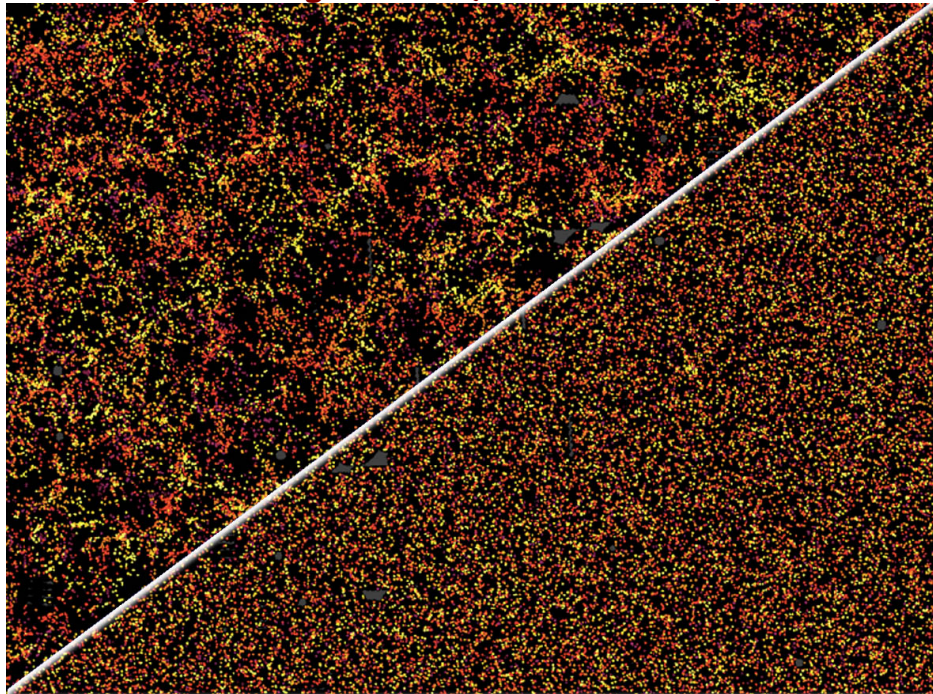


Cosmology with eBOSS and DESI

N. Palanque-Delabrouille, CEA-Saclay (IRFU)

500 deg² BOSS galaxies (0.50 < z < 0.55)



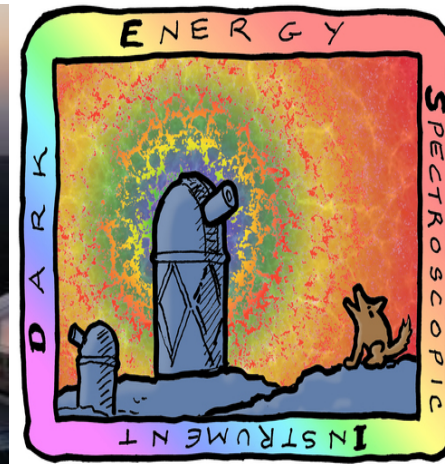
500 deg² random (0.50 < z < 0.55)

BOSS, eBOSS, DESI:
a clustering saga

- BAO (dark energy)
- RSD (gravity)
- Free-streaming cut-offs ($\Sigma m\nu$, WDM)

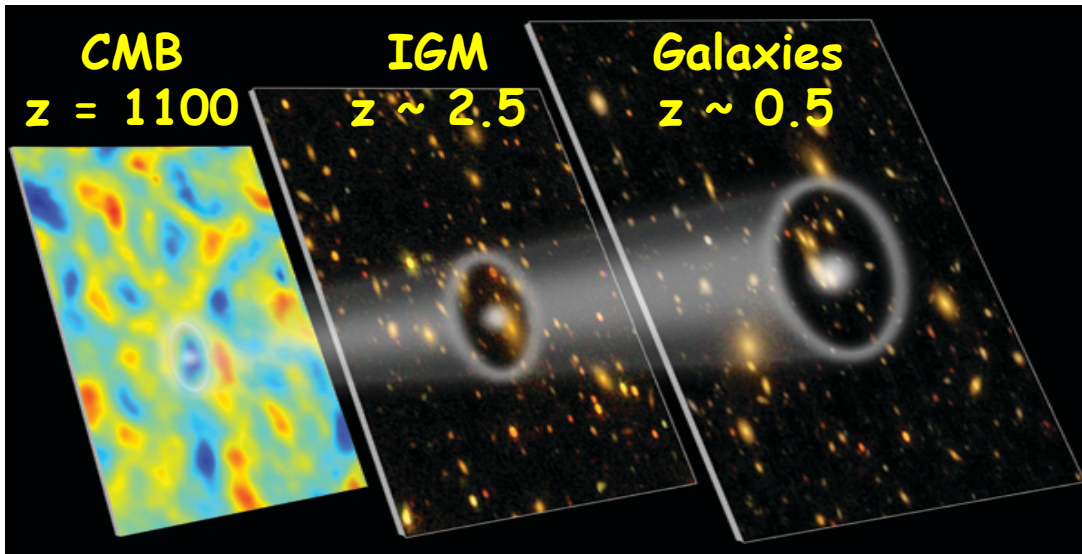


SDSS



DESI

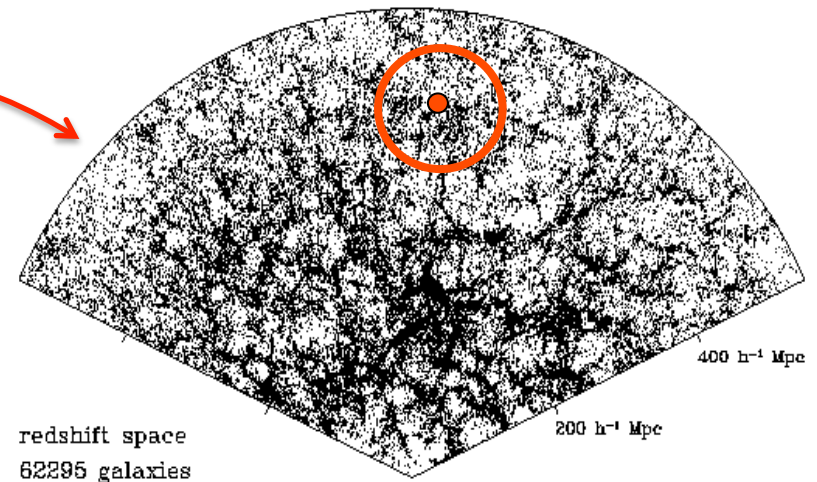
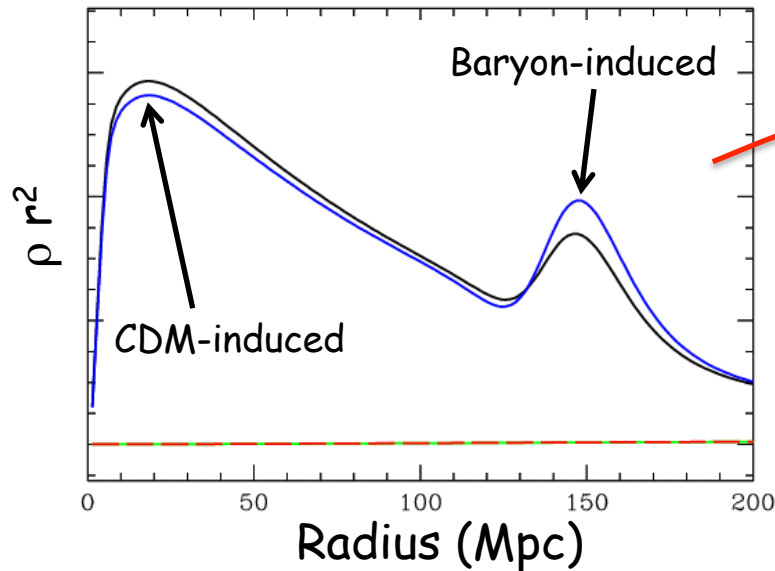
Baryon Acoustic Oscillations (BAO)



Propagation of baryon-photon overdensity wave in plasma

Wave frozen at recombination, at comoving $r_s \sim 150$ Mpc

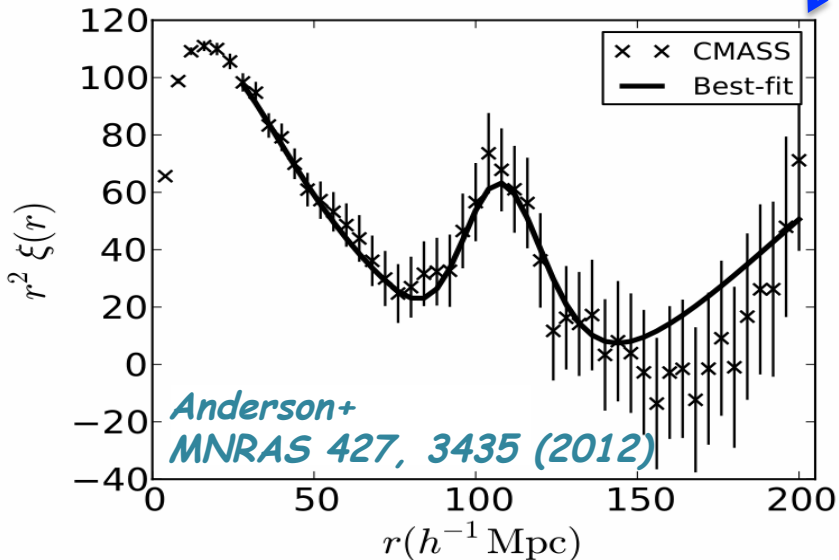
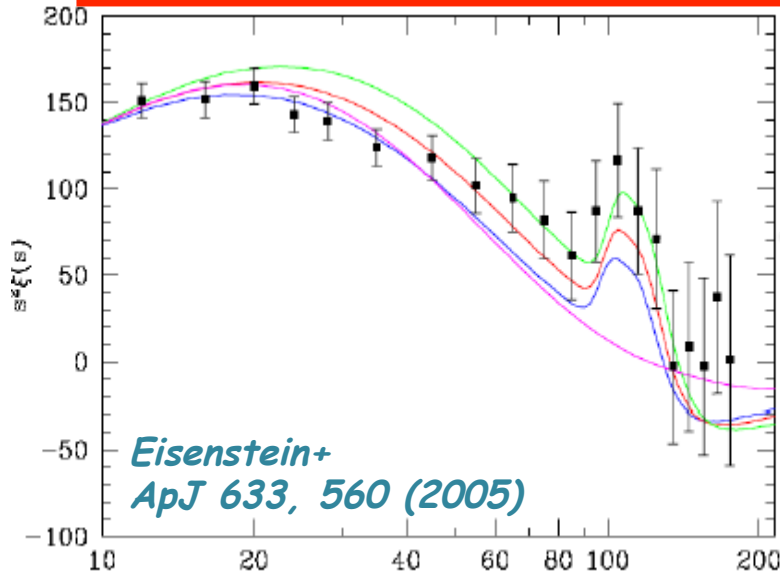
Standard ruler in LSS
A preferred 3D scale



Baryon Acoustic Oscillations (BAO)

Observations

- 2005: First detection of BAO peak
- 2012: 5σ confirmation by BOSS
- 2014: First 3D measurements of BAO



Baryon Acoustic Oscillations (BAO)

Observations

2005: First detection of BAO peak

2012: 5σ confirmation by BOSS

2014: First 3D measurements of BAO

Transverse direction

$$\Delta\theta = r_s / [(1+z) D_A(z)]$$

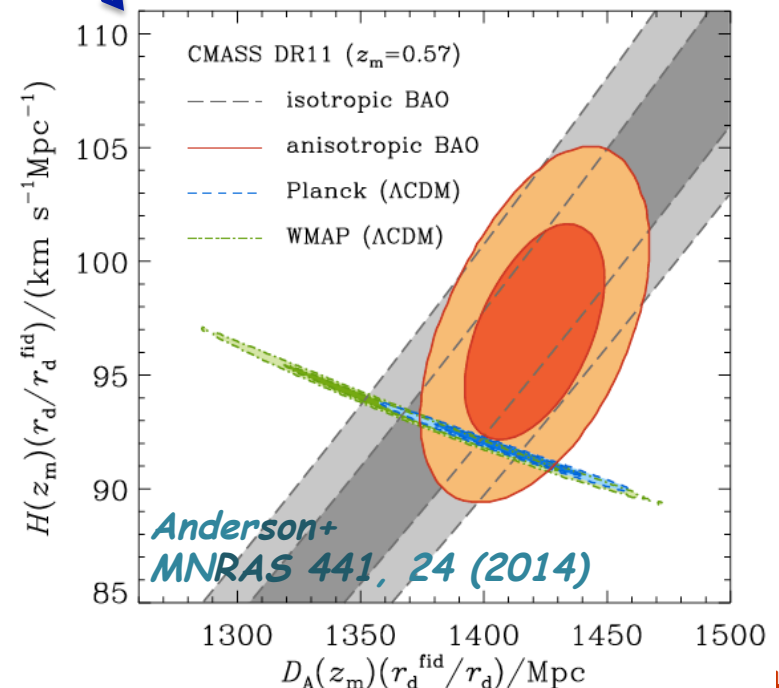
\Rightarrow Angular distance $D_A(z)$

as SNIa: $D_L(z) = (1+z)^2 D_A(z)$

Radial direction (along line of sight)

