



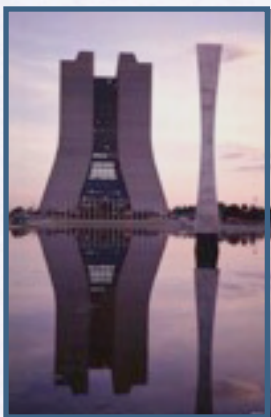
Initial results on ν_e appearance in MINOS

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MINOS in a nutshell

- Produce a high intensity beam of muon neutrinos at Fermilab.
- Measure background at the Near Detector and use it to predict the Far Detector spectrum.
- If neutrinos oscillate we will observe a distortion in the data at the Far Detector in Soudan.



← long baseline →
735 km

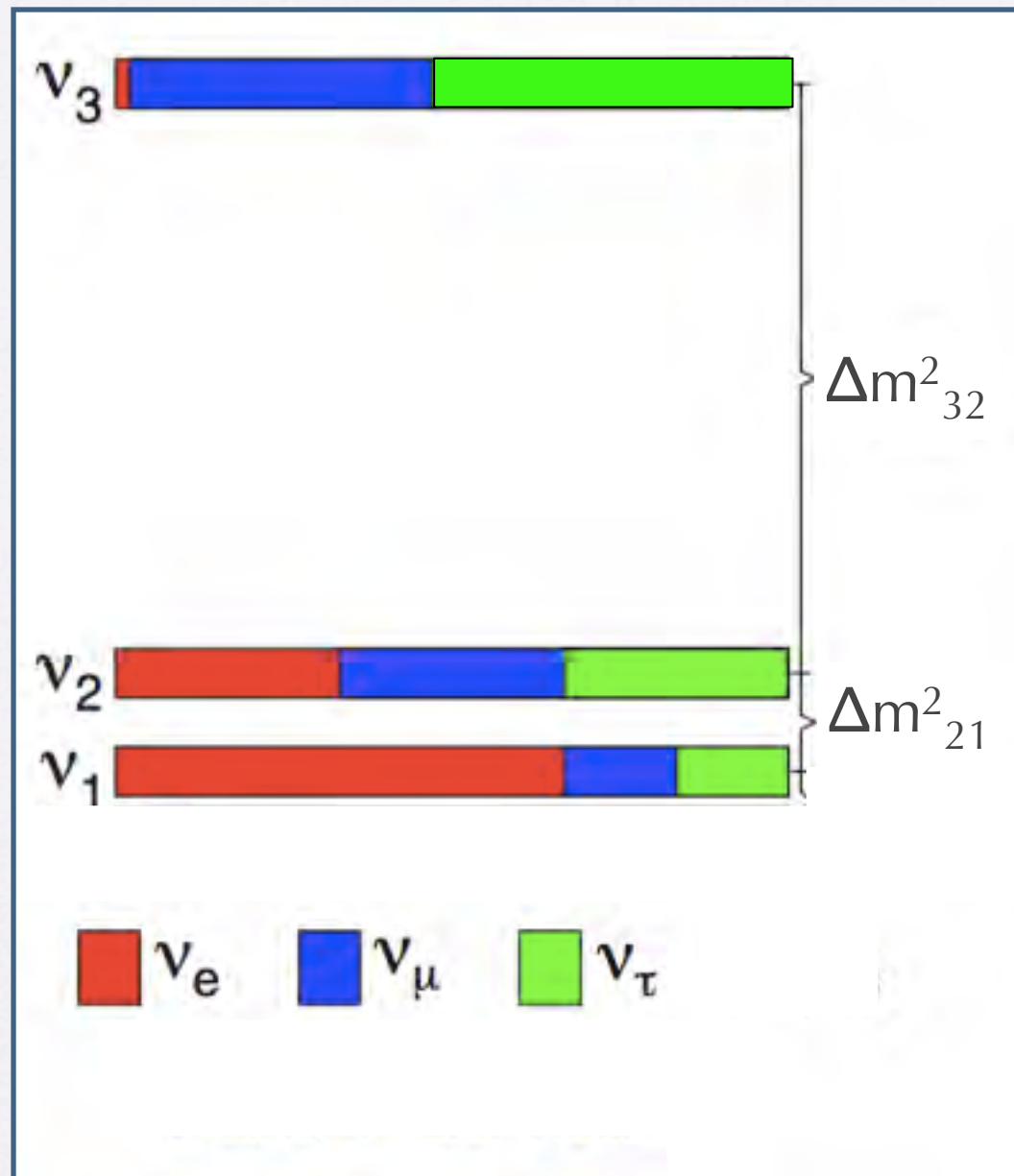


Main Injector Neutrino Oscillation Search



Neutrino masses and mixing

What is the current experimental picture?



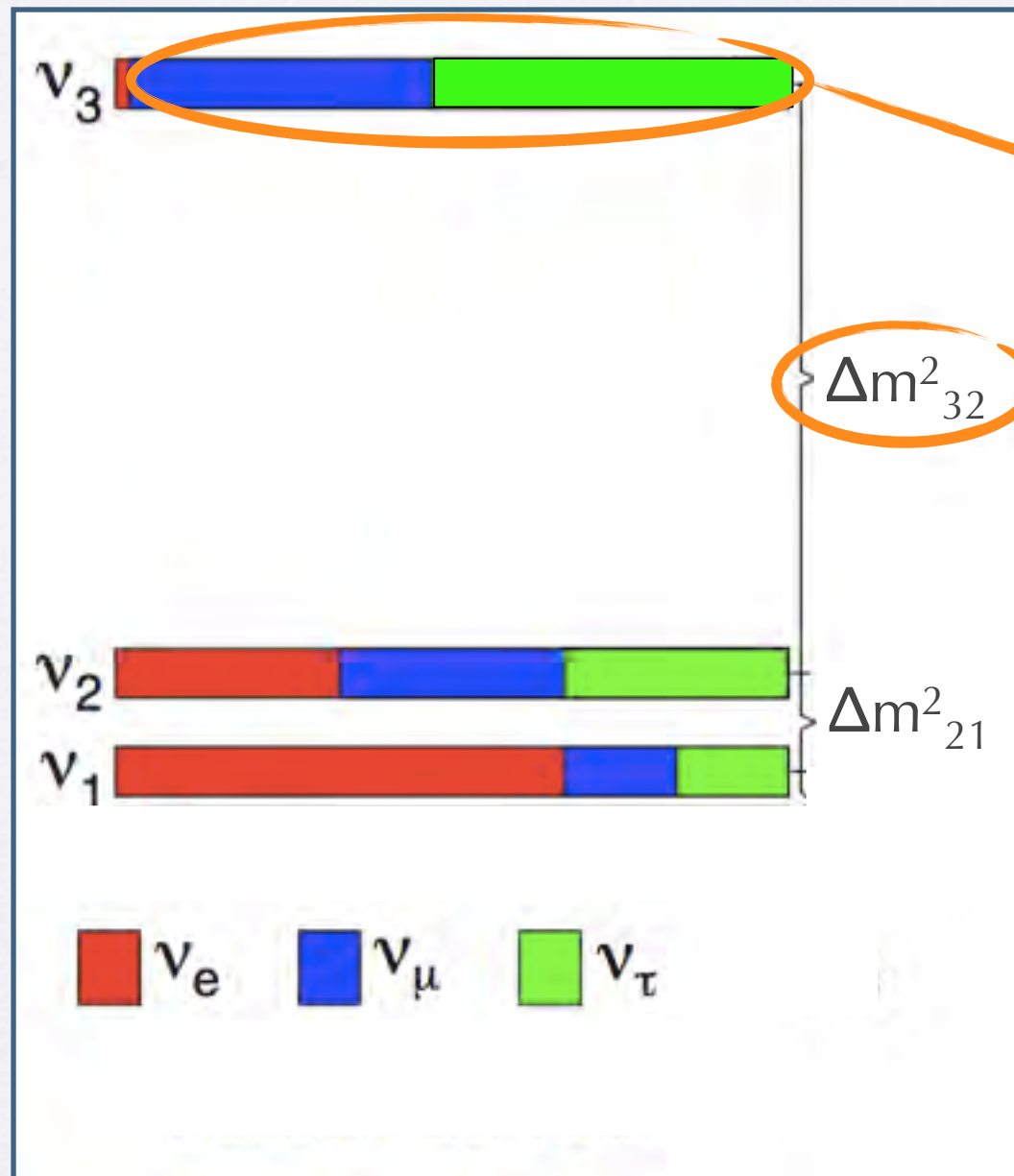
- Two mass scales:
 - *the atmospheric mass scale: Δm^2_{32}*
 - *the solar mass scale: Δm^2_{21}*
- *Large mixing angle for atmospheric neutrino oscillations.*
- Solar neutrino oscillations are subject to matter effects. Non maximal mixing angle.
- *Mass ordering known for the solar mass scale. Not known for the atmospheric.*

MINOS recent results

study “atmospheric” neutrino oscillation parameters

✓ Study ν_μ disappearance as a function of energy:

✓ Precision measurements of Δm^2_{32} and $\sin^2(2\theta_{23})$.



$$|\Delta m^2_{32}| = 2.43 \pm 0.13 \times 10^{-3} \text{ eV}^2 \text{ (68\% CL)}$$
$$\sin^2 2\theta_{23} > 0.90 \text{ (90\% CL)}$$
$$\chi^2/\text{ndf} = 90/97$$

✓ Mixing to sterile neutrinos at $\Delta m^2_{34} \sim \Delta m^2_{32}$:

✓ Fraction of neutrinos that oscillate into the sterile state.

