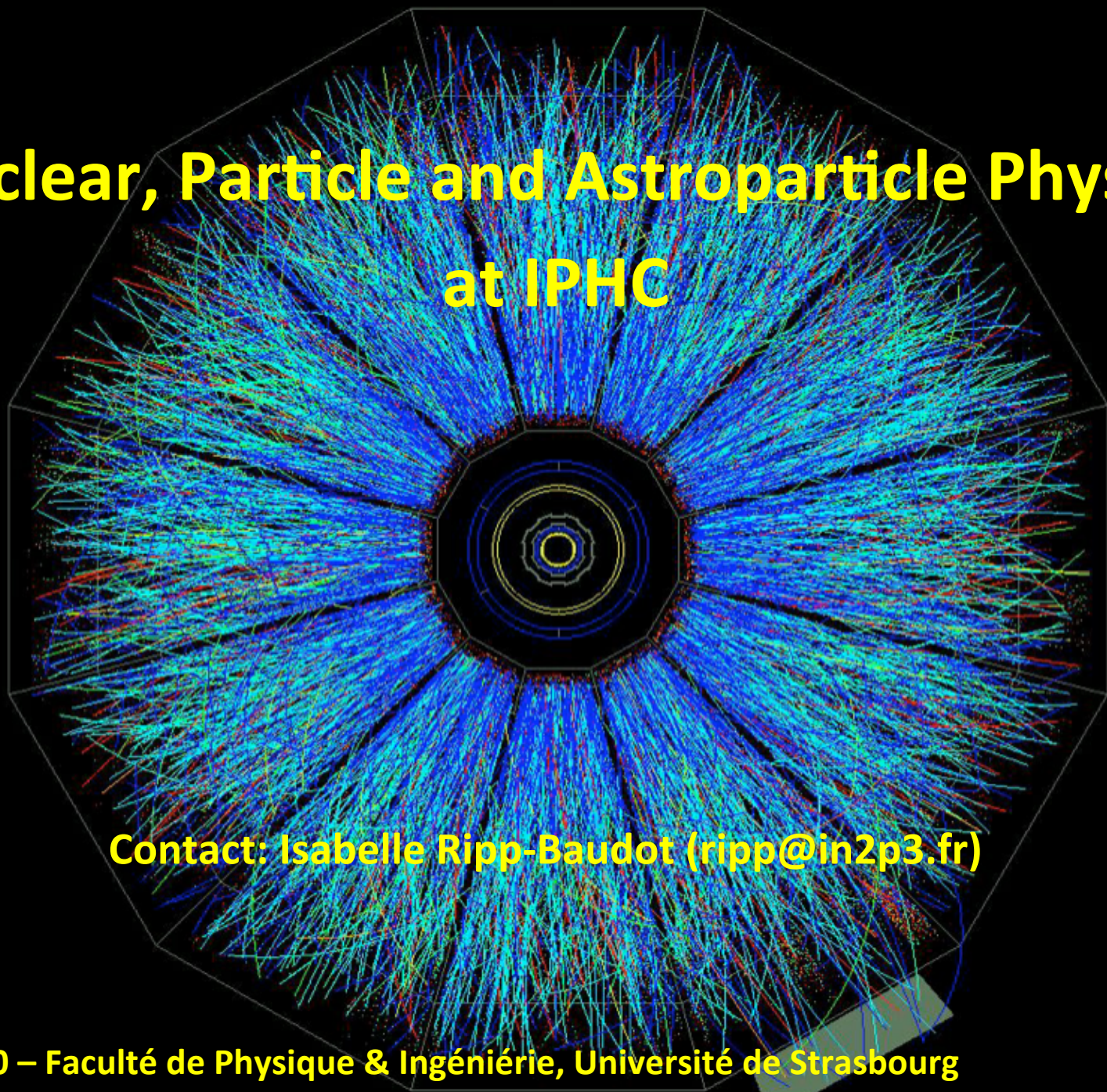


Nuclear, Particle and Astroparticle Physics at IPHC



Contact: Isabelle Ripp-Baudot (ripp@in2p3.fr)

Institut Pluridisciplinaire Hubert Curien



IPHC
Institut Pluridisciplinaire
Hubert CURIEN
STRASBOURG

IN2P3
INC
INEE
INSB

UNIVERSITÉ DE STRASBOURG

<http://www.iphc.cnrs.fr/>

Identity of IPHC

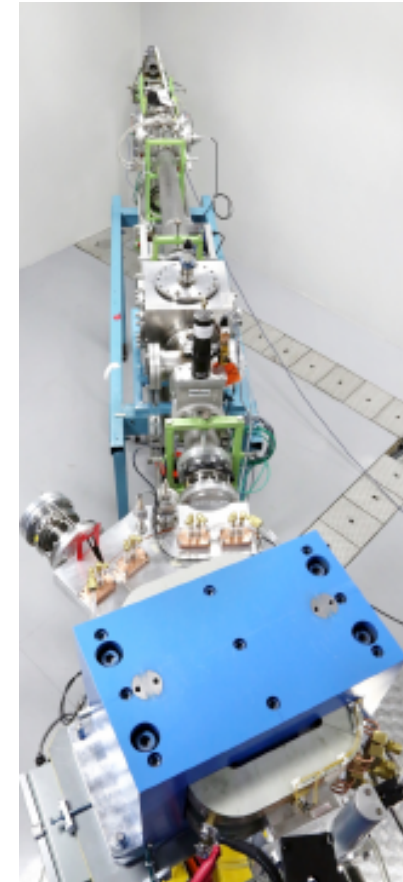


□ Two supporting institutions:

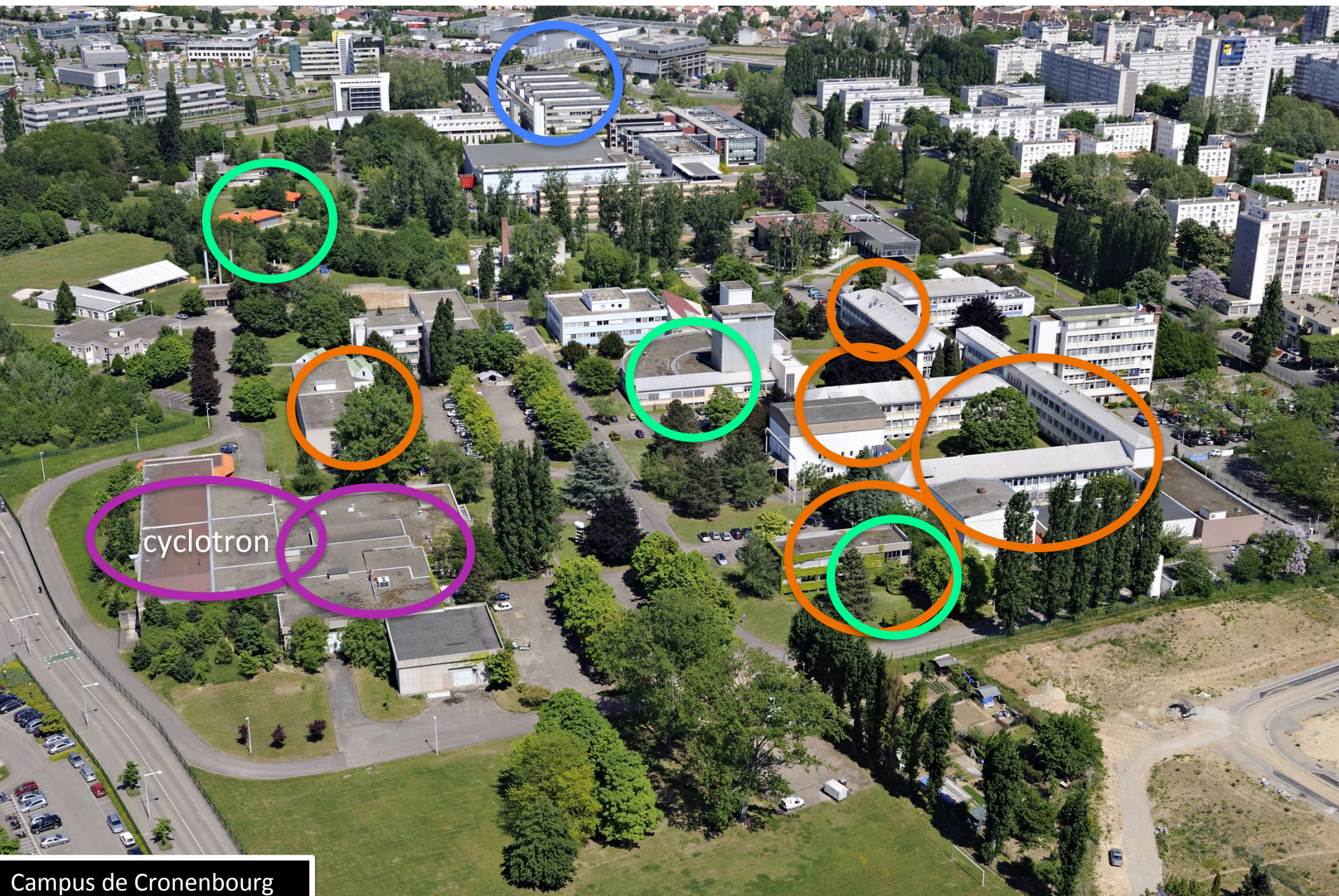
- **CNRS** : National Center for Scientific Research
with **IN2P3** = Institute of Nuclear Physics and Particle Physics
in CNRS including nowadays also Cosmology and Astroparticles.
- **University of Strasbourg**.
- Workforce of IPHC: **~400 personnel**.

