The state-of-the-art of the Hadronic Uncertainties in $B o K^*\mu\mu$

Bernat Capdevila

Institut de Física d'Altes Energies (IFAE)

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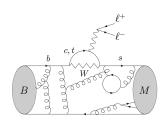
Current Trends in Flavour Physics (Paris)

In collaboration with: S. Descotes-Genon, L. Hofer, J. Matias & J. Virto Based on 1701.08672 JHEP (2017)



- 1. Introduction
- 2. Updated global fit results
- 3. Anatomy of power corrections
- 4. Long-distance $c\bar{c}$ loops from fits to $B o K^* \mu \mu$ data
- 5. Conclusions

Theoretical framework



 $A \sim C_i$ (short dist.)

× Hadronic Matrix Elements (long dist.)

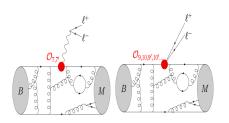
$b \to s \gamma^{(*)}$ Effective Hamiltonian

$$\mathcal{H}_{\Delta F=1}^{\mathsf{SM}} \propto V_{ts}^* V_{tb} \sum_{\cdot} C_i \mathcal{O}_i$$

$$C_{10}^{\text{SM}}(\mu_b) = -4.3 \qquad (\mu_b = m_b)$$

- ⇒ In this picture, New Physics (NP) effects can enter through two mechanisms:
 - Extra contributions to the WCs.
 - Additional effective operators: \mathcal{O}'_i , \mathcal{O}_S , \mathcal{O}_P , \mathcal{O}_T ,...

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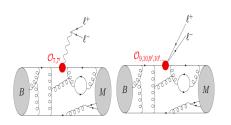
$$C_7^{\text{SM}}(\mu_b) = -0.29$$
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 \Rightarrow QCD factorisation (QCDF): Large recoil symmetries ($E_{K^*} \to \infty$ and HQL),

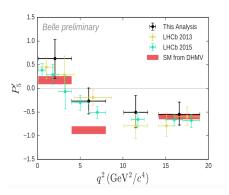
$$\{V, A_0, A_1, A_2, T_1, T_2, T_3\} \sim \{\xi_{\perp}, \xi_{\parallel}\}$$

up to $\mathcal{O}(\alpha_s)$ and $\mathcal{O}(\Lambda/m_B)$.



The P_5' anomaly

 $b o s \ell \ell$ driven processes have provided some interesting anomalies during the recent years.



 P_5' was proposed in DMRV, JHEP 1301(2013)048

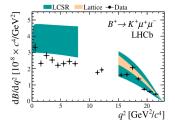
$$P_5' = \sqrt{2} \frac{\text{Re}(A_0^L A_\perp^{L*} - A_0^R A_\perp^{R*})}{\sqrt{|A_0|^2(|A_\perp|^2 + |A_\parallel|^2)}}$$

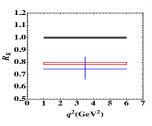
Optimized Obs.: Soft form factor (ξ_{\perp}) cancellation at LO.

- **2013**: 1fb^{-1} dataset LHCb found 3.7σ
- **2015**: 3fb^{-1} dataset LHCb found 3σ in 2 bins.
- Belle confirmed it in a bin [4,8] few months ago.

Other tensions beyond P_5'

	bin	SM	EXP	Pull
$10^7 \times BR(B_s \to \phi \mu^+ \mu^-)$	[2,5] [5,8] [15,18.8]	$\begin{array}{c} 1.55 \pm 0.33 \\ 1.89 \pm 0.40 \\ 2.20 \pm 0.17 \end{array}$	$\begin{array}{c} 0.77 \pm 0.14 \\ 0.96 \pm 0.15 \\ 1.62 \pm 0.20 \end{array}$	+2.2 +2.2 +2.2
$10^7 \times BR(B^0 \to K^0 \mu^+ \mu^-)$	[15,19]	0.91 ± 0.12	0.67 ± 0.12	+1.4
$10^7 \times BR(B^0 \to K^{*0} \mu^+ \mu^-)$	[16,19]	1.66 ± 0.15	1.23 ± 0.20	+1.7
${10^7 \times BR(B^+ \rightarrow K^{*+}\mu^+\mu^-)}$	[15,19]	2.59 ± 0.25	1.60 ± 0.32	+2.5





- Deviations in $BR(B_s \to \phi \mu \mu)$.
- Several systematic low-recoil small tensions in BR_{μ} .
- $BR(B o K\mu\mu)$ small compared to SM predictions.
- $\begin{array}{c} \blacksquare \ \ R_{K} = 0.745^{+0.090}_{-0.074} \\ \pm 0.036 \end{array}$
 - \Rightarrow SM Prediction: $R_K = 1 \ (2.6\sigma)$.
 - ⇒ Possible lepton-flavour universality violation.