Assuming no electron neutrino appearance:

$$f_s = 0.28^{+0.25}_{-0.28} \text{ (stat.+syst.)}$$

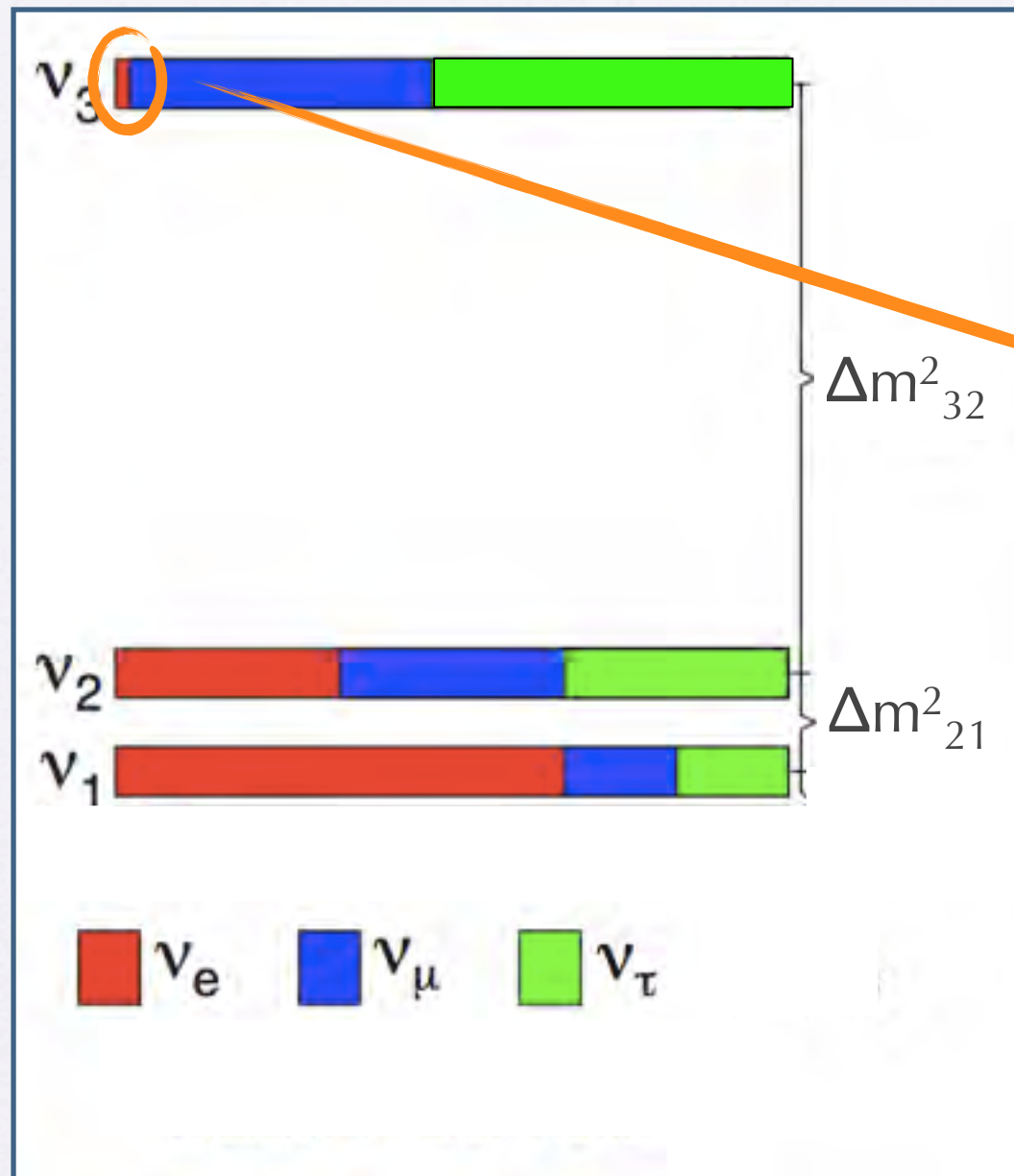
$$f_s < 0.68 \text{ (90\% CL)}$$

PRL **101** 131802 (2008)

PRL **101** 221804 (2008)

MINOS new results

study “atmospheric” neutrino oscillation parameters



♦ Search for subdominant oscillations i.e. $\sin^2(2\theta_{13}) > 0$

NEW!

Searching for θ_{13} in MINOS

- The probability of ν_e appearance in a ν_μ beam:

$$A \equiv \frac{G_f n_e L}{\sqrt{2} \Delta} \approx \frac{E}{11 \text{ GeV}}$$

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2(A-1)\Delta}{(A-1)^2} \\ + 2\alpha \sin \theta_{13} \cos \delta \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \cos \Delta \\ - 2\alpha \sin \theta_{13} \sin \delta \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \sin \Delta$$

$$\Delta \equiv \frac{\Delta m_{31}^2 L}{4E}$$

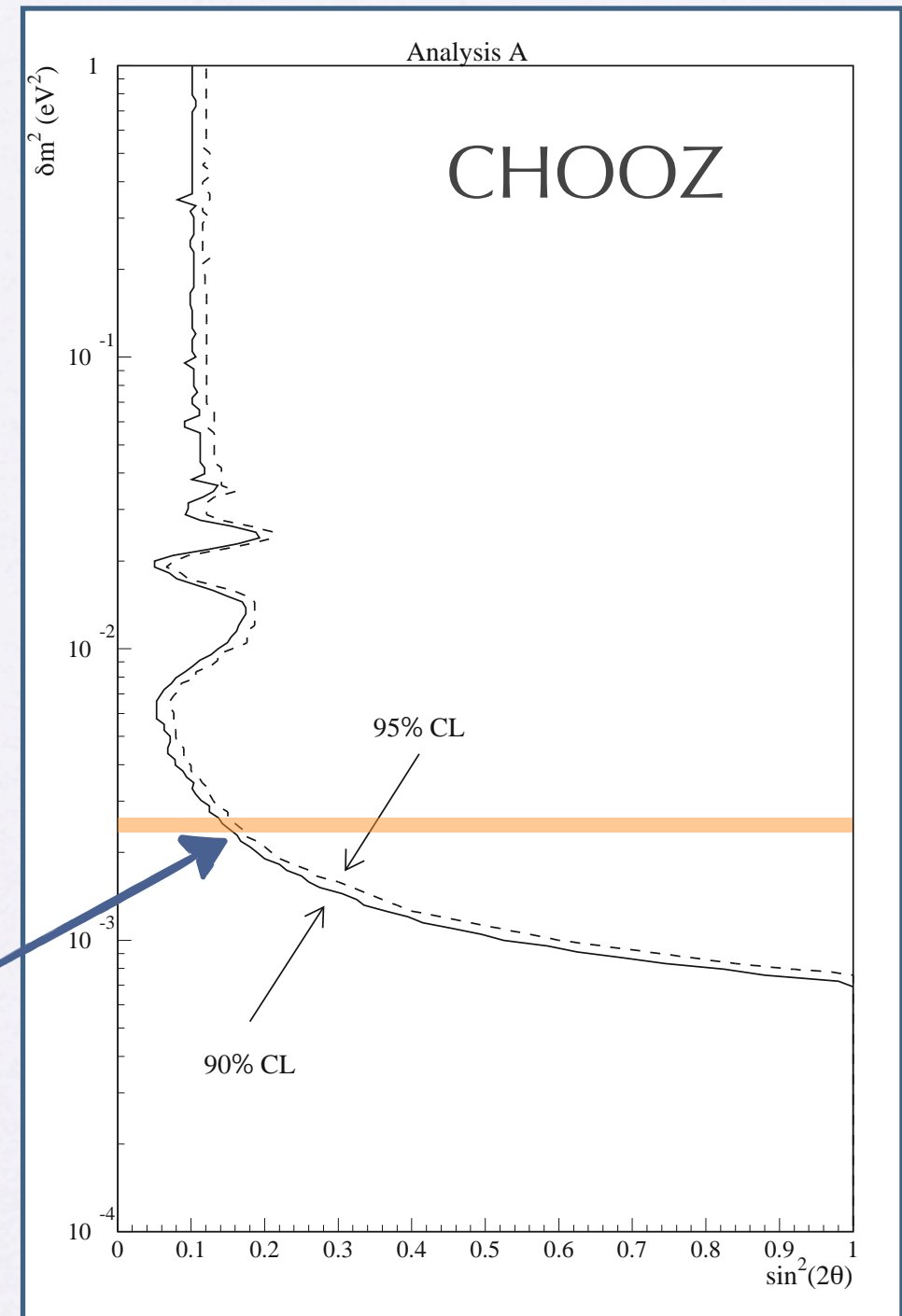
- Searching for ν_e events in MINOS, we can access $\sin^2(2\theta_{13})$.
- Probability depends not only on θ_{13} but also on δ_{CP} .
- Probability is enhanced or suppressed due to matter effects which depend on the mass hierarchy i.e. the sign of $\Delta m_{31}^2 \sim \Delta m_{32}^2$.

Relevant oscillation parameters

- The CHOOZ experiment published a limit in $\sin^2(2\theta_{13})$.
- Since then MINOS has measured Δm^2_{32} very precisely.
- Thus for this talk:

MINOS best fit
 $|\Delta m^2_{32}| = 2.43 \times 10^{-3} \text{ eV}^2$
 $\sin^2 2\theta_{23} = 1.00$

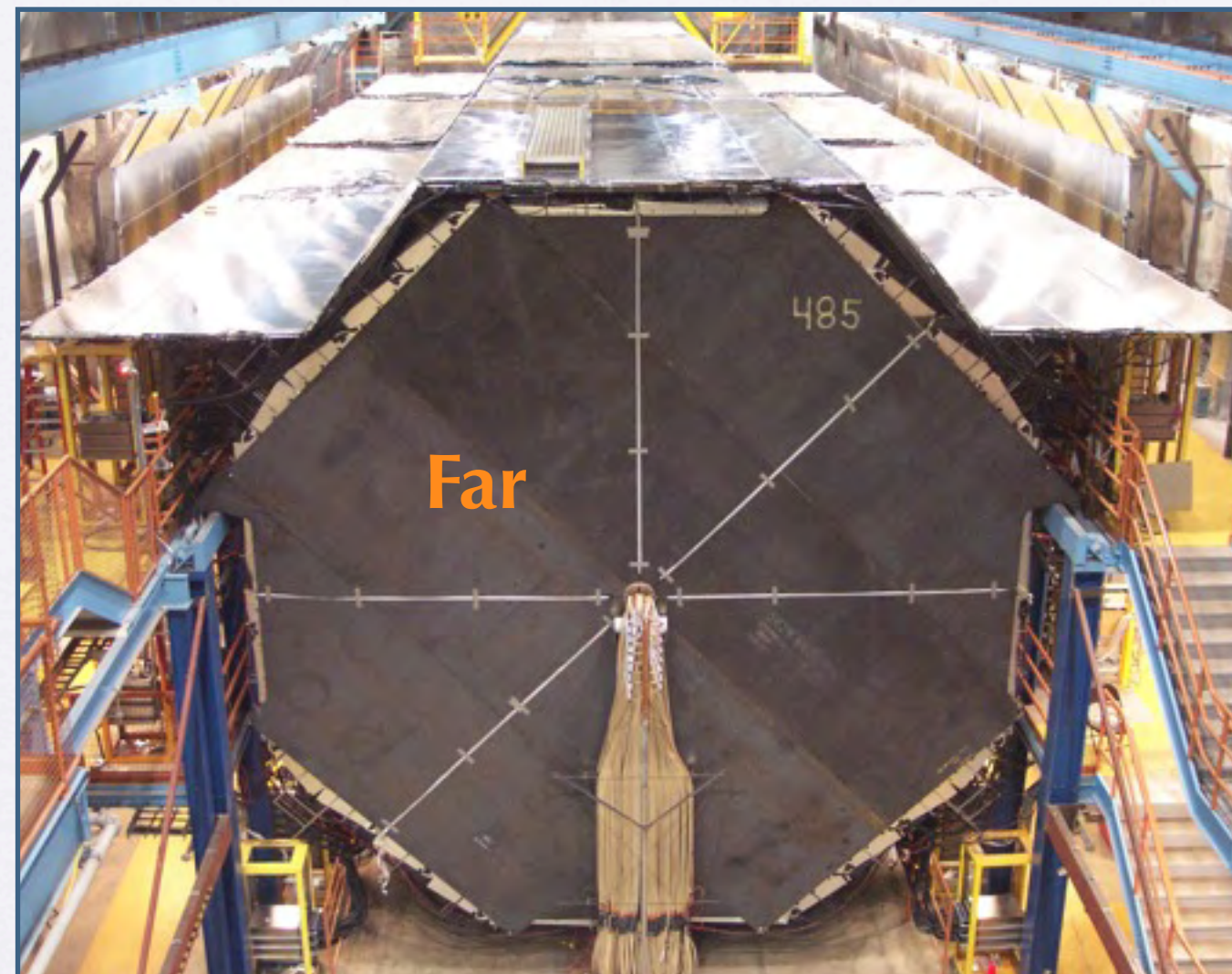
CHOOZ limit (90%CL)
 $\sin^2 2\theta_{13} = 0.15$



There are no measurements for δ_{CP} or the mass hierarchy.