□ **Multidisciplinary** research at IPHC:

- Present laboratory created in 2006 based on 3 former labs :
Biology - Chemistry - Subatomic Physics.
- **Cyclotron** built at IPHC in 2012: also Medical Doctors joined
IPHC to participate to research in Molecular Imaging,
Hadrontherapy and Radiobiology.



DRS DEPE DSA DRHIM



Campus de Cronenbourg

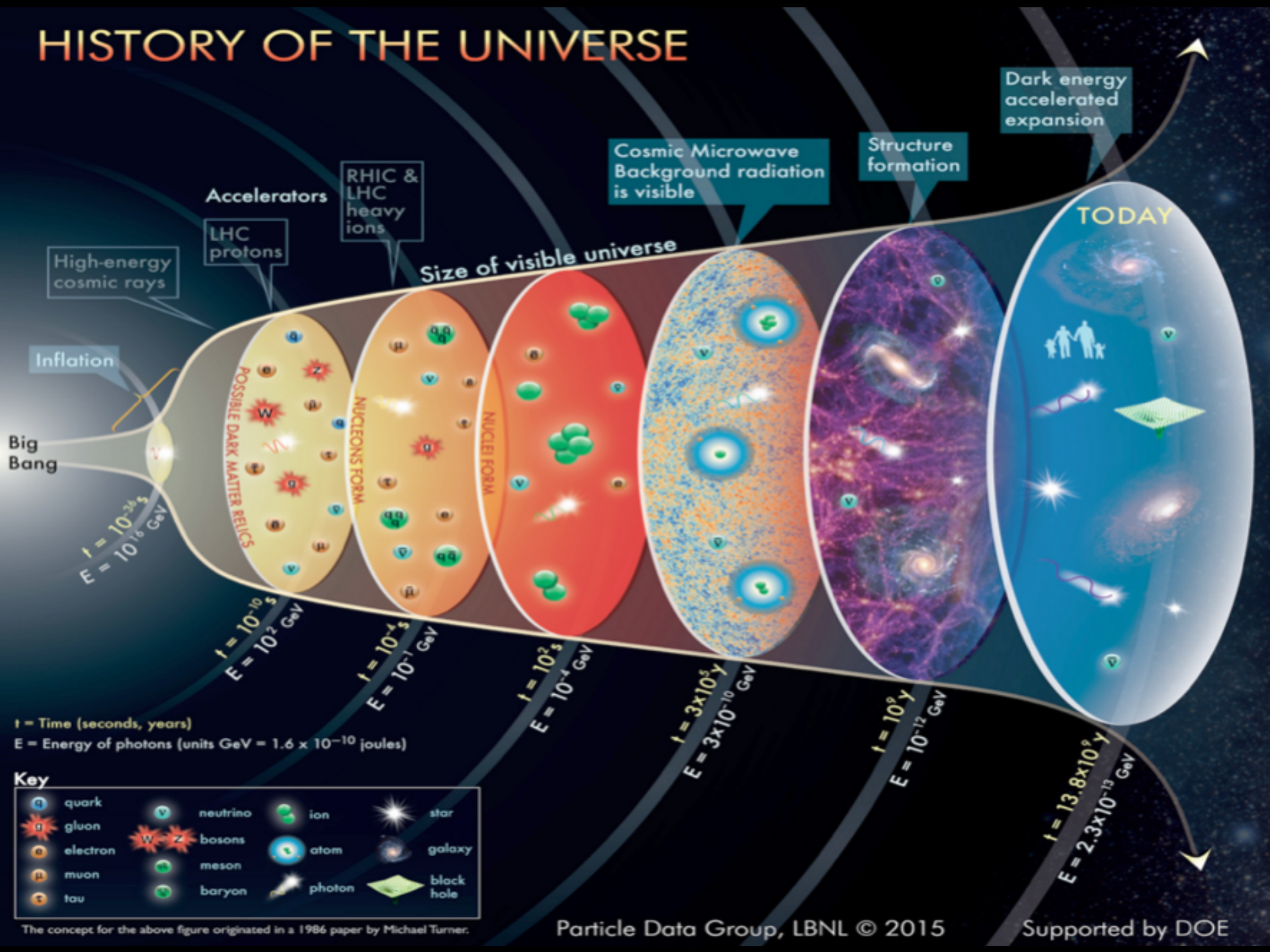


The Department of Subatomic Research

Subatomic Physics in Strasbourg: very old laboratory created after WWII.

~200 personnel: 70 researchers + 100 engineers + 30 students.

HISTORY OF THE UNIVERSE



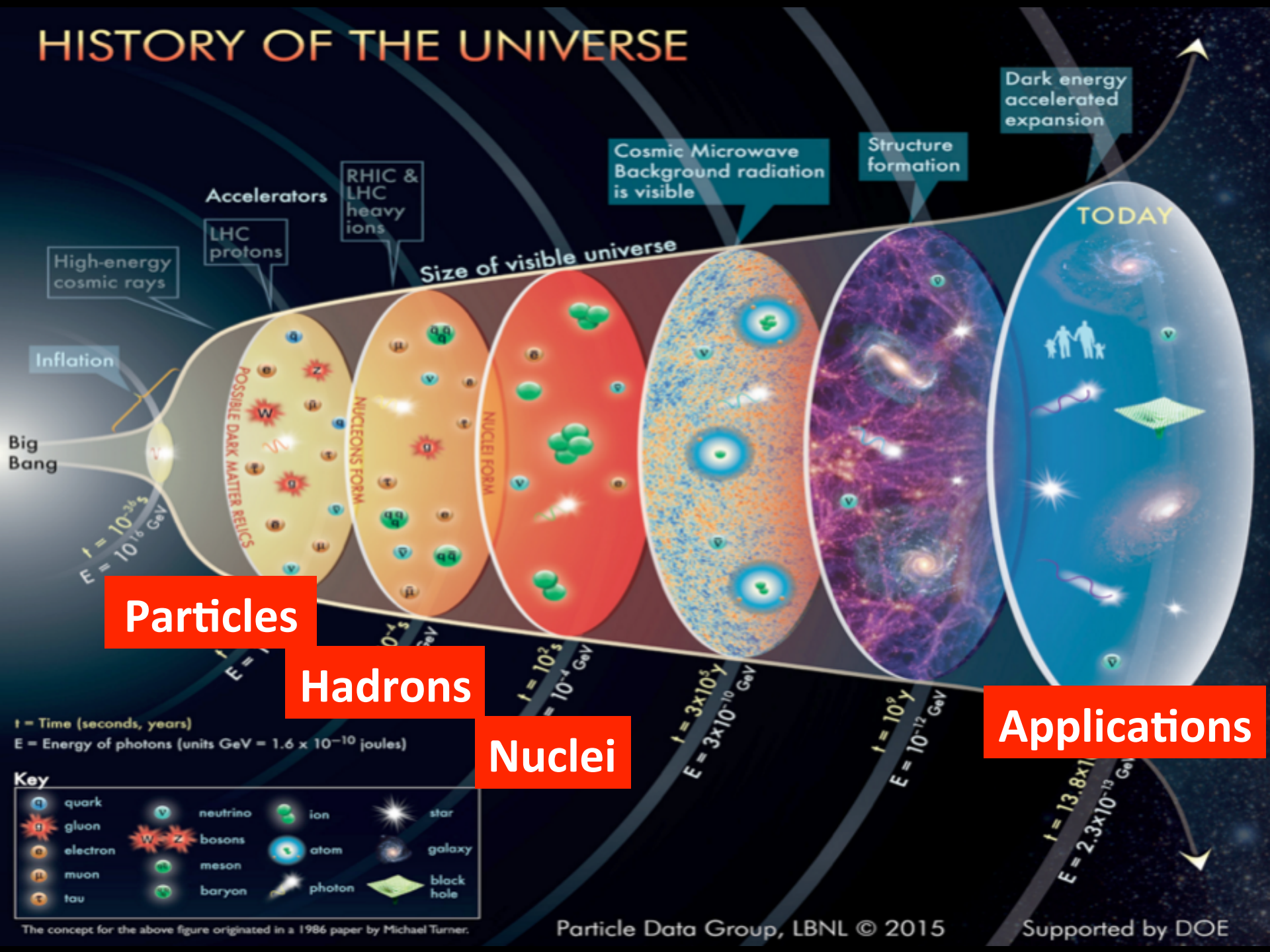
t = Time (seconds, years)
 E = Energy of photons (units GeV = 1.6×10^{-10} joules)

Key

quark	neutrino	ion	star
gluon	bosons	atom	galaxy
electron	meson	photon	black hole
muon	baryon		
tau			

The concept for the above figure originated in a 1986 paper by Michael Turner.