Updated fits

Includes updated $BR(B \to K^* \mu^+ \mu^-)$ + corrected BSZ for $B_s \to \phi \mu^+ \mu^-$. $P_5^{\prime \mu \rm BELLE}$ adds from +0.1 to $+0.3\sigma$.

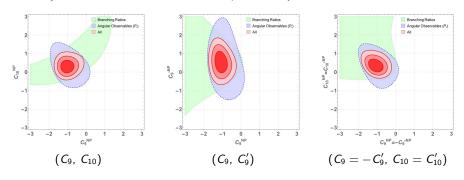
Coefficient	Best fit	1σ	$\mathrm{Pull}_{\mathrm{SM}}$	p-value (%)
$\mathcal{C}_7^{\mathrm{NP}}$	-0.02	[-0.04, -0.00]	1.1	13.0
$\mathcal{C}_{9}^{ ext{NP}}$	-1.05	[-1.25, -0.85]	4.7	61.0
$\mathcal{C}_{10}^{ ext{NP}}$	0.55	[0.34, 0.77]	2.8	24.0
$\mathcal{C}^{ ext{NP}}_{7'}$	0.02	[-0.00, 0.04]	0.9	13.0
$\mathcal{C}_{9'}^{ ext{NP}}$	0.06	[-0.18, 0.30]	0.3	12.0
$\mathcal{C}_{10'}^{ ext{NP}}$	-0.03	$\left[-0.20, 0.14\right]$	0.2	12.0
$\mathcal{C}_9^{ ext{NP}}=\mathcal{C}_{10}^{ ext{NP}}$	-0.18	[-0.36, 0.02]	0.9	13.0
$\mathcal{C}_{9}^{\mathrm{NP}} = -\mathcal{C}_{10}^{\mathrm{NP}}$	-0.59	[-0.74, -0.44]	4.3	51.0
$\mathcal{C}_{9'}^{ ext{NP}} = -\mathcal{C}_{10'}^{ ext{NP}}$	0.03	[-0.08, 0.13]	0.2	12.0
$\mathcal{C}_{9}^{ ext{NP}} = -\mathcal{C}_{9'}^{ ext{NP}}$	-1.00	[-1.20, -0.78]	4.4	54.0
$\mathcal{C}_{9}^{ ext{NP}} = -\mathcal{C}_{10}^{ ext{NP}} \\ = -\mathcal{C}_{9'}^{ ext{NP}} = -\mathcal{C}_{10'}^{ ext{NP}}$	-0.61	[-0.45, -0.45]	4.3	50.0

Global fit to \sim 100 obs. (radiative + $b \to s \ell^+ \ell^-$, $\ell = e, \mu$)

All deviations add up constructively

- New Physics contribution to $C_{9,\mu}^{\rm NP}$ =-1.1 alleviates all tensions.
- At the moment, NP contributions to the rest of Wilson coefficient are still not significantly different from zero.

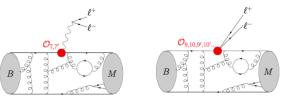
Fitting to data hypothesis where two Wilson coefficients are allowed to vary freely, we obtain several scenarios with pull-SM beyond 4σ :



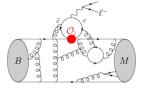
 \Rightarrow Both the confidence regions drawn from BR and $\langle P_i \rangle$ data, suggest $C_9^{\rm NP} \simeq -1$ in all the most significant scenarios.

Theory predictions receive different types of QCD corrections.

Factorisable Corrections: corrections that can be absorved into the definition of the (full) form factors.



Non-factorisable Corrections: corrections that cannot be absorved into the definition of the (full) form factors.



CDHM17: "We provide detailed arguments showing that factorisable power corrections cannot account for the observed anomalies and that an explanation through long-distance charm contributions is disfavoured."