$$\Delta z = r_s H(z) / c$$

\Rightarrow Hubble parameter $H(z)$



Sloan Digital Sky Survey

- 2.5m telescope
(New Mexico)
- 7 500 deg² eBOSS
10 000 deg² BOSS
- 1000 fibers

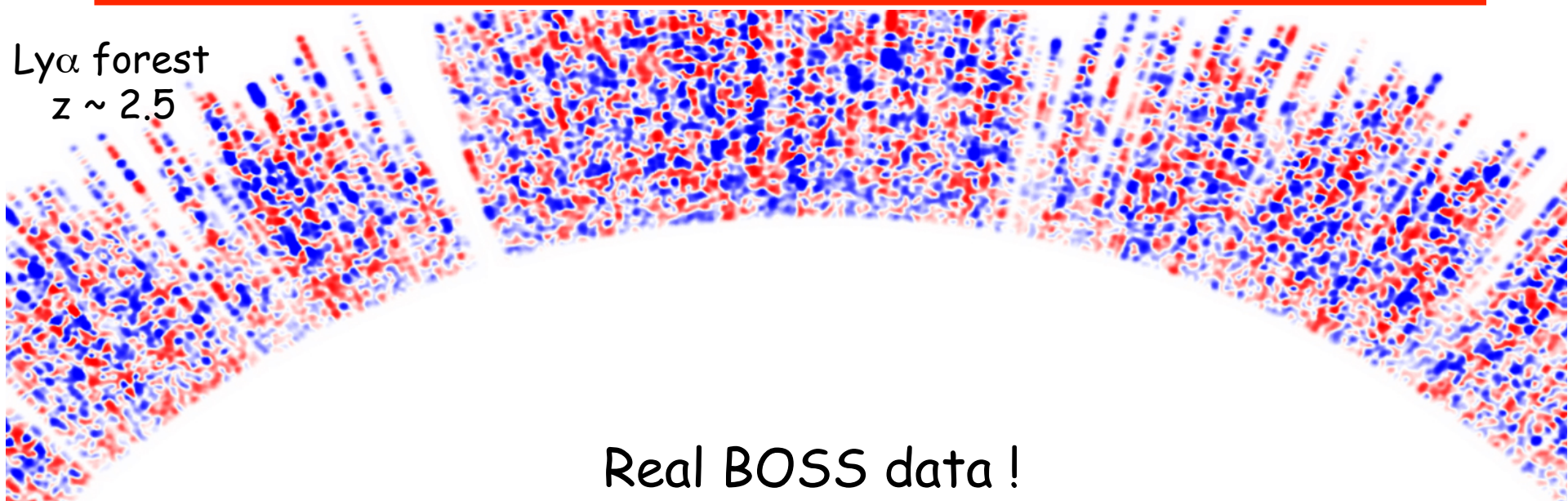
Latest results & predictions

- BAO
- RSD
- Ly α and neutrinos
(active M_ν or sterile ν)

BOSS 2009-2014
eBOSS 2014-2020

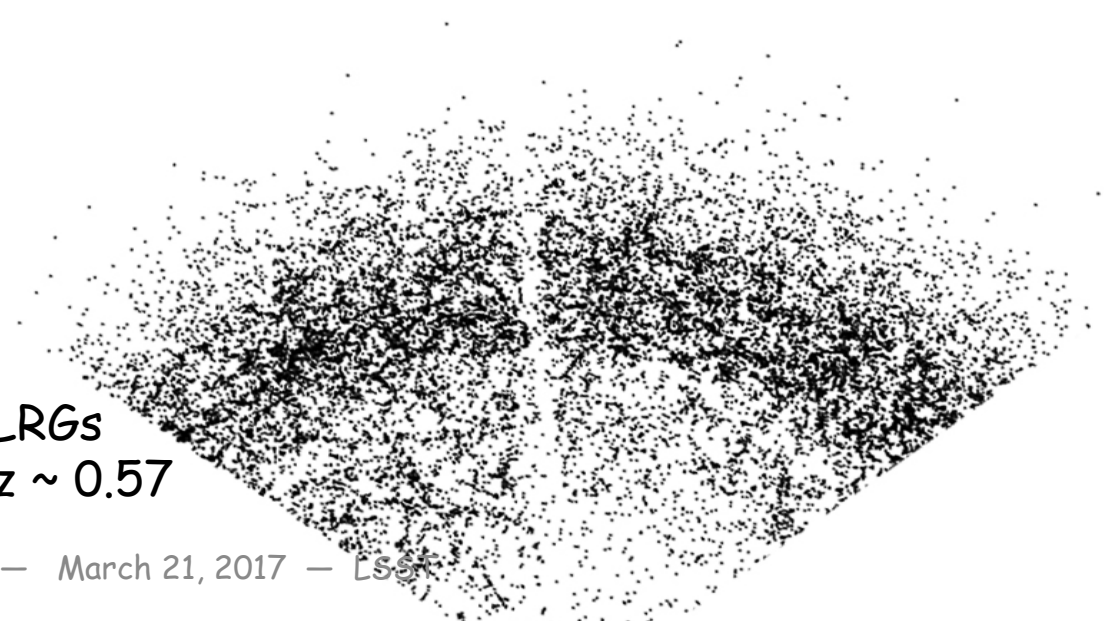
Baryon Acoustic Oscillations (BAO)

Ly α forest
 $z \sim 2.5$

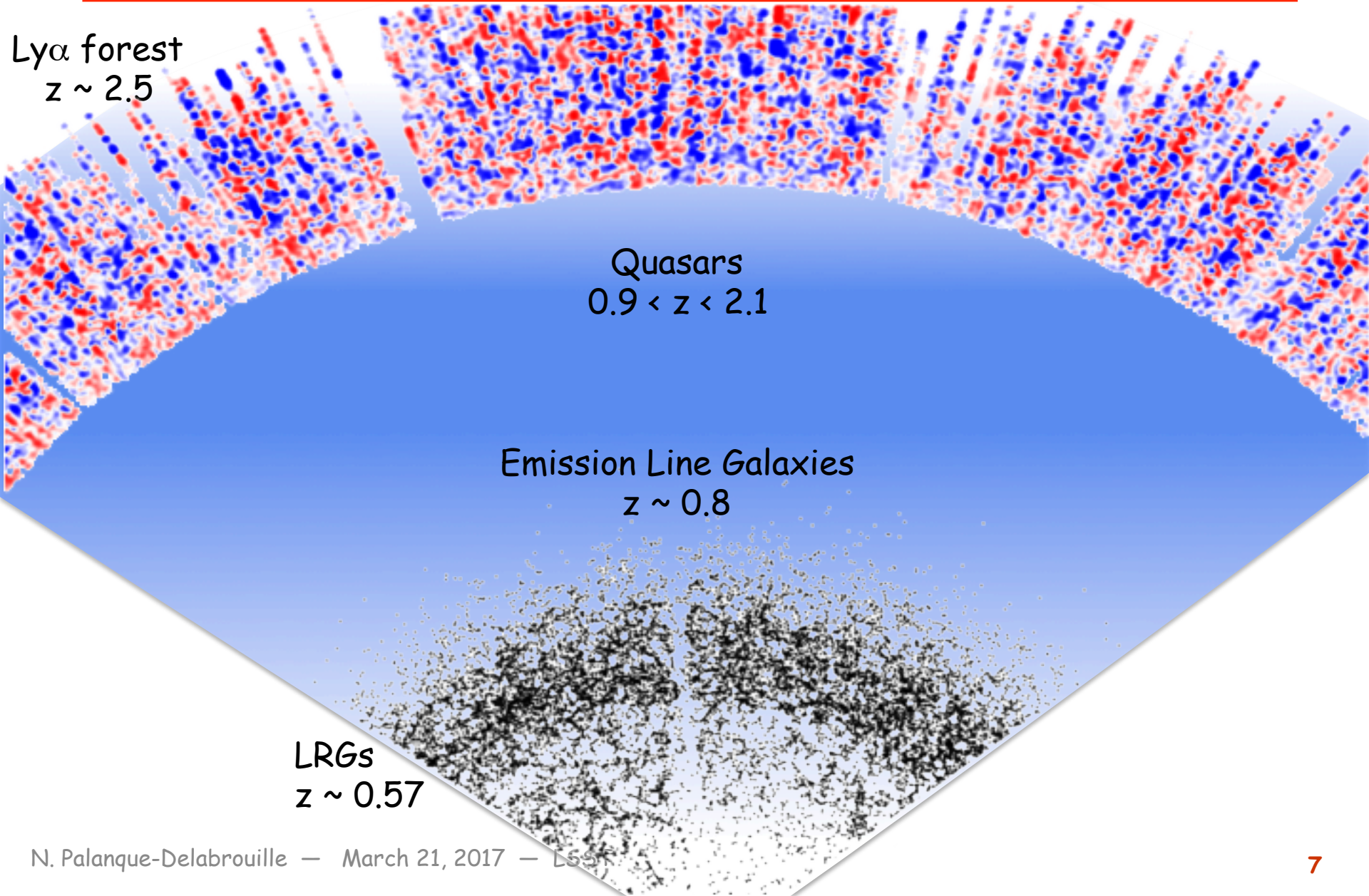


Real BOSS data !

LRGs
 $z \sim 0.57$



Baryon Acoustic Oscillations (BAO)



eBOSS overview

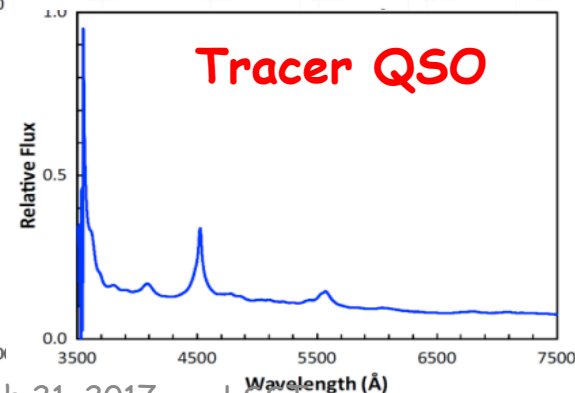
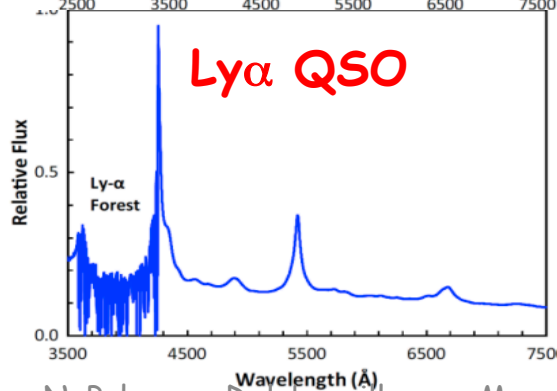
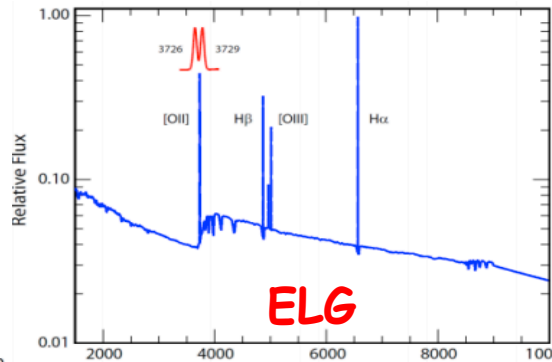
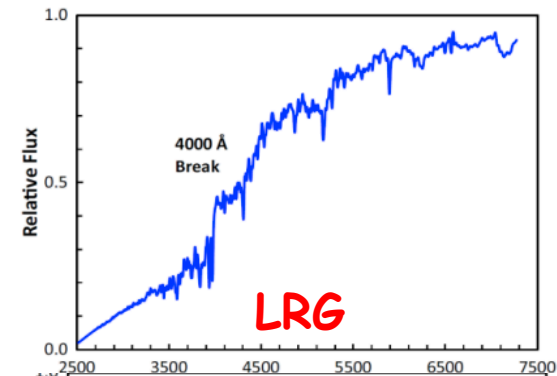
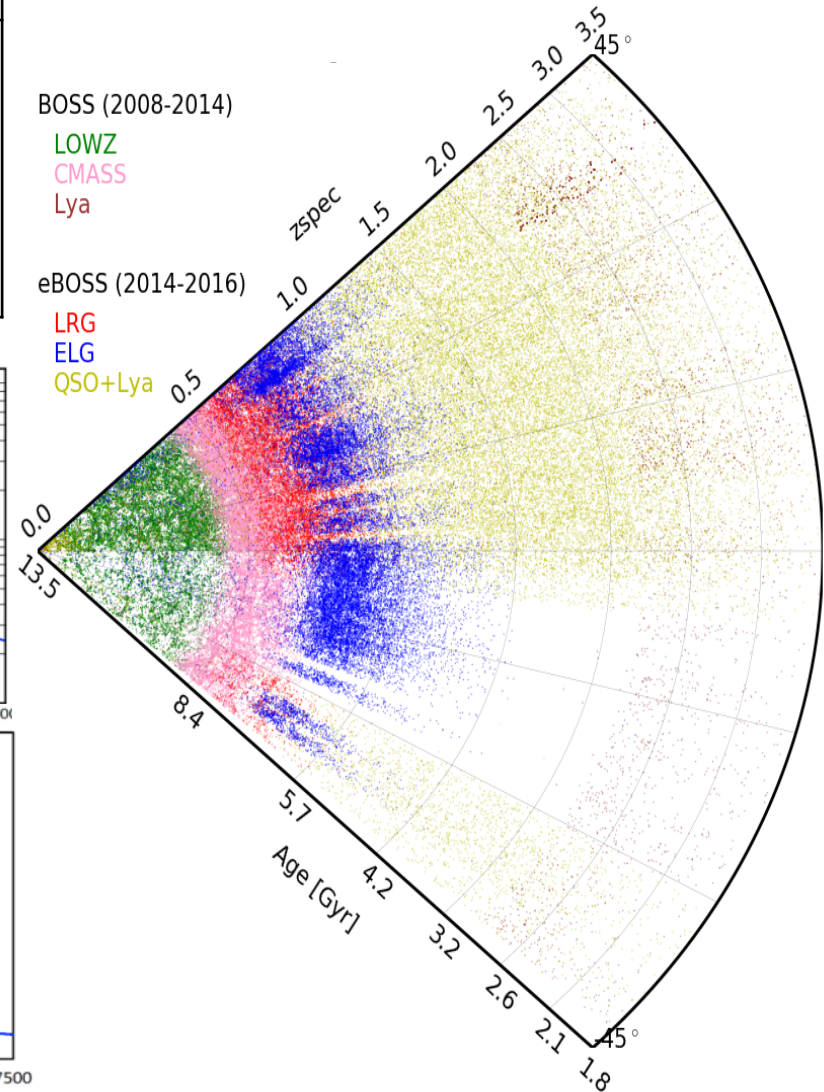
	BOSS		eBOSS	
LRG	$1.2 \cdot 10^6$	$z \sim 0.57$	$+ 250 \cdot 10^3$	$z \sim 0.72$
ELG	—	—	$195 \cdot 10^3$	$z \sim 0.87$
Quasars	—	—	$500 \cdot 10^3$	0.9 - 2.1
Lyα	$180 \cdot 10^3$	$z > 2.1$	$+ 60 \cdot 10^3$	$z > 2.1$