The MINOS detectors

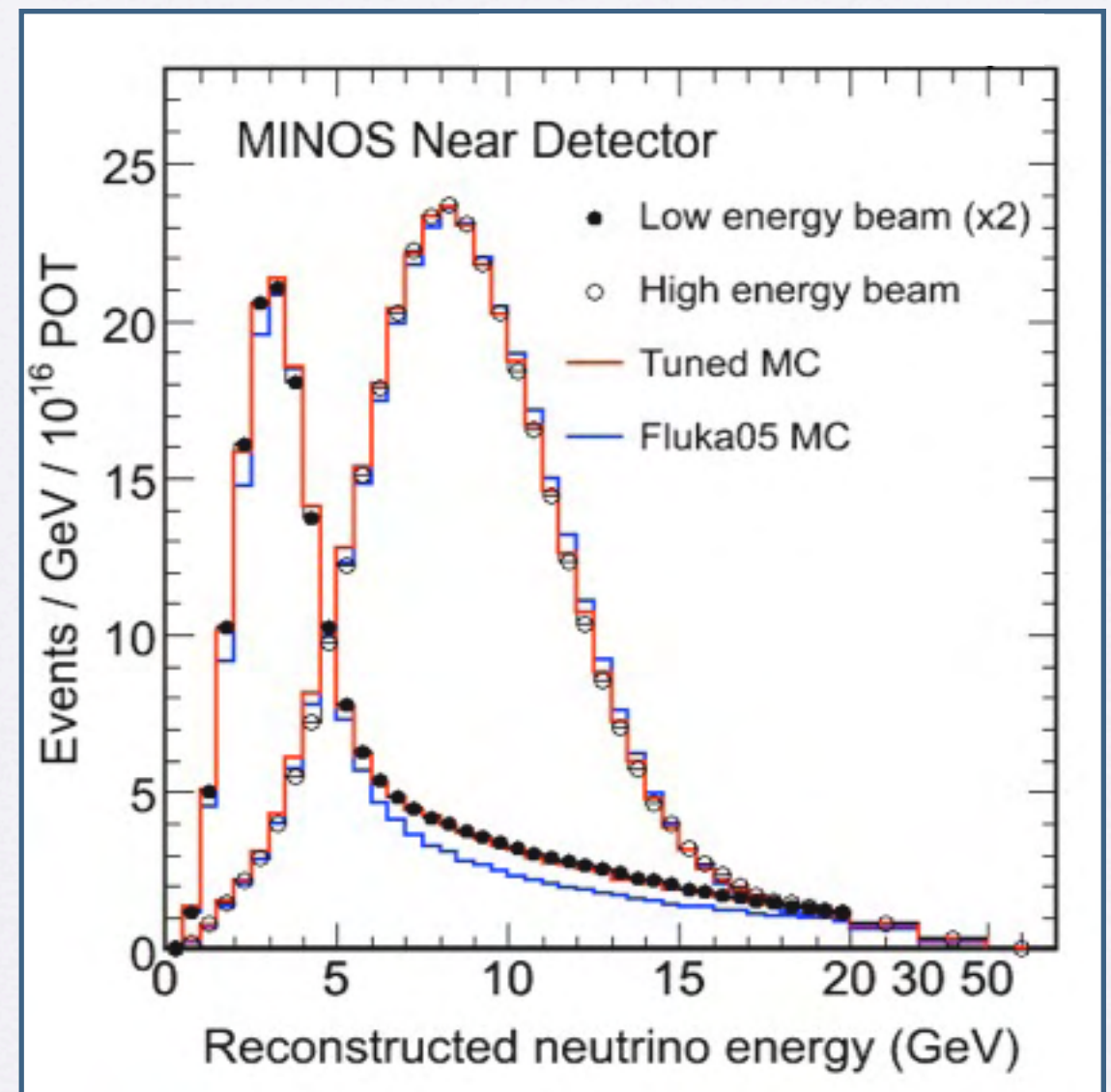
- Functionally identical: **Near and Far detectors**
- Octagonal steel planes (2.54cm thick $\sim 1.44X_0$)
- Alternating with planes of scintillator strips (4.12cm wide, Moliere rad ~ 3.7 cm).
 - **Near (ND)**: ~ 1 kton, 282 steel squashed octagons. Partially instrumented.
 - **Far (FD)**: 5.4 kton, 486 (8m/octagon) fully instrumented planes.



The NuMI Beam

- NuMI is primarily a ν_μ beam.
 - 1.3% of ν_e contamination from pion and kaon decays.
- Neutrino spectrum changes with target position with respect to focusing horns.
 - We use ν_μ CC events in ND to constrain flux.
- Region of interest dominated by events from secondary muon decays,
 - Constrained by ν_μ CC spectra.
- Uncertainties on the beam ν_e flux in the region of interest are $\sim 10\%$.

