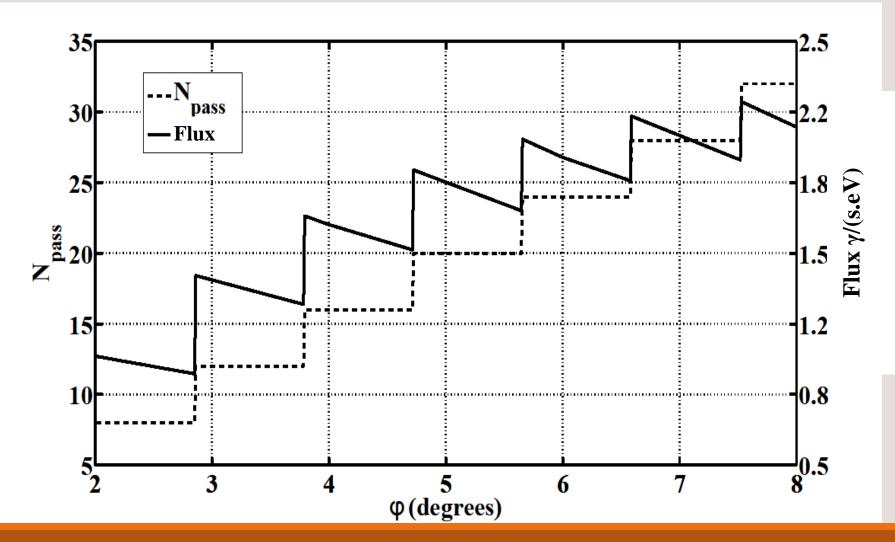
### Polarisation



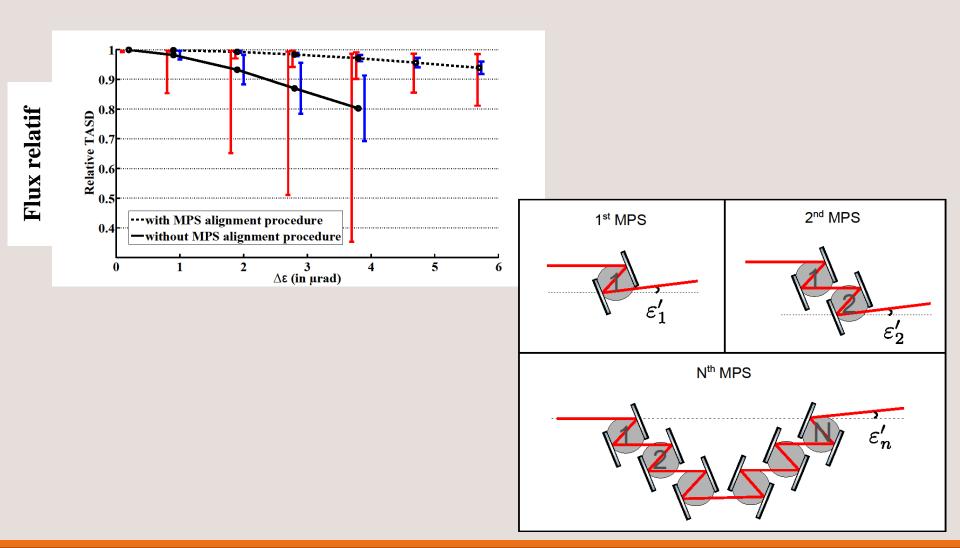
- Simulation with multilayer coatings and coating birefringence
- Polarization preserved during circulation (>99%)
  - Linear
  - Circular



