

High Performance Computing for gamma ray detection

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Versailles/Saclay

- Nahid Emad
- LI-PaRAD
- Maison de la simulation

LAPP

- Gilles Maurin
- Jean Jacquemier
- Thomas Vuillaume
- Florian Gaté

CTA

- Data analysis

ASTERICS

- CTA
- SKA
- LSST
- KM3Net
- E-ELT
- LIGO

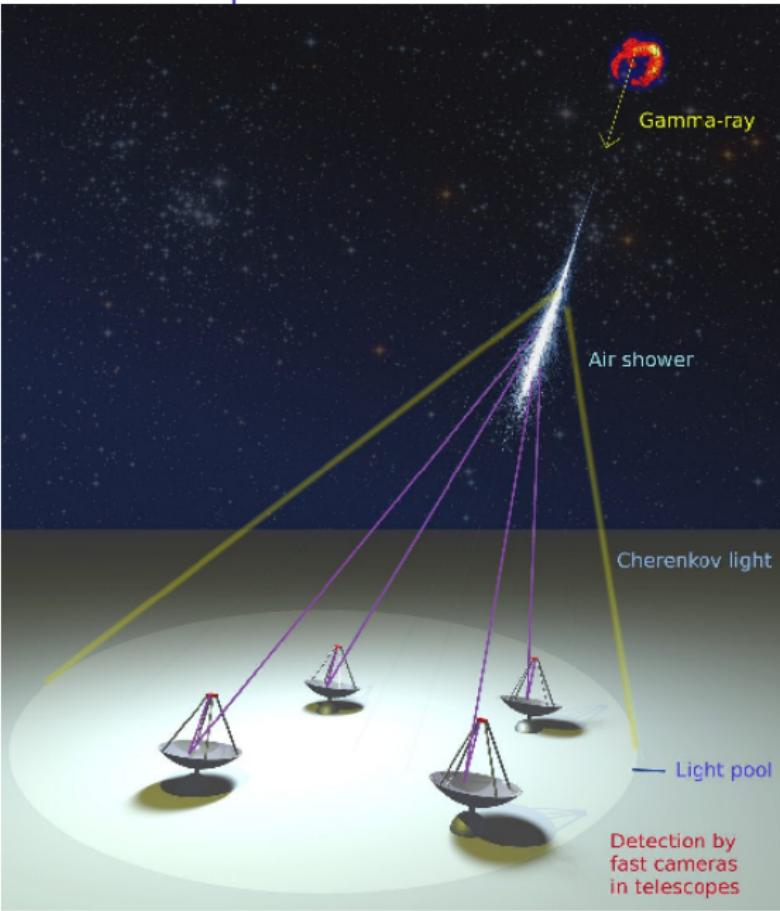
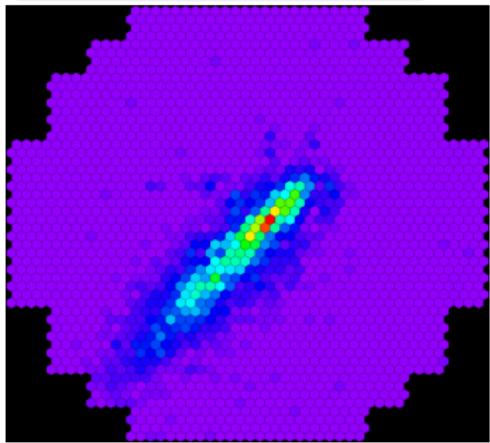
LAPP

- ASTERICS
 - ▶ CTA
 - ▶ LSST
- COMUE IDEX UGA
 - ▶ Call for project Pôle-PAGE
 - ▶ Data@UGA project

Detection of particle showers in the atmosphere

Cosmic rays

- 89% : protons
- 10% : heavy nuclei
- 1% : electrons
- 1‰ : gamma rays



Telescopes

- 4 small, 13 m of diameter
- 1 large, 28 m of diameter
- 2 types of camera

Data flow

- 200 frames per second
- 3 MB/s

Monthly time

- 100 h observation
- 1000 h analysis

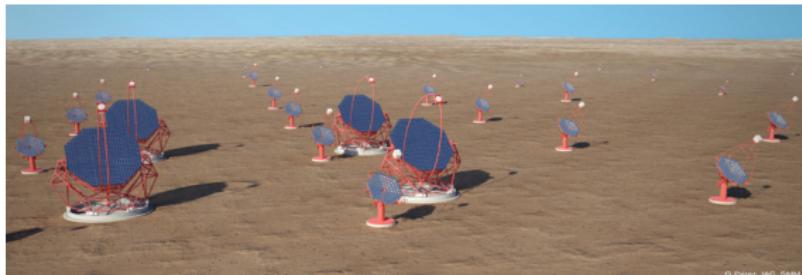


Observatory

- Provide/Store raw data to the international community
- Create/Maintain/Provide analysis to the international community
- Accessible via a web portal

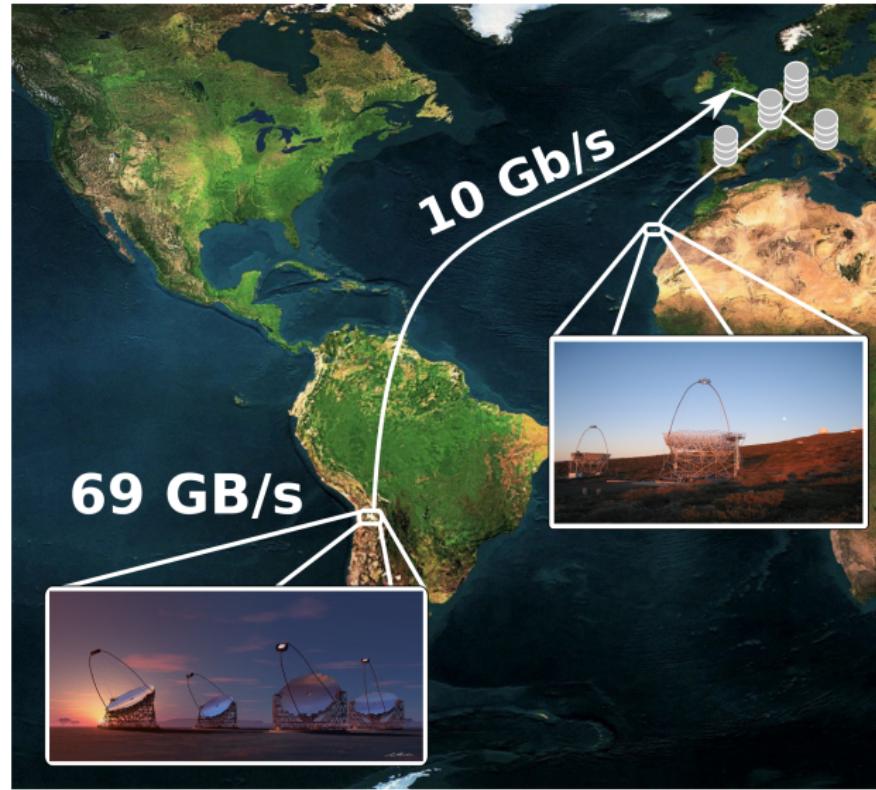
Telescopes

- 126 telescopes
- 3 telescope sizes
- 7 camera types

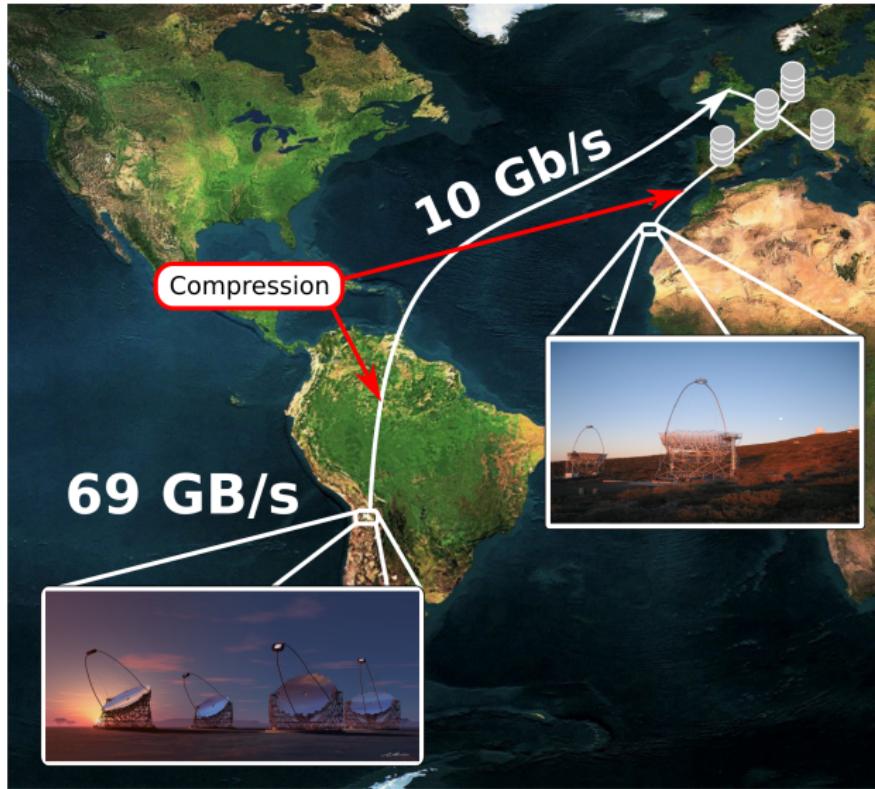


Data rate

- 15 000 frames per second per large telescope
- $69 \text{ GB/s} \implies 2 \text{ PB/yr}$
- Optic fiber line : 10 Gb/s
 - ▶ Because 100 Gb/s line $\implies 23\text{M}\text{\euro}$ per year



First challenge : the Data compression



Raw data Compression Ratio and speed (Lossless compression)

Test file run 497 (Paranal, Gamma Monte-Carlo CTA PROD_3)

- 475 MB
- More than 99% ADC values

	Compression ratio	Time	File size (MB)
LZMA (7z)	4.84	7 min 48.636 s	98
Advanced Polynomial Compression	3.74	3.7 s	127

Raw data Compression Ratio and speed (Lossless compression)