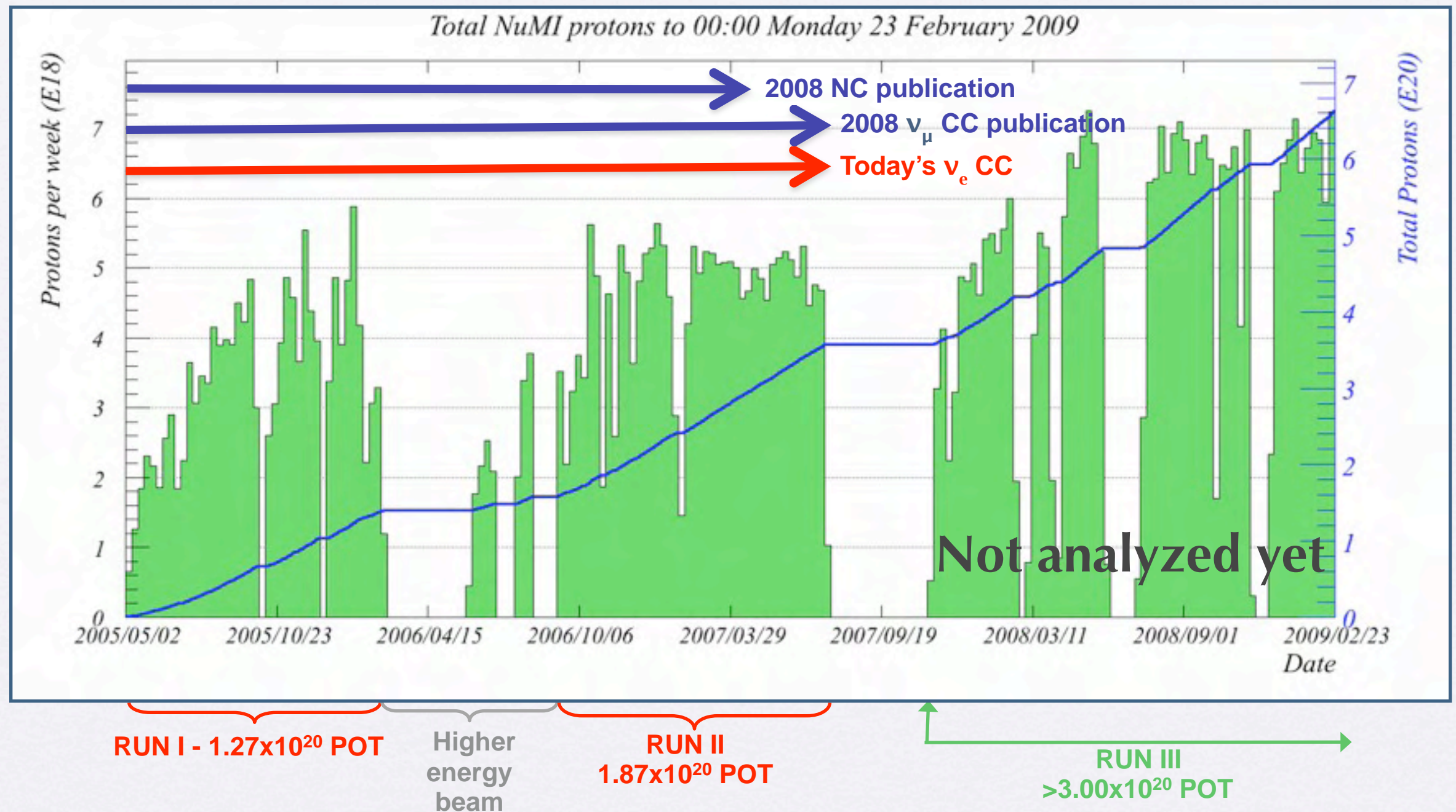
NeUtrinos from the Main Injector



Measured ν_μ CC events

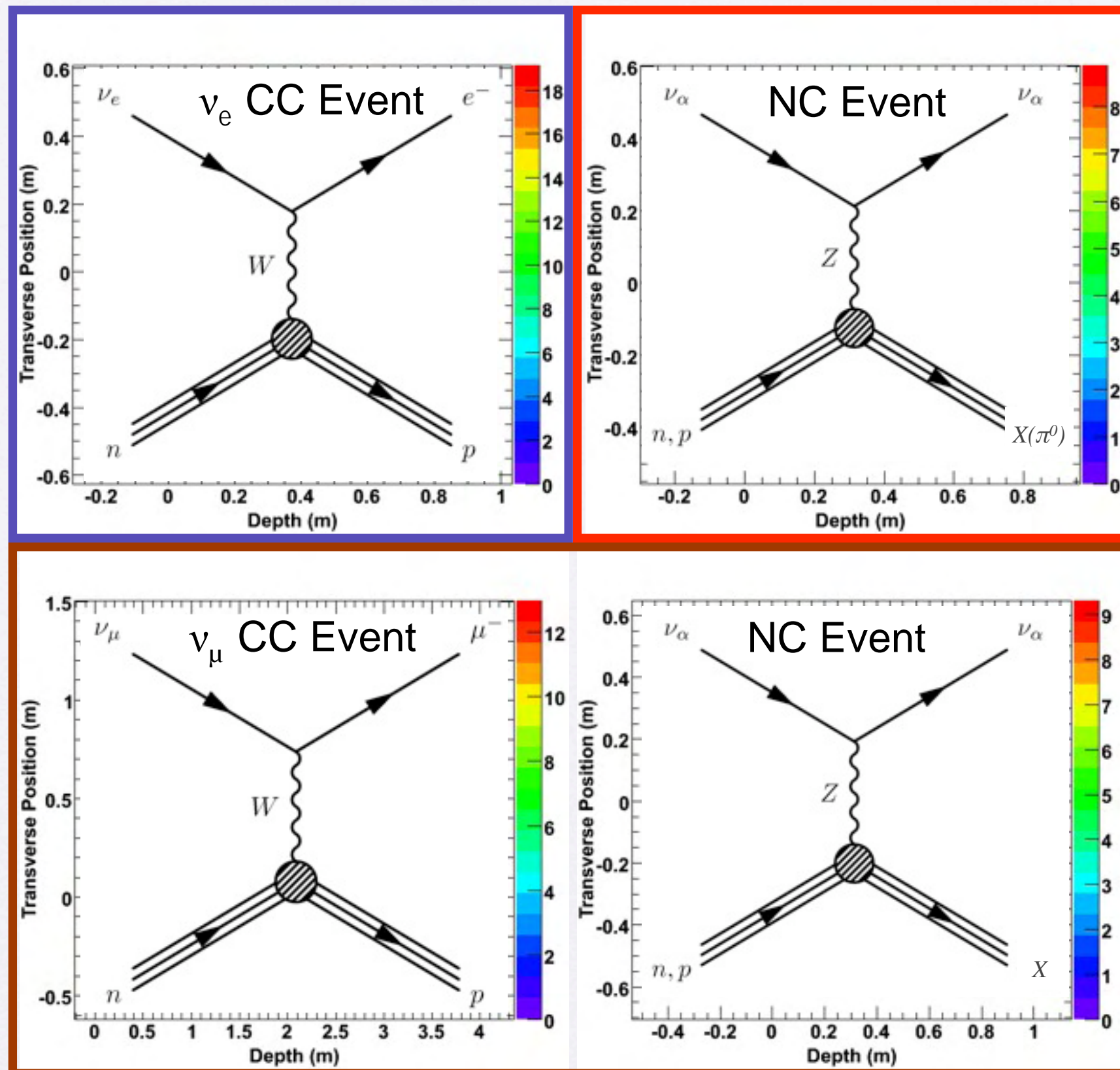
MINOS data

Current results on data through Run II



Results shown for run I+II - 3.14×10^{20} POT

Neutrino Event Topologies

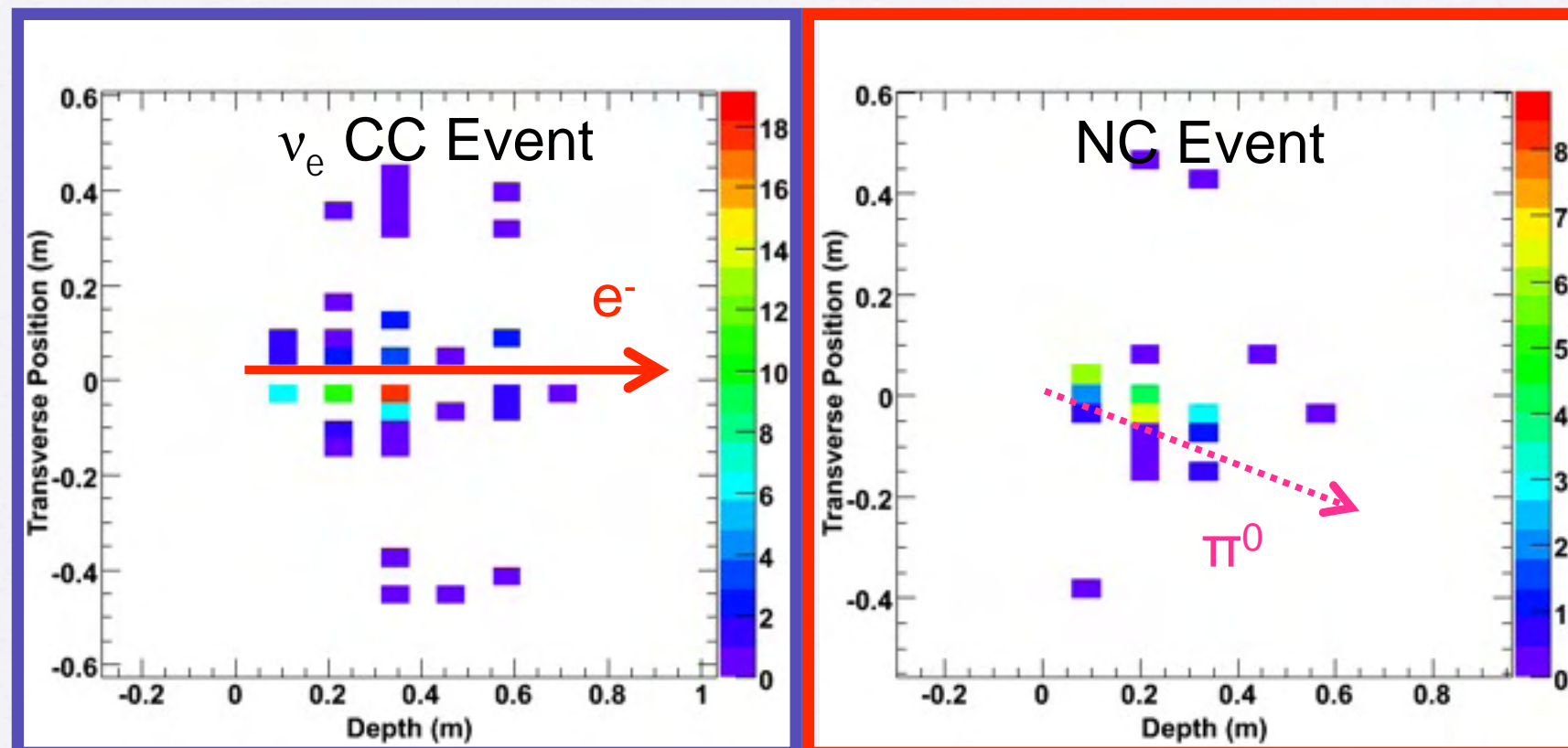


Neutrino Event Topologies

To select ν_e CC we focus on finding compact showers.