HISTORY OF THE UNIVERSE



Particles

Hadrons

Nuclei

Applications

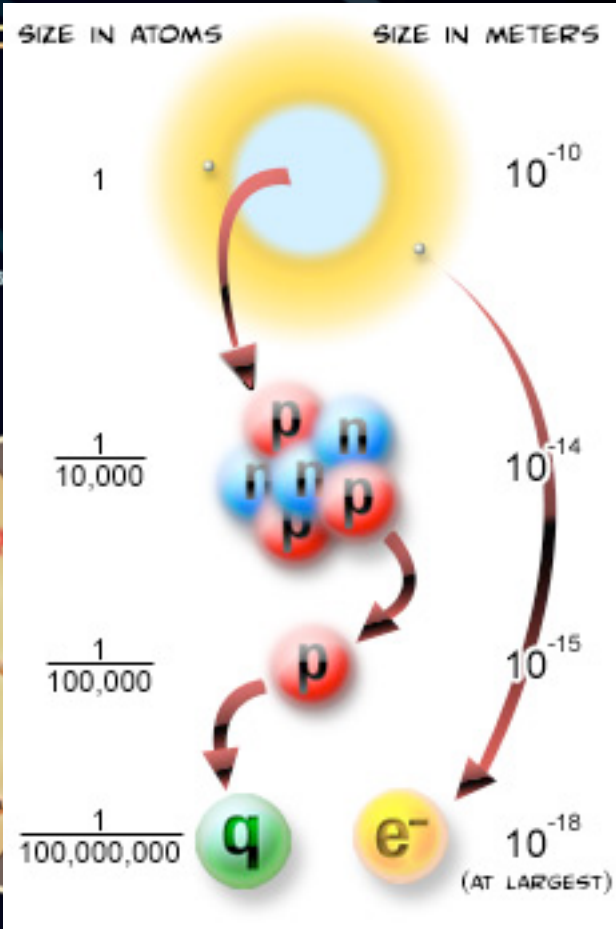
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	quark		neutrino		ion		star
	gluon		bosons		atom		galaxy
	electron		meson		photon		black hole
	muon		baryon				
	tau						

The concept for the above figure originated in a 1986 paper by Michael Turner.

HISTORY OF



Low energy radiation

Structure formation

Dark energy accelerated expansion

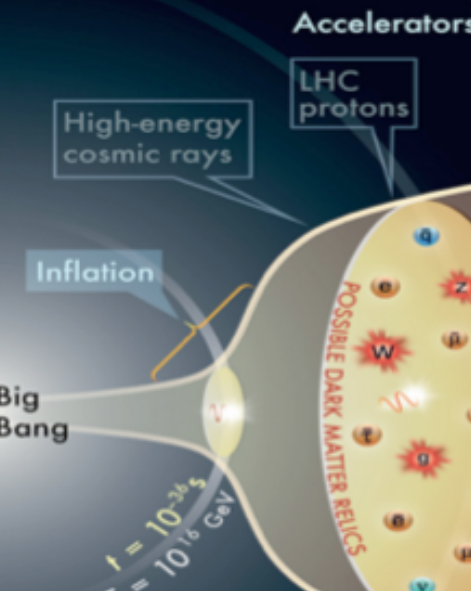
TODAY

Particles

Hadrons

Nuclei

Applications



t = Time (seconds, years)
E = Energy of photons (units GeV = 1.6×10^{-10} joules)

Key

quark	neutrino	ion	star
gluon	bosons	atom	galaxy
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tau			

The concept for the above figure originated in a 1986 paper by Michael Turner.

Particle physics (1)

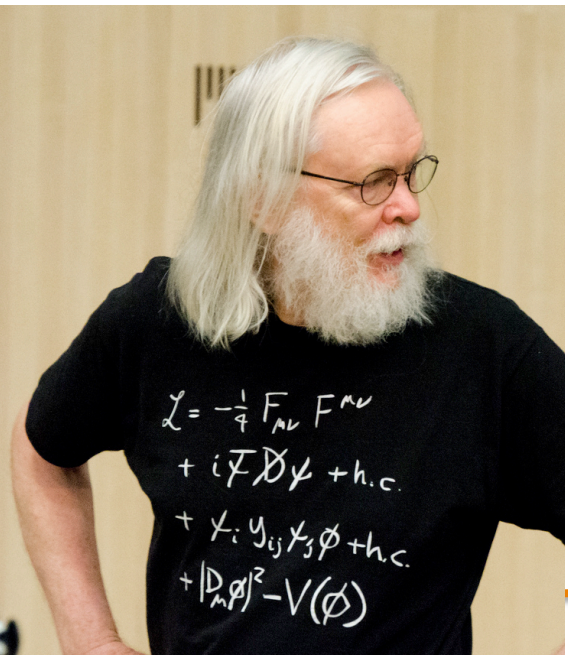
- Goal: understand the **content** (particles) and the fundamental **symmetries** of the **early universe**.
- Very precise description achieved with a **field theory** based on **special relativity** and **quantum mechanics**.



" who ordered that?"

(by I. Rabbi according to [Phys.Rept. 532 (2013) 27-64])

- Important concept: **effective theory**, valid in a given range (energy, distance, time).



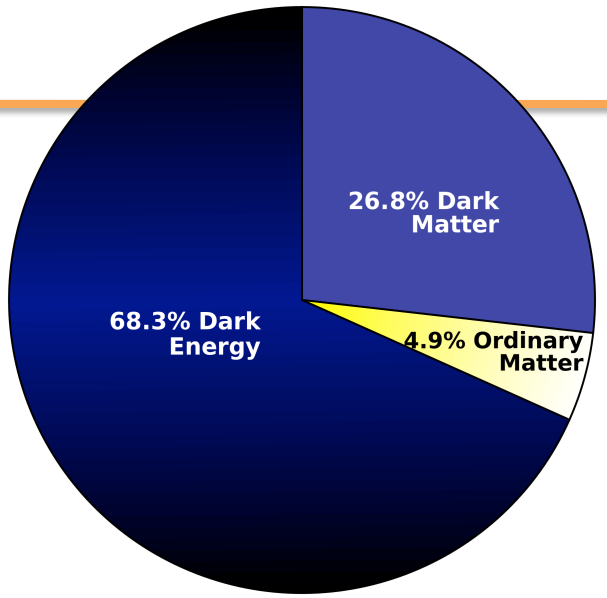
Particle physics (2)

What theory describes nature at higher energies

(younger universe)?

Still many questions, e.g.:

- Why did antimatter disappear?
- Nature of Dark Energy?
- Nature of Dark Matter?
- Particles of matter: why 3 families?
- How to include Gravitation in the theory?
- ...



	1 ^{ERE} GÉNÉRATION	2 ^{EME} GÉNÉRATION	3 ^{EME} GÉNÉRATION		
masse →	≈2.3 MeV/c ²	≈1.275 GeV/c ²	≈173.07 GeV/c ²	0	≈126 GeV/c ²
charge →	2/3	2/3	2/3	0	0
spin →	1/2	1/2	1/2	1	0
	u up	c charm	t top	g gluon	H boson de Higgs
	d down	s strange	b bottom	γ photon	
	e électron	μ muon	τ tau	Z boson Z	
	ν_e neutrino électronique	ν_μ neutrino muonique	ν_τ neutrino tauique	W[±] bosons W [±]	

QUARKS (left side of the table)

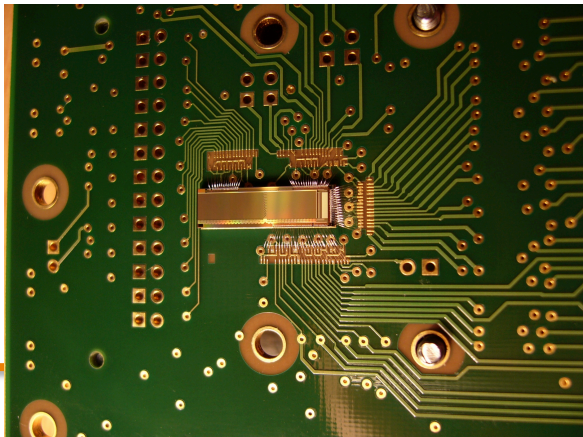
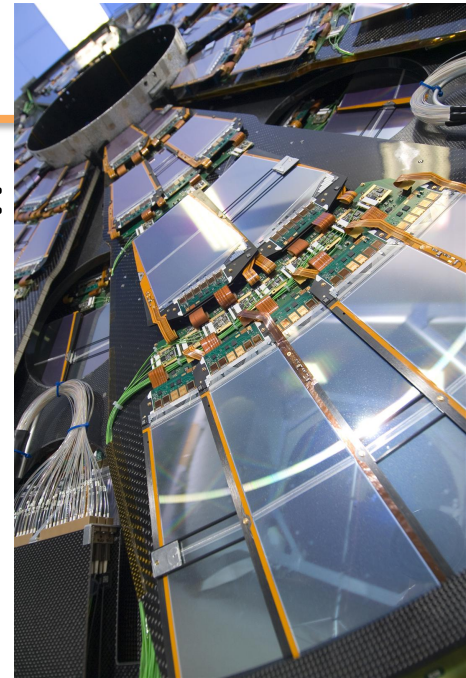
LEPTONS (left side of the table)

BOSONS DE JAUGE (right side of the table)

Particle physics (3)