BOSS (2008-2014)

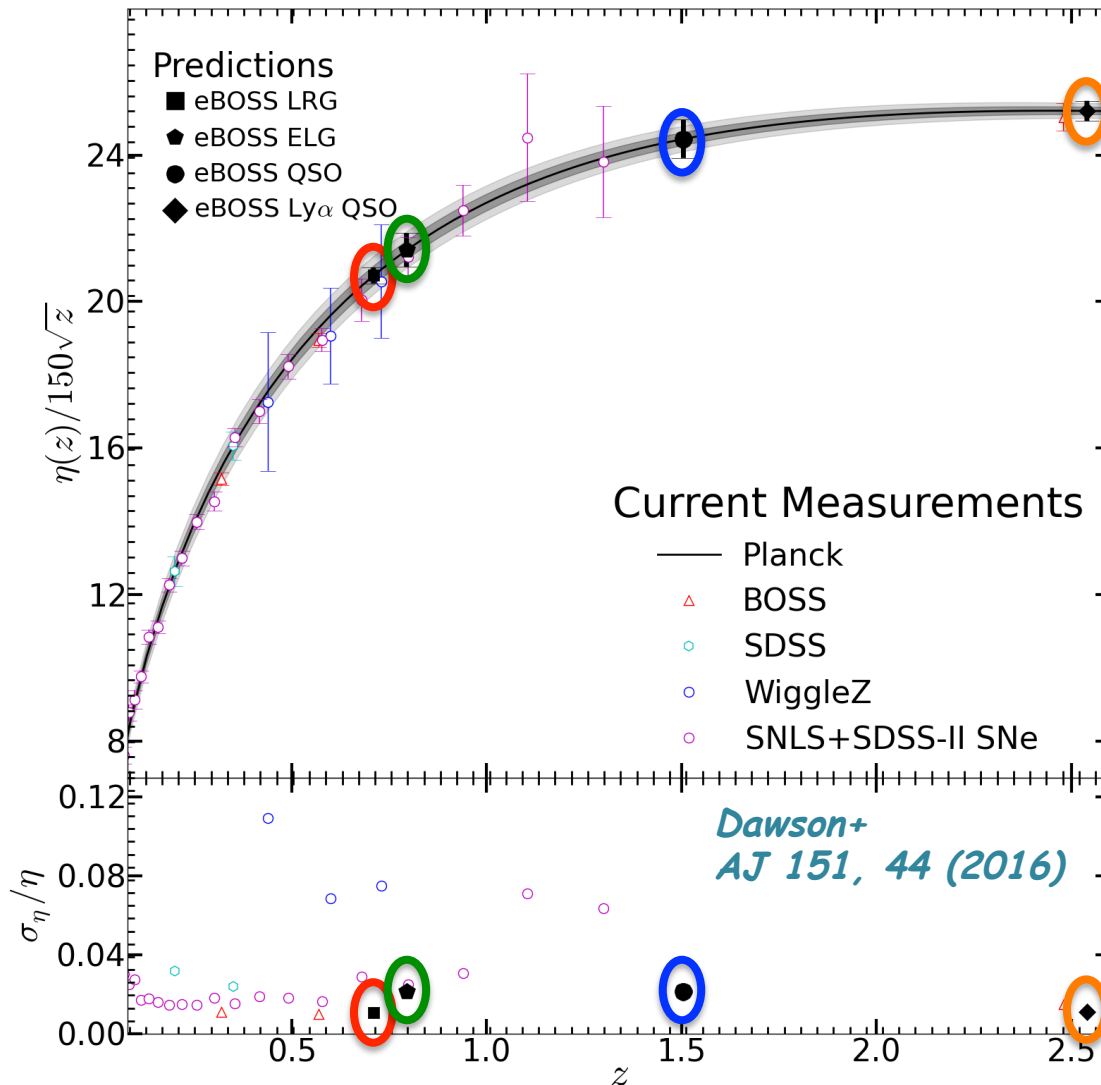
LOWZ
CMASS
Lya

eBOSS (2014-2016)

LRG
ELG
QSO+Lya



eBOSS overview - BAO



Expect 1-2%
distance precision
on all tracers

LRG: 0.8%

ELG: 2.0%

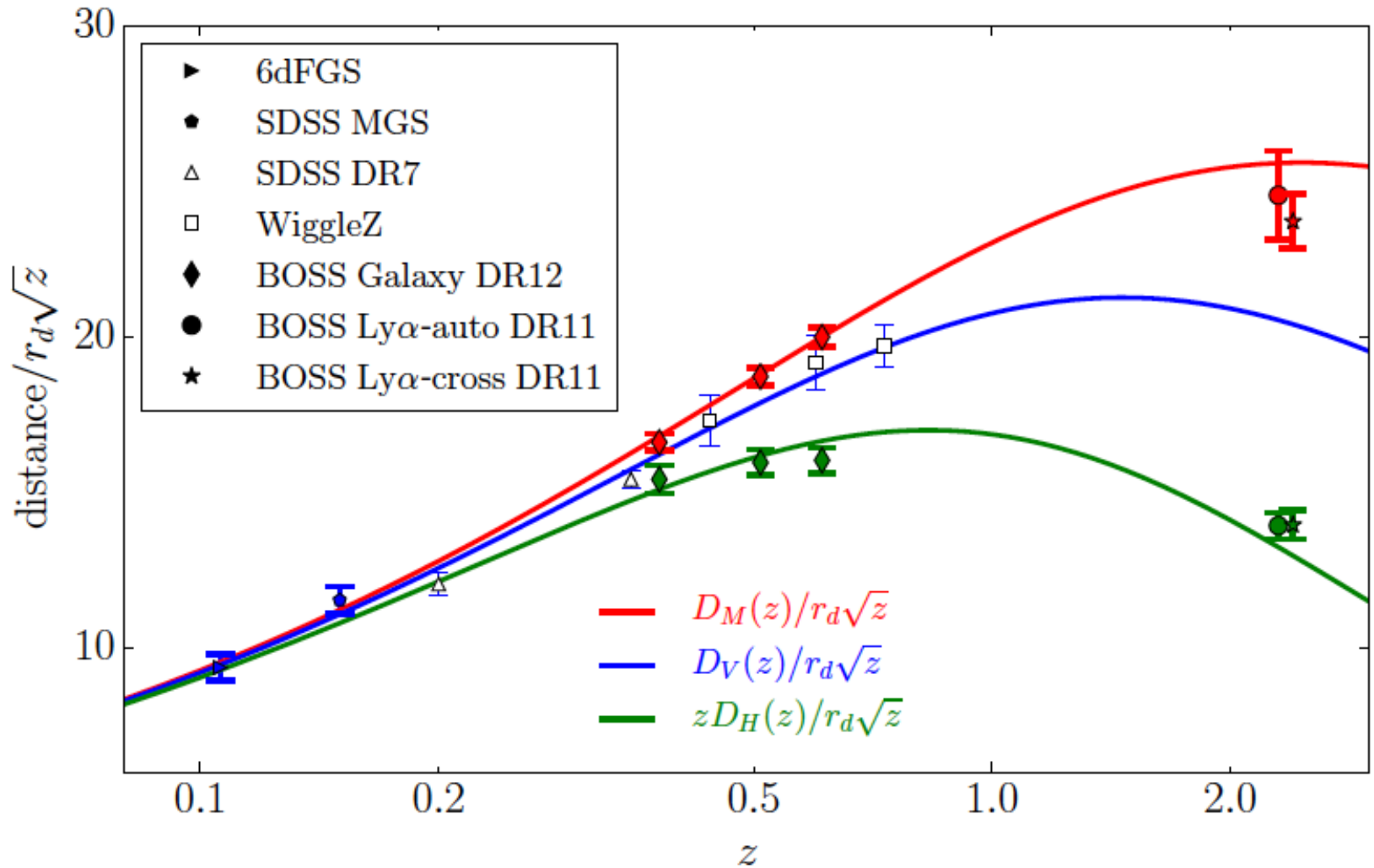
QSO: 1.2%

Ly α :

1.4% on $H(z)$

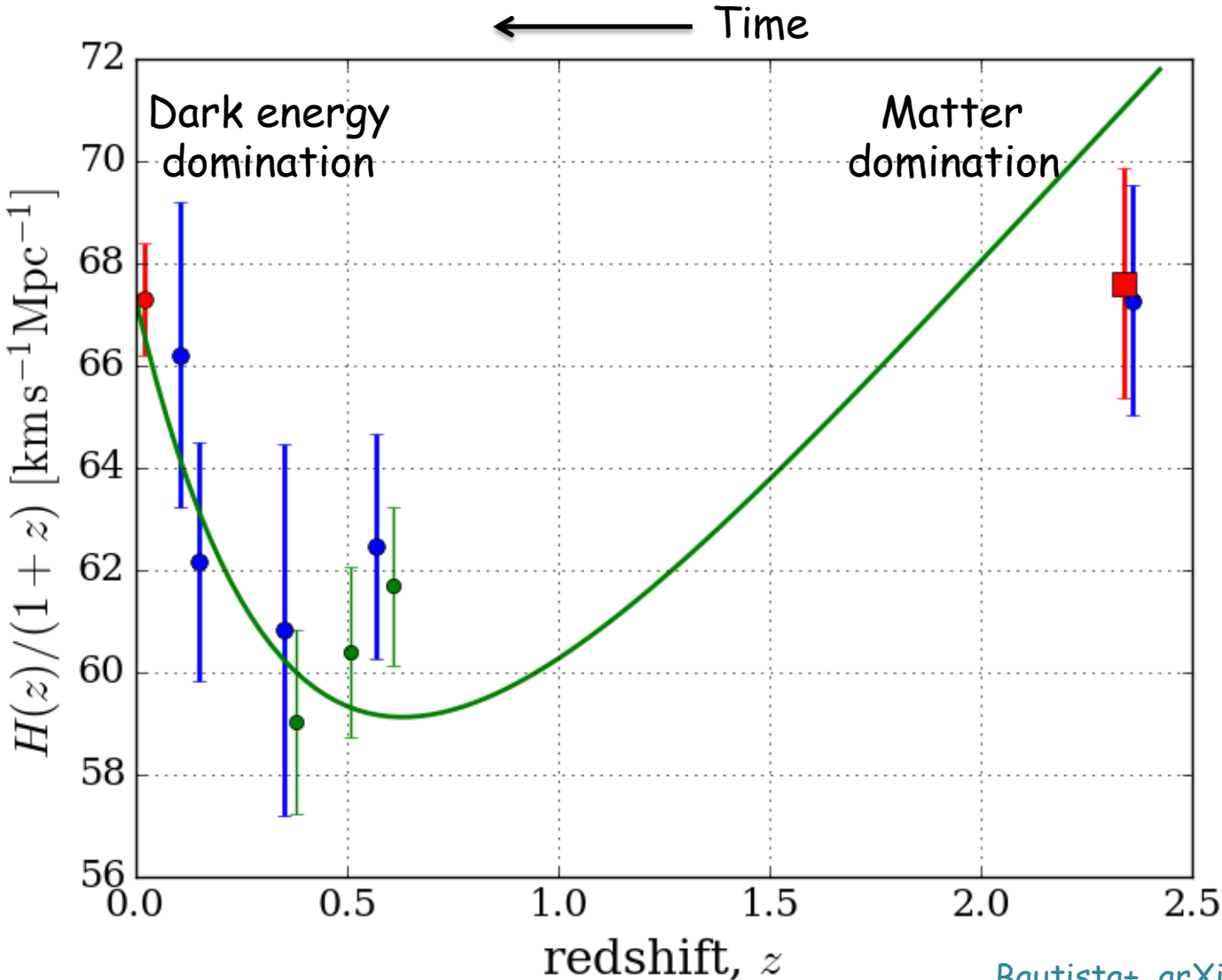
1.7% on $d_A(z)$

Latest results - BAO (BOSS DR12)