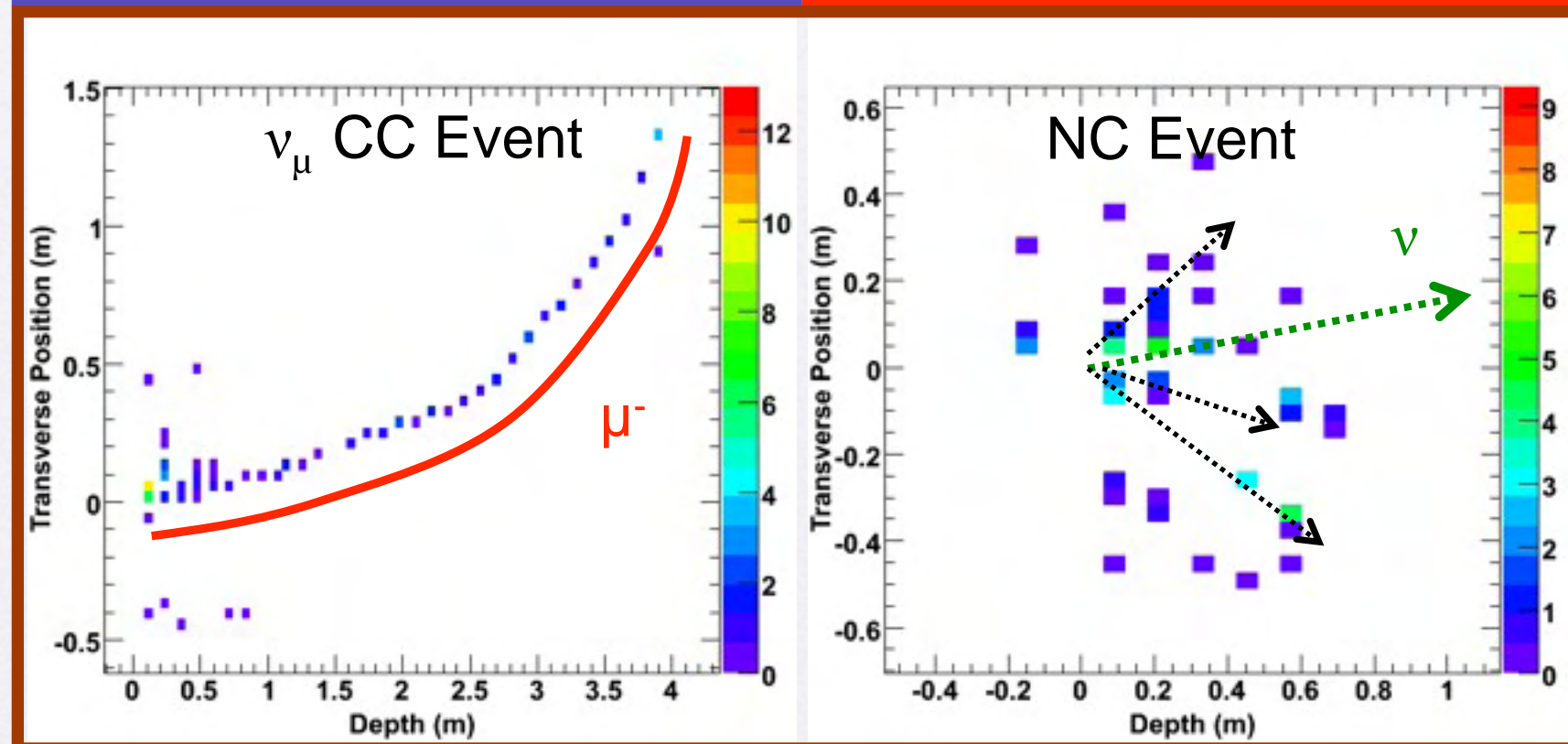
MC events

Signal



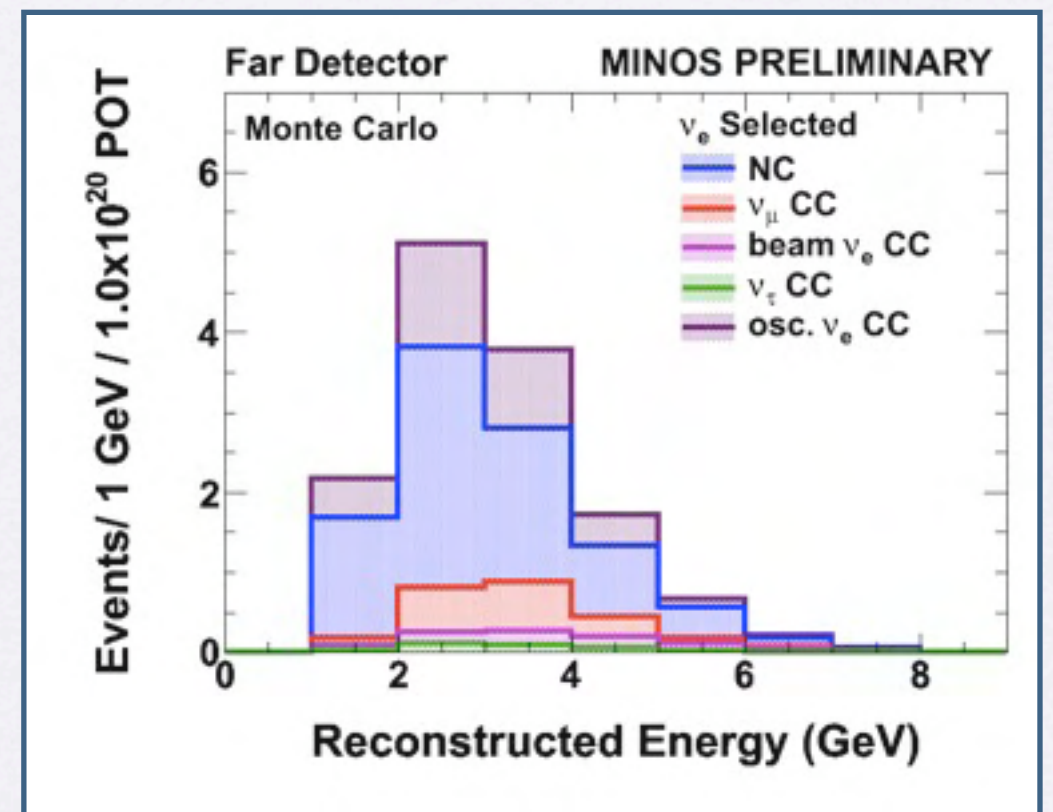
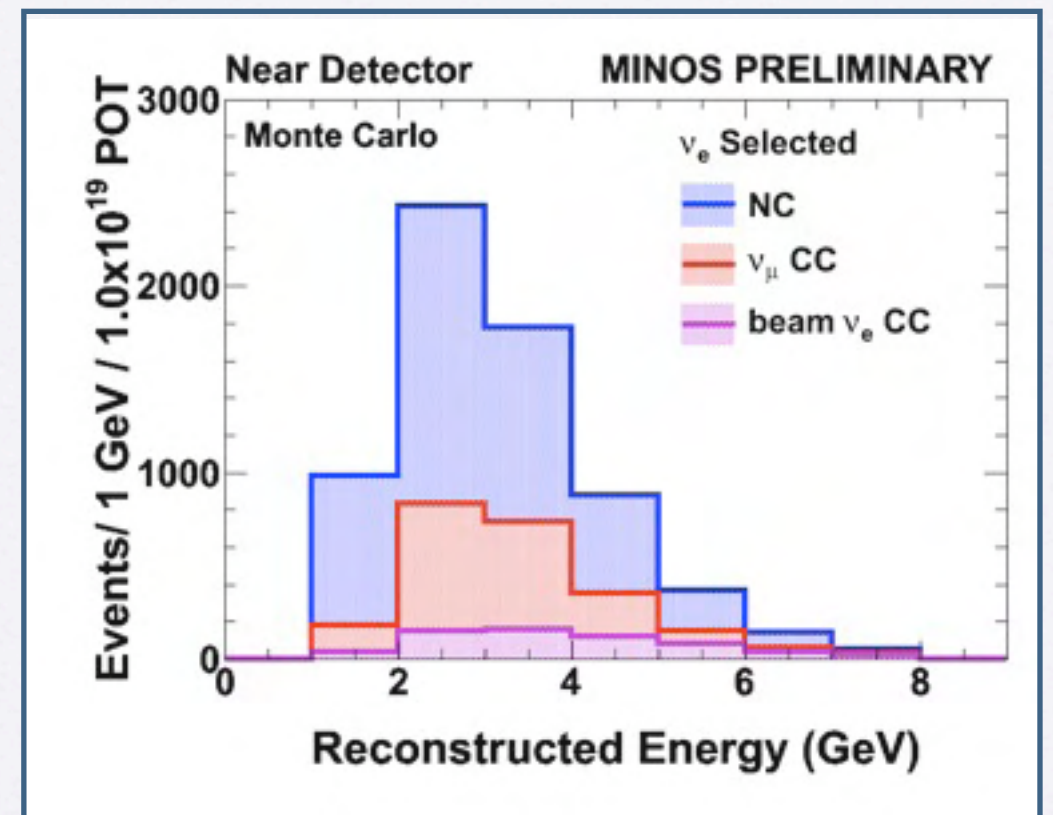
“Irreducible”
Background

Reducible
Background



ν_e appearance in MINOS

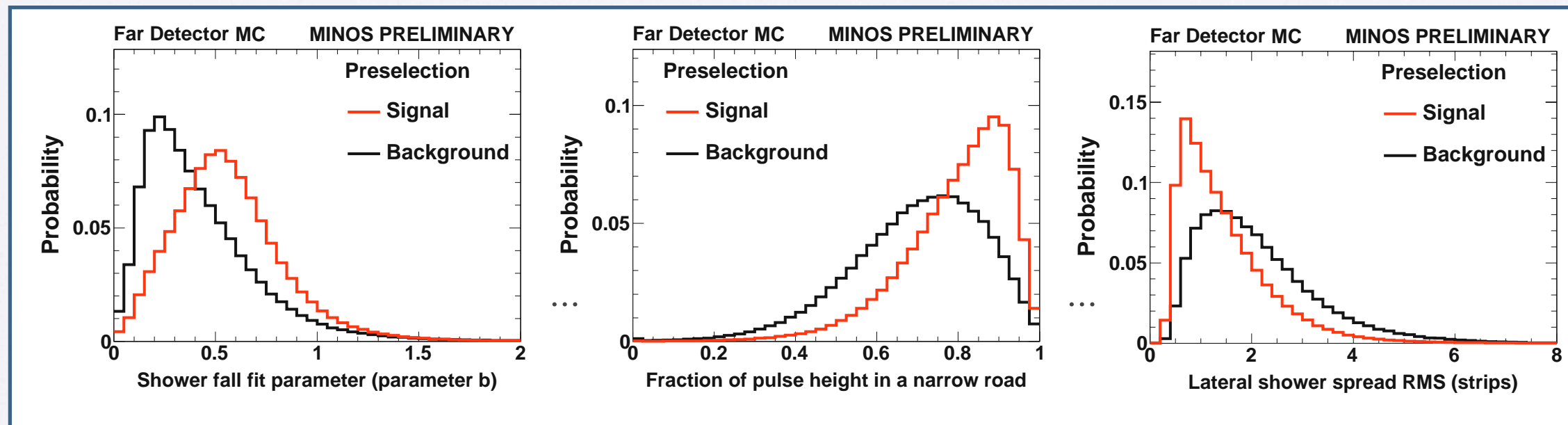
- **Select ν_e events** by finding electron candidates in the MINOS Detectors.
- **Measure the background** from events passing ν_e selection in the Near Detector.
 - **Separate the main background components** NC, ν_μ CC and beam ν_e CC since they extrapolate differently.
- **Extrapolate each background type** to the Far Detector taking into account ν_μ to ν_τ oscillations.
- **Look for an excess of ν_e events** in the Far Detector data. Cut and count events.



Selecting ν_e events

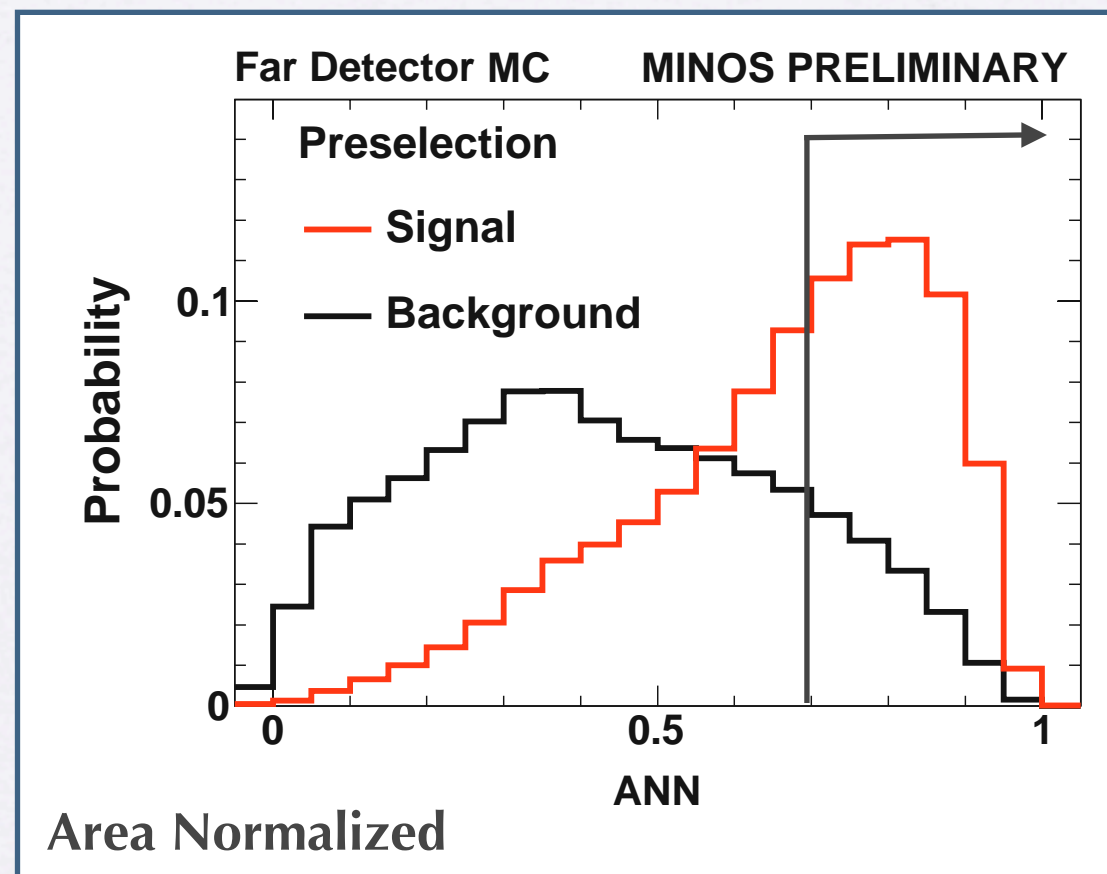
After selecting a shower dominated sample in signal energy region