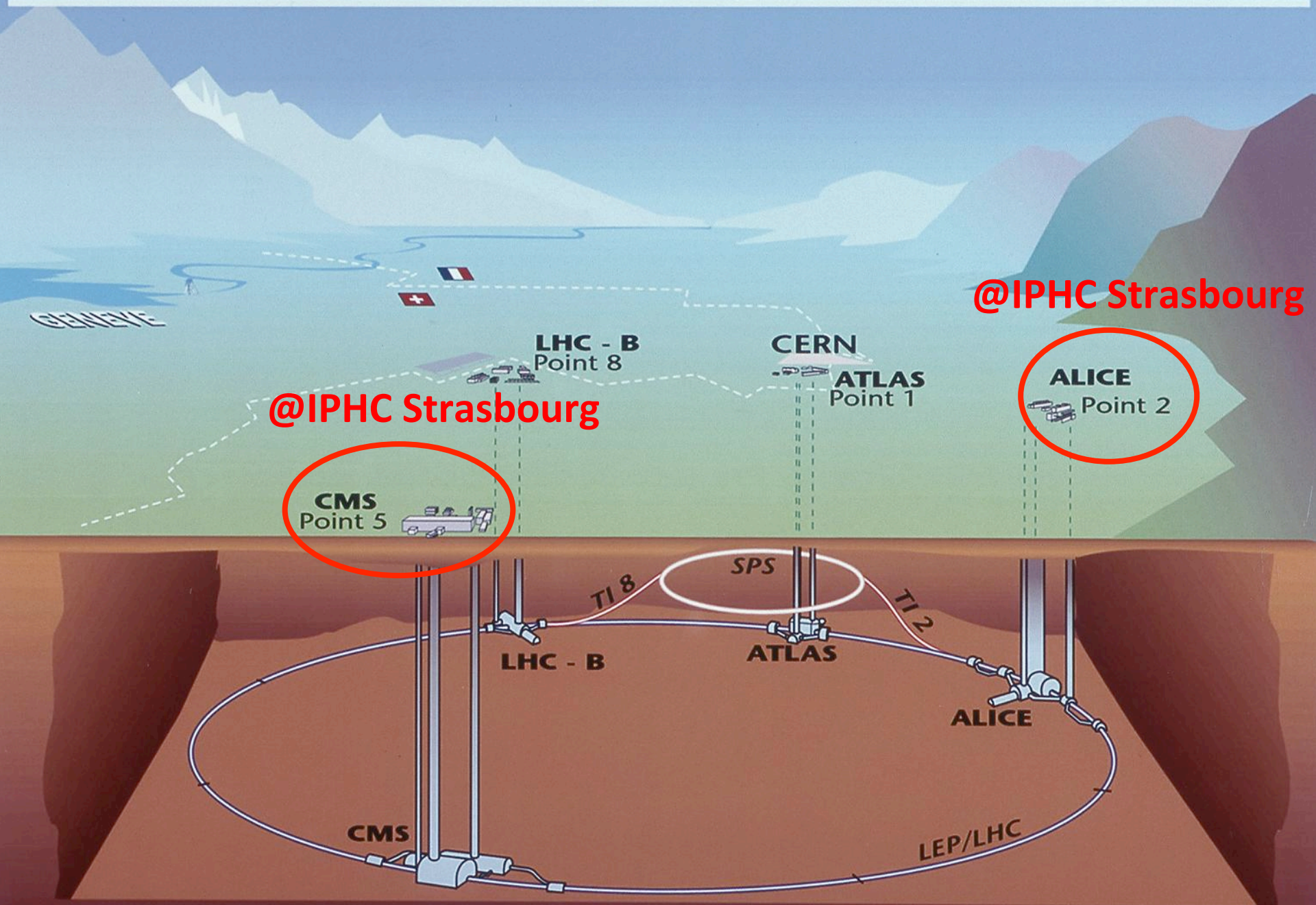
□ Different approaches at IPHC to answer these questions:

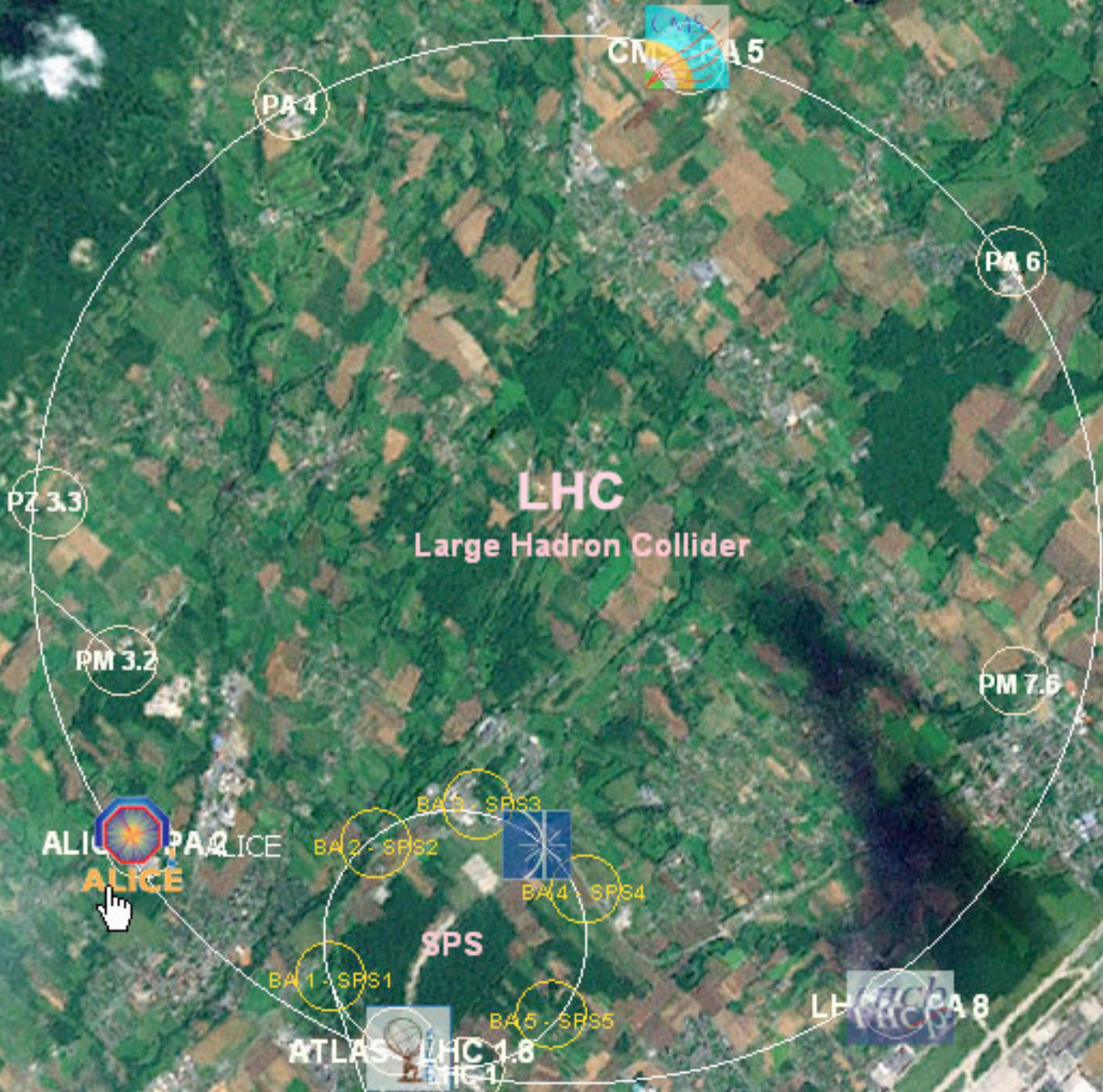
- Collisions at the **energy frontier** (LHC pp collisions at CERN): **relativistic production** of unknown particles.
- Collisions at the **intensity frontier** (SuperKEKB e^+e^- collisions in Japan): sensitive to **quantum manifestations** of unknown particles.
- **Neutrino physics**: very elusive particle, also used as a **messenger of the universe**.
- Study of the **Quark and Gluon Plasma** with Heavy Ion collisions at LHC: new state of matter resulting from Quantum Chromodynamics (strong interaction).



- **R&D of new instruments** to improve sensitivity to very rare processes and spatial and time precision.

Overall view of the LHC experiments.