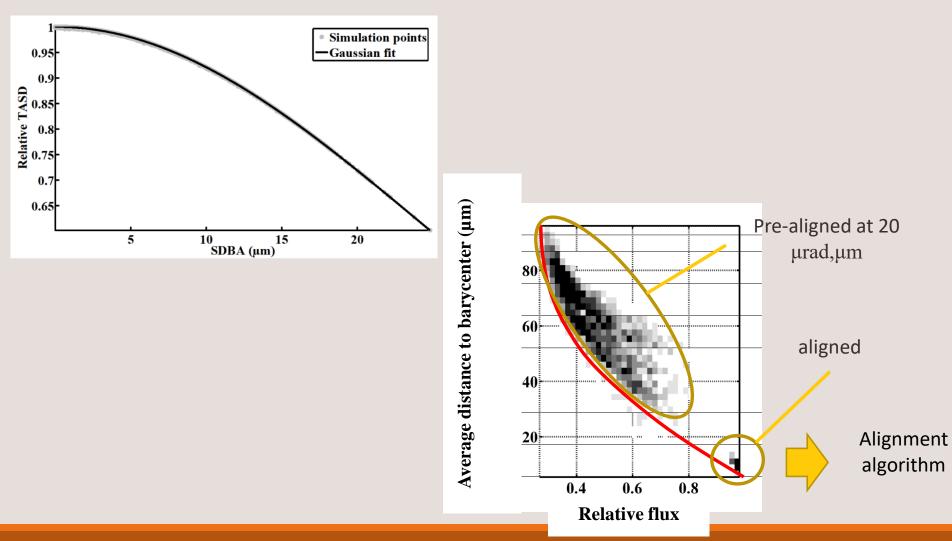
#### **Optimization No. passes**

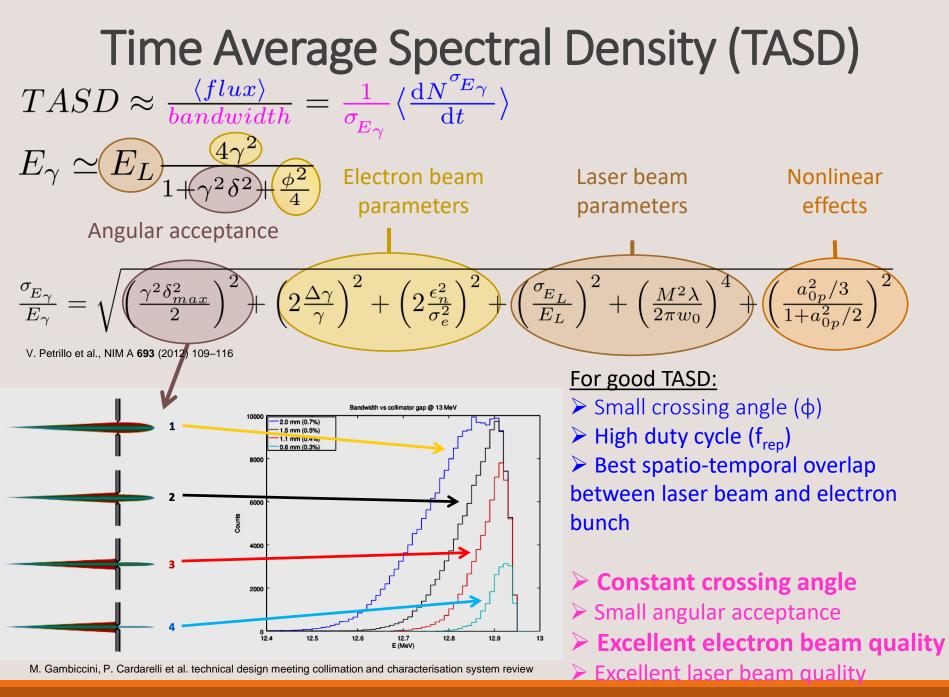


## MPS parallelism



## Alignment





## ELI-NP gamma beam source (laser system)

ACCELERATOR CHARACTERISTICS

CONSTRAINED BY ELECTRONS BEAM QUALITY

Train at 100 Hz

Few bunches = ~30 bunches max (photoinjector limit)

250 pC (low charge)

~750 MeV max

LASER REQUIRED FOR COMPTON INTERACTION POINT

Laser average power > 1 kW green light,00 Hz

#### **Solution**

#### → optical system !

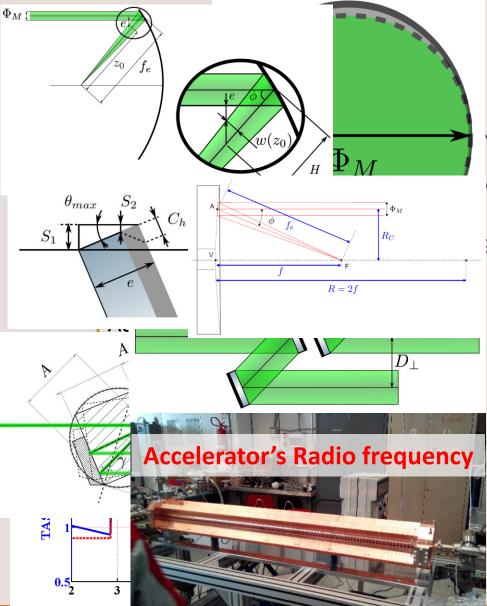
- Resonant cavity = too much R&D risk for 100 Hz
- Multipass (non resonant)
  Laser beam circulator