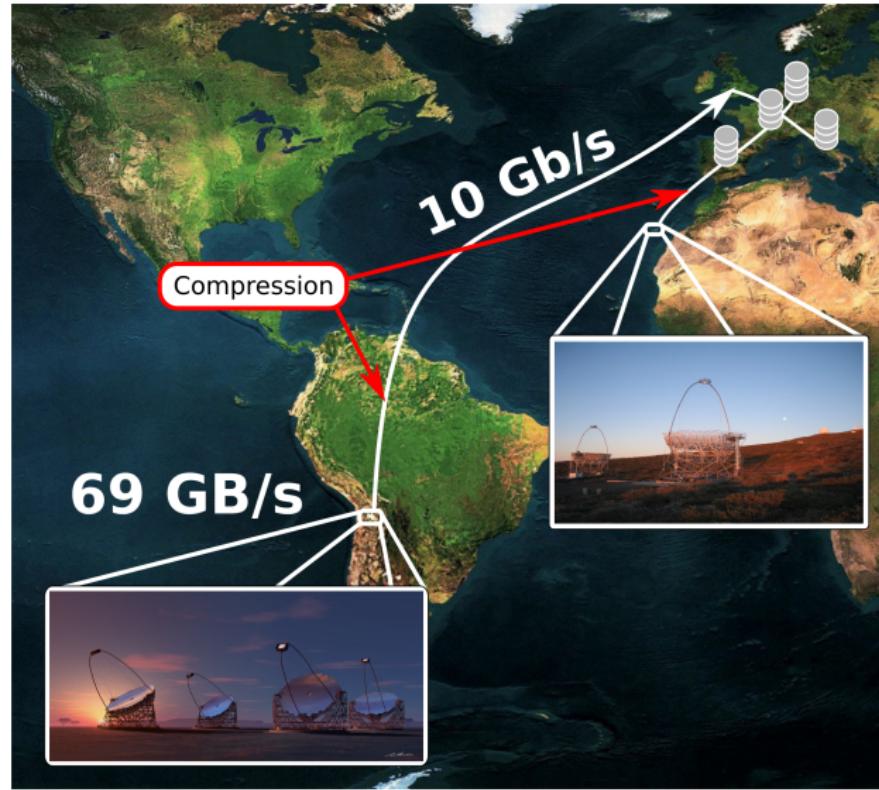
Test file run 497 (Paranal, Gamma Monte-Carlo CTA PROD_3)

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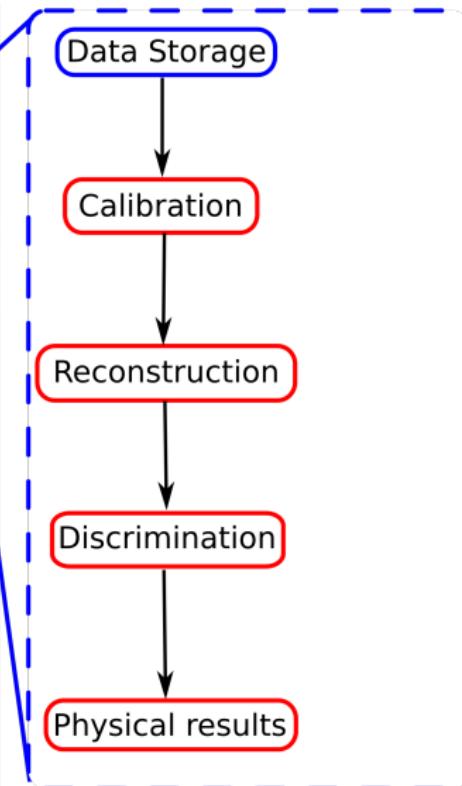
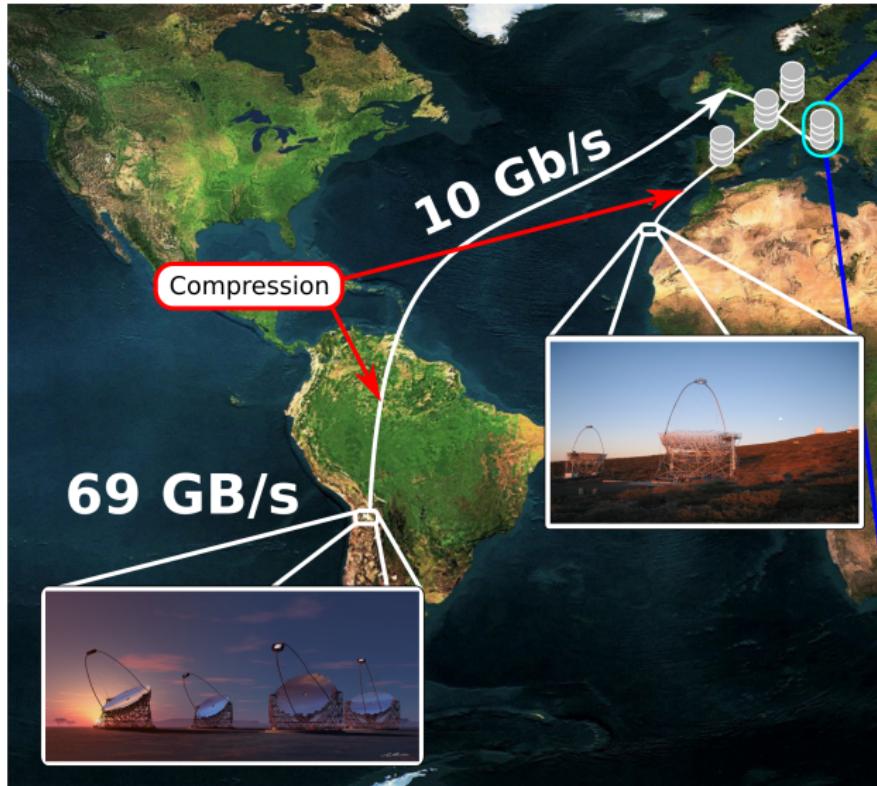
	Compression ratio	Time	File size (MB)
LZMA (7z)	4.84	7 min 48.636 s	98
Advanced Polynomial Compression	3.74	3.7 s	127
Advanced Polynomial Compression + LZMA	4.84	24.646 s	98

Same ratio but faster ($\times 19$) compression !

- Can be used for any data format
- Publication ongoing



The CTA data analysis



Data format

What is good for an HPC data format ?

- Allows CPU data pre-fetching
 - ▶ With contiguous data
 - ▶ Guaranty data locality
 - ▶ Cache friendly
- Allows vectorization
 - ▶ With aligned data for vectorizable computation
- Must be as simple as possible (user friendly)
 - ▶ Simple types (`float`, `int`, `double`, `unsigned int`, ...)
 - ▶ Tables
 - ▶ Allows the simpler solution for complex problem

LAPP Sorted HPC Autotuned Data Optimized Kernel format

Generated Data format for CTA

Describes :

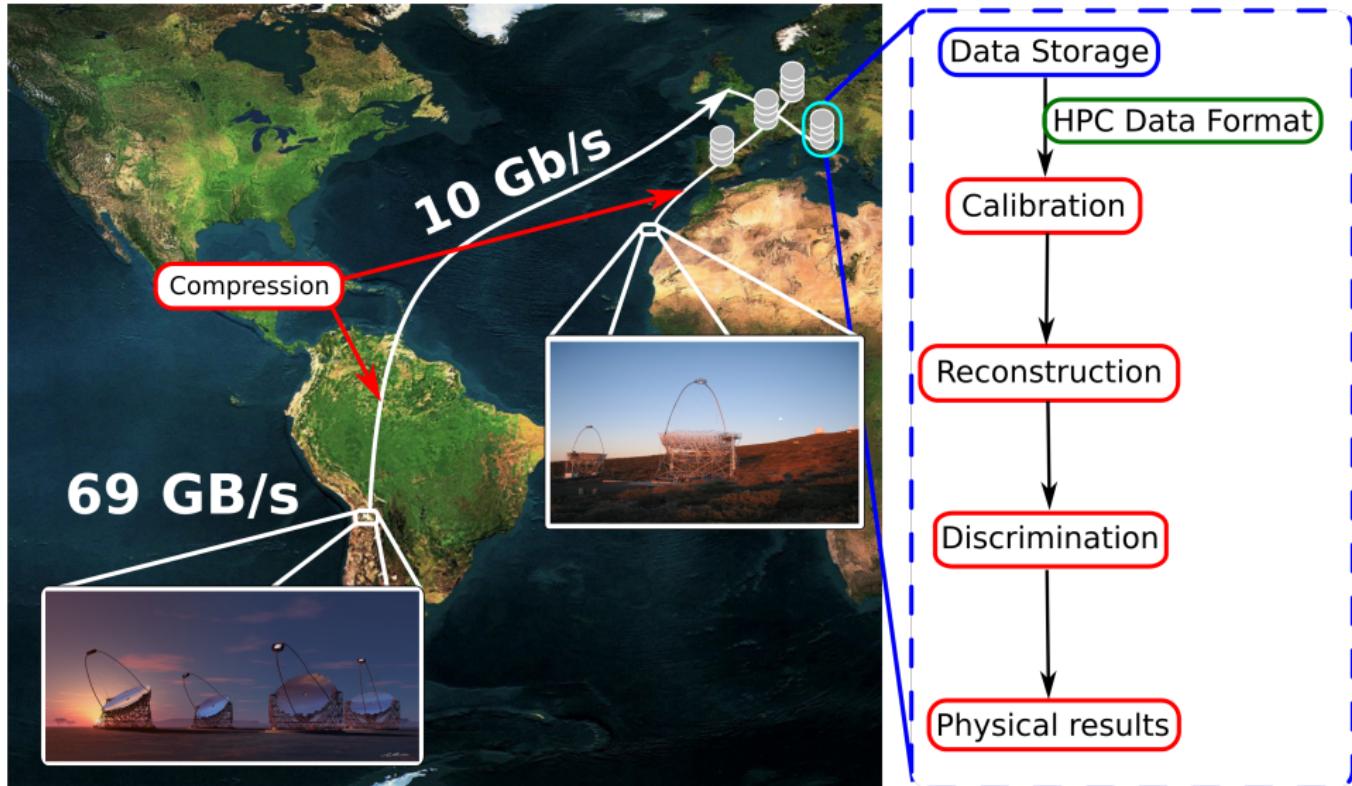
- All the telescopes
- All the Cameras
- The CTA PROD_3 Monte-Carlo
- Allows Layouts and Masks
- DL0, DL1, DL2

Allows :

- Allows CPU data pre-fetching
 - ▶ With contiguous data
 - ▶ Guaranty data locality
 - ▶ Cache friendly
- Vectorization of algorithms
 - ▶ Calibration, Cleaning, Hillas