CM PA 5

PA 4

PA 6

LHC

Large Hadron Collider

PZ 3.3

PM 3.2

PM 7.6

ALICE PA 2

ALICE

BA 2 - SPS2

BA 3 - SPS3

BA 4 - SPS4

SPS

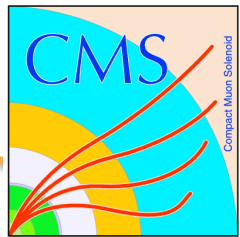
BA 1 - SPS1

BA 5 - SPS5

ATLAS LHC 1.8

LHC PA 8





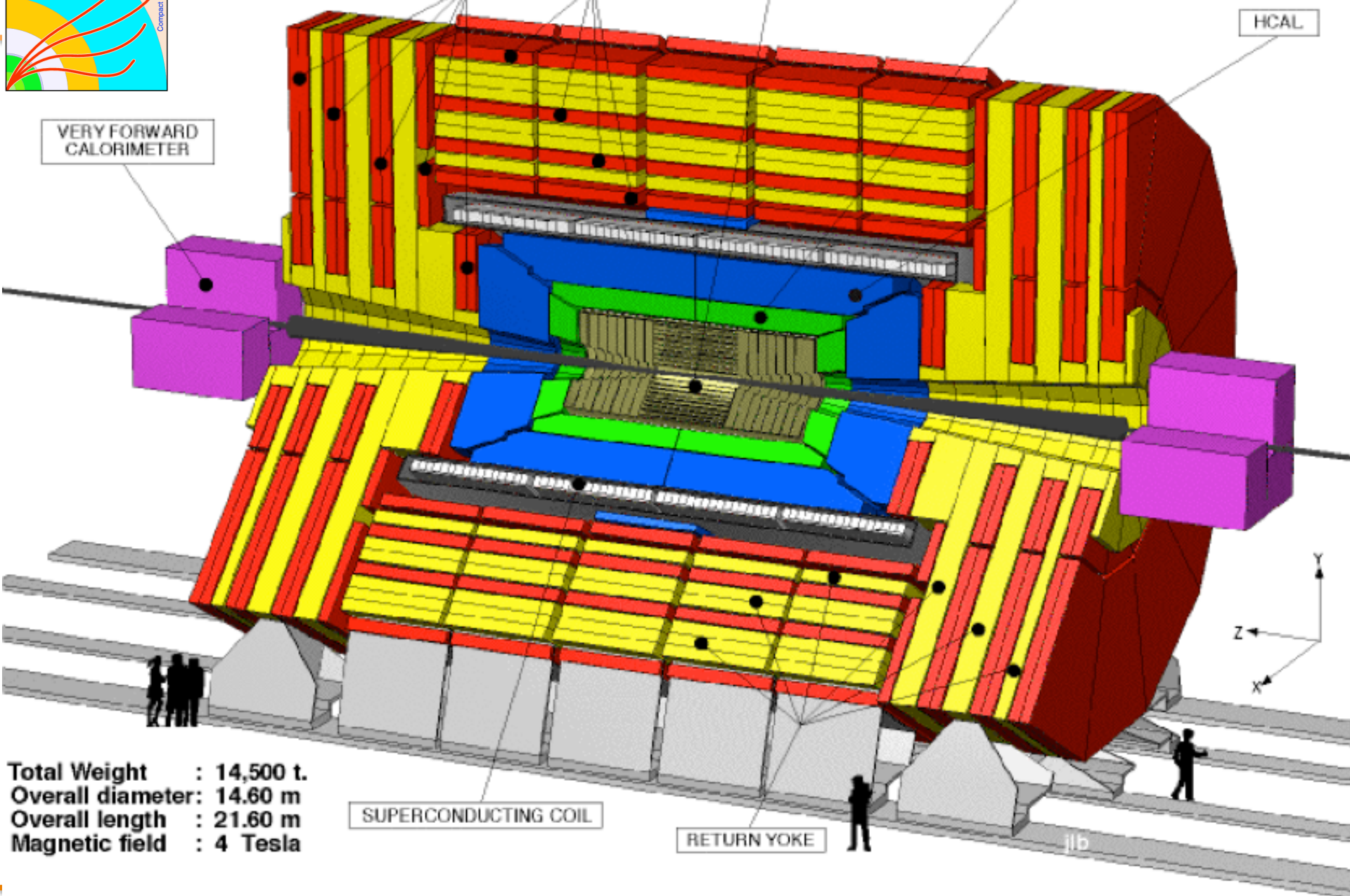
VERY FORWARD CALORIMETER

MUON CHAMBERS

INNER TRACKER

CRYSTAL ECAL

HCAL



SUPERCONDUCTING COIL

RETURN YOKE

Total Weight : 14,500 t.
Overall diameter: 14.60 m
Overall length : 21.60 m
Magnetic field : 4 Tesla