Alam+ arXiv:1607.03155

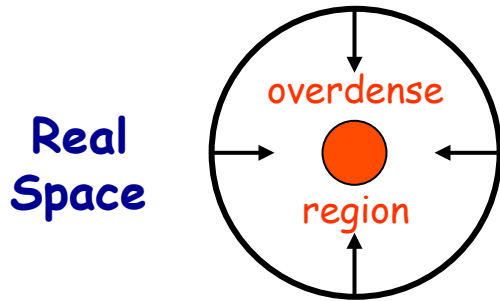
Latest results - BAO (BOSS DR12)



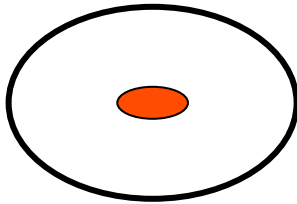
Bautista+, arXiv:1702.00176

eBOSS overview - RSD

→ Peculiar velocity



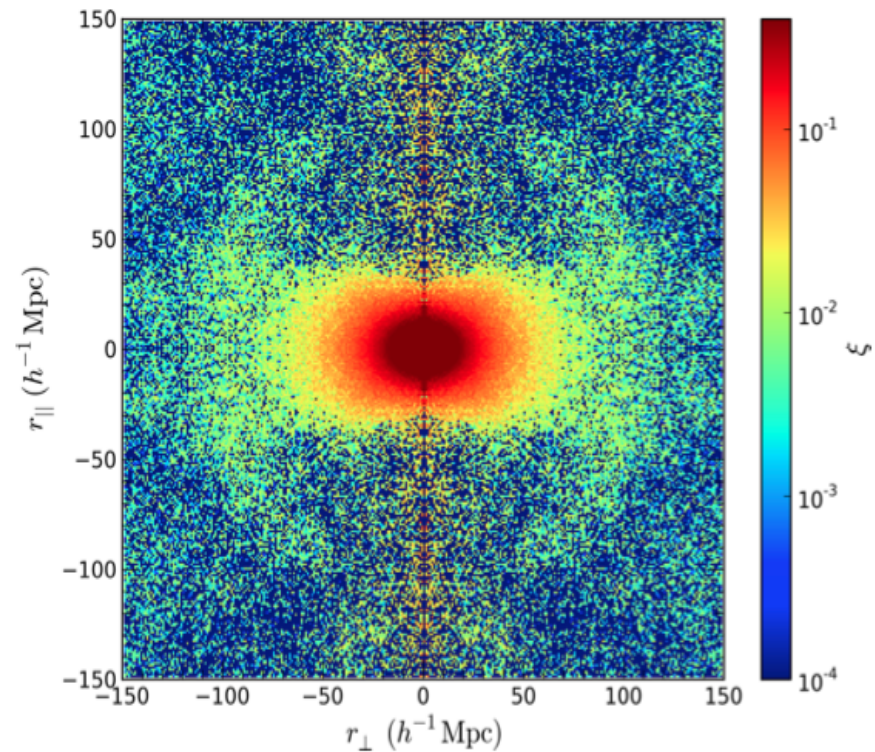
Redshift Space



Measure of gravitational growth

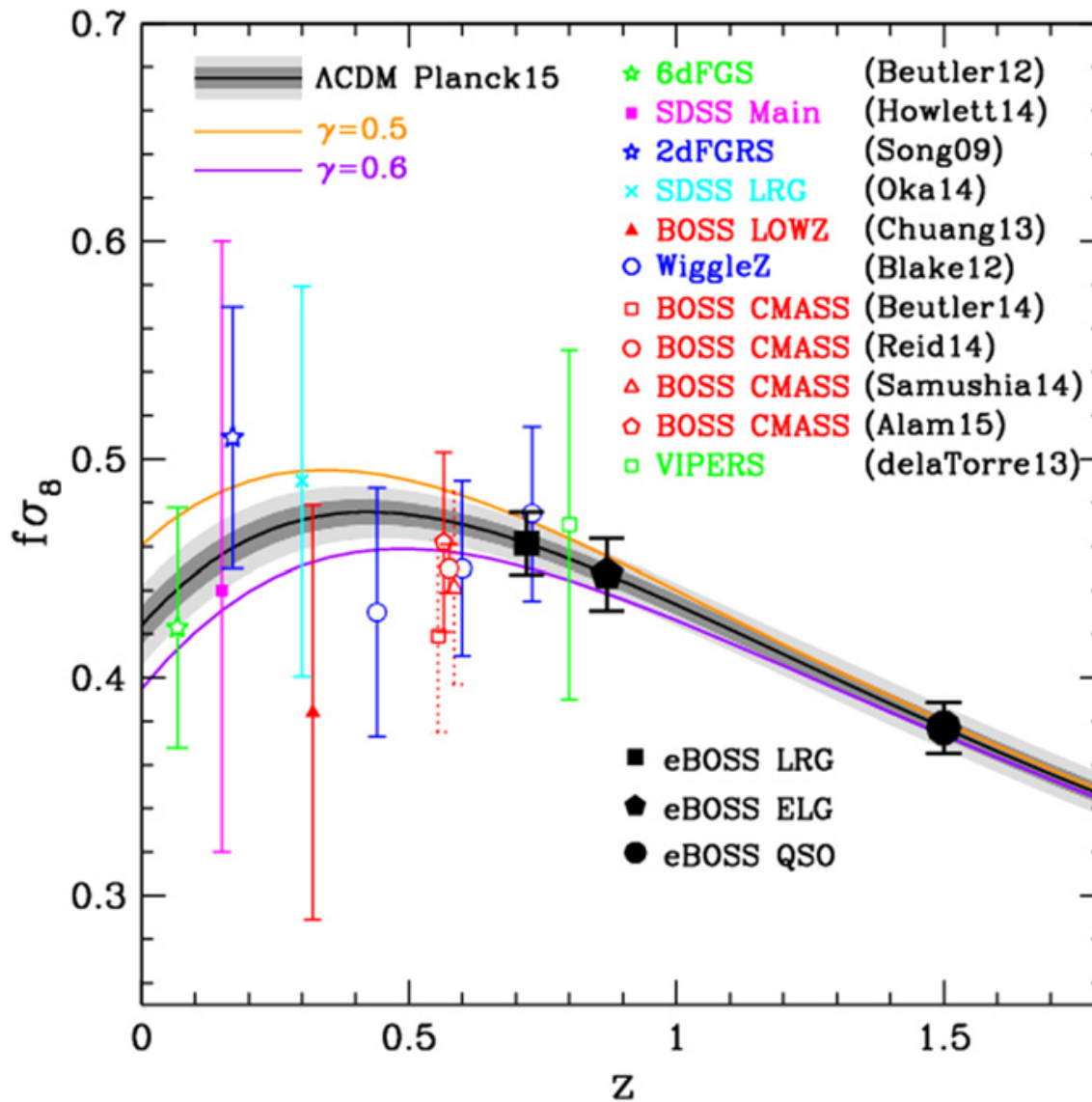
$$P_F(k) = b_F^2 \times \left[1 + \beta \cos(\theta)^2 \right]^2 \times P_L(k)$$

$$\beta \rightarrow f\sigma_8$$



Samushia+ (2014)

eBOSS overview - RSD



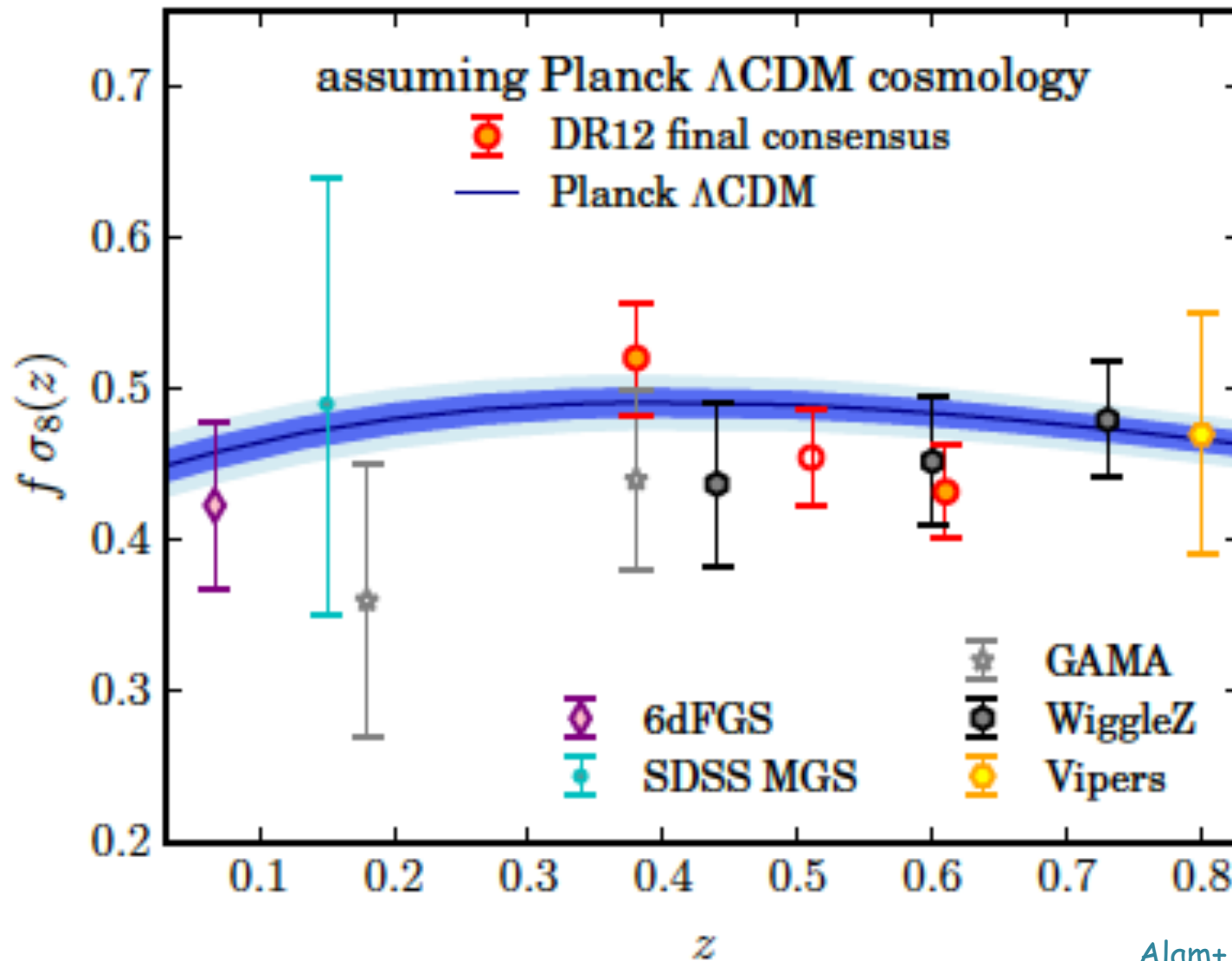
Expect 2.6-3.8%
precision

on $f \sigma_8$

for all tracers

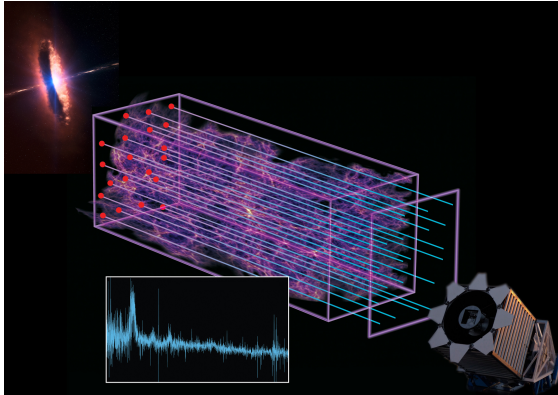
Dawson+
AJ 151, 44 (2016)

Latest results - RSD (BOSS DR12)