And it's user friendly

Second challenge : the data analysis



Second challenge : the data analysis

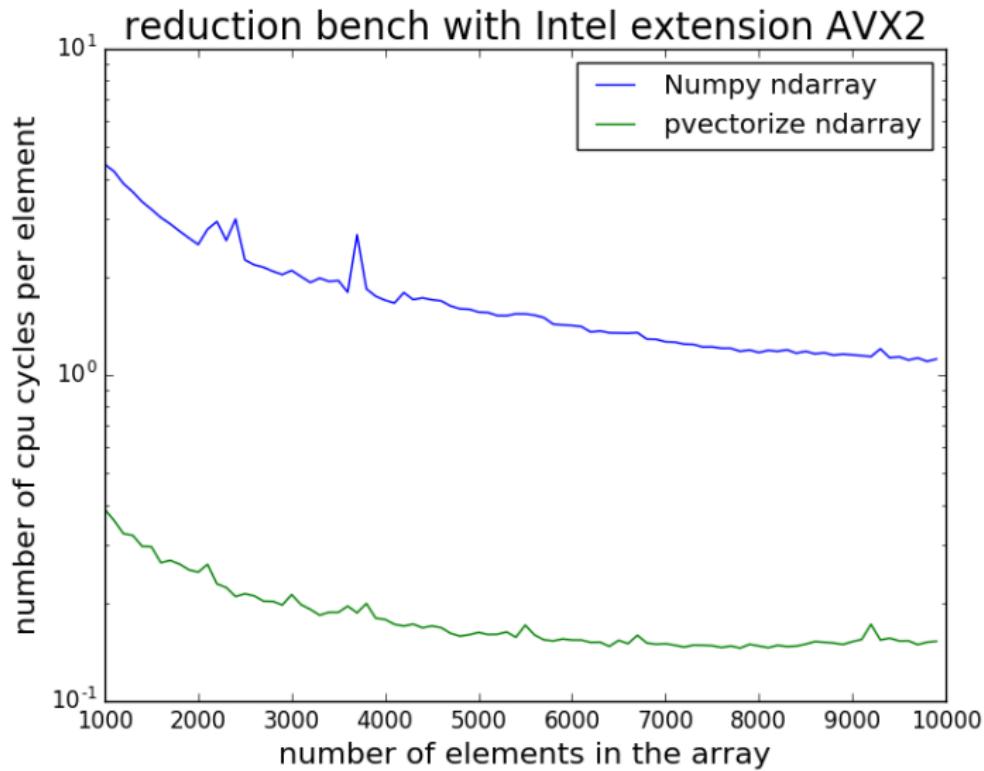
Programming language for CTA

- Python
 - ▶ Good to develop tests
 - ▶ Bad to analyse data in production

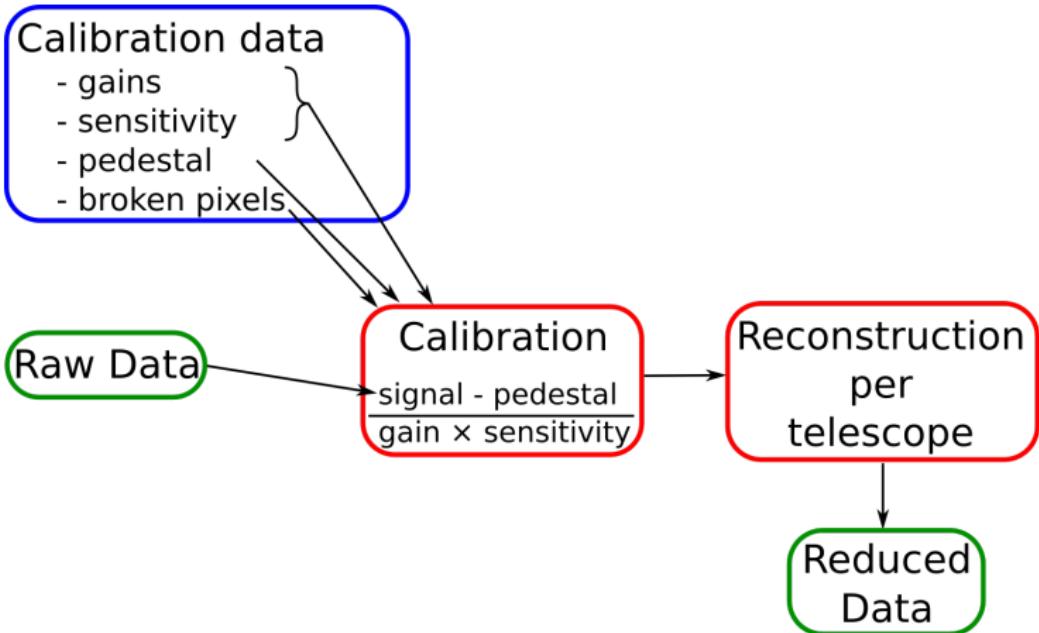
Optimization

- Fast C++ library
- Wrapper python

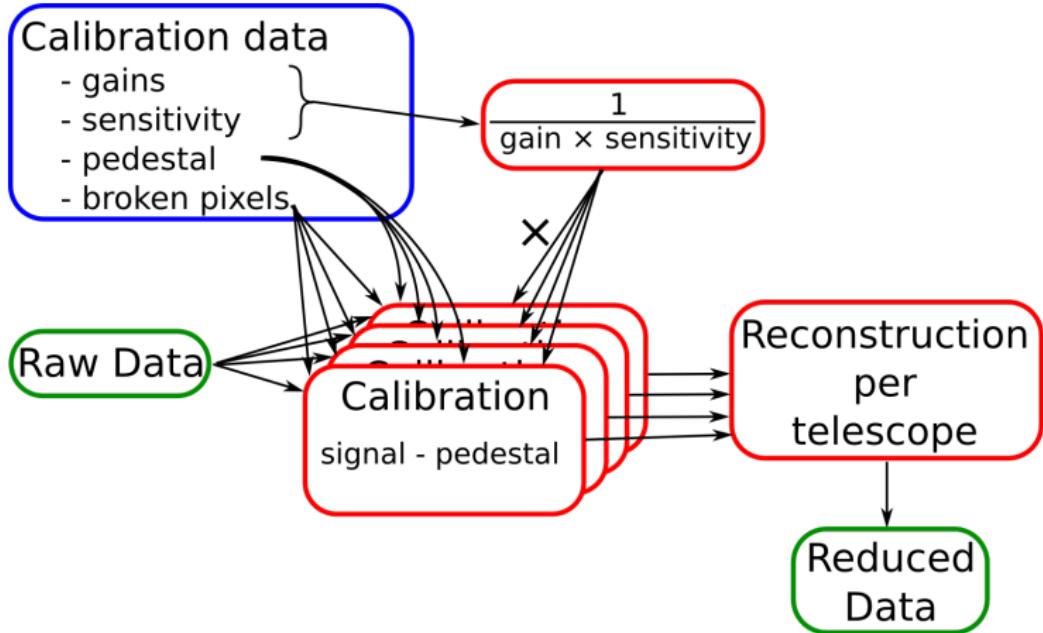
Second challenge : the data analysis



Raw data calibration



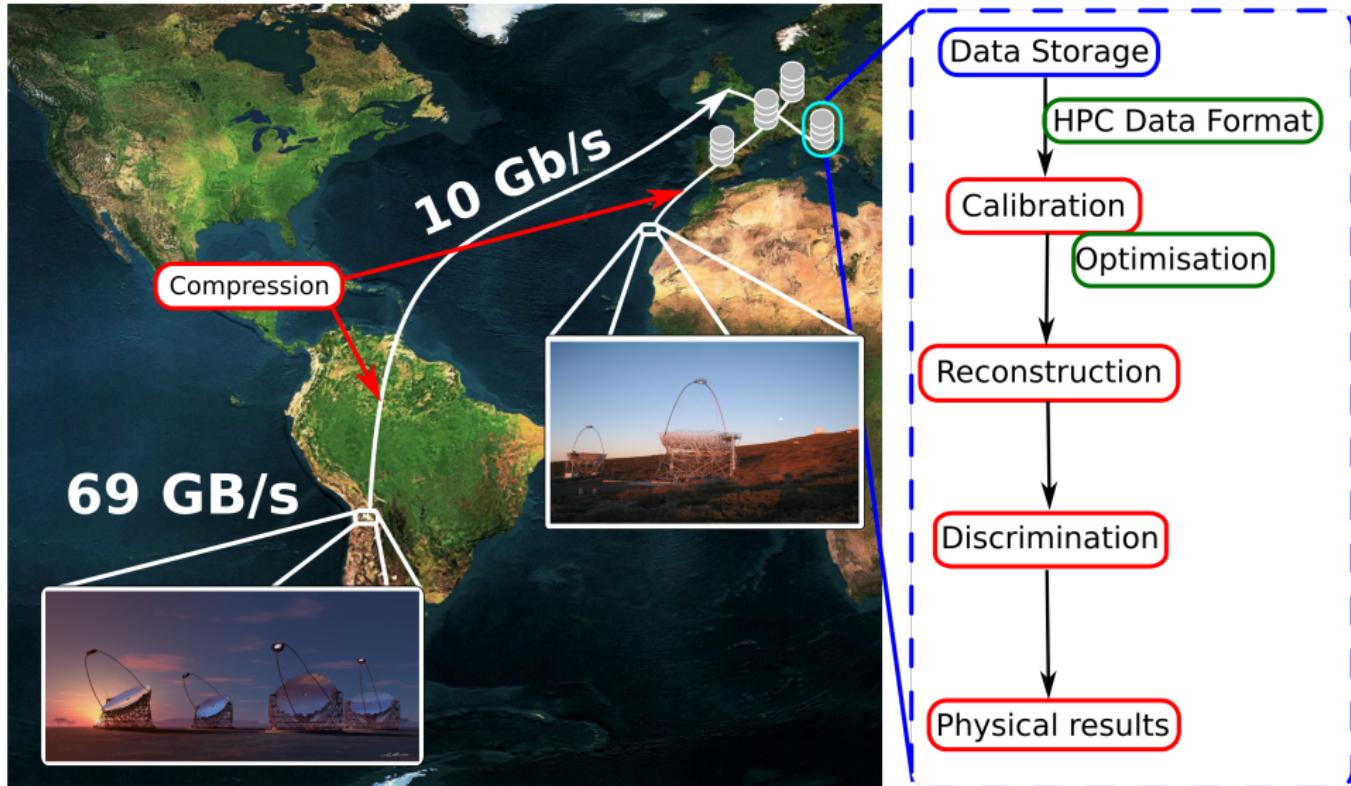
Raw data calibration



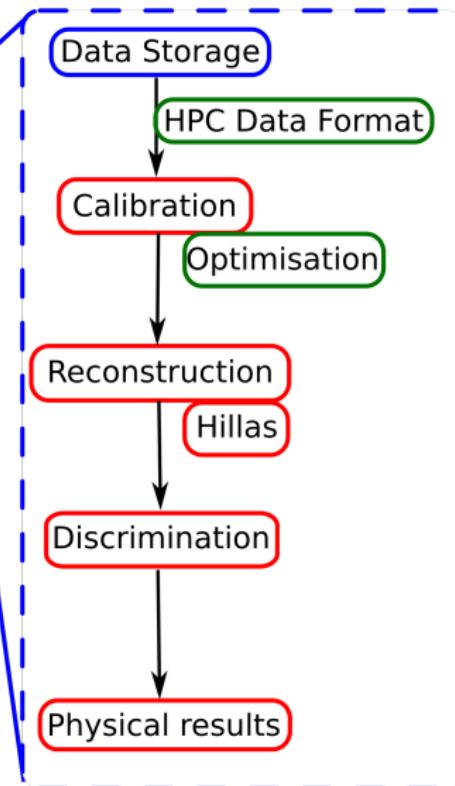
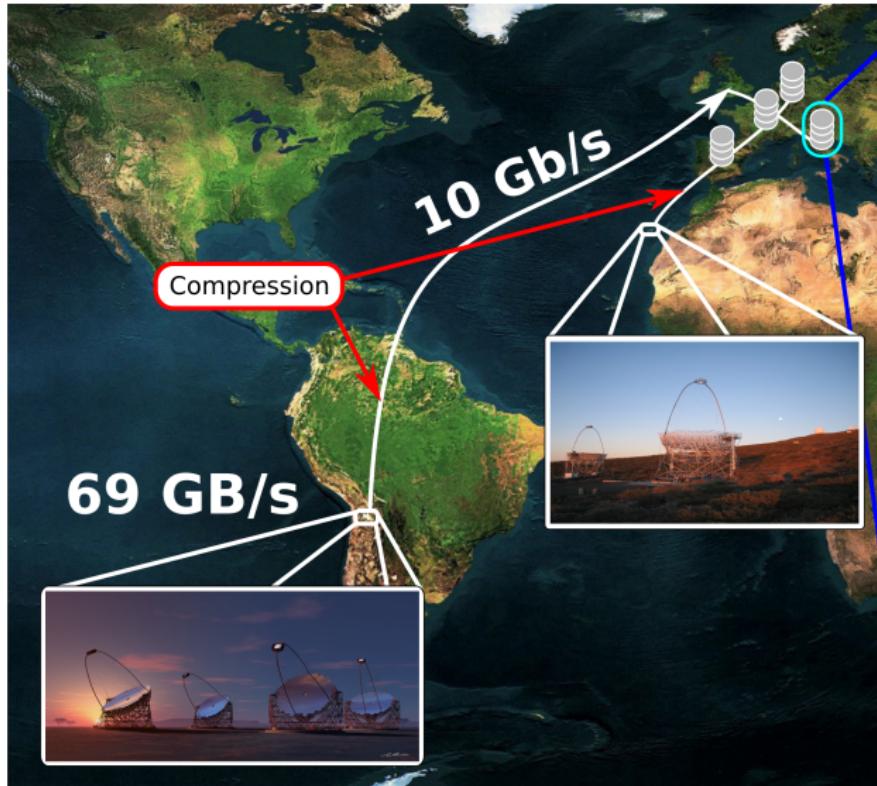
Optimization

- Clearly vectorizable (contiguous data, no back dependencies)
- Parallelization non efficient because the pictures are too small

Second challenge : the data analysis



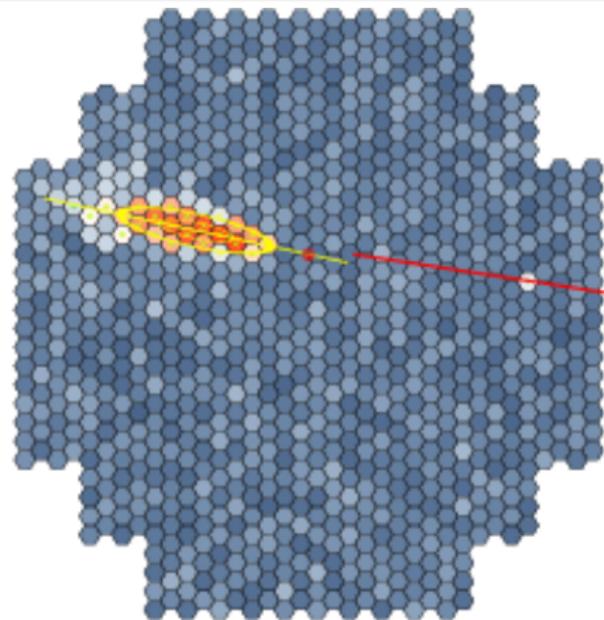
Second challenge : the data analysis



Hillas

Hillas parameters calculation

- Reduction : Sum of all pixels photoelectrons signal per camera
- First momentum : ellipse's position
- Second momentum : ellipse's orientation



Hillas with optimized data format

Optimizations

- No ROOT
- Allows CPU data-pre-fetching

Hillas

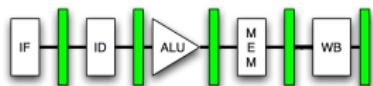
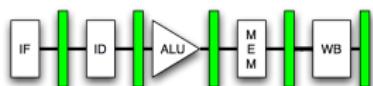
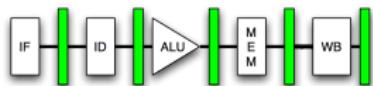
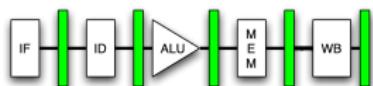
	Speed (cy/el)	Speed up
Initial data format	2125.5	1
Optimized data format	53.1375	40.0

CPU Recent Architectures

SSE4

4 floats

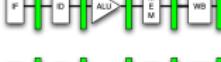
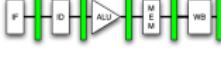
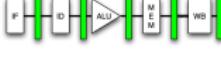