### **Optimization constraints**

Synchronization (< few 100fs) for each pass

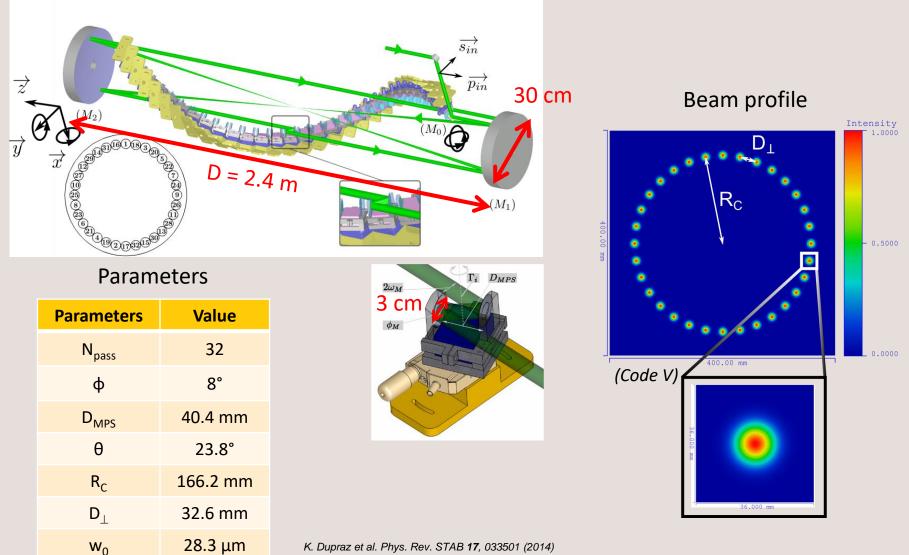
- → Mirror-Pair System (MPS) rotation
- Mechanical tolerances
- Laser damage threshold (laser beam size)
- Vignetting of the beam
- Accelerator's Radio frequency
- Beam focalization
- Mirror shape
- Parabolic mirror parameters
- Circulator length (accelerator hall limit)
- Maximum number of passes (photoinjecto limit)
- TASD (Time Average Spectral Density)