jlb

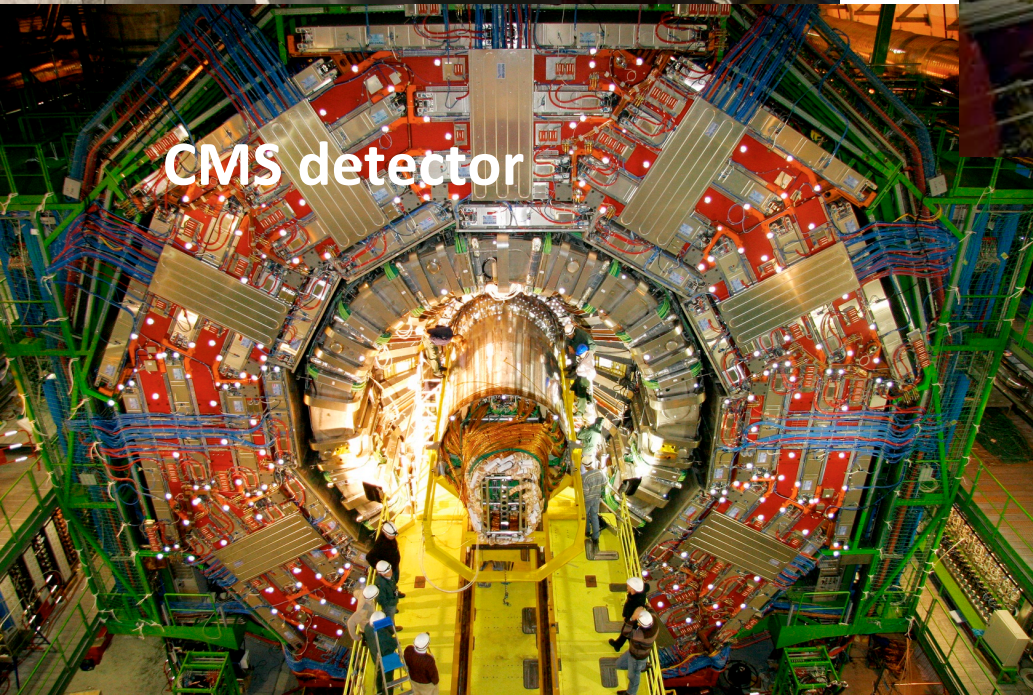
Collider physics



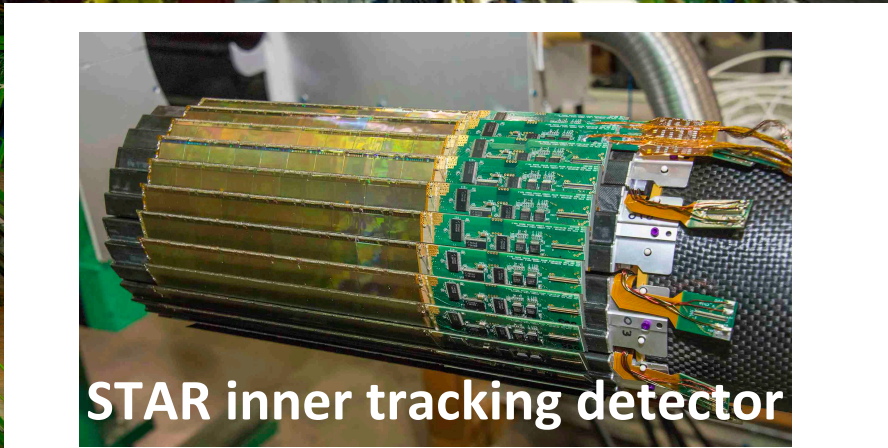
LHC collider



ALICE detector

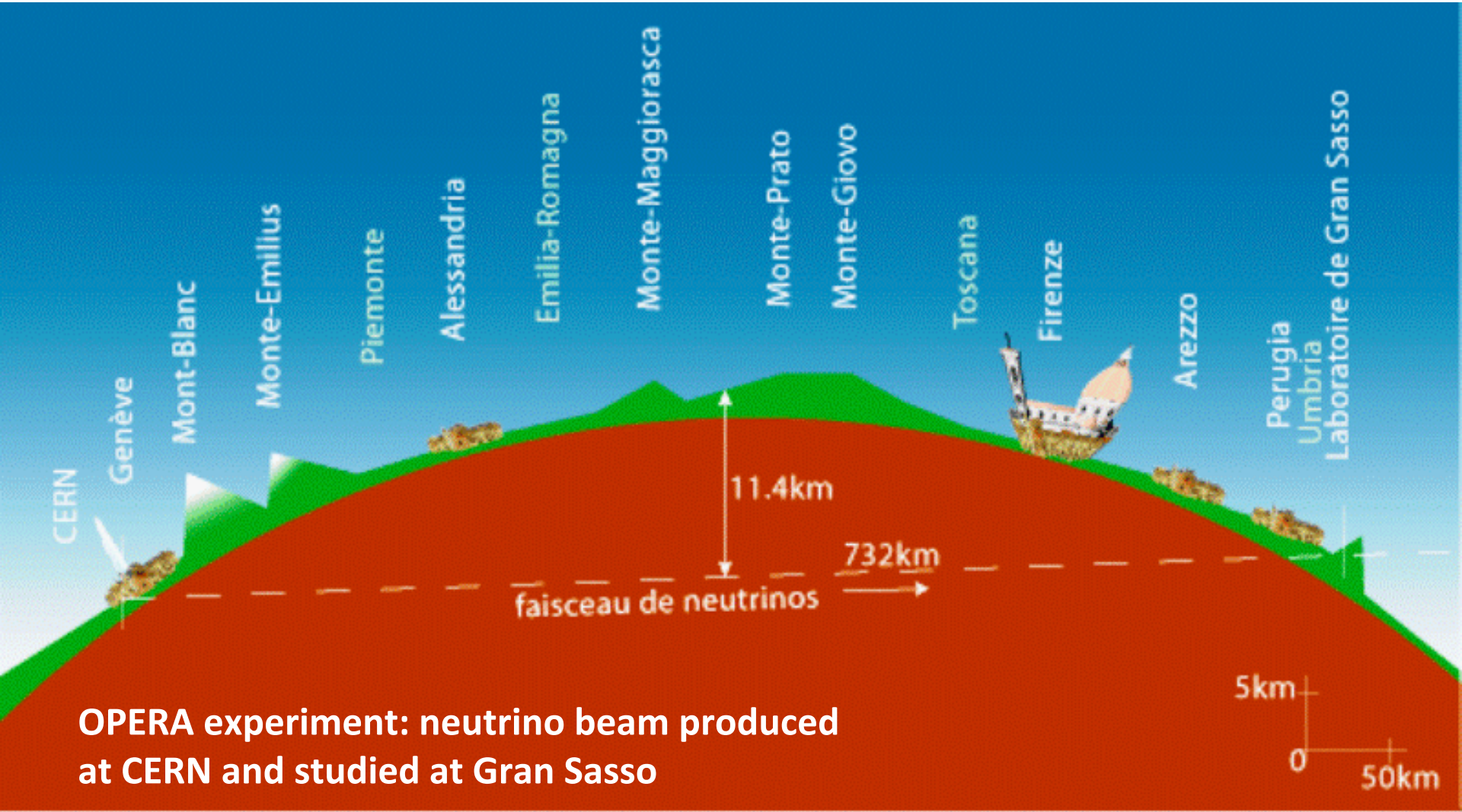


CMS detector



STAR inner tracking detector

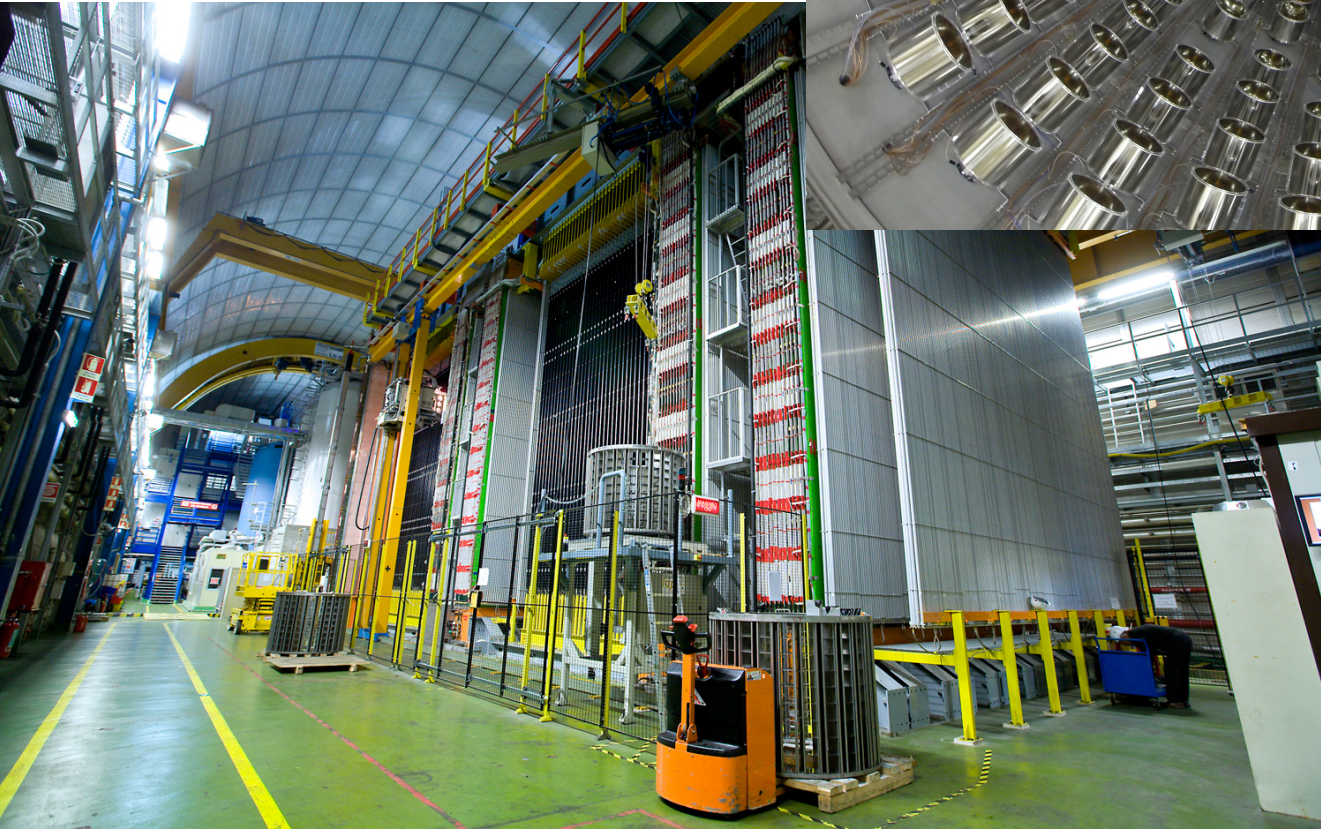
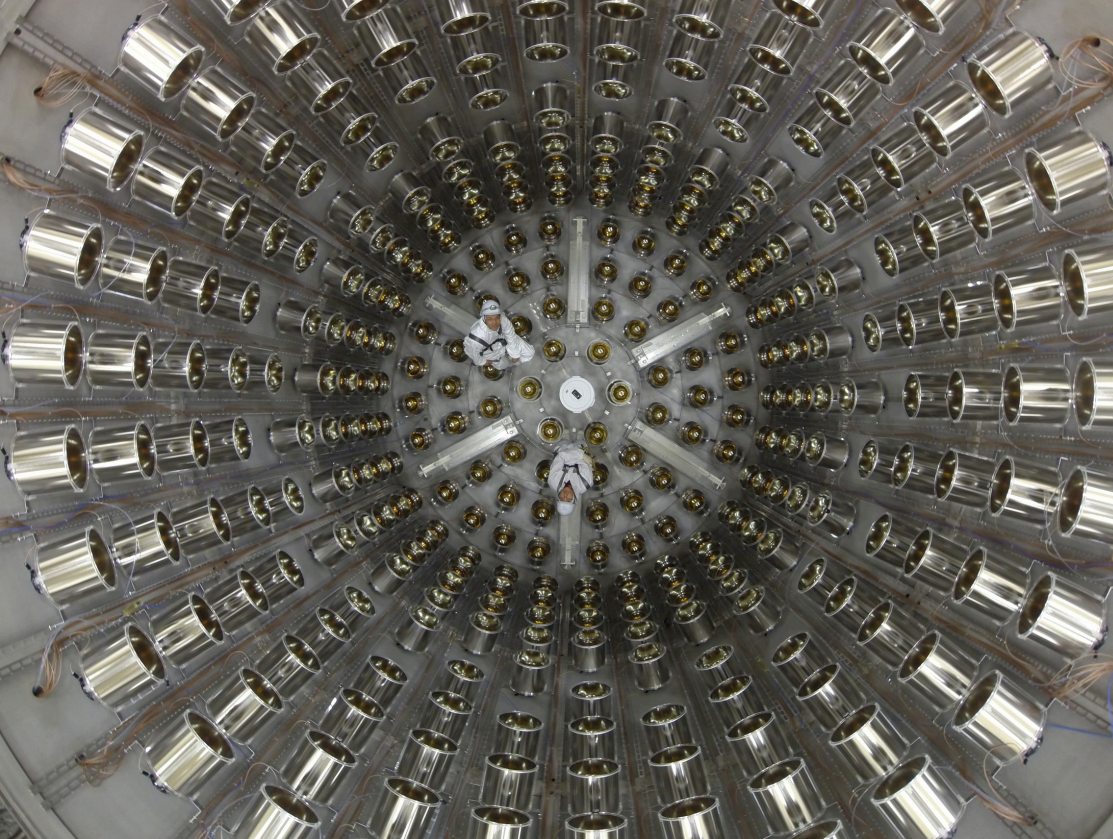
Neutrino physics