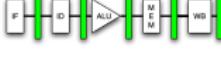
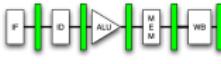
Instruction set : 2006
CPU : 2007



AVX

8 floats

Instruction set : 2008
CPU : 2011



AVX 512

16 floats

Instruction set : 2013
CPU : 2016



Data format

Efficient only if data
are contiguous

Reduction optimization

Reduction Sum of all pixels photoelectrons signal per camera

	Speed (cy/el)	Speed up
Classical	2.69842	1
Vectorized (GCC, SSE4)	0.702845	3.8

Reduction optimization

Reduction

Sum of all pixels photoelectrons signal per camera

	Speed (cy/el)	Speed up
Classical	2.69842	1
Vectorized (GCC, SSE4)	0.702845	3.8
Intrinsics Vectorized SSE4	0.226675	11.9
Intrinsics Vectorized AVX	0.11379	23.7

Hillas with optimized format and vectorization

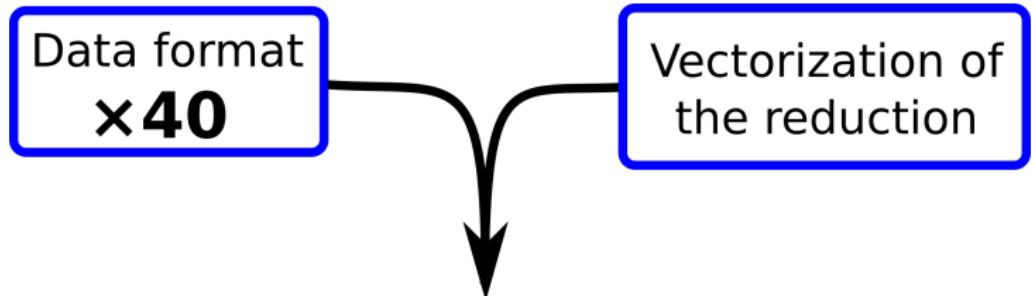
Data format
×40

Vectorization of
the reduction

Hillas

	Speed (cy/el)	Speed up
Standard data format	2125.5	1
New data format	53.1375	40.0
+ intrinsics vectorization SSE4	6.39931	332
+ intrinsics vectorization AVX	2.98499	712

Hillas with optimized format and vectorization



Hillas

	Speed (cy/el)	Speed up
Standard data format	2125.5	1
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+ intrinsics vectorization SSE4	6.39931	332
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The Hillas parameters calculation costs approximately one reduction
(2.69842 cy/el)

The more complex is your calculation, the higher speed up you can have

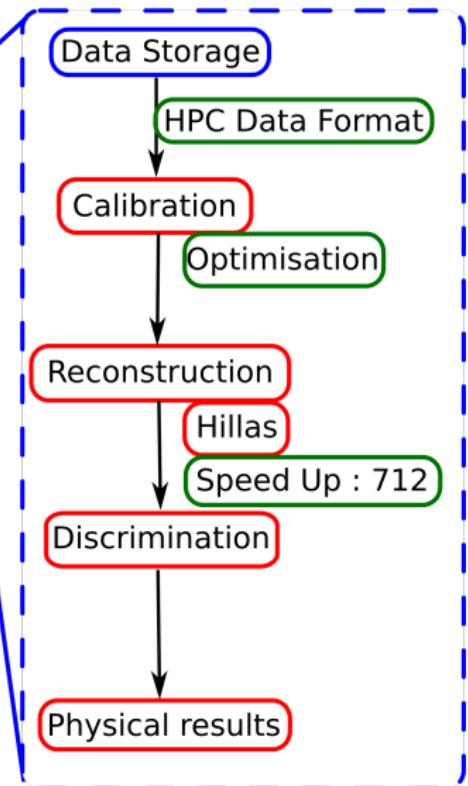
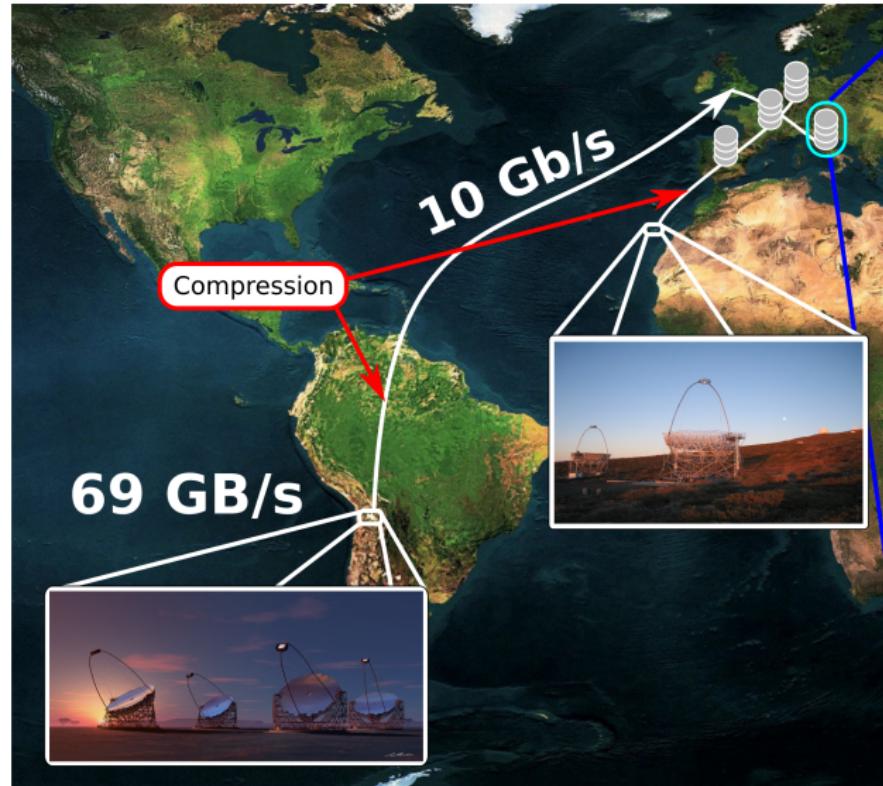
Next analysis optimization