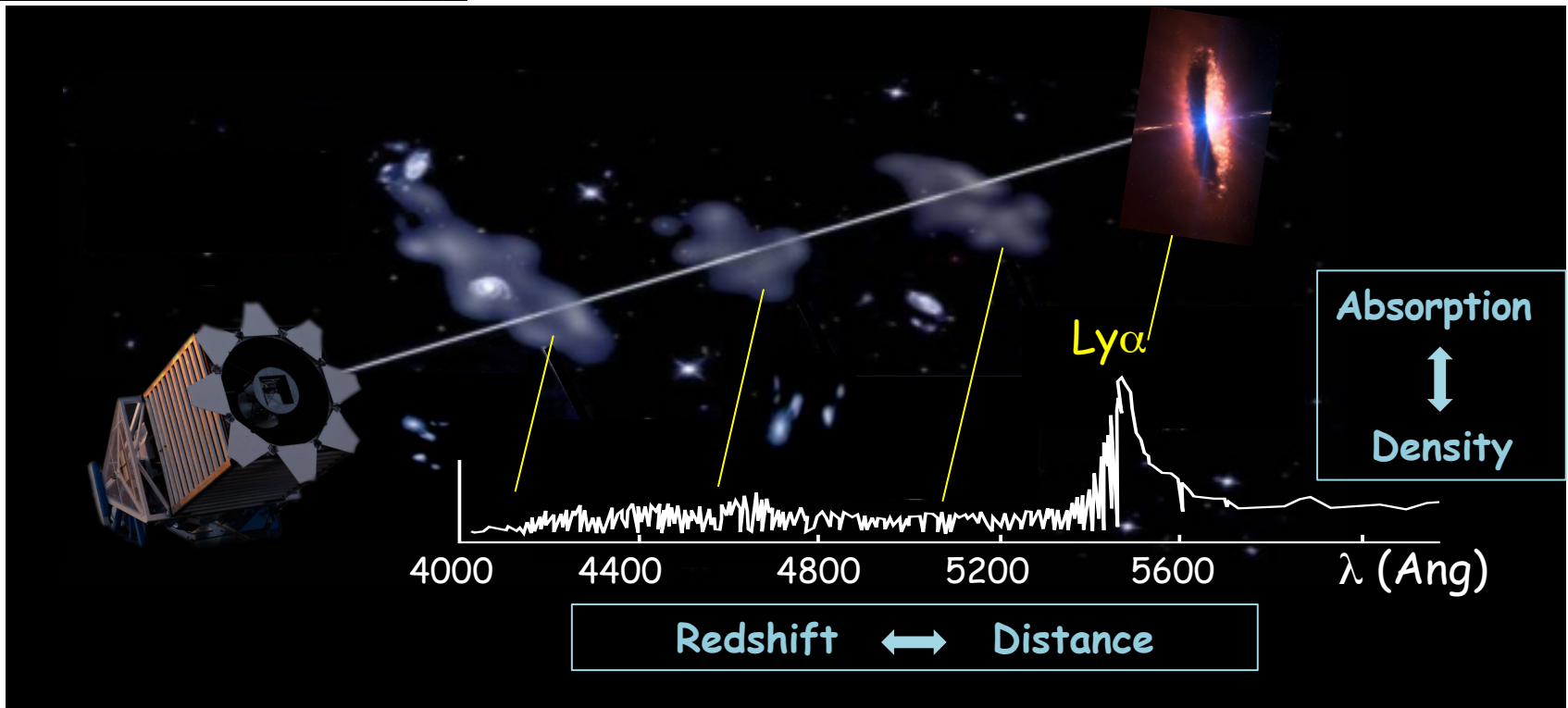


Alam+ arXiv:1607.03155

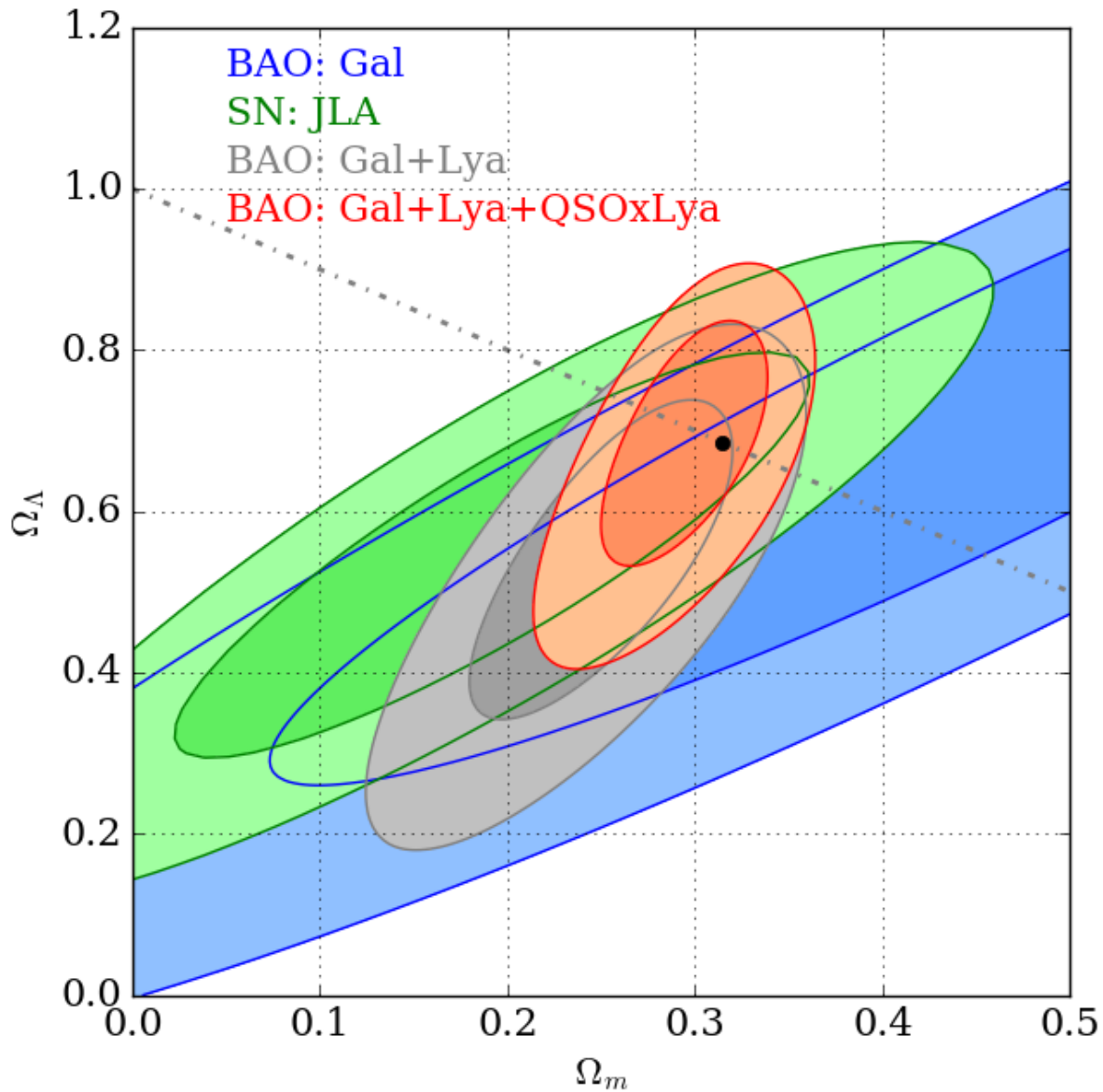
Lyman- α overview



- Quasars visible to **high redshift** (>5)
- Absorption by neutral H (IGM) along line-of-sight
- IGM probes **matter** density
- Matter distribution **on small scales** (v , v_s)



Lyman- α overview



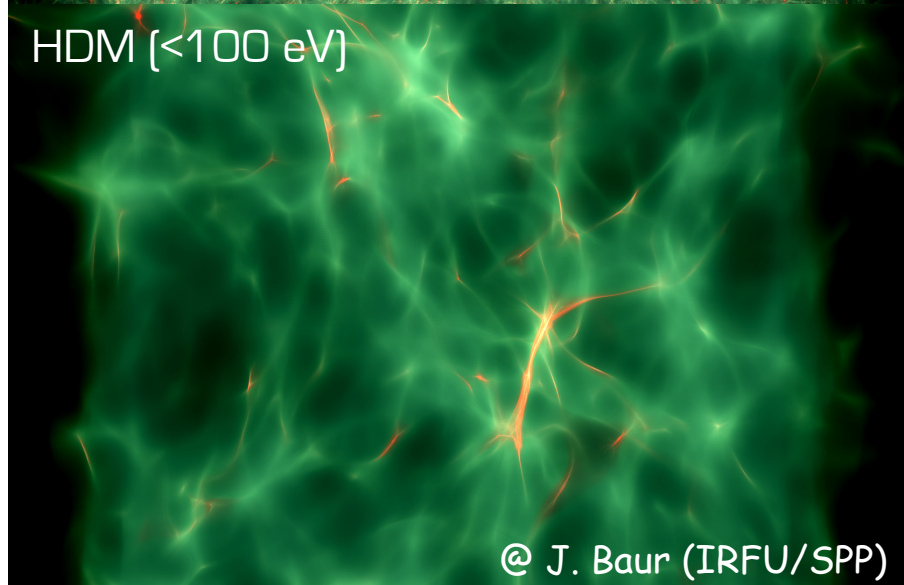
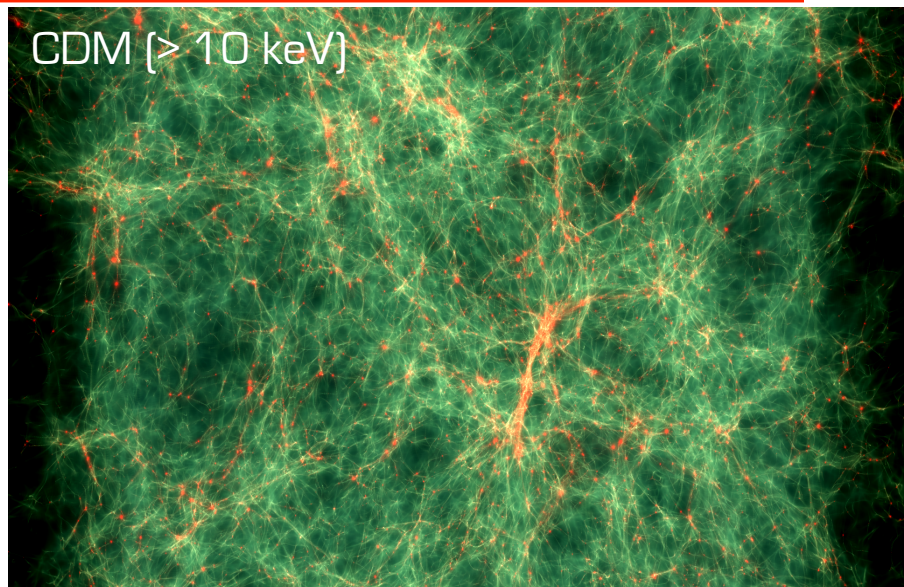
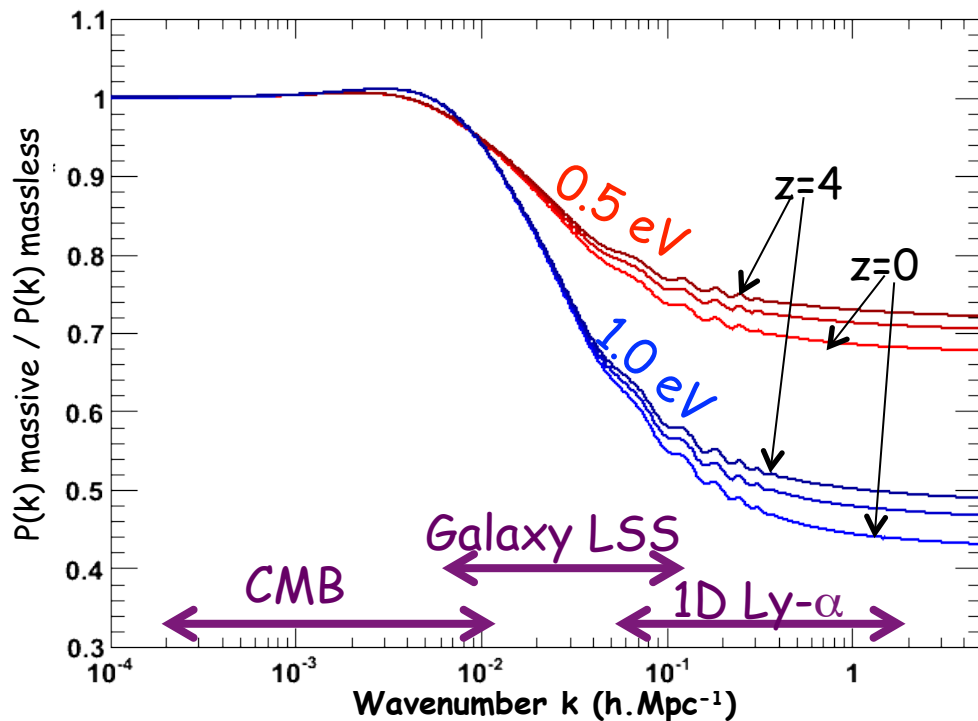
Spatial flatness
independently
of CMB

Ly α & M ν

Neutrinos have mass but m_ν is small

→ ν 's are relativistic during most of history of Universe

→ ν 's free stream and slow-down clustering

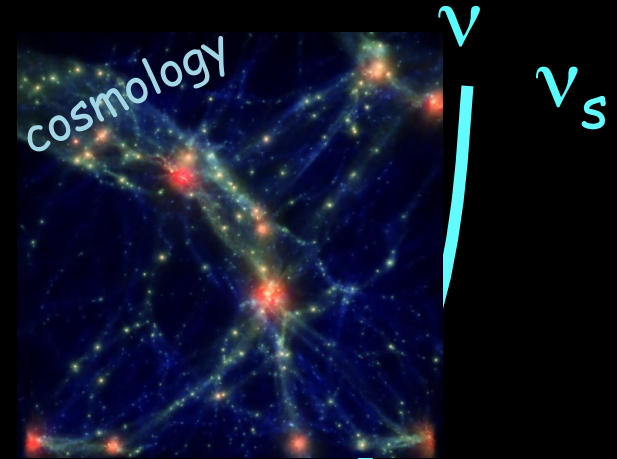
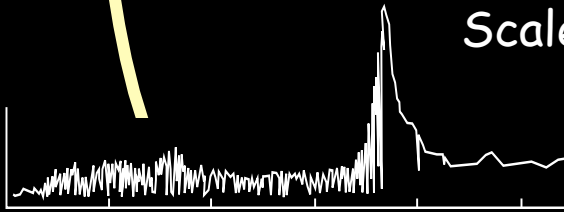


© J. Baur (IRFU/SPP)