Improved QCDF

Improved QCDF (iQCDF) Approach: General decomposition of a full form factor (FF)

$$\textit{F}^{\text{Full}}(\textit{q}^2) = \textit{F}^{\infty}(\xi_{\perp}(\textit{q}^2), \xi_{\parallel}(\textit{q}^2)) + \Delta \textit{F}^{\alpha_{\text{s}}}(\textit{q}^2) + \Delta \textit{F}^{\Lambda}(\textit{q}^2)$$

where F stands for any form factor, either from the helicity or transversity basis.

- Large recoil symmetries: low- q^2 and at LO in α_s and Λ/m_B
 - \Rightarrow Soft FF $\{\xi_{\perp}, \xi_{\parallel}\}$ (KMPW or any other FF parametrization)
 - ⇒ **Dominant correlations** automatically taken into account
- $\mathcal{O}(\alpha_s)$ corrections \Rightarrow QCDF
- $O(\Lambda/m_B)$ corrections \Rightarrow cannot be explicitly computed within QCDF

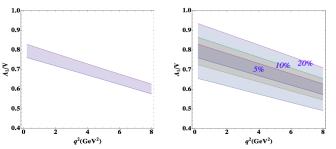
Parametrization of ΔF^{Λ} [JC12]

$$\Delta F^{\Lambda}(q^2) = a_F + b_F \frac{q^2}{m_B^2} + c_F \frac{q^4}{m_B^4} + \dots$$



Improved QCDF (vs full FF approach)

- How to estimate ΔF^{Λ} ?
 - \Rightarrow Central values for a_F , b_F , c_F from **fit to full LCSR FF**.
 - ⇒ Error estimate: assign **uncorrelated** ~ 100% errors to $a_F, b_F, c_F = \mathcal{O}(\Lambda/m_B) \times F = 10\% \times F$
- Is our estimation of errors conservative?
 FF ratio A₁/V (that controls P'₅): BSZ (including correlations) vs iQCDF for different size of power corrections.



Already a 5% power corrections (right) reproduces the BSZ full FF approach errors (left).

Scheme dependence

- Different possibilities for what to take as input for the two independent soft FFs $\{\xi_{\perp}, \xi_{\parallel}\}$
 - e.g. scheme 1 [DHMV] $\{V, a_1A_1 + a_2A_2\}$ or scheme 2 [JC] $\{T_1, A_0\}$ or...
 - ⇒ choice defines input scheme.
- Observables are scheme independent if and only if all the correlations among FF are included.
 - \Rightarrow also correlations among $\Delta a_F, \Delta b_F, \ldots$!
 - \Rightarrow Uncorrelated errors in $\Delta F^{\Lambda} \Rightarrow$ scheme dependence at $\mathcal{O}(\Lambda/m_B)$.
- Input FF do not receive power corrections
 - \Rightarrow Appropriate scheme choices reduce the impact of ΔF^{Λ} .
 - \Rightarrow Non-optimal schemes can artificially inflate the errors due to ΔF^{Λ} .

An illustrative example: $BR(B \to K^*\gamma)$

- How a non-optimal scheme can artificially inflate the errors?
 - \Rightarrow Take $BR(B \to K^*\gamma)$ as an example:

$$BR(B \to K^* \gamma) \propto T_1(0)$$

 \Rightarrow Natural choice: Choose an scheme where T_1 is used as input,

$$\textit{T}_{1}(0) = \textit{T}_{1}^{LCSR}(0) \pm \Delta \textit{T}_{1}^{LCSR}(0) \ \Rightarrow \ \Delta \textit{BR}(\textit{B} \rightarrow \textit{K}^{*}\gamma) \propto \Delta \textit{T}_{1}^{LCSR}(0)$$

 \Rightarrow "Wrong" choice: Use any other FF related to T_1 as input (e.g. T_2),

$$T_1(0) = \left(T_2^{\mathsf{LCSR}}(0) + a_{\mathcal{T}_1}\right) \pm \left(\Delta T_1^{\mathsf{LCSR}}(0) + \Delta a_{\mathcal{T}_1}\right)$$

$$\Rightarrow \Delta BR(B \to K^* \gamma) \propto \Delta T_1^{\mathsf{LCSR}}(0) + \Delta a_{\mathcal{T}_1}$$

 Unnatural scheme choices generate extra contributions in error computations.