The Hillas method

- Publication **soon**

The Model++ method

- Simulation of the telescopes
- Images comparison



Model++ Method

Idea

Modelization of what is observable with the telescopes

- Particle shower simulation (KASCADE, CORSIKA)
- Telescopes answer simulation
- Raw Data compared to the simulated pictures
- Get the physical parameters of the best shower

Model parameters

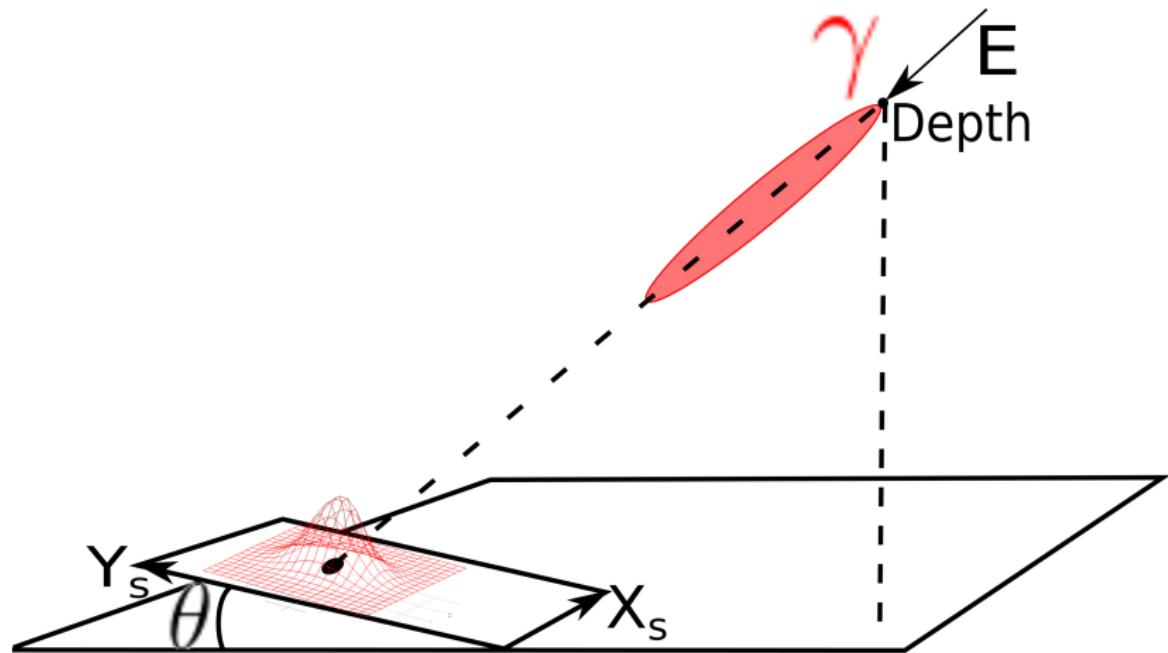
- θ : zenith angle of the telescopes
- E : primary particle energy
- $Depth$: first interaction high
- ρ : distance from the telescope to the shower impact

Adjusting the model

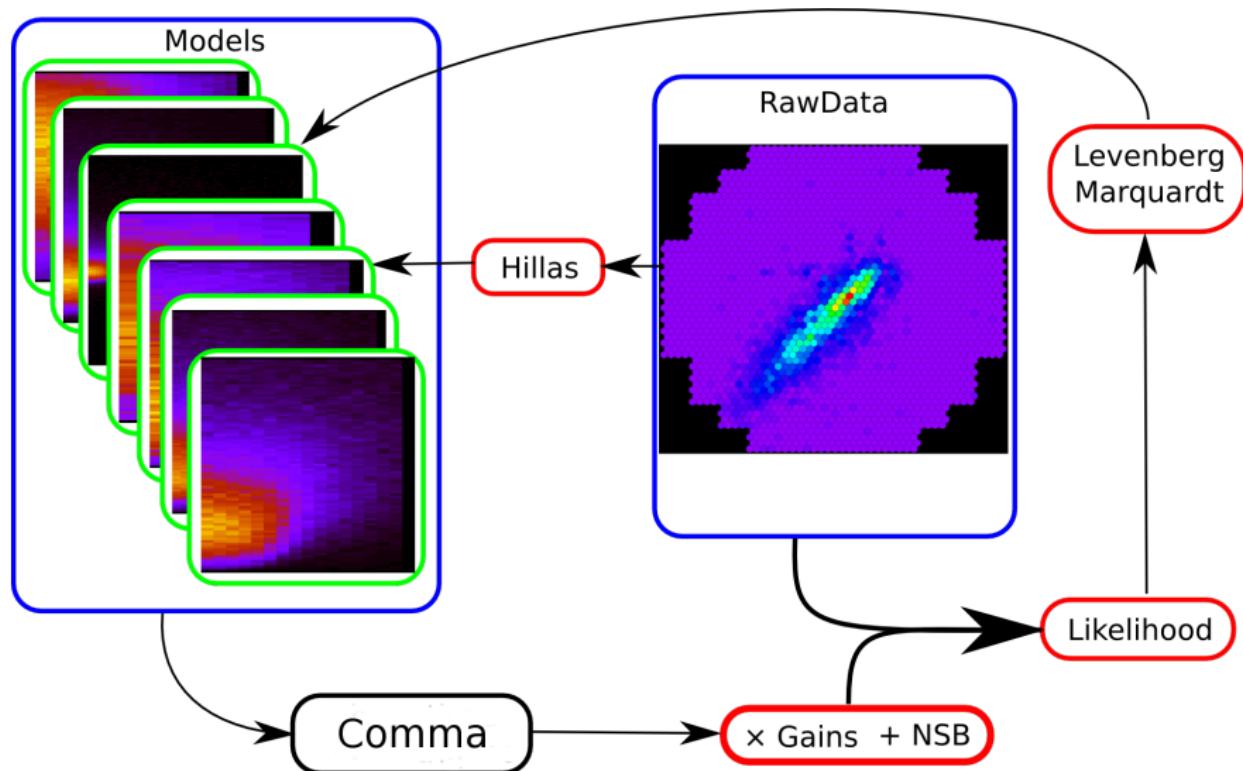
- (x_s, y_s) : model barycentre position in the telescope's camera
- ϕ : model rotation in the telescope's camera

Model++ Method

Model creation



Model++ Algorithm



Model++ Algorithm

Algorithm requiristies

- Stereoscopy (Telescope of the same event in RAM)
- Enought RAM to contain all the telescope's data
- **Telescopes** orderd by **Event**

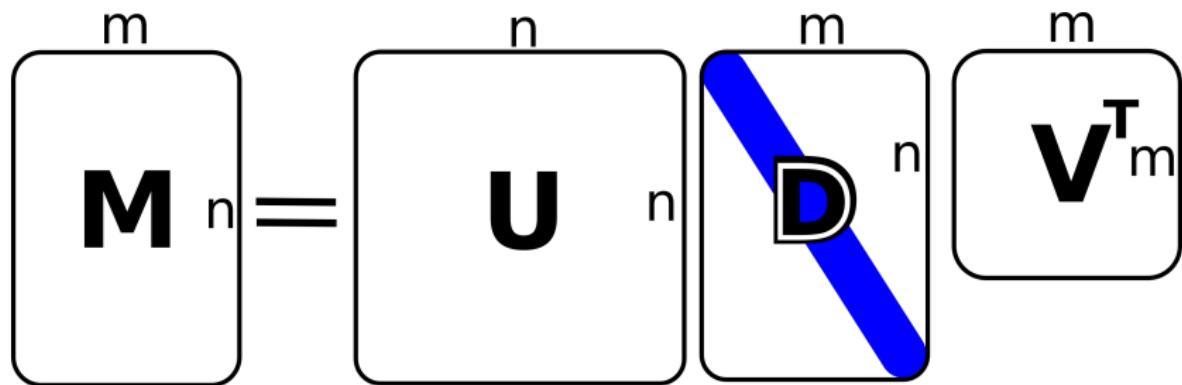
Issues

- Data Scale \Rightarrow 126 Telescopes, 69 GB/s
 - ▶ Too much data for one single computer
 - ★ Lot of communications for the many-core approach
- Data Order :
 - ▶ **Event** orderd by **Telescopes** **Needs sorting**
 - ★ Sort a huge data volume \Rightarrow Huge RAM **Data access time unrealistic**