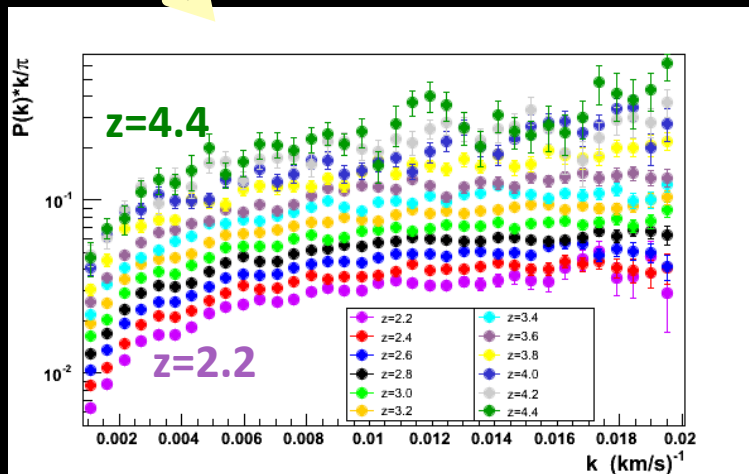
Latest results - Ly α & Mv (BOSS DR9)

Spatial correlation of absorption features (power spectrum)

Scales ~ 1 to 100 Mpc

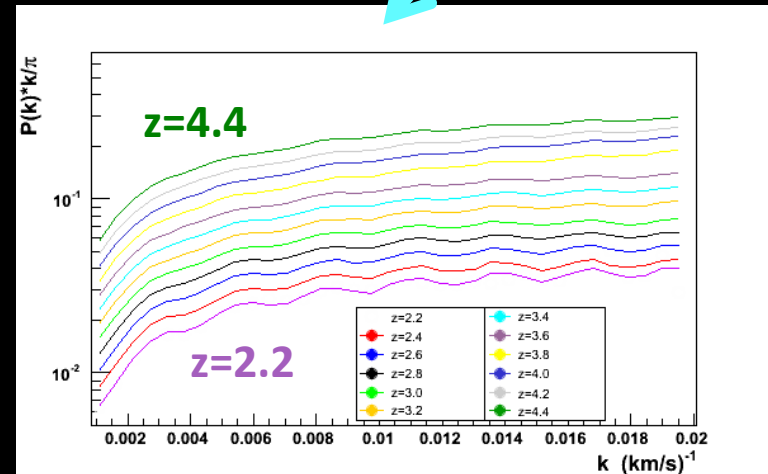


Data - PD+ A&A 559 85 (2013)

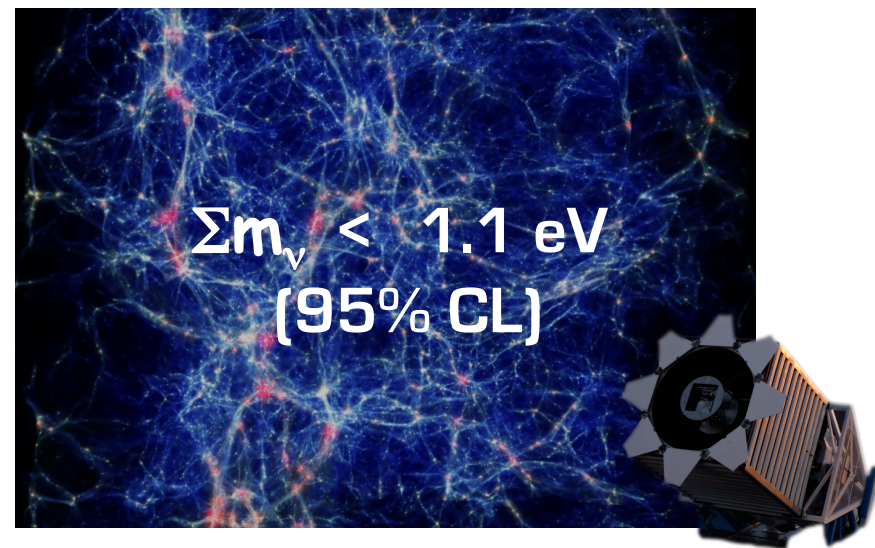
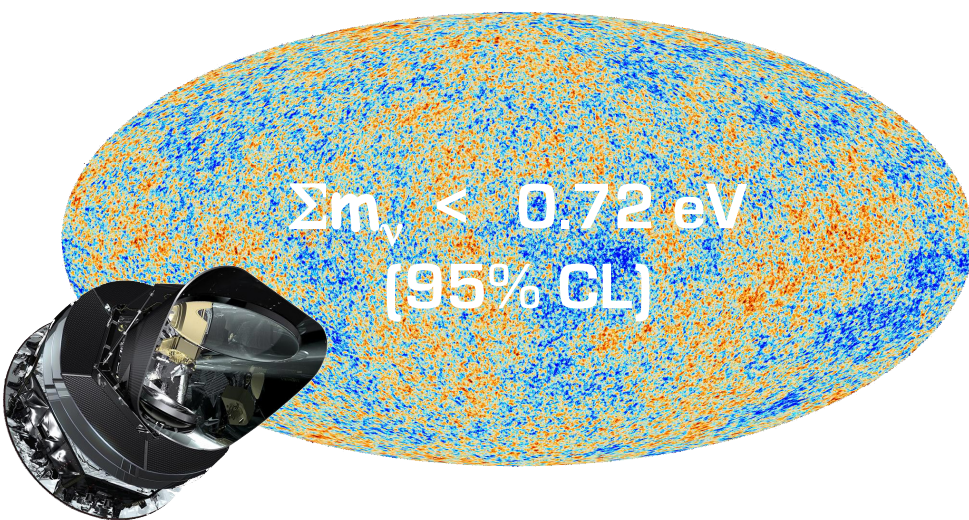


Simulations

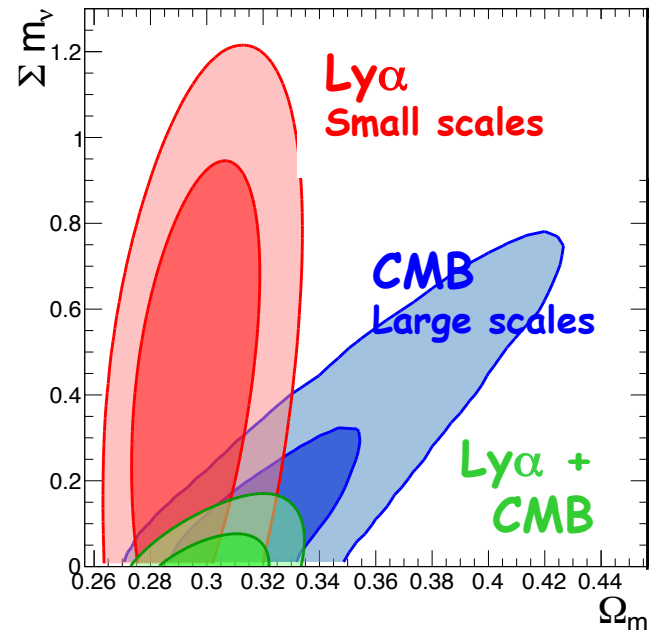
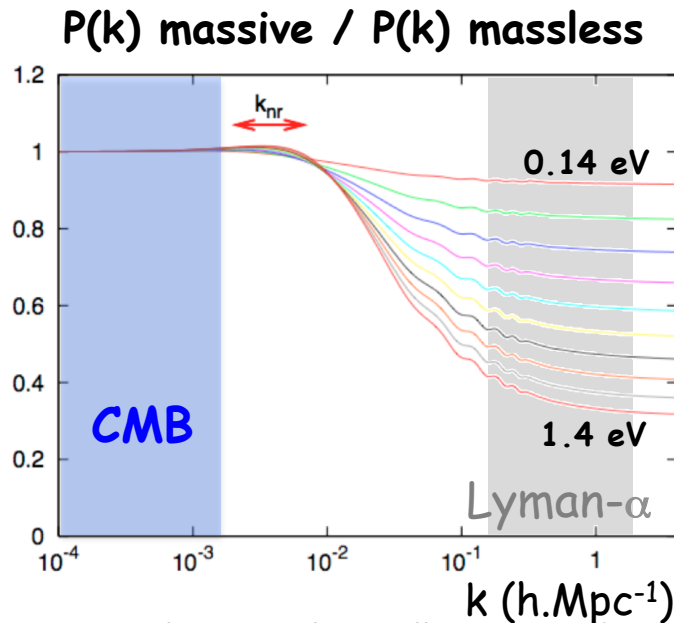
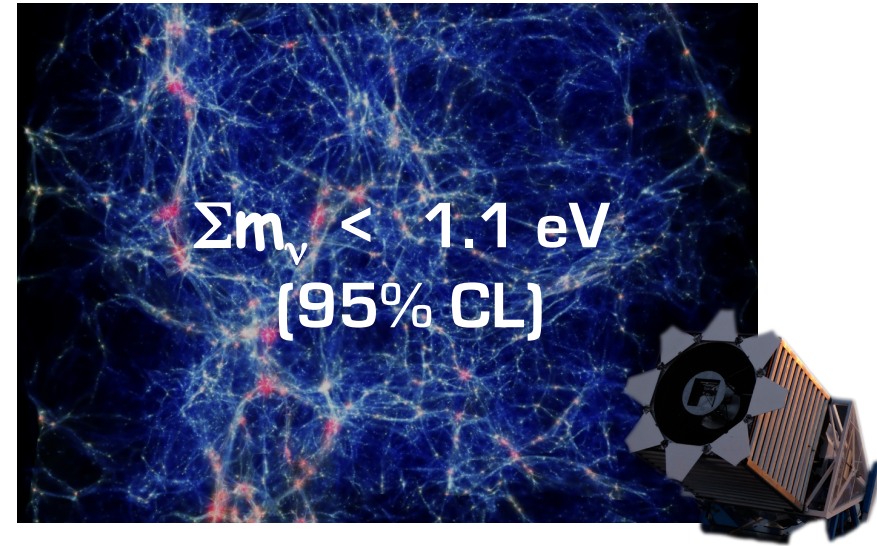
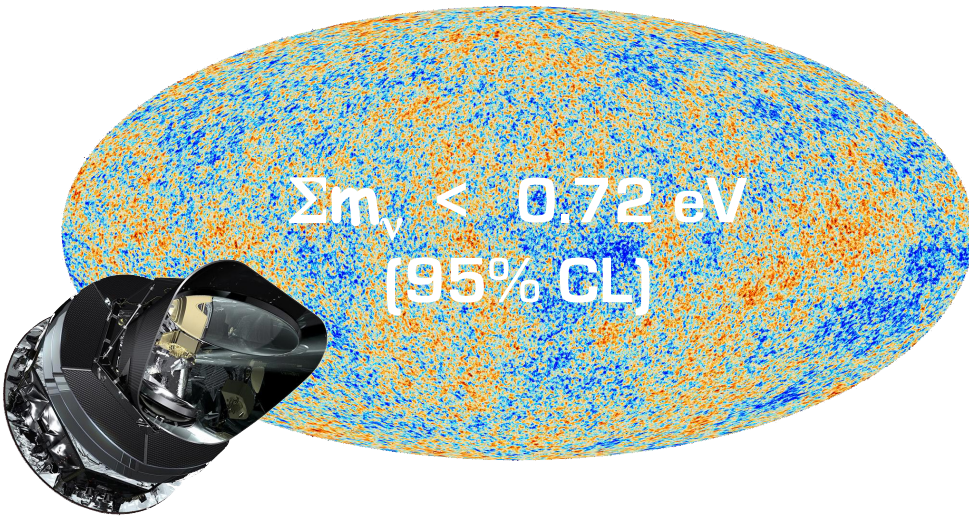
Borde+
JCAP 7 5 (2014)
Rossi+
A&A 567 79 (2014)



Latest results - $Ly\alpha$ & M_V (BOSS DR9)



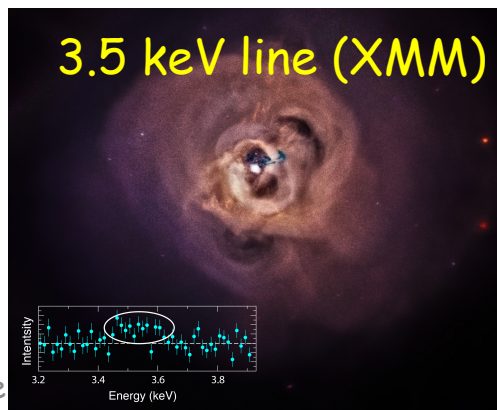
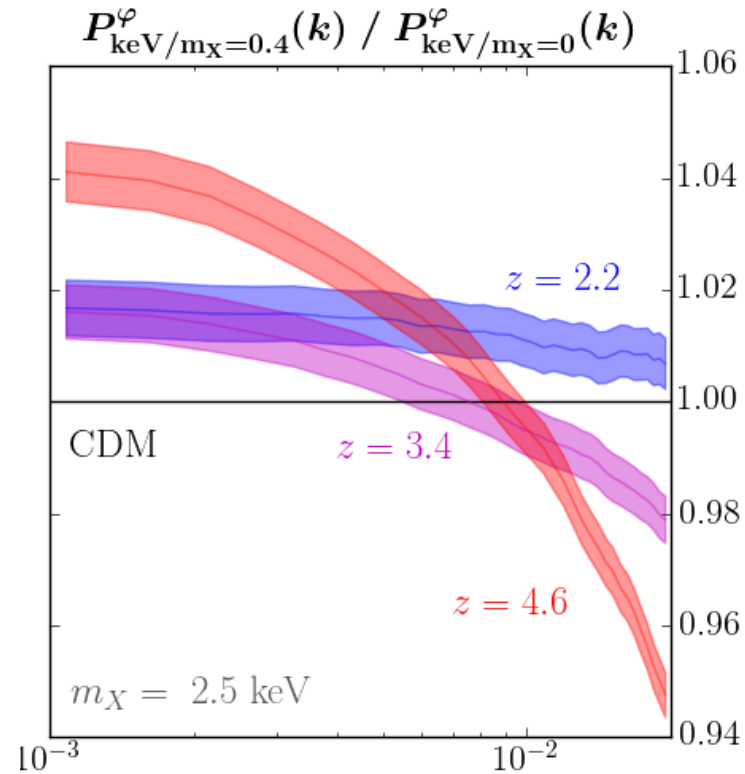
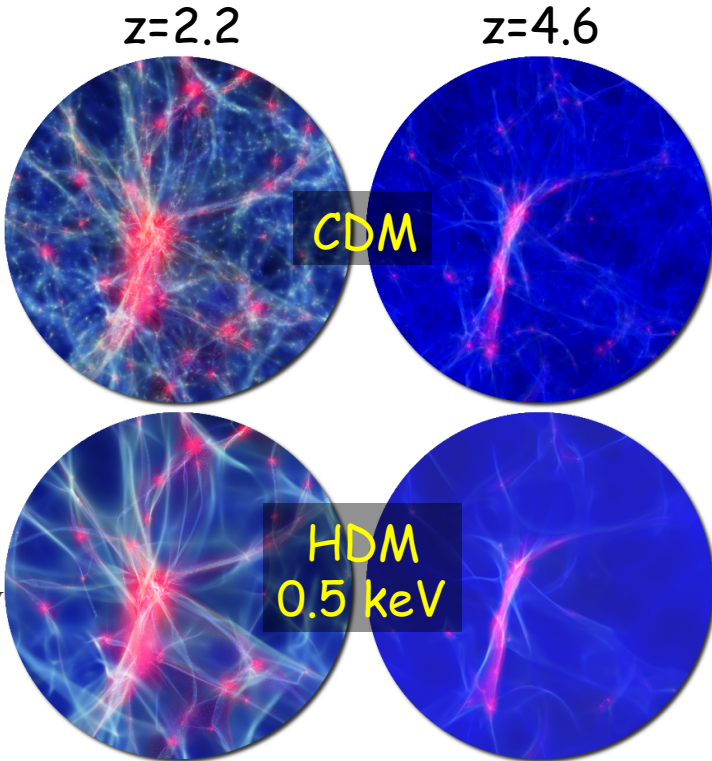
Latest results - $Ly\alpha$ & M_V (BOSS DR9)