OPERA experiment: neutrino beam produced at CERN and studied at Gran Sasso

Neutrino physics

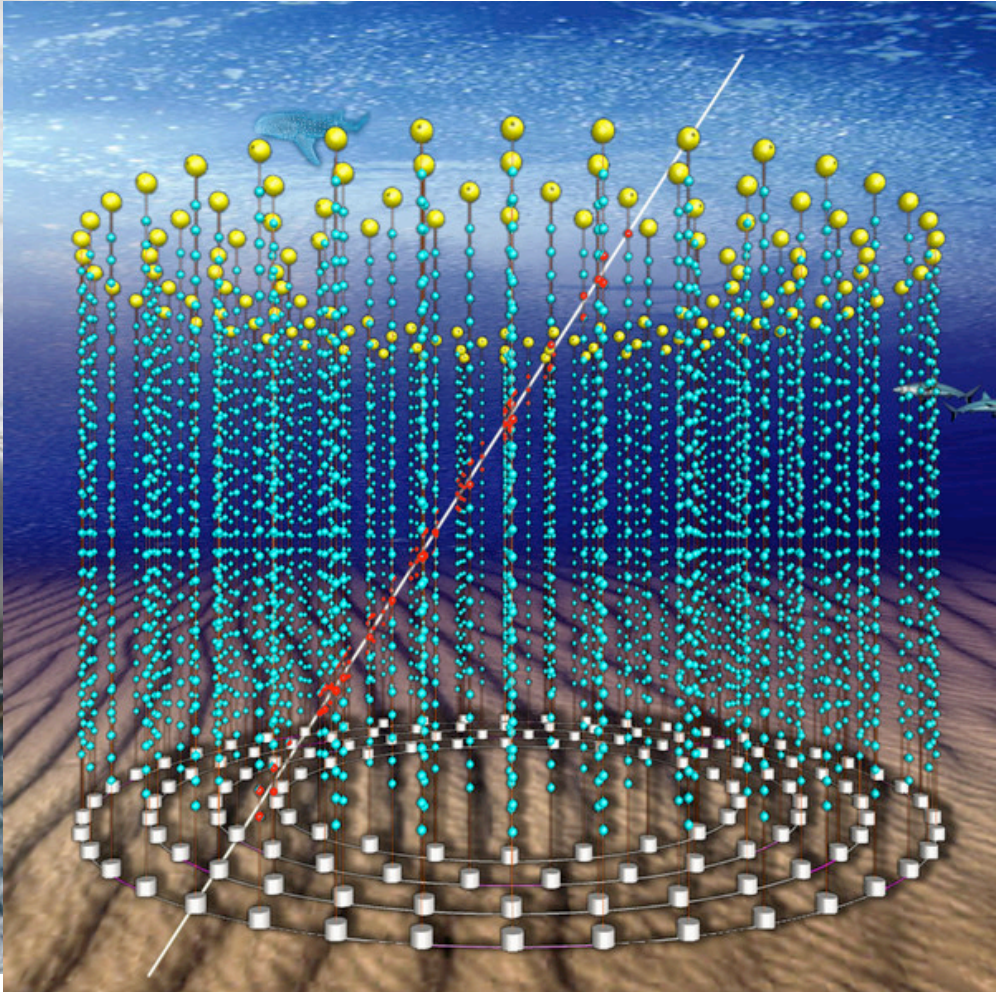
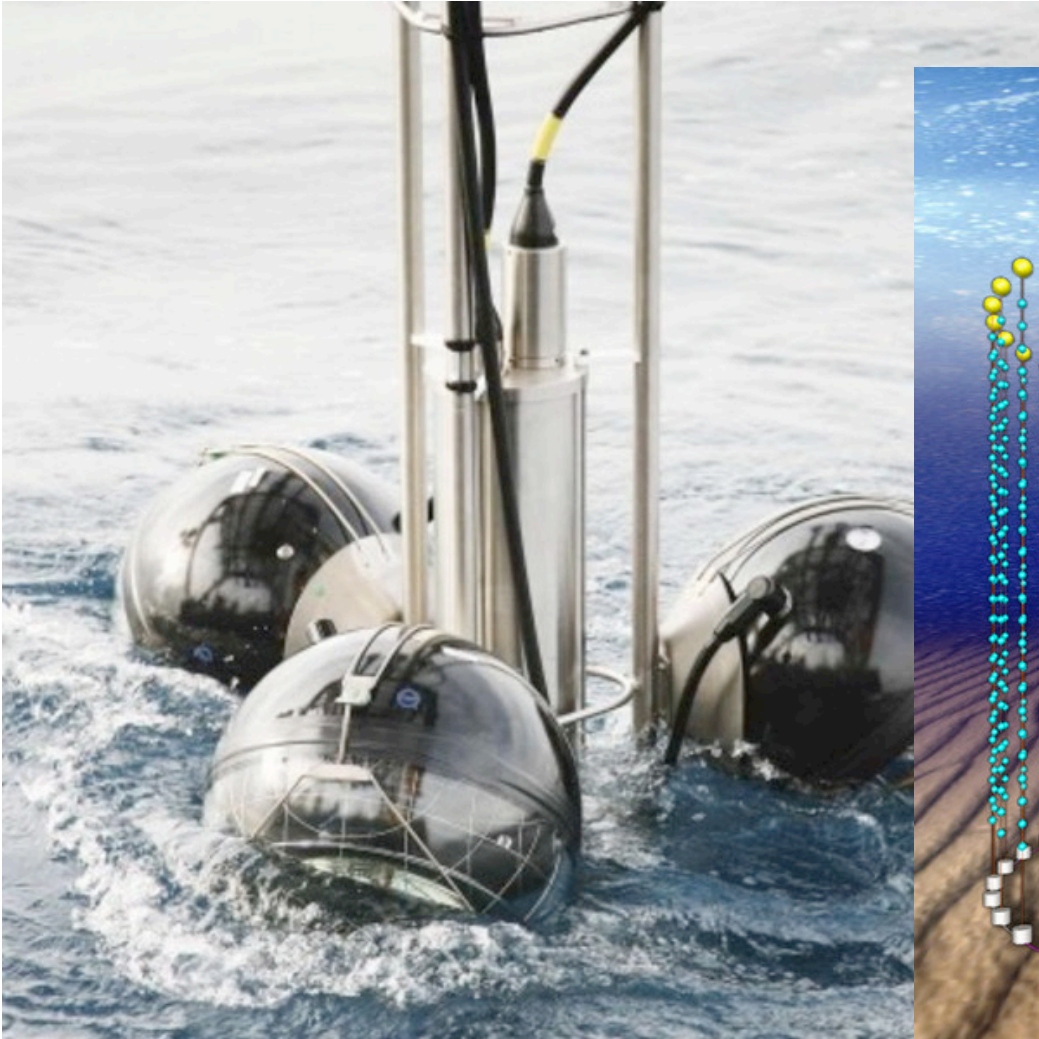
Double-Chooz detector of neutrinos produced by the Chooz nuclear plant. →



← OPERA detector of neutrinos produced at CERN, after a 723 km travel in earth.

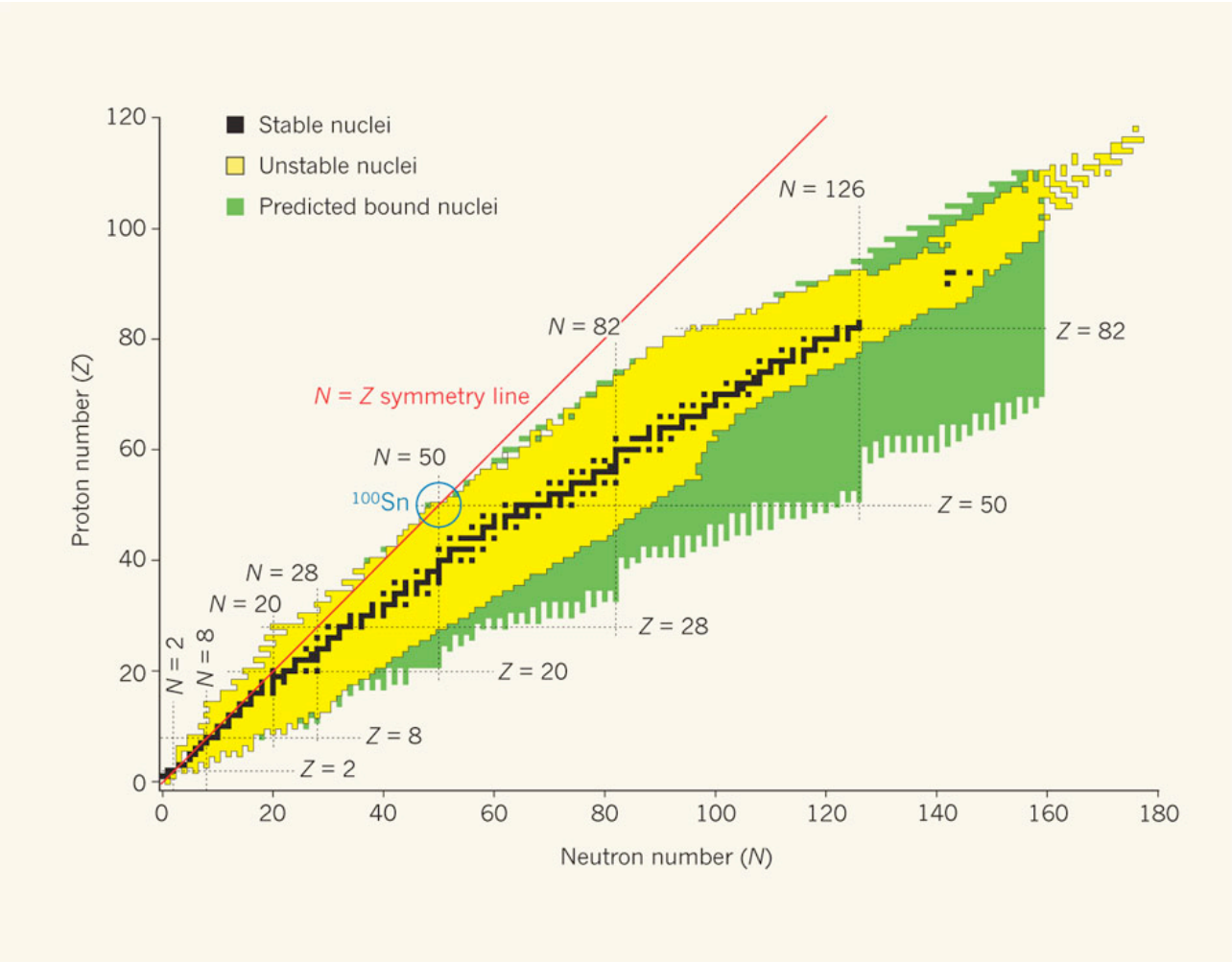
Neutrino physics

ANTARES and KM3NeT submarine telescopes to detect atmospheric neutrinos.