The SVD method

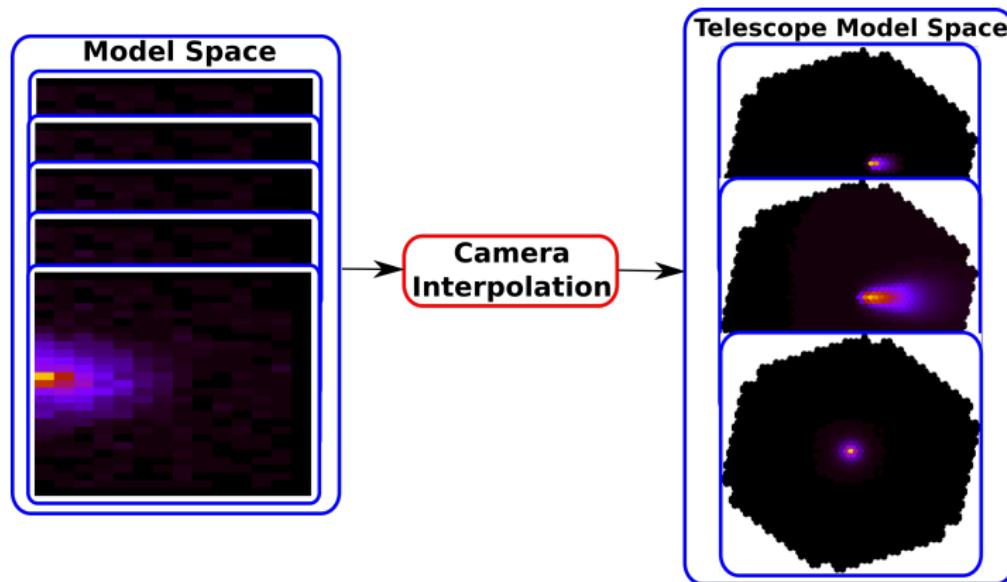
The Single Value Decomposition

$$M = UDV^T$$



The SVD method

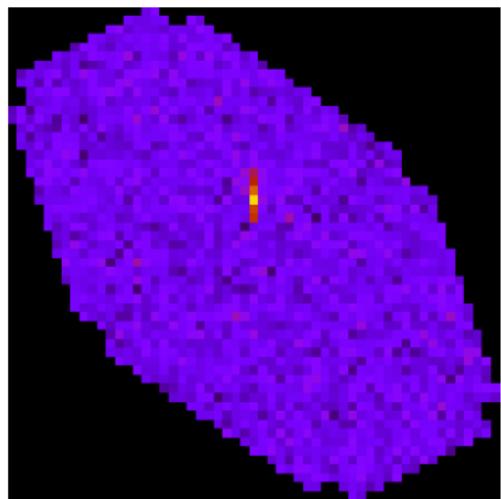
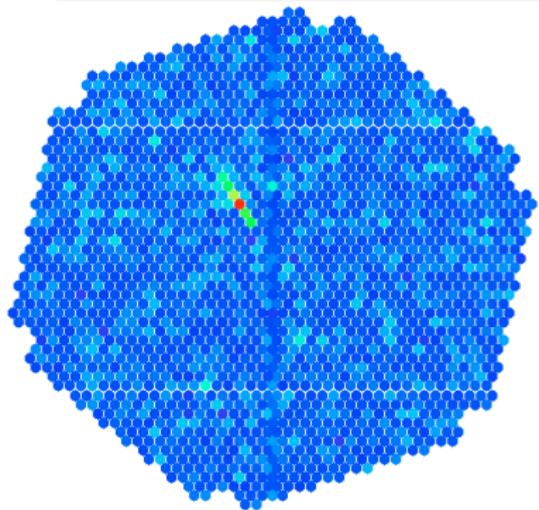
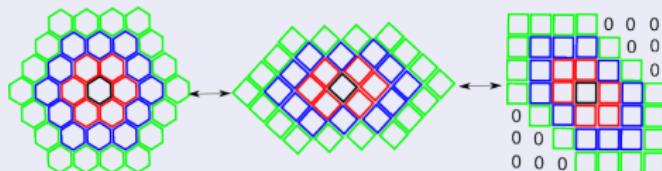
Model conversion : Dragon