K. Dupraz et al. Phys. Rev. STAB 17, 033501 (2014)



#### **Geometry parameters**

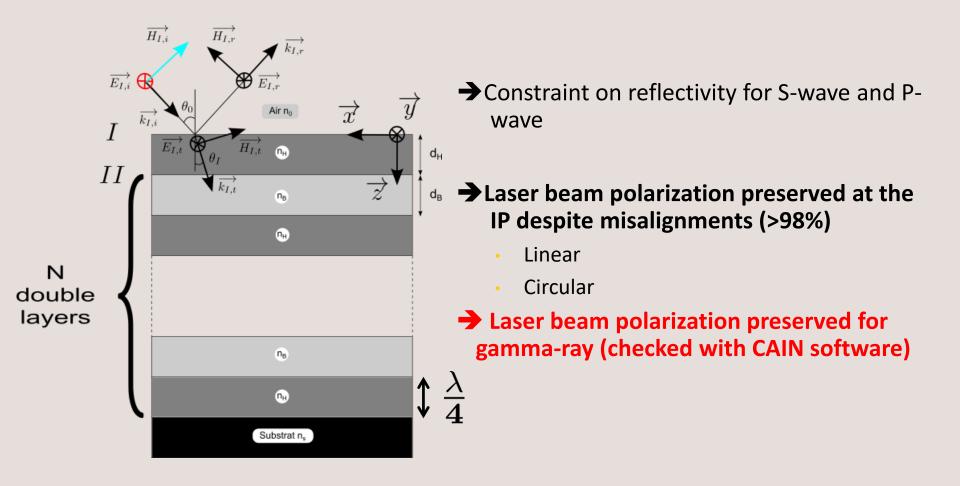
#### Geometry



K. Dupraz et al. Phys. Rev. STAB 17, 033501 (2014)

## **Coatings and Polarization**

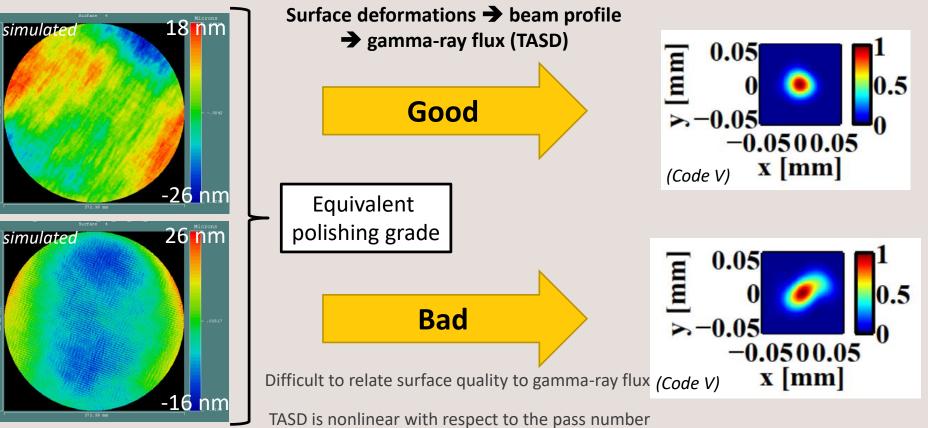
Simulation with multilayer coatings (interferential coatings)