Nuclear physics (1)

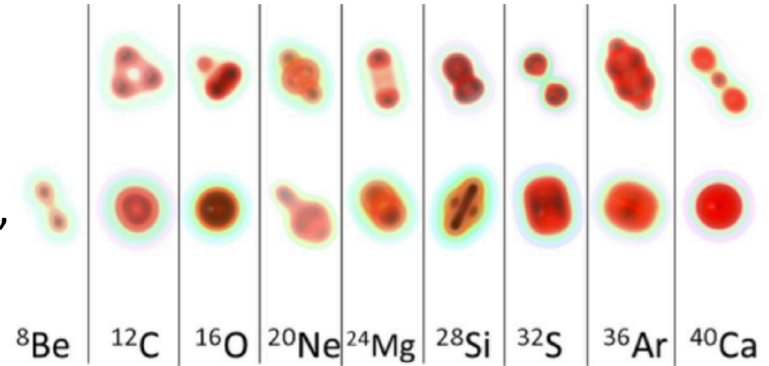
□ Goal: why are **nuclei stable**? How were they **created**?



Nuclear physics (2)

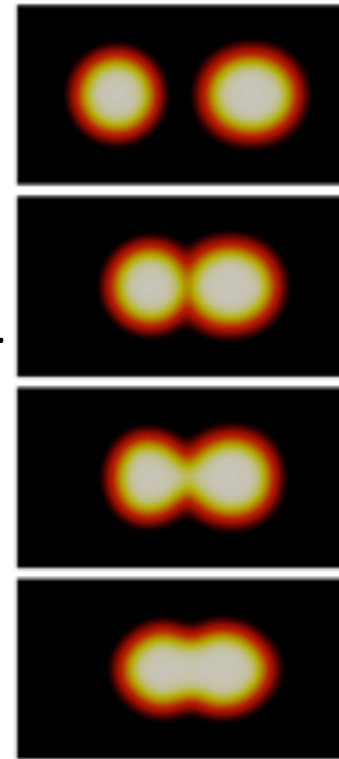
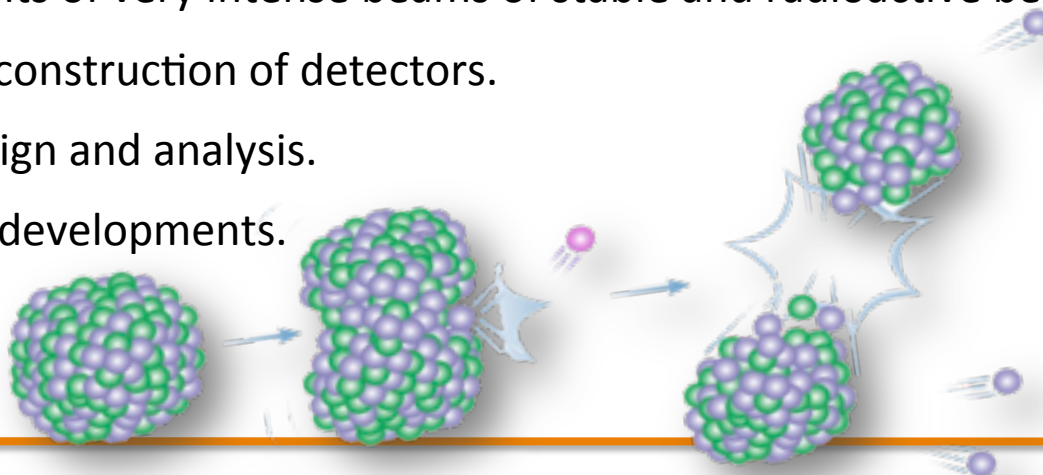
□ Main questions:

- Limits of stability as a function of the nucleus **shape**, its **mass** (study of superheavy elements), the **asymmetry** between numbers of neutrons vs. protons.
- How are the **light nuclei produced** in the early universe?
- How are the **heavy nuclei produced** in stars?



□ Research activities:

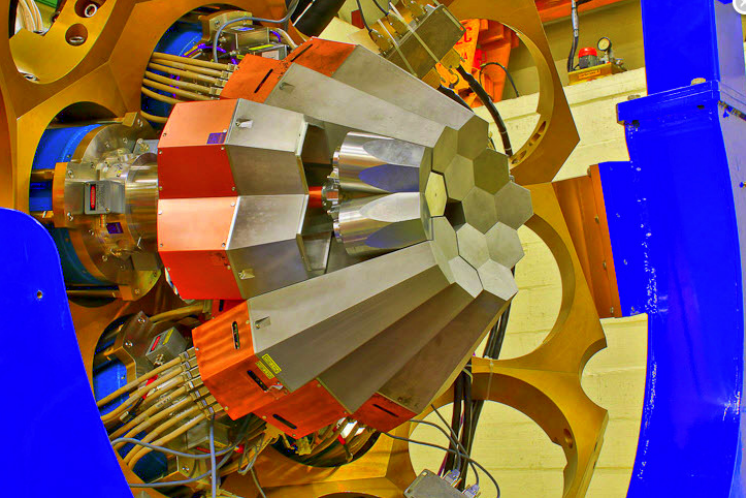
- Developments of very intense beams of stable and radioactive beams.
- Design and construction of detectors.
- Data campaign and analysis.
- Theoretical developments.



Nuclear physics (3)



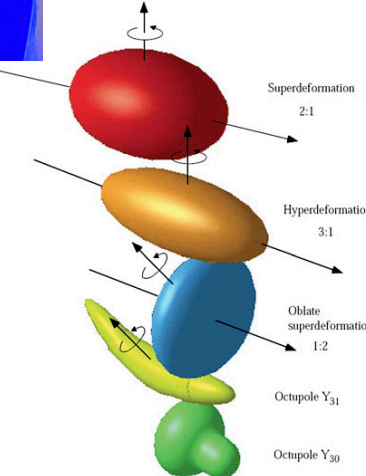
← Development of new intense beams of ^{50}Ti to produce superheavy nuclei.



↑
AGATA 4π gamma-ray spectrometer.



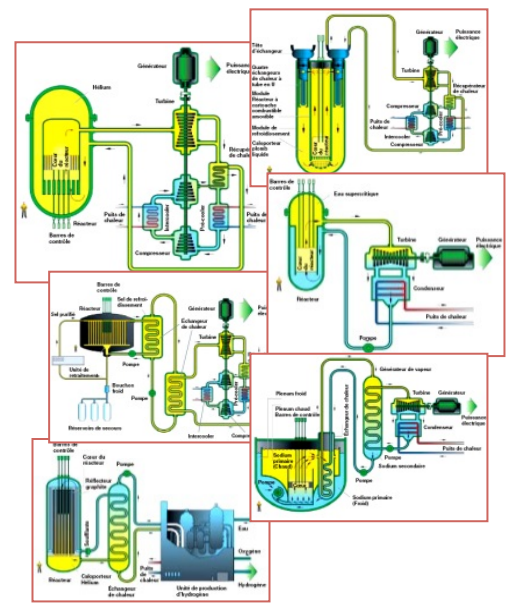
← Andromède accelerator (Orsay).



Applications to Energy, Environment and Health

□ Use of fundamental knowledge in subatomic physics and radiochemistry to answer societal issues:

- Nuclear energy: nuclear data needed to simulate and design the next generation of nuclear reactors.
- Dosimetry and metrology of ionising radiations: radioactive pollution, radiation worker safety, ionised food, dose monitoring during hadrontherapy.
- Radiochemistry: ground pollution, chemical modification and impact of ionising radiation on organic matter.



The Master of Subatomic and Astroparticle Physics in Strasbourg

UNIVERSITY OF STRASBOURG

MASTER OF SCIENCE SUBATOMIC AND ASTROPARTICLE PHYSICS

2nd year of master programme in Physics

Particles

Nuclei

Cosmology

Astroparticle

Prepare for PhD studies at world-class facilities, like:

- LHC collider at CERN in Switzerland,
- SPIRAL at GANIL in France,
- the large telescope HESS in South-Africa.
- the Fermi satellite...

Follows lectures on

- Quantum Field Theory
- Radiation interactions

Learns proactively
The Master thesis (

Conditions for applications: successful Master 1 in physics or equivalent.

Contacts: Prof. Jérôme BAUDOT (baudot@in2p3.fr)

Student with high academic achievements can apply for a grant of excellence.



master-psa.u-strasbg.fr



Hosted by IPHC

- Hosting laboratory: IPHC / DRS.
- Content:
 - Common and chosen lectures,
 - Project in Physics: 1 month,
 - Research internship: 4 months,
 - Both theoretical and experimental points of view.
- After the Master:
 - PhD thesis (possible grants).
 - Industry: nuclear power, Big Data, nuclear metrology, simulation, ...

Faculty of physics & engineering
3 rue de l'Université
67000 Strasbourg
France



Thank you for your attention



Cosmology

DARK MATTER

Neutrino ~~C~~

PROTON DECAY

AXIONS