The SVD method

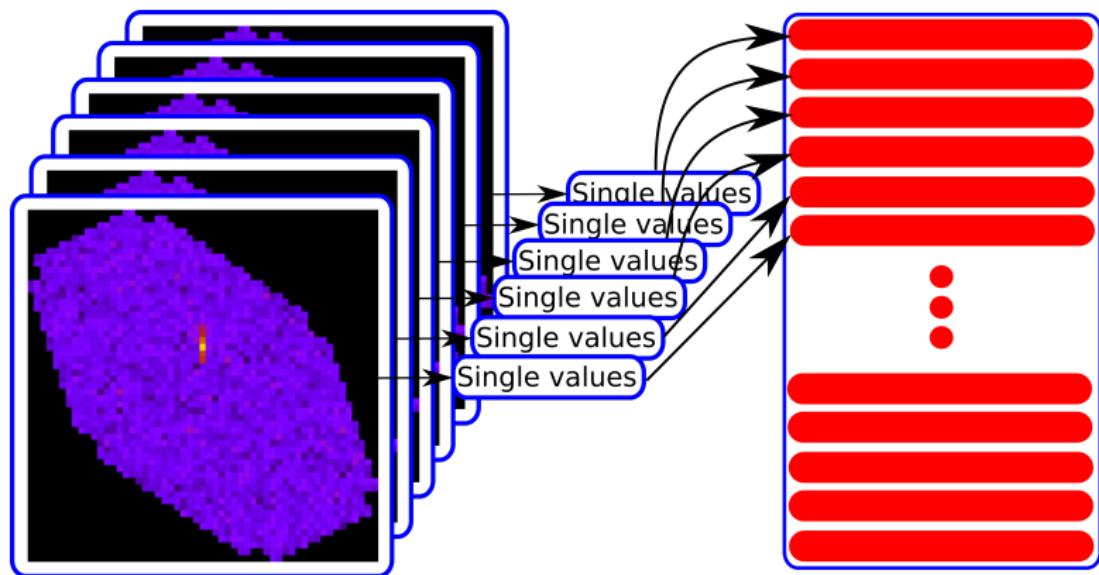
Dragon camera to matrix

Camera to matrix conversion



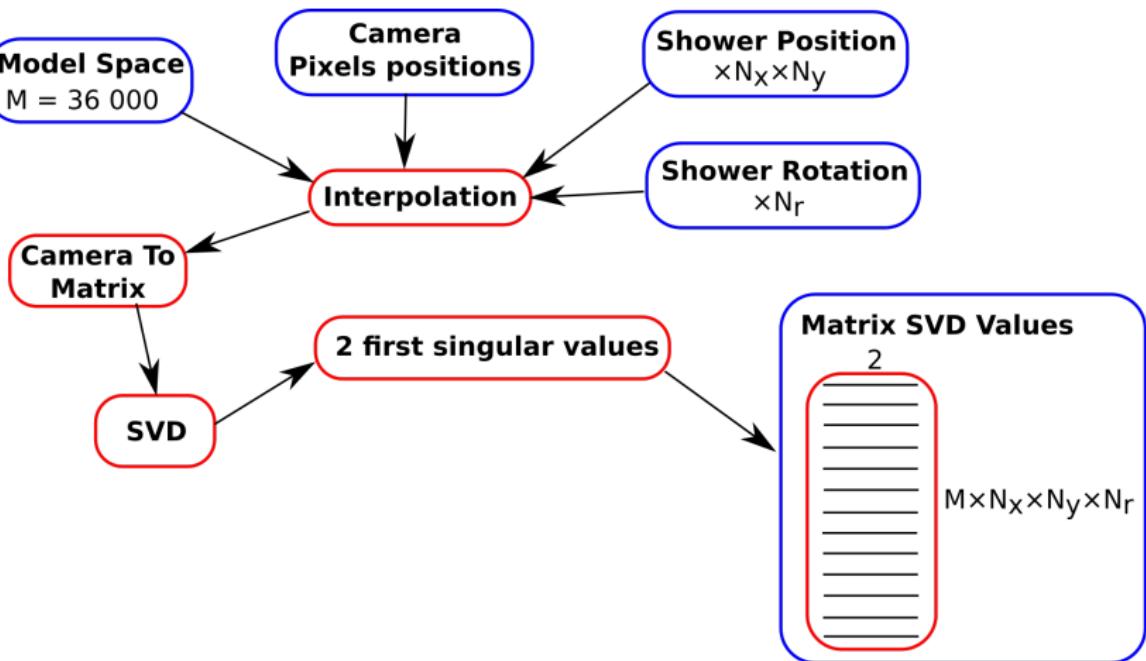
The SVD method

SVD on camera : Dragon



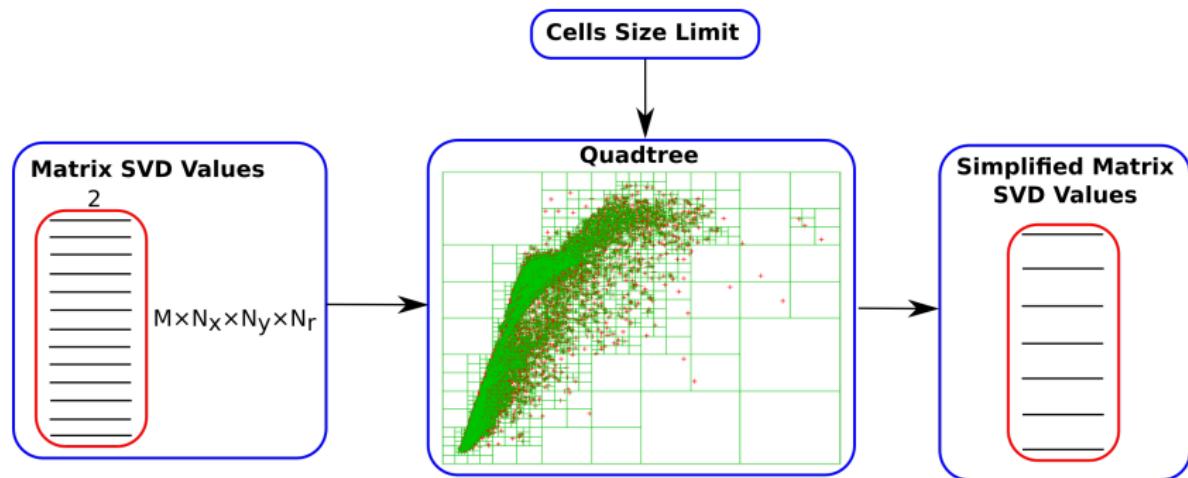
The SVD method

From Model Space to SVD matrix



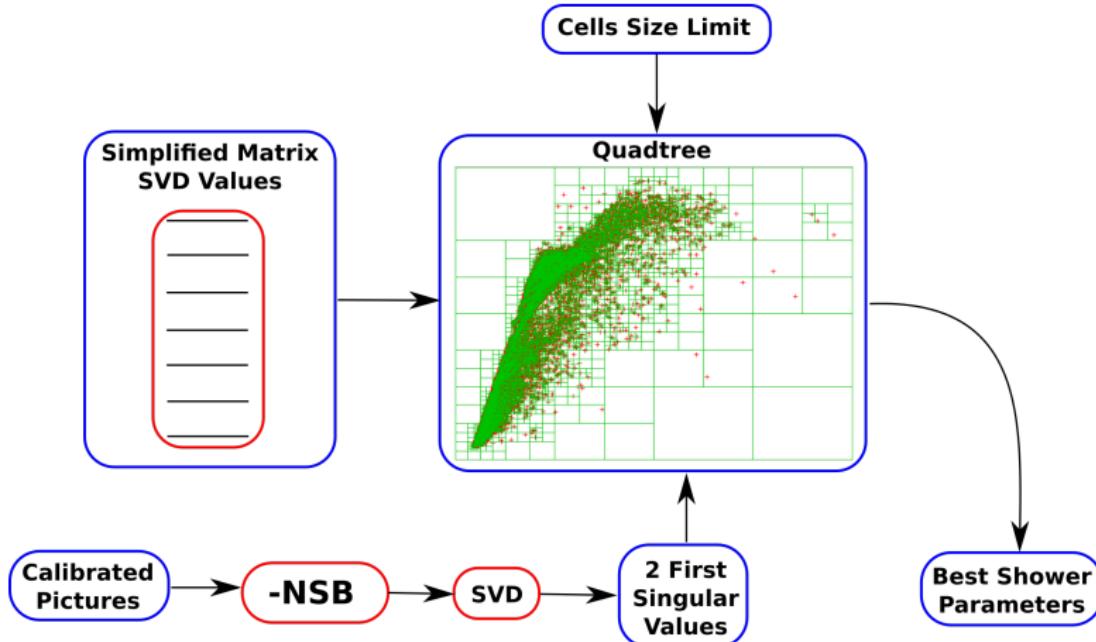
The SVD method

SVD matrix simplification



The SVD method

SVD Reconstruction



Performances of several Cores/CPUs/Nodes

Parallelization

- Load-balancing
- Optimization of CPU use

Official ctapipe Parallelization system

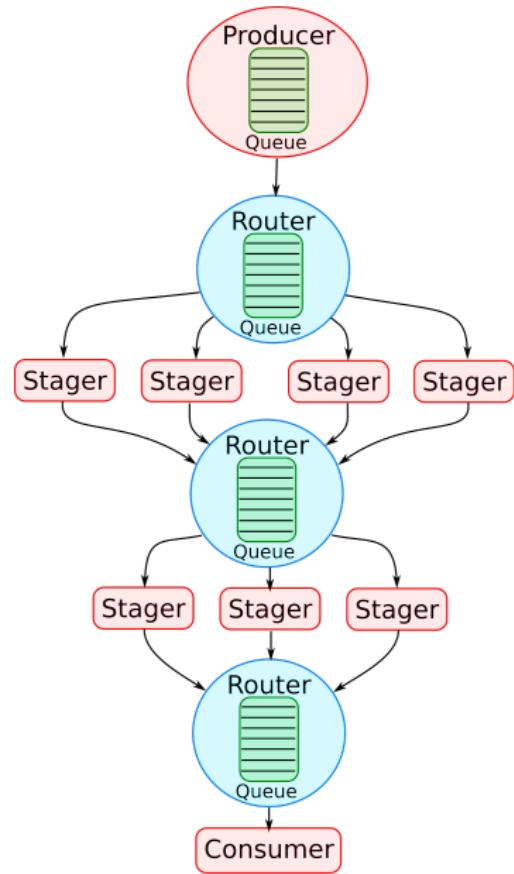
Functionnalities

- Flexible
- Chains the analysis steps
- Parallelizes the analysis steps
- Load-balances the analysis steps

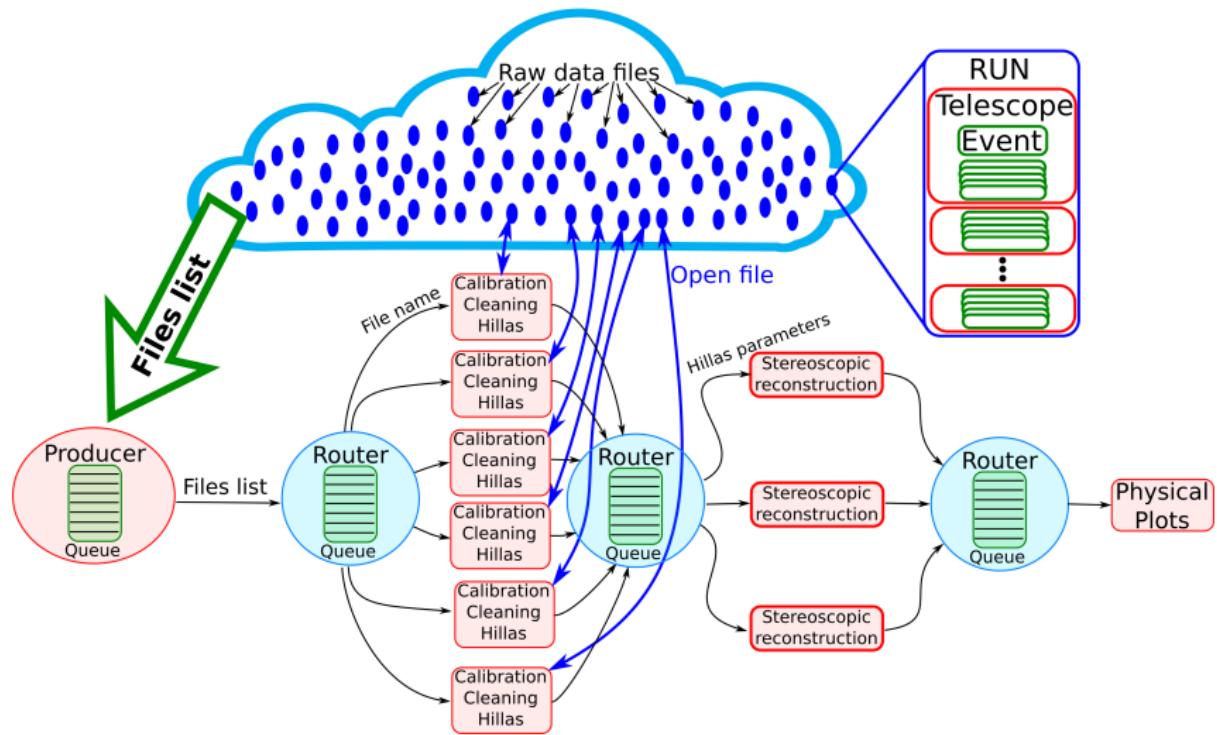
ZeroMQ

- Communication between stager and router
- Independent of core/node/computer
- No All-to-all communications

Jean Jacquemier



CTA pipe for analysis



Implementation

- C, C++, Python (wrapper)
- Intrinsics (vectorization by hand)
- Loop unrolling
- Loop interleaving

Framework

- Open source
- Licence CeCILL-C
- https://gitlab.in2p3.fr/CTA-LAPP/PLIBS_8

Conclusion

Data transfert (69 GB/s \Rightarrow 10 Gb/s)

- Data compression : Publication **ongoing**

Data Analysis (69 GB/s)

- Optimized data format
- Vectorization
 - ▶ Hillas parameters calculation (speed up 712, publication **soon**)
- Other algorithms
 - ▶ Model++ optimization **ongoing**
 - ▶ SVD Model optimization **ongoing**

Backups

CPU Architecture

CPU Architecture

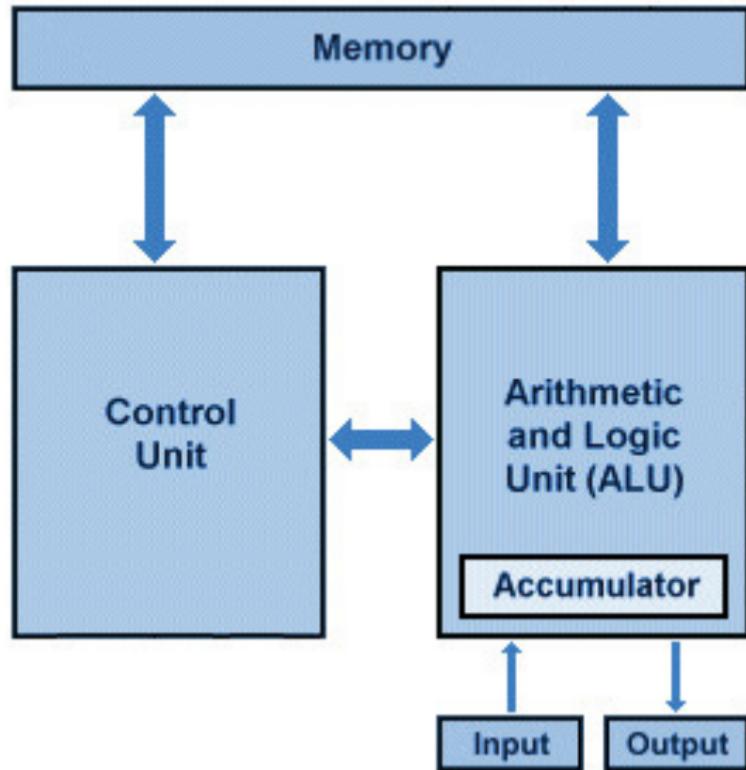
Von Neumann architecture 1945

Definition

Cycle : basis unit of time in a CPU

Time

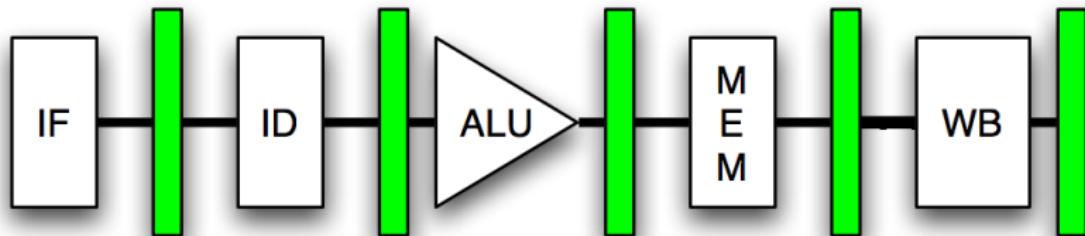
- 1 cycle per elementary operation (load, store, add, ...)
- 4 cycles per whole operation ($c = a + b$)