Scheme dependence of P_5'

Explicit analytic formulae for the power corrections to P'_5 [CDHM]:

Helicity basis,

$$\begin{split} P_5' &= P_5'|_{\infty} \left(1 + \frac{2a_{V_{-}} - 2a_{T_{-}}}{\xi_{\perp}} \frac{C_7^{\text{eff}}(C_{9,\perp} C_{9,\parallel} - C_{10}^2)}{(C_{9,\perp} + C_{9,\parallel})(C_{9,\perp}^2 + C_{10}^2)} \frac{m_b m_B}{q^2} \right. \\ &\quad + \frac{2a_{V_0} - 2a_{T_0}}{\tilde{\xi}_{\parallel}} \frac{C_7^{\text{eff}}(C_{9,\perp} C_{9,\parallel} - C_{10}^2)}{(C_{9,\perp} + C_{9,\parallel})(C_{9,\parallel}^2 + C_{10}^2)} \frac{m_b}{m_B} - \frac{2a_{V_{+}}}{\xi_{\perp}} \frac{C_{9,\parallel}}{C_{9,\perp} + C_{9,\parallel}} + \dots \right) \end{split}$$

- \Rightarrow We recovered the expression in JC12 + an additional term
- Transversity basis,

$$\begin{split} P_5' &= P_5'|_{\infty} \left(1 + \frac{a_{A_1} + a_V - 2a_{T_1}}{\xi_{\perp}} \frac{C_7^{\text{eff}}(C_{9,\perp} C_{9,\parallel} - C_{10}^2)}{(C_{9,\perp} + C_{9,\parallel})(C_{9,\perp}^2 + C_{10}^2)} \frac{m_b m_B}{q^2} \right. \\ & \left. - \frac{a_{A_1} - a_V}{\xi_{\perp}} \frac{C_{9,\parallel}}{C_{9,\perp} + C_{9,\parallel}} - \frac{a_{T_1} - a_{T_3}}{\tilde{\xi}_{\parallel}} \frac{C_7^{\text{eff}}(C_{9,\perp} C_{9,\parallel} - C_{10}^2)}{(C_{9,\perp} + C_{9,\parallel})(C_{9,\parallel}^2 + C_{10}^2)} \frac{m_b}{m_{K^*}} + \dots \right) \end{split}$$

with
$$C_{9,\perp}=C_9^{\rm eff}+rac{2m_bm_B}{a^2}C_7^{\rm eff}$$
 and $C_{9,\parallel}=C_9^{\rm eff}+rac{2m_b}{m_B}C_7^{\rm eff}$



Scheme dependence of P_5'

- The FF ratio A_1/V dominates P'_5 ,
 - \Rightarrow Convenient: scheme 1 [DHMV] { $V, a_1A_1 + a_2A_2$ }
 - \Rightarrow Inconvenient: scheme 2 [JC] { T_1, A_0 }
- Evaluating the expression for the power corrections to P_5' at $q^2=6$ GeV² (around the anomaly),

$$P_5'(6\,\text{GeV}^2) = P_5'|_{\infty}(6\,\text{GeV}^2) \left(1 + 0.18\frac{a_{A_1} + a_V - 2a_{T_1}}{\xi_{\perp}} - 0.14\frac{a_{T_1} - a_{T_3}}{\tilde{\xi}_{\parallel}} - 0.73\frac{a_{A_1} - a_V}{\xi_{\perp}}\right)$$

- \Rightarrow Scheme 1: $P_5'(6 \text{ GeV}^2) \simeq P_5'|_{\infty}(6 \text{ GeV}^2) \left(1 0.73 \frac{a_{A_1}}{\xi_{\perp}}\right) \Rightarrow$ reduced errors.
- \Rightarrow Scheme 2: $P_5'(6\,\mathrm{GeV^2}) \simeq P_5'|_{\infty}(6\,\mathrm{GeV^2}) \left(1 0.73\frac{a_{A_1} a_V}{\xi_{\perp}}\right) \Rightarrow \mathrm{increased}$ errors.



Correlations and scheme dependence of P_5'

Assessing the impact of the correlations among power corrections (PC) \pm scheme dependence,

- 1 PC Analysis
- correlations from large-recoil sym. $\Rightarrow \xi_{\perp,\parallel}, \Delta F^{\Lambda}$ uncorr.

- 2 PC Analysis
 - ΔF^{Λ} from fit to LCSR [BSZ].
 - **correlations** from large-recoil sym. $\Rightarrow \xi_{\perp,\parallel}$, ΔF^{Λ} uncorr.