Examples



- 11 variables describing length, width and shower shape.
- ANN algorithm achieves:
 - signal efficiency 41%
 - NC rejection >92.3%
 - CC rejection >99.4%
 - signal/background 1:4

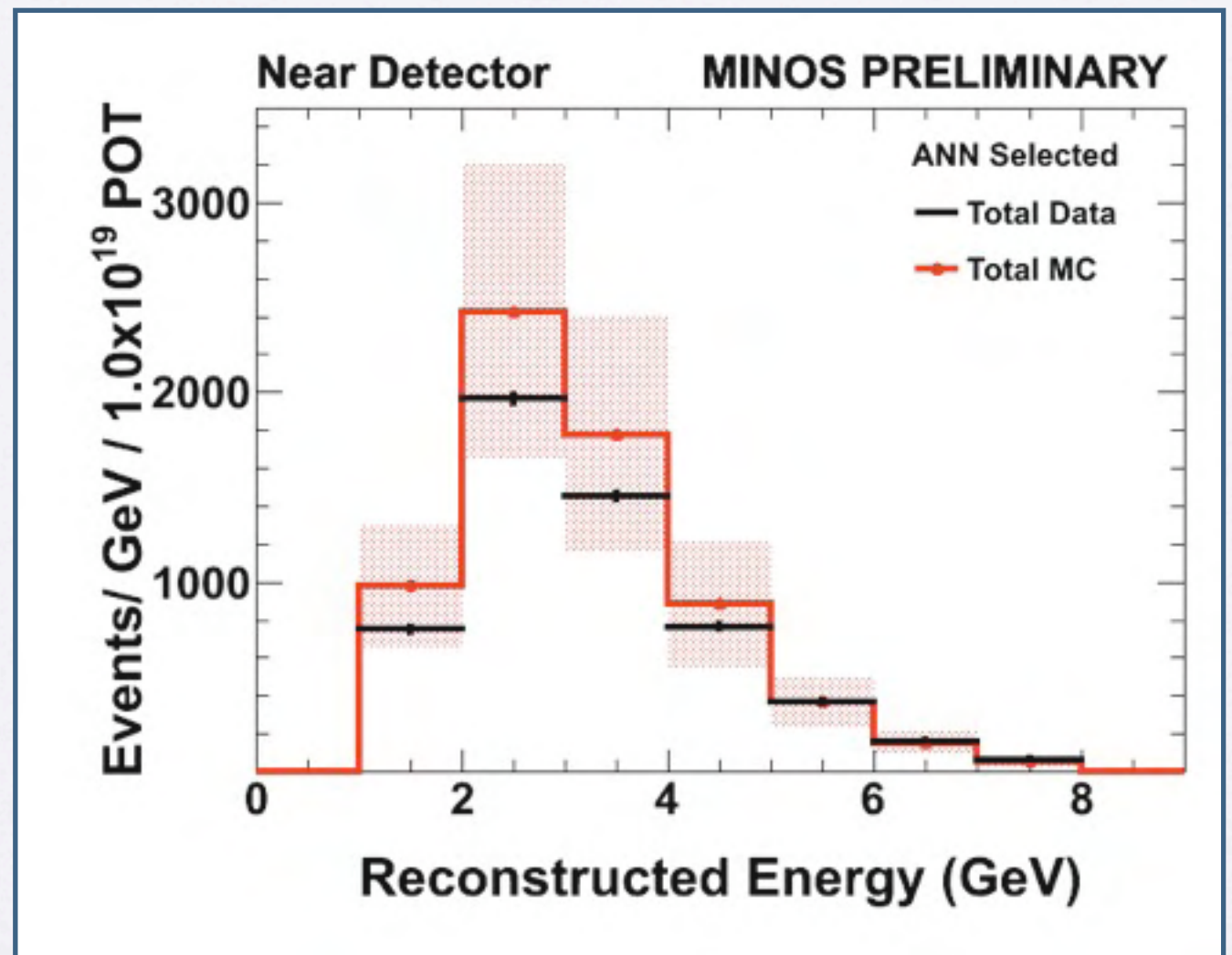
Secondary selection also studied.
Better signal efficiency and background rejection. Different systematics.



$$\Delta m_{32}^2 = 0.0024 \text{ eV}^2, \sin^2 \theta_{23} = 1.0$$

Estimating the background in Near Detector data

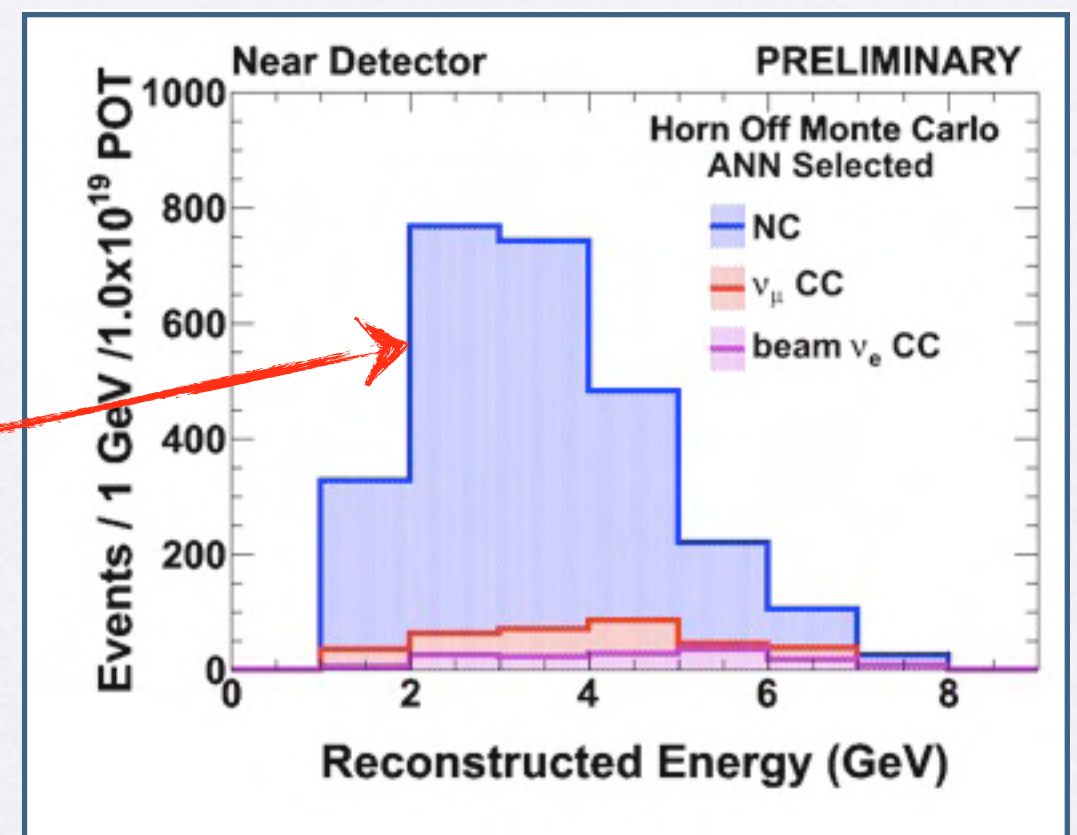
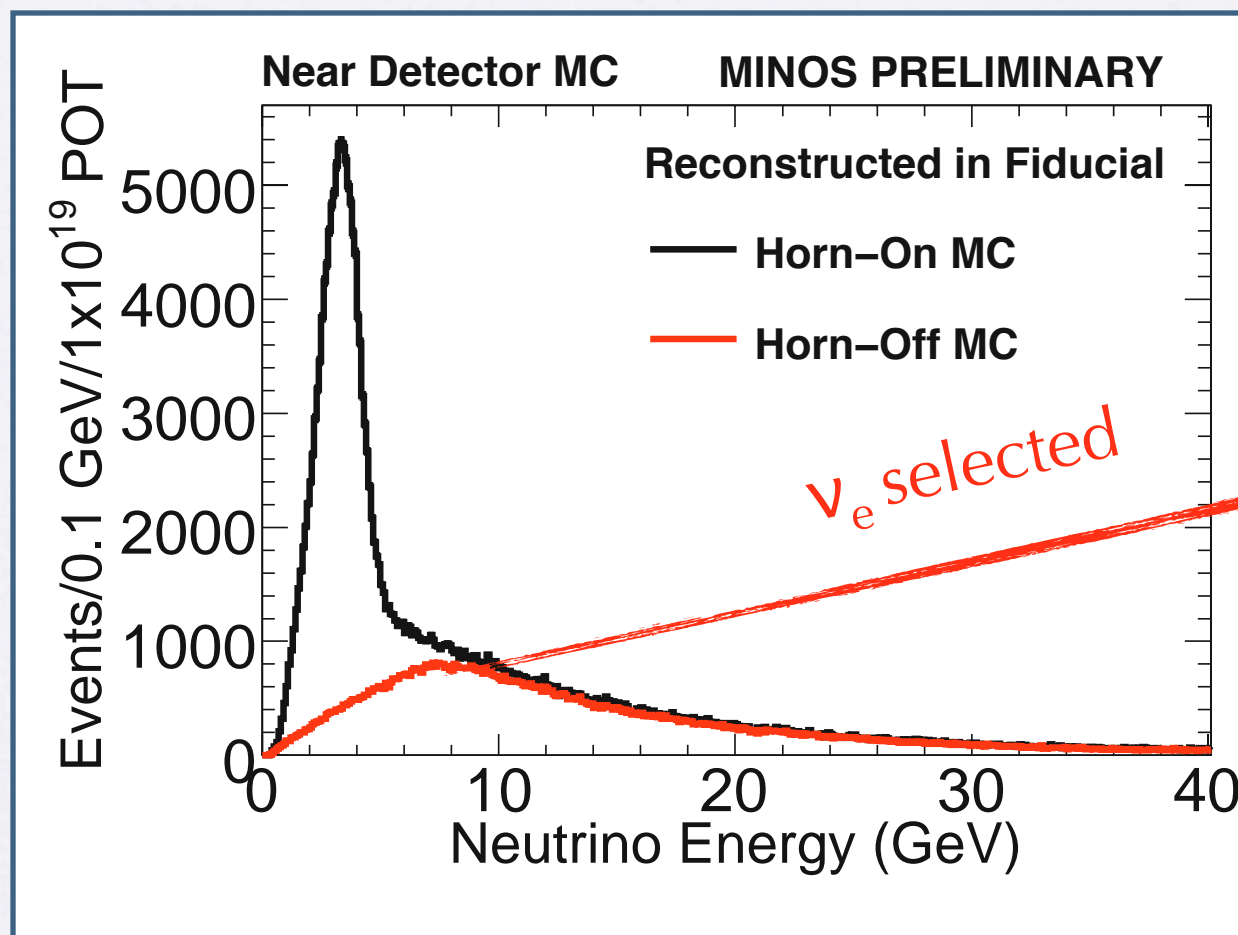
- Disagreement is within the large uncertainties of the model.
 - Mostly due to modeling of hadronic shower.
- We have developed **two data-driven methods**.
 - To measure the different background contributions in the Near Detector.