## Optical quality and beam profile

TASD = Time Average Spectral Density MPS = Mirror-Pair System

#### SURFACE DEFORMATIONS

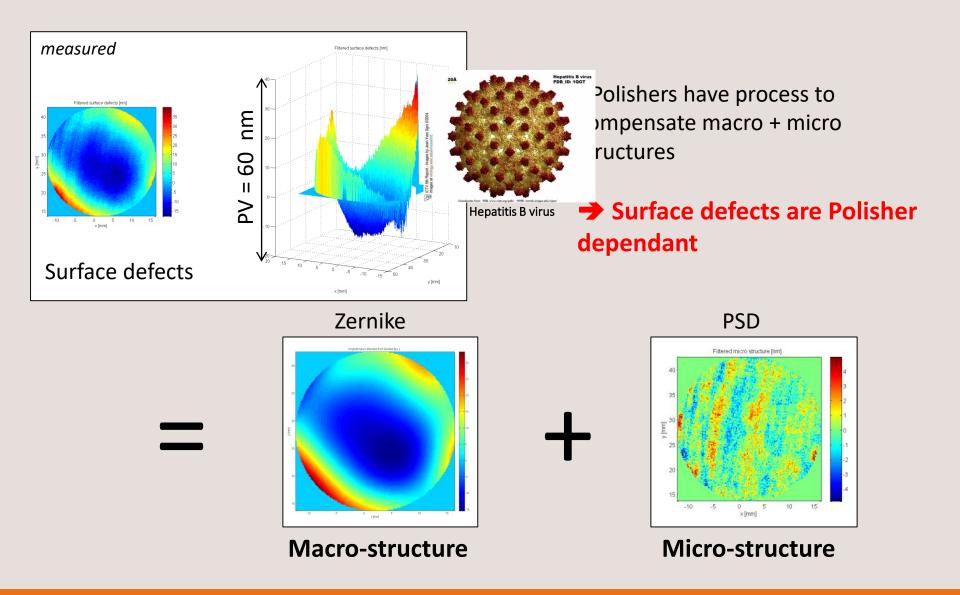


→ the 32 passes have to be simulated

## Parabolic mirrors deformations < λ/80 RMS</li> MPS mirrors < few nm of residual focus</li>

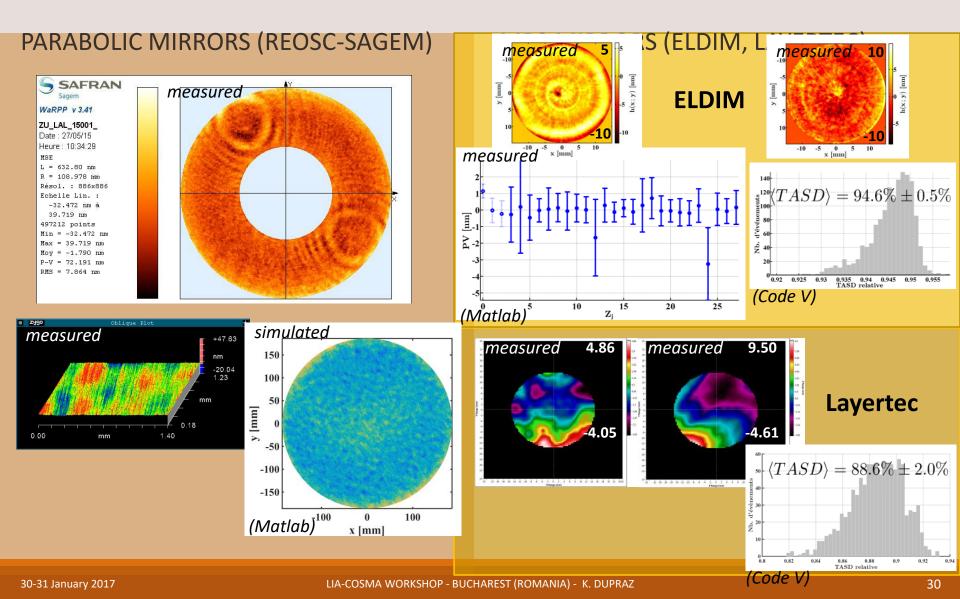
IP BEAM PROFILE

#### Surface defects representation



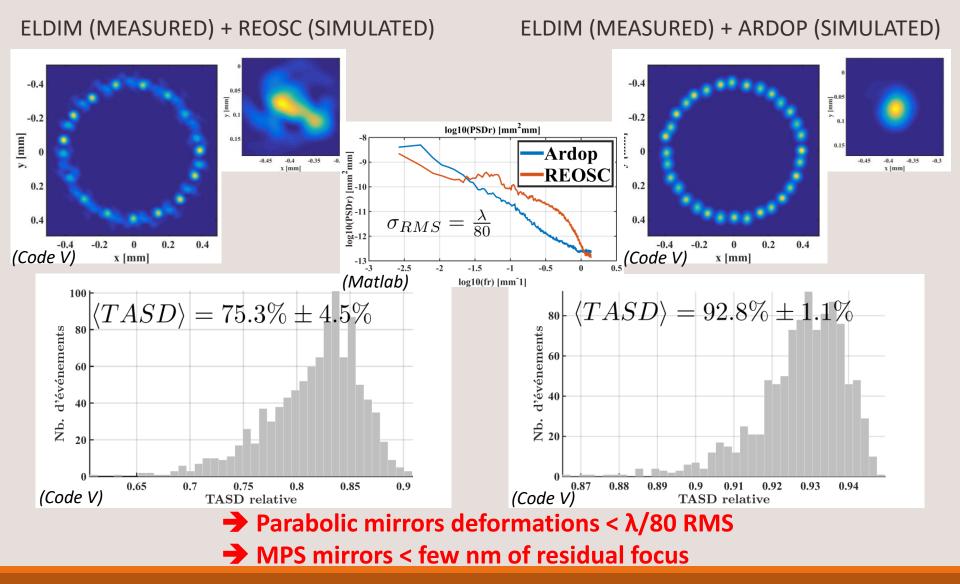
## **Real mirrors**<sub>TASD = Time Average Spectral Density</sub>

MPS = Mirror-Pair System

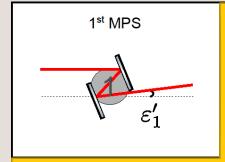


### Results with the real mirrors

TASD = Time Average Spectral Density



#### MPS parallelism

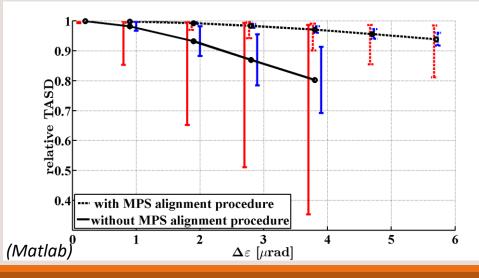


$$\Delta \varepsilon' = \sqrt{\sum_{i=1}^{N} \left(\varepsilon'_{i}\right)^{2}}$$