- 3 PC Analysis
 - ΔF^{Λ} from fit to LCSR [BSZ].
 - **correlations** from LCSR [BSZ] $\Rightarrow \xi_{\perp,\parallel}$, ΔF^{Λ} corr.

P' ₅ [4.0, 6.0]	scheme 1 [CDHM]	scheme 2 [JC]	
1	$-0.72 \pm extbf{0.05}$	-0.72 ± 0.12	
2	-0.72 ± 0.03	-0.72 ± 0.03	
3	-0.72 ± 0.03	-0.72 ± 0.03	
full BSZ	-0.72 + 0.03		

errors only from pc with BSZ form factors



Non-factorisable hadronic corrections

There are two different types of non-factorisable hadronic corrections

- \blacksquare α_s -corrections from hard gluon exchange $(\mathcal{O}_{1-6}, \mathcal{O}_8 \text{ topologies}) \Rightarrow \mathsf{QCDF}$
- $\mathcal{O}(\Lambda/m_B)$ corrections involving $c\bar{c}$ loops
 - \Rightarrow [KMPW] LCSR + dispersion relations (only th. calculation)
 - \Rightarrow Non-factorisable $\mathcal{O}(\Lambda/m_B)$ power corrections (charm loops) yield q^2 and hadronic-dependent contributions with $\mathcal{O}_{7,9}$ structures that may
 mimic NP:

$$C_9^{\text{eff } i}(q^2) = C_9^{\text{eff SM}}(q^2) + C_9^{\text{NP}} + C_9^{\tilde{c}\tilde{c}i}(q^2) \quad i = \perp, \parallel, 0$$

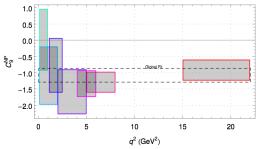
Disentangling cc loops from New Physics

NP and hadronic effects have different signatures on C_9 :

- NP effects: universal and q^2 -independent.
- Hadronic effects: transversity dependent and (most likely) q^2 -depedent.

Testing the q^2 dependence of the contributions to C_9 by means of data,

 C_9^{NP} bin-by-bin fit to $b \to s\ell\ell$ data (assuming KMPW-like $C_9^{c\bar{c}i}(q^2)$):



 \Rightarrow Excellent agreement with a q^2 -independent $C_9 \simeq -1$.

Fitting a charm-loop parametrization to data

Following Ciuchini et al., we performed a fit of the charm loop contributions to data using a polinomial paramatrization,

$$A_{L,R}^0 = A_{L,R}^0(Y(q^2)) + rac{N}{q^2} \left(h_0^{(0)} + rac{q^2}{1 \, GeV^2} h_0^{(1)} + rac{q^4}{1 \, GeV^4} h_0^{(2)} + rac{q^6}{1 \, GeV^6} h_0^{(3)}
ight)$$

- Non-zero $h_{\lambda}^{(2),(3)}$ $(\lambda=+,-,0)$ introduce q^2 -dependent terms in C_9 .
 - \Rightarrow Disclaimer: $C_{7,9}^{NP}$ contribute to $h_i^{(2),(3),...} \Rightarrow C_i^{NP} \times F(q^2)$.
- Frequentist fit of $h_{\lambda}^{(i)}$ to $B \to K^* \mu \mu$ data: using KMPW FF and without including any charm-loop estimate to $C_{7,9}$.
- Comparing hypothesis with increasing orders of the h_{λ} polinomials (n = 0, 1, 2, 3), we conclude [CDHM]:
 - \Rightarrow Hypotheses with linear h_{λ} polinomials are the ones with better improvement of the fit.
 - \Rightarrow Setting $C_9^{\rm NP}=-1.1$ significantly improves the fit (already with $h_\lambda=0$).

Conclusions

- Global fits: $C_{9,\mu}^{\text{NP}} = -1.1$ with a pull-SM of 4.7σ alleviates all the tensions.
- Analysis of the two main sources of hadronic uncertainties:
 - ⇒ Factorisable power corrections can artificially inflate the errors when considered as uncorrelated by means of non-optimal scheme choices.
 - \Rightarrow Non-factorisable $c\bar{c}$ -loops: our tests do not show signs of large q^2 contributions to C_9 .