- The Horn on/off method uses the difference in background composition of the two horn configurations.
- The MRCC method uses muon removed ν_μ CC to study the hadronic showers and correct MC.

Separate background components NC vs CC

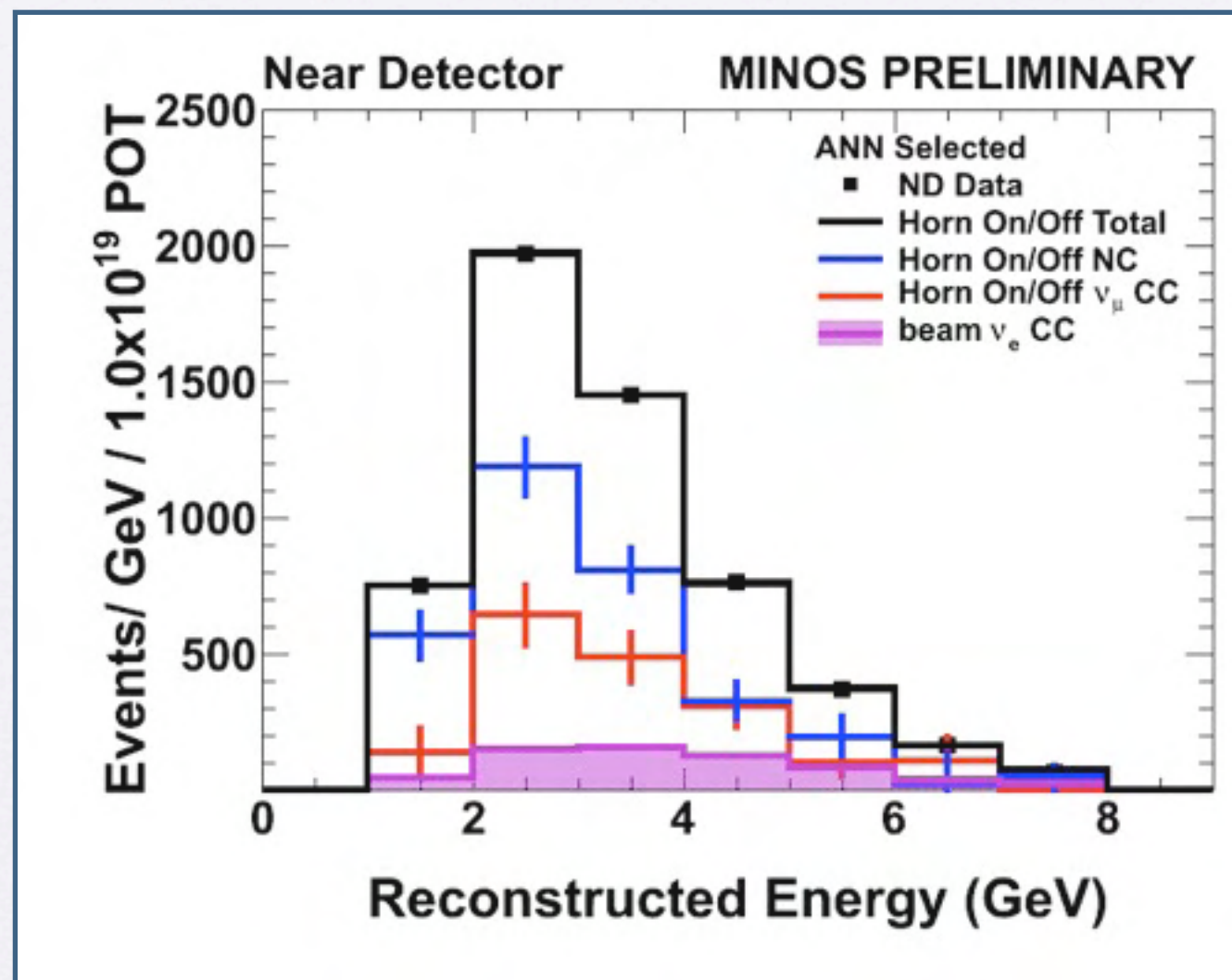
- Normal beam configuration has a peak which is focused by the horns.
- When beam horns are turned off, the parent pions do not get focused result in no peak.



- The consequence is a spectrum dominated by NC arising from the long tail in true neutrino energy that is ν_e selected in the region of interest in visible energy.

Separate background components NC vs CC

- We calculate the NC and ν_μ CC fractions by correcting the measurement using the horn off/on ratios for each components.
 - These ratios agree well with the data.



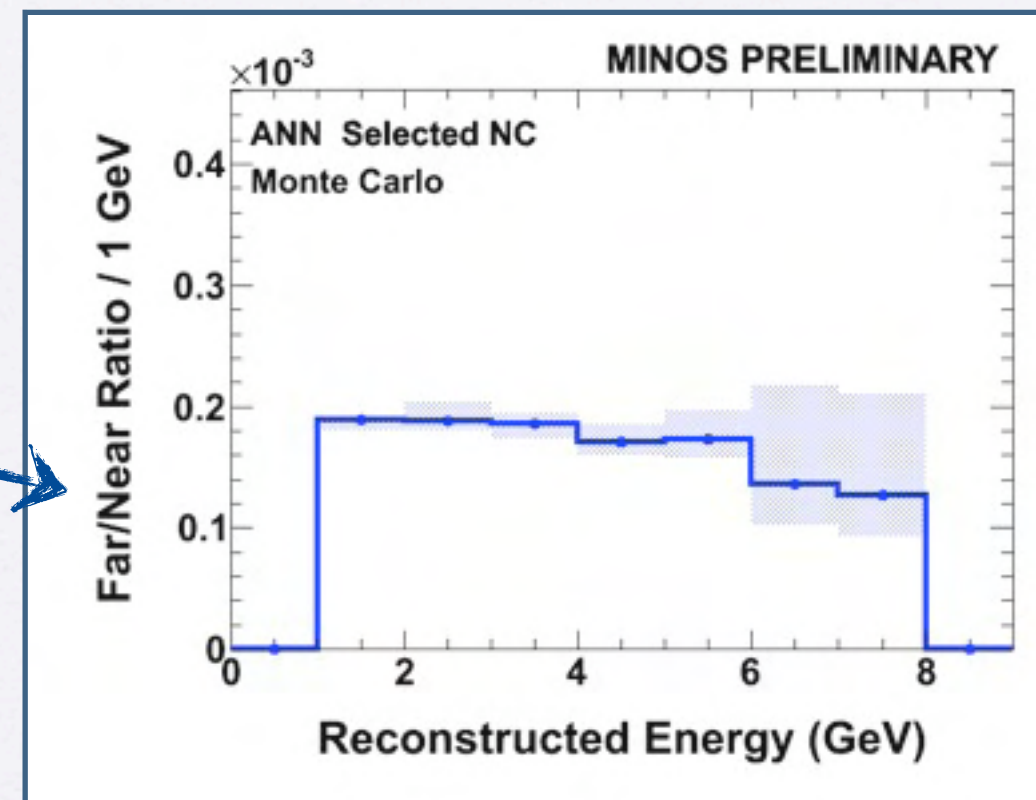
The two data-driven methods are in excellent agreement.

Predicting the FD background

- The Near Detector ν_e selected **NC** and ν_μ **CC** background components are corrected by the Far/Near MC ratio.

$$FD_i^{predicted} = \frac{FD_i^{MC}}{ND_i^{MC}} ND_i^{Data}$$

$\sim 10^{-4}$ expected from geometry and fiducial volume ratio alone



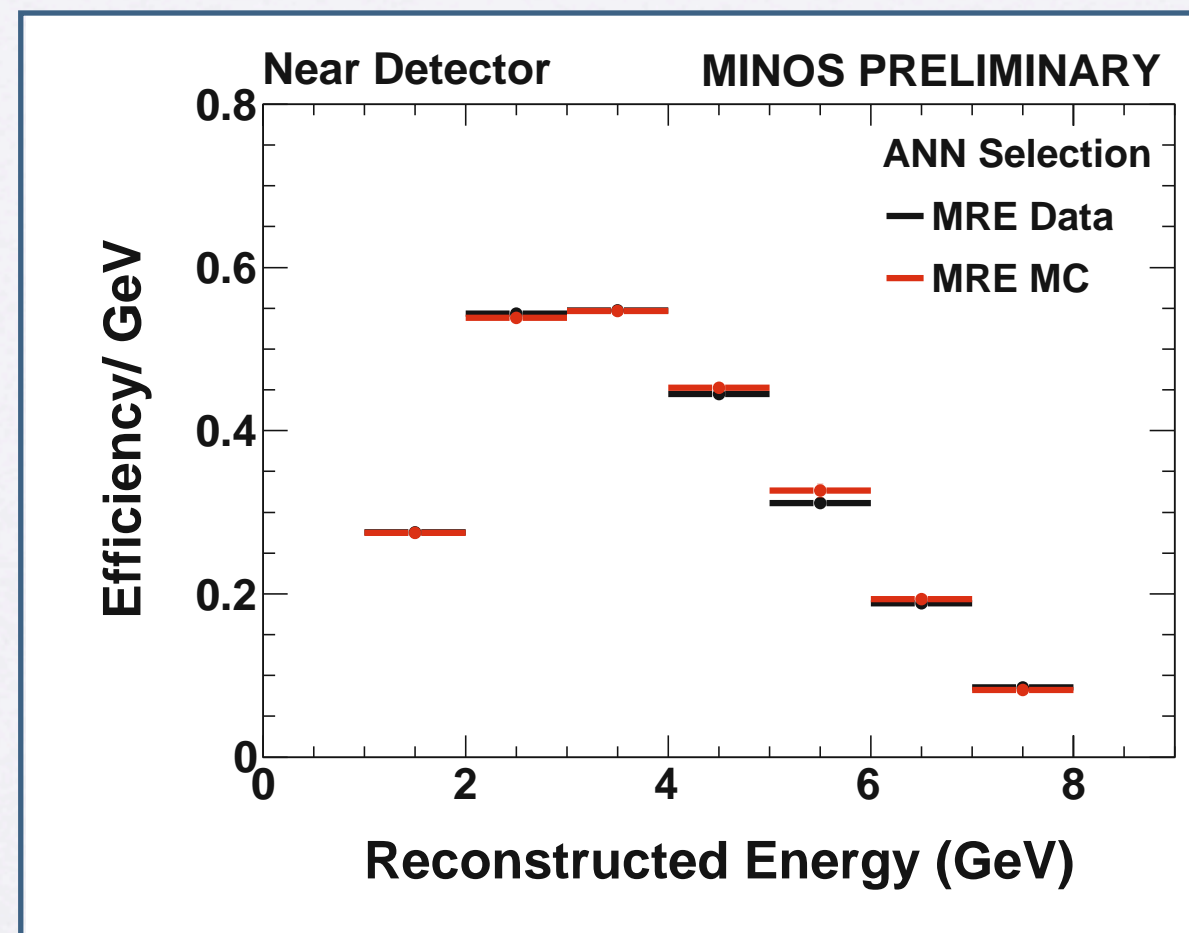
- Far/Near ratio accounts for geometry, fiducial volume ratio, intensity, detector differences and oscillations.
- The **signal** ν_e and the ν_τ **CC** from ν_μ oscillations are corrected using the extrapolation of the ν_μ CC spectrum. The **beam** ν_e in the Far Detector are taken from the MC.