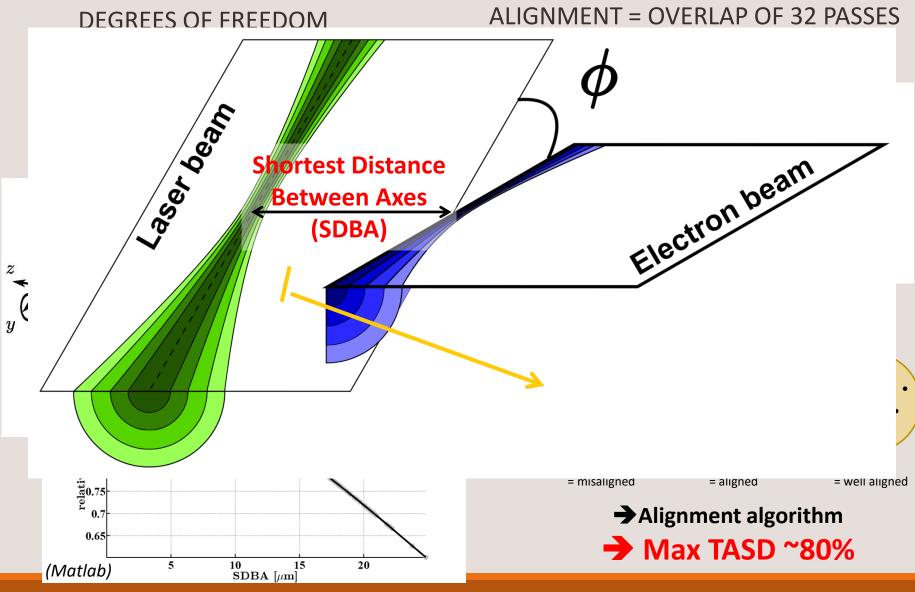
#### Big accumulation effect → dedicated

parallelism alignment procedure:

Each stage beam angular deflection <  $\epsilon'$ 



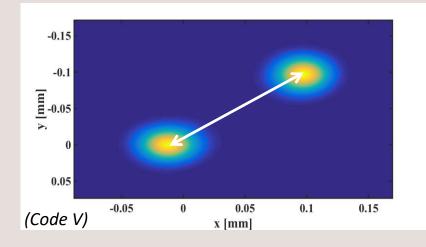
## Alignment

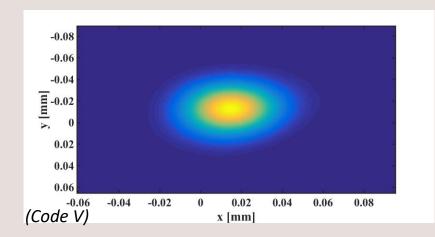


#### Observables

#### **ON-IP** (ELDIM (MEASURED) + ARDOP (SIMULATED))

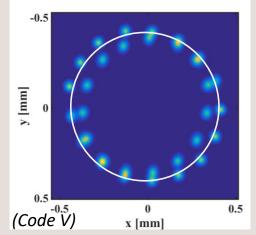
#### Superimpose spots together



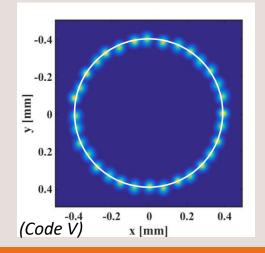


#### OFF-IP (ELDIM (MEASURED) + ARDOP (SIMULATED))

#### Superimpose spots on a circle

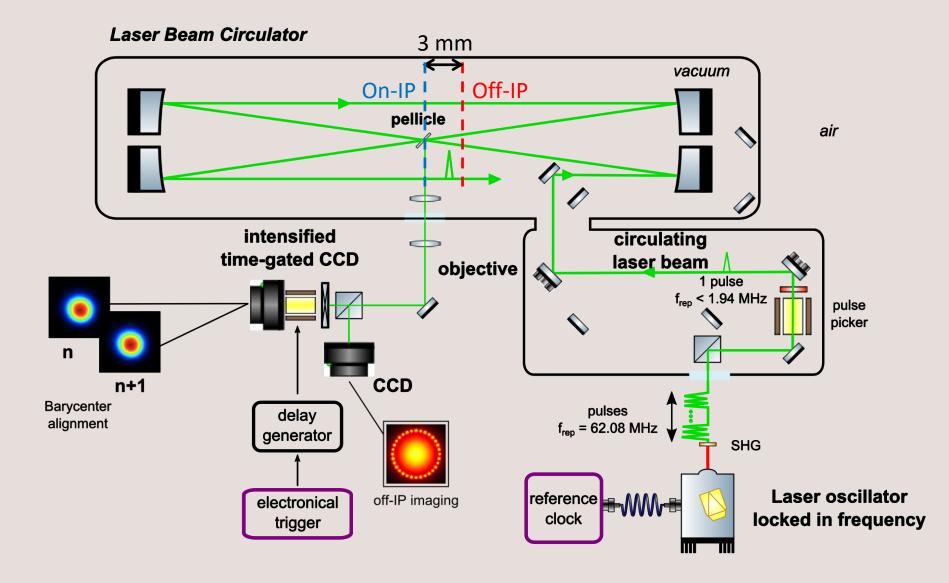


#### Before alignment



#### After alignment

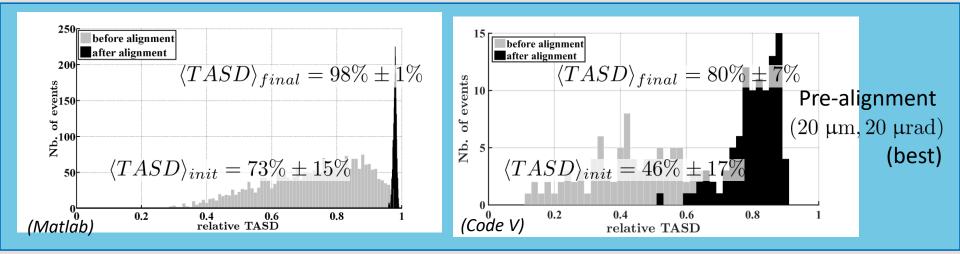
## Alignment tool

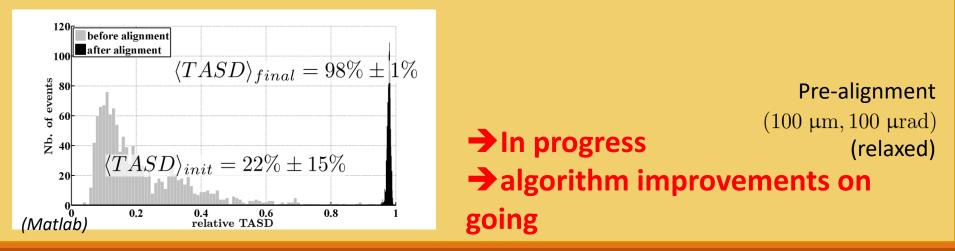


#### Results

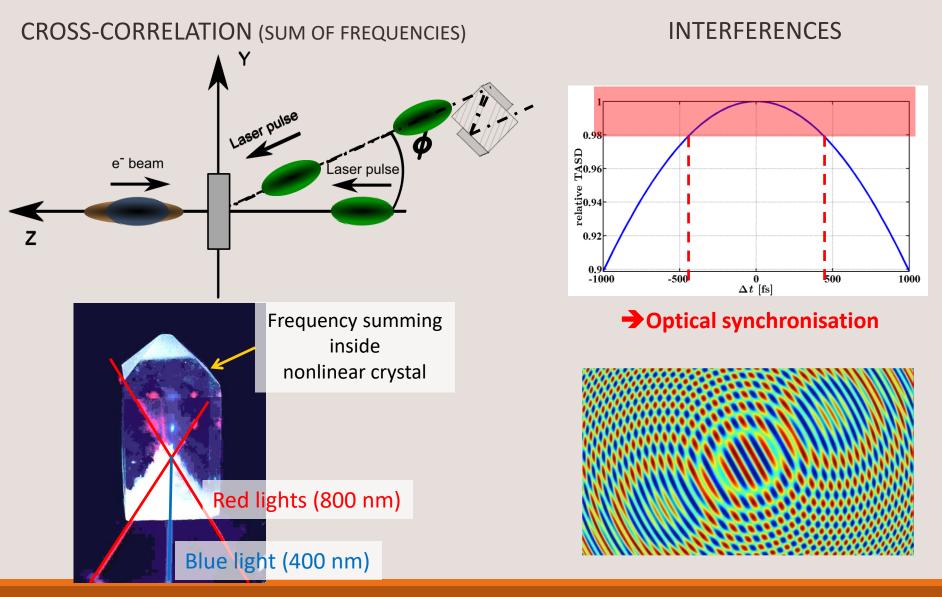
#### WITHOUT MIRROR SURFACE DEFECTS

#### WITH MIRROR SURFACE DEFECTS (ELDIM + ARDOP)

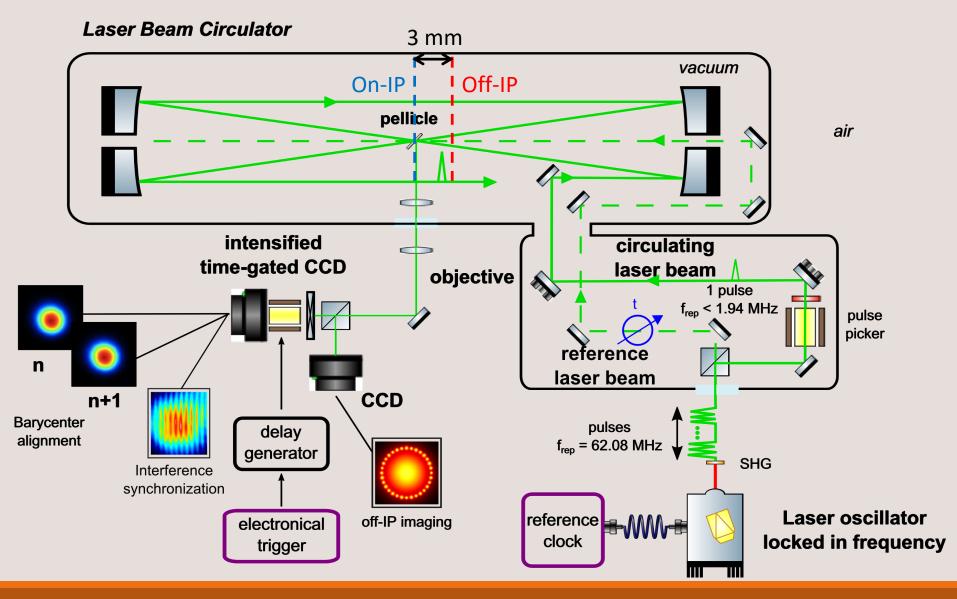




#### Synchronization

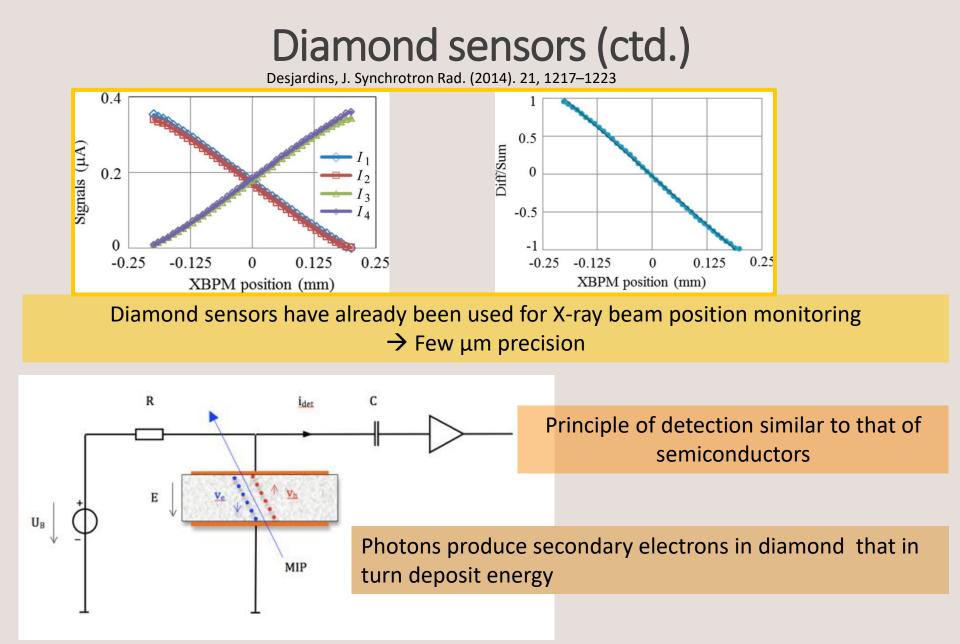


## Synchronization tool



## Synchronization (Proof of principle)

Intensity of the second Harmonic Synchronization max at: 9790µm CCD unit data < 500fs fitted curve cross-correlation intentisty S.H. 6.5 delay [fs] 5.5 7.5 8.5  $x 10^4$ Synchronization OK ! 10000 11000 12000 7000 8000 9000 Mira (800nm) Delay line position [µm] → Could be improved ~75Mhz (mode ps) S.H.G crystal interferences µ-objective pellicle 7.5° Comparison between CCD Contrast intensity max at: 9822µm <u>الم</u> 2 algorithms vers la cavité data 8 (arbitrary delay line 1 fitted curve téléscope algo 1 intentisty algo 2 f(τ) [arbitrary unit] M1 Contrast 1.182.874 5.5 6.5 delay [fs] x 10 lentille 1m M2 < 10000 12000 7000 8000 9000 11000 50 200 250 [arbitrary unit] 150 300 350 Delay line position [µm]



#### **Energy response/Detection efficiency**

#### Simulated (GEANT4) efficiency about few per mille