$\Sigma m_\nu < 0.12 \text{ eV}$

PD+ JCAP 11 11 (2015)

Latest results - Ly α & v_s (BOSS DR9)



Ly- α from SDSS/BOSS

$\Rightarrow m_{\text{WDM}} > 4.1 \text{ keV}$ (95% CL)

$m_{\text{sterile}} > 24.4 \text{ keV}$ (95% CL)

Baur+ JCAP 8 12 (2016)

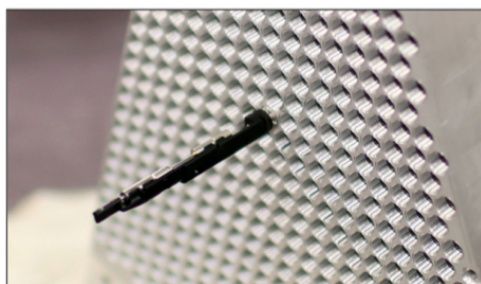
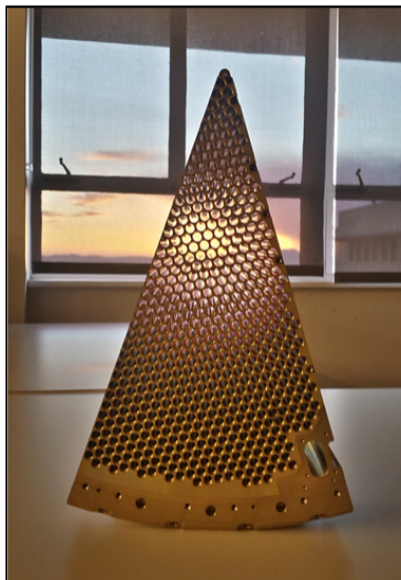
DESI (2019 - 2024)

- 4m telescope
(Arizona)
- 14 000 deg² survey
20 million spectra
 $0 < z < 5$
- 5000 fibers
(robot positioner)

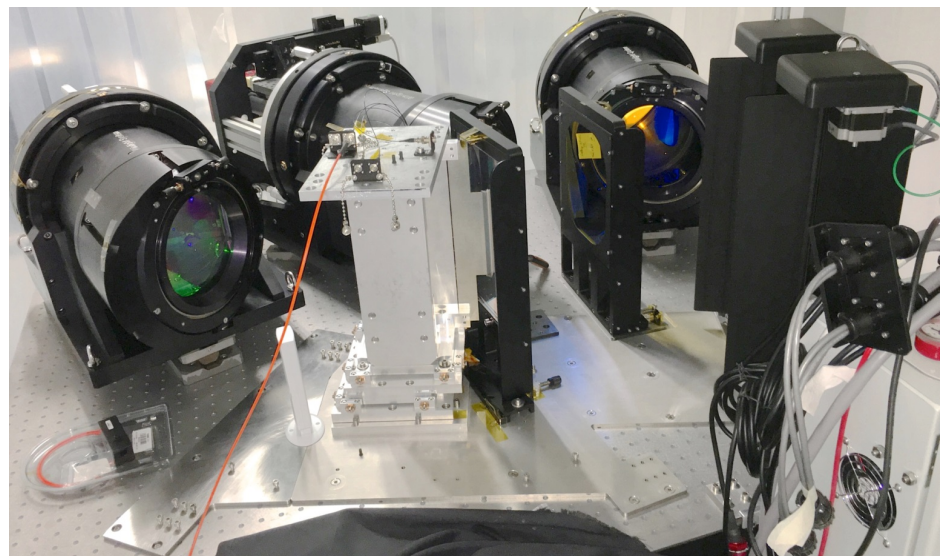


DESI overview

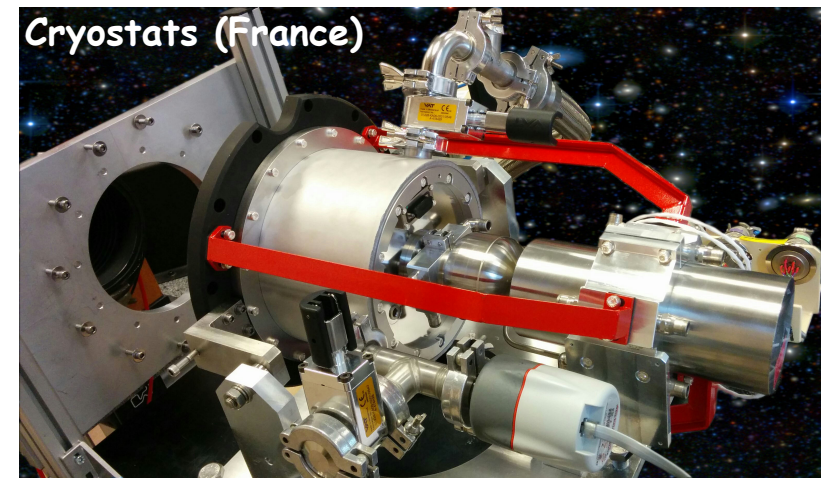
Petal (1/10 focal plane) and 1st fiber positioner



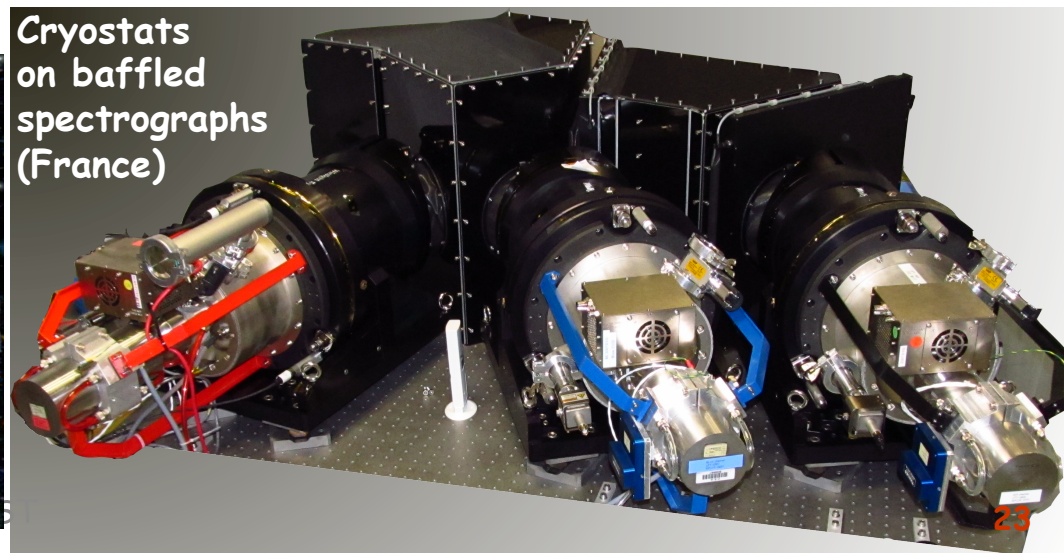
Spectrographs (France)



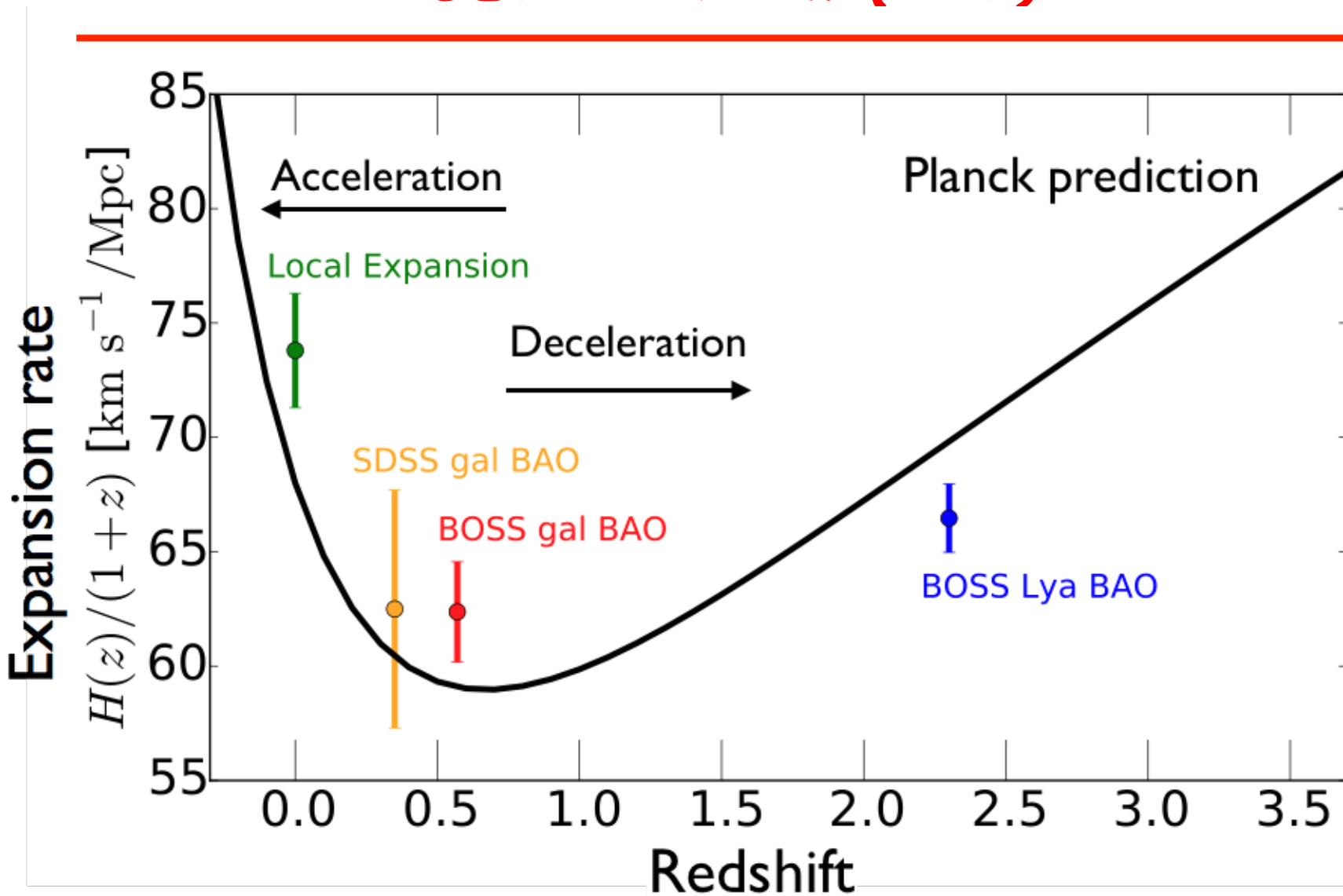
Cryostats (France)