The background prediction at 3.14×10^{20} POT is:
 $27 \pm 5(\text{stat}) \pm 2(\text{sys})$

Estimating the signal efficiency

- The ν_e selection is applied to muon removed ν_μ CC with electron added data and MC.
- The ratio of data/MC is applied as a correction to the signal efficiency.
- Expected signal depends on δ_{CP} and the mass hierarchy.

**The signal prediction at the CHOOZ limit (3.14×10^{20} POT)
ranges from: 6 to 12 events.**



ν_e appearance result:

MINOS PRELIMINARY

ν_e appearance result:

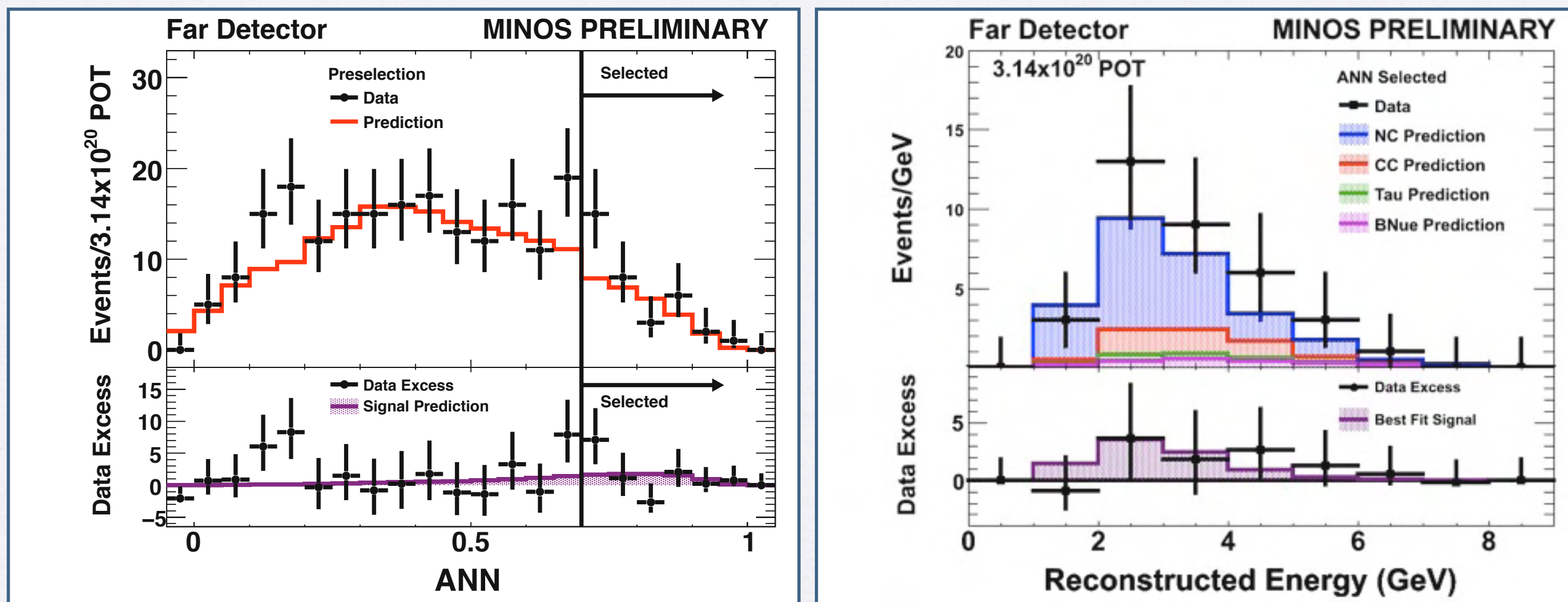
Observation 35 events

Expected Background $27 \pm 5(\text{stat}) \pm 2(\text{sys})$
for 3.14×10^{20} POT

MINOS PRELIMINARY

ν_e Selected Far Detector Data

- Blind analysis done. Background/signal predictions and systematic errors finalized before looking at data in Far Detector.

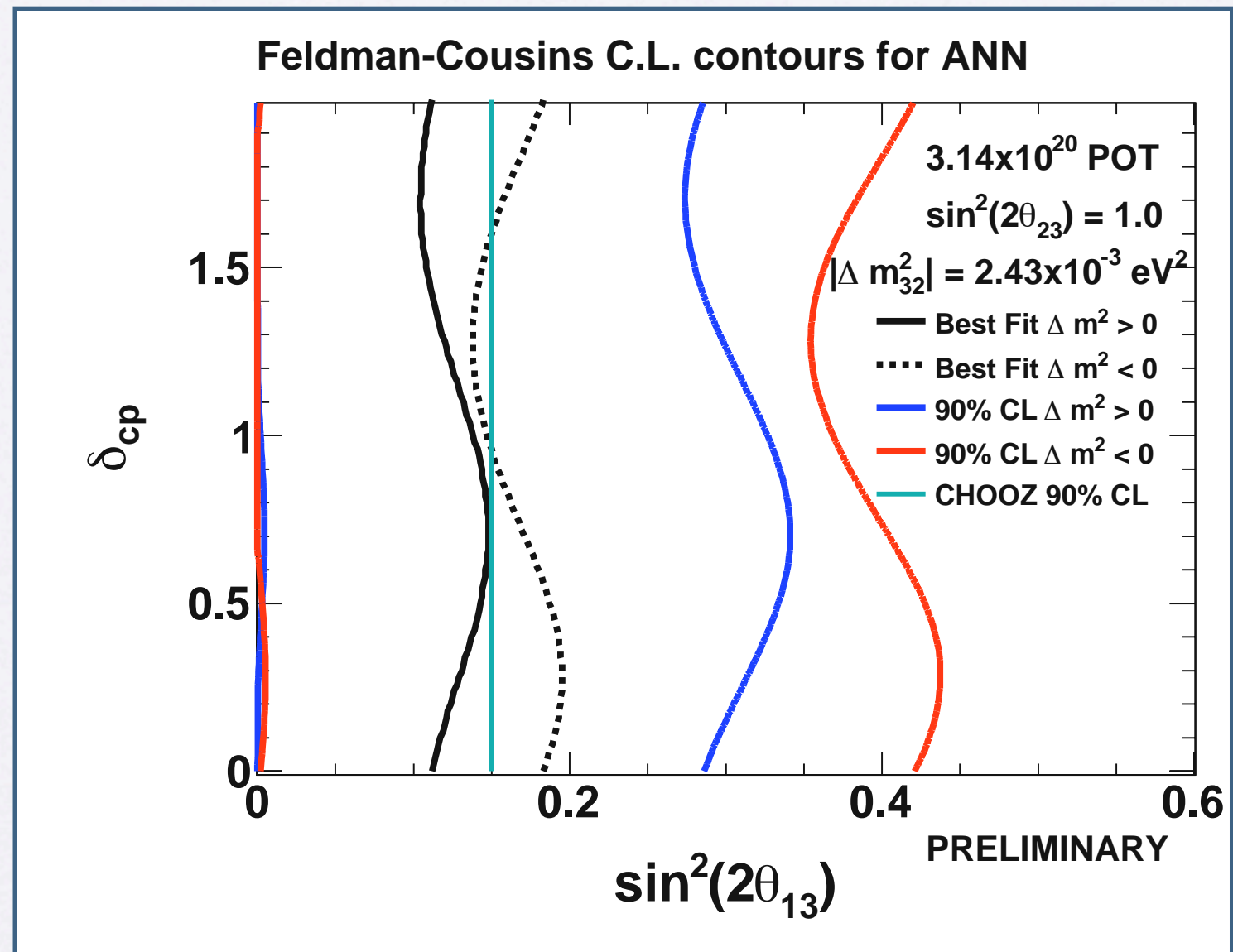


- Sidebands studied in Far Detector: muon removed and below signal region samples. Both consistent within 1-2 σ .

MINOS 90% CL in $\sin^2 2\theta_{13}$

Fitting the oscillation hypothesis to our data

- Plot shows 90% CL limits in δ_{CP} vs. $\sin^2 2\theta_{13}$
 - shown at the MINOS best fit value for Δm^2_{32} and $\sin^2 2\theta_{23}$.
 - for both mass hierarchies
- A Feldman-Cousins method was used.
- Results are for primary selection and primary separation method.



Results are consistent with secondary selection and secondary separation method

Summary

- We have completed an **initial search for ν_e appearance** in the MINOS data.
 - Developed two ν_e selections and two data-driven background estimates.
- We observe a total of **35 events** and expect **$27 \pm 5(\text{stat}) \pm 2(\text{sys})$** background events for 3.14×10^{20} POT.
- If fitted to a oscillation hypothesis we obtain the limits at the MINOS best fit for Δm^2_{32} and $\sin^2(2\theta_{23})$:
 - normal hierarchy, $\delta_{CP} = 0$: **$\sin^2(2\theta_{13}) < 0.29$ (90% CL)**
 - inverted hierarchy, $\delta_{CP} = 0$: **$\sin^2(2\theta_{13}) < 0.42$ (90% CL)**

We are close to doubling these data in current running!
Expect next results with $\sim 7 \times 10^{20}$ POT