CPU Architecture

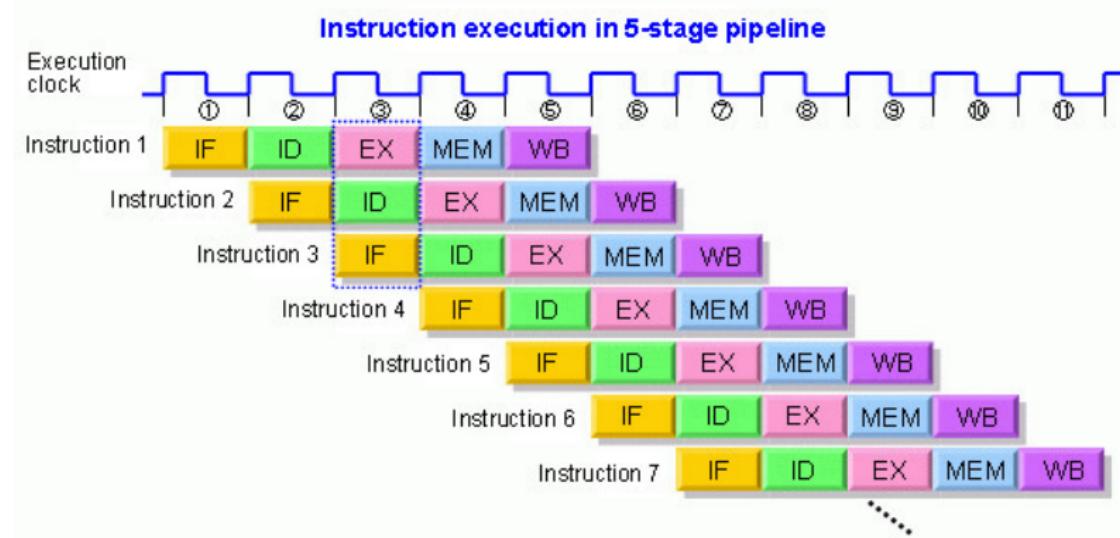
Pipeline approach

- IF : Instruction Fetch
- ID : Instruction Decode
- ALU : Execution
- MEM : Memory
- WB : Write Bytes



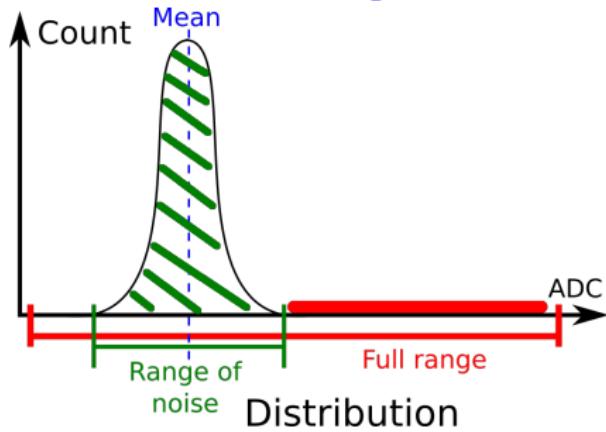
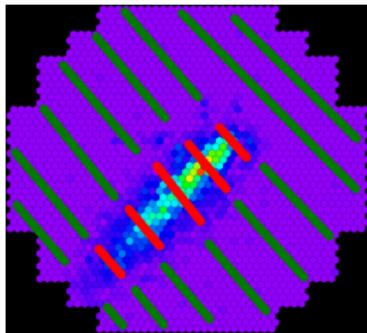
CPU Architecture evolution

Pipeline using

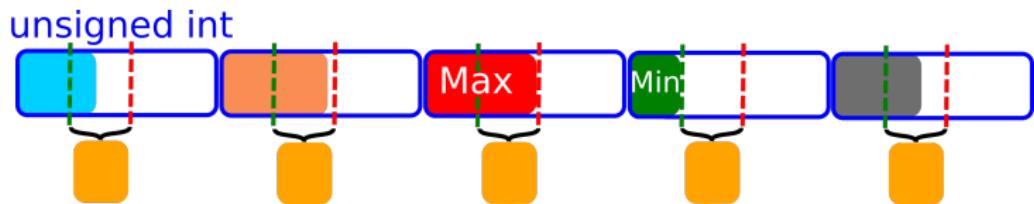
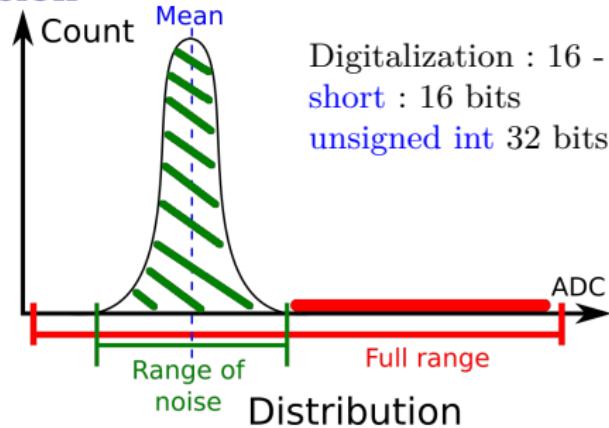
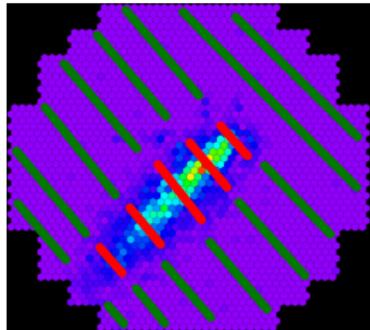


Data lossless compression

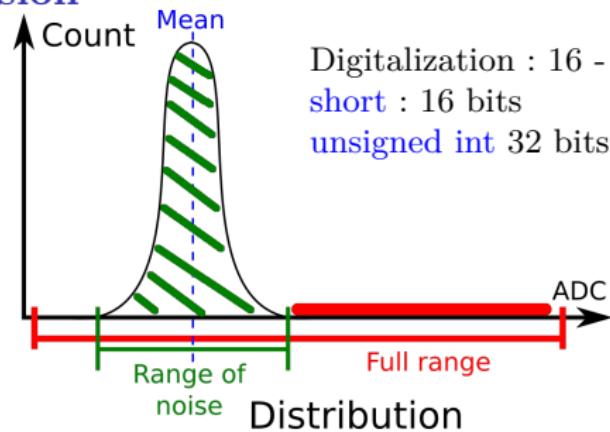
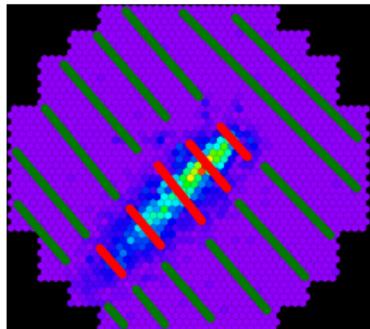
Digitalization : 16 - 24 bits
short : 16 bits
unsigned int 32 bits



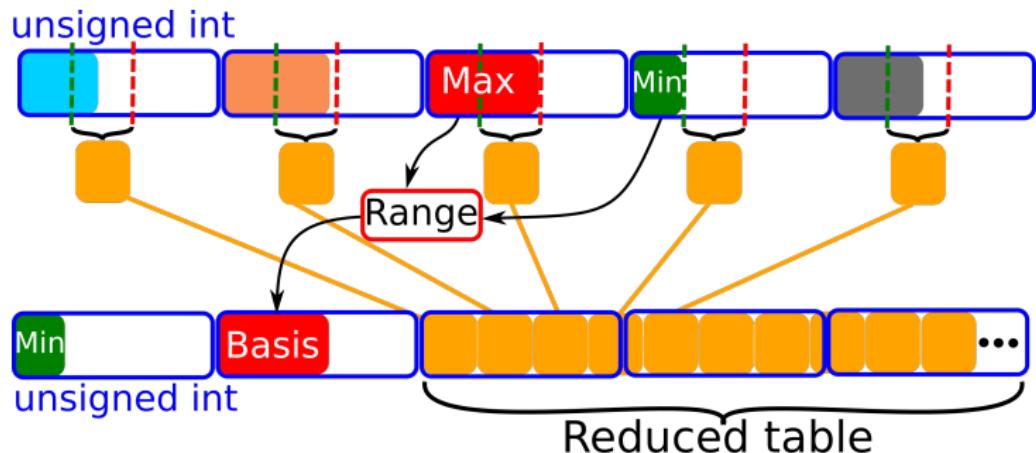
Data lossless compression



Data lossless compression

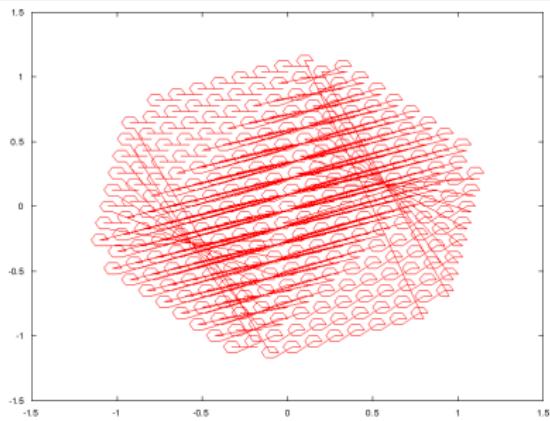


Digitalization : 16 - 24 bits
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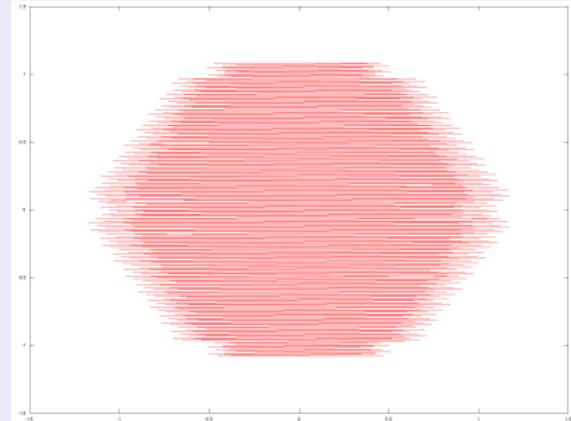


Sorting the camera's pixel to guaranty data locality

Before pixels sorting



After pixels sorting



Advantages

- Guaranty data locality
- Allows fast pixel's neighbours search
 - ▶ Faster Hillas cleaning
- Helps the data lossless compression