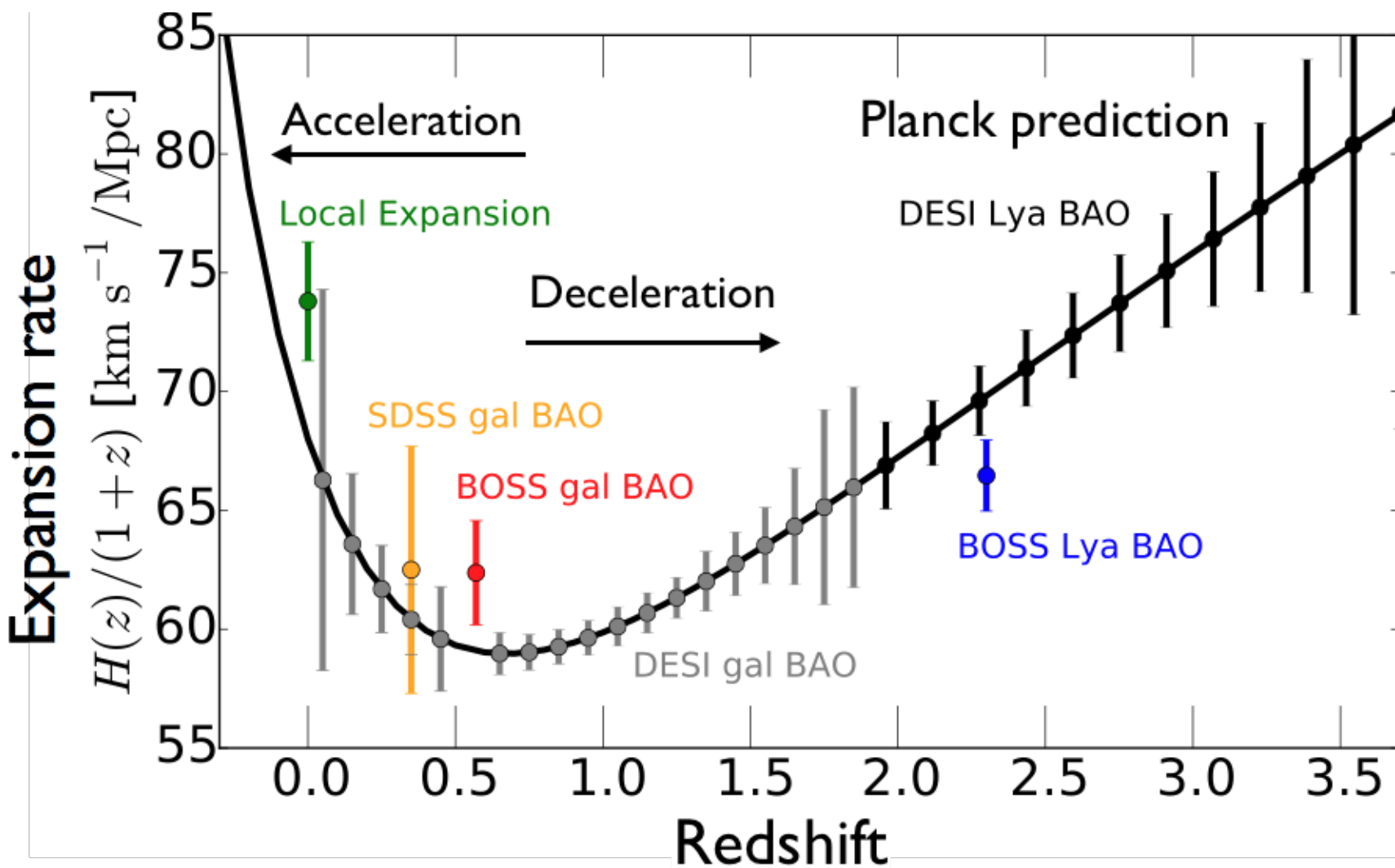
Cryostats on baffled spectrographs (France)



DESI overview (BAO)

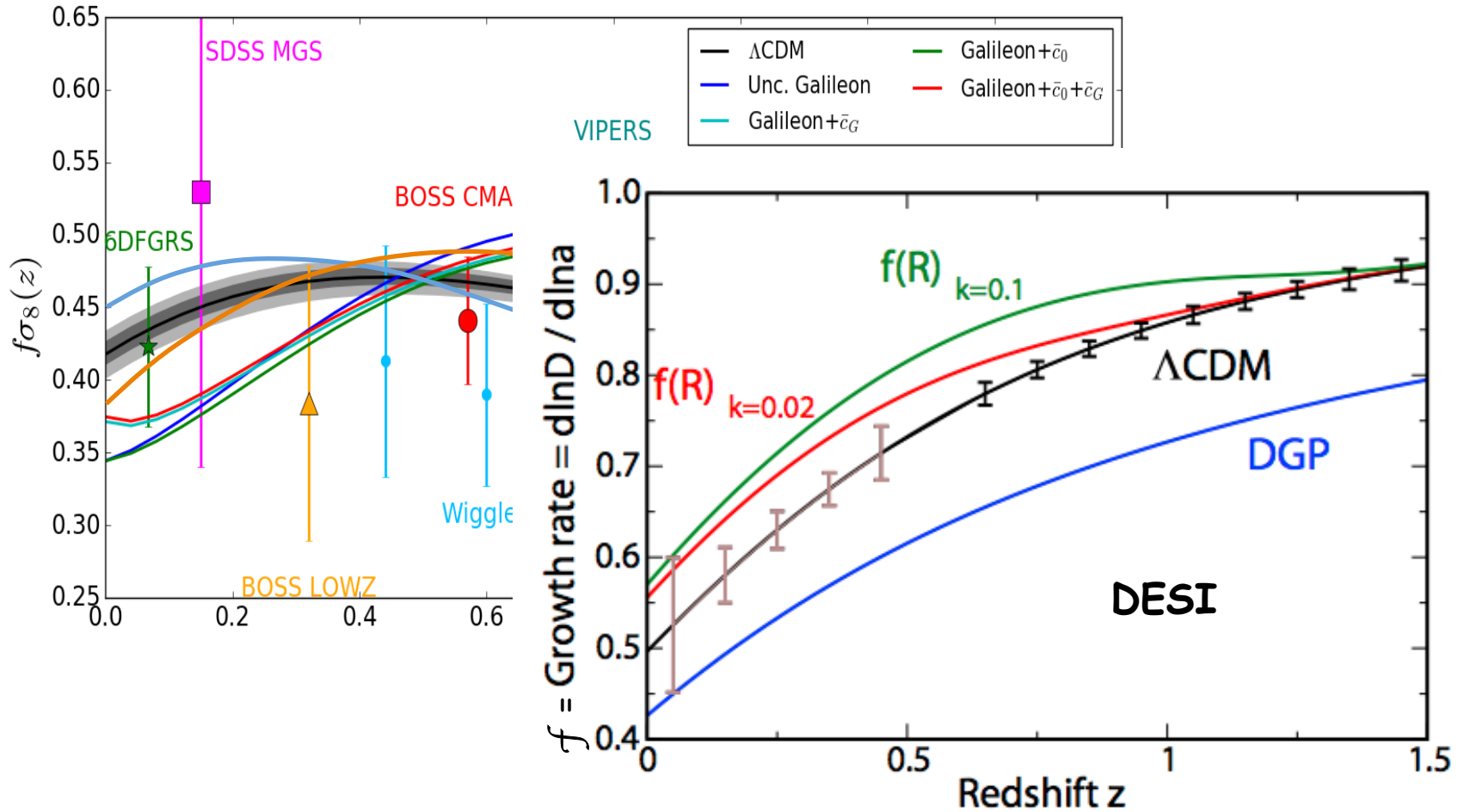


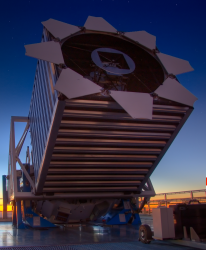
DESI overview (BAO)



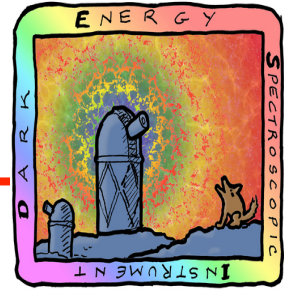
DESI projections Font-Ribera+ (2014)

DESI overview - RSD





Conclusions

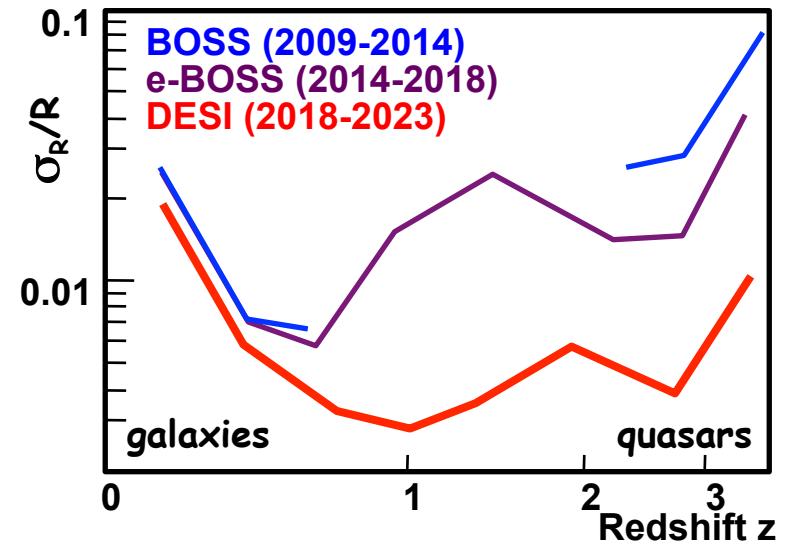


eBOSS well under way

- Percent-level constraints on $H(z)$ & $d_A(z)$ up to $z \sim 3.5$
- BAO data impose spatial flatness, independently of CMB
- Tight constraints from $\text{Ly}\alpha$ on
 - Neutrino masses $\Sigma m_\nu < 0.12 \text{ eV}$ (95% CL) from $\text{Ly}\alpha + \text{CMB}$
 - Sterile neutrinos

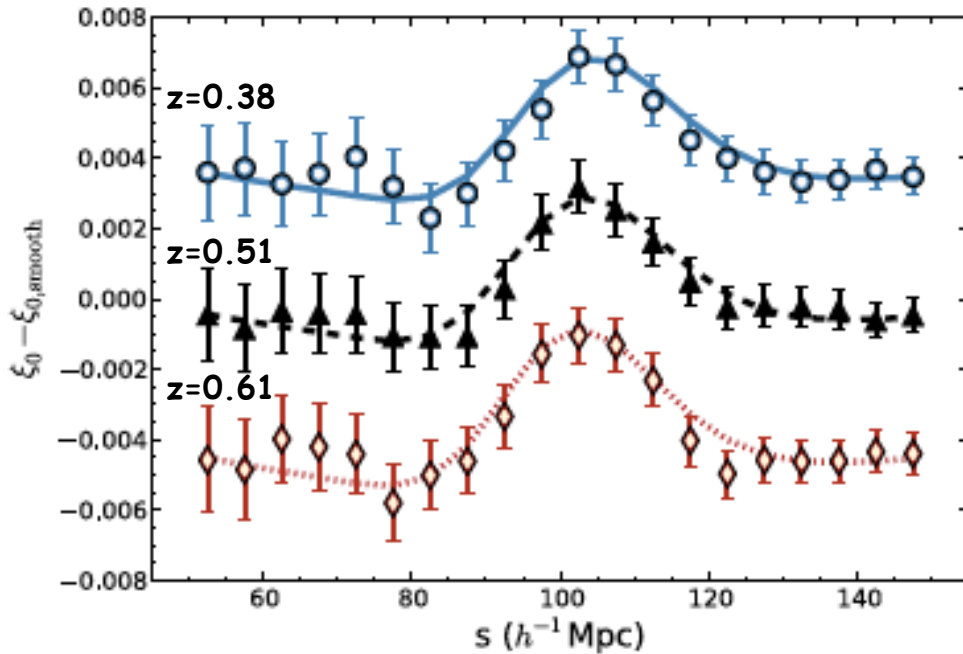
DESI coming up

- x2-10 improvement on **BAO**
- x5 improvement on Σm_ν
24 meV from galaxy clustering

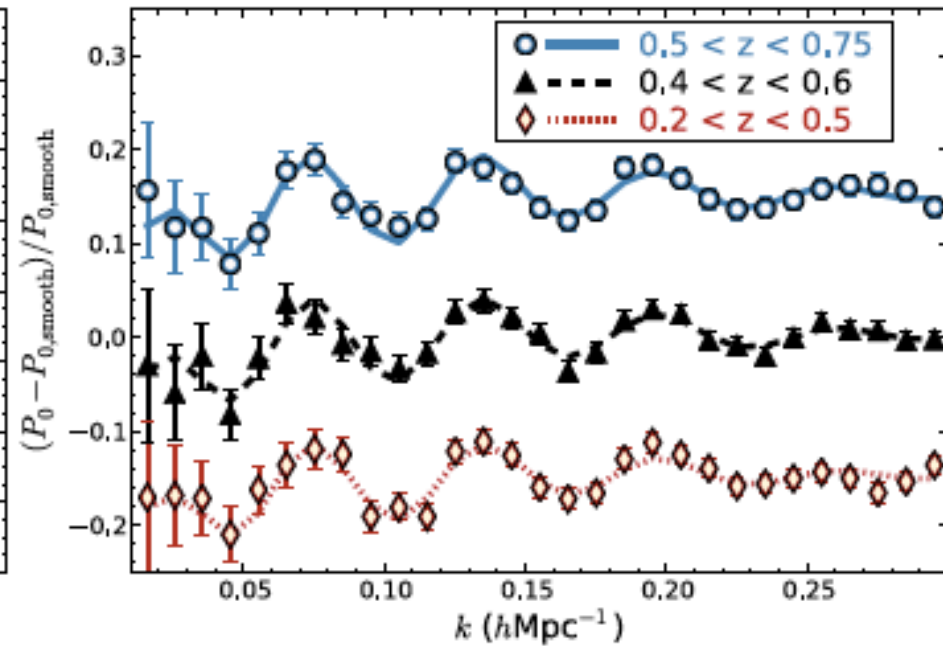


Latest results - BAO (BOSS DR12)

Not only in the correlation function



But also in the power spectrum



Alam+ arXiv:1607.03155