



# Mass-Radius Relation of Neutron Stars in a Scalar-Tensor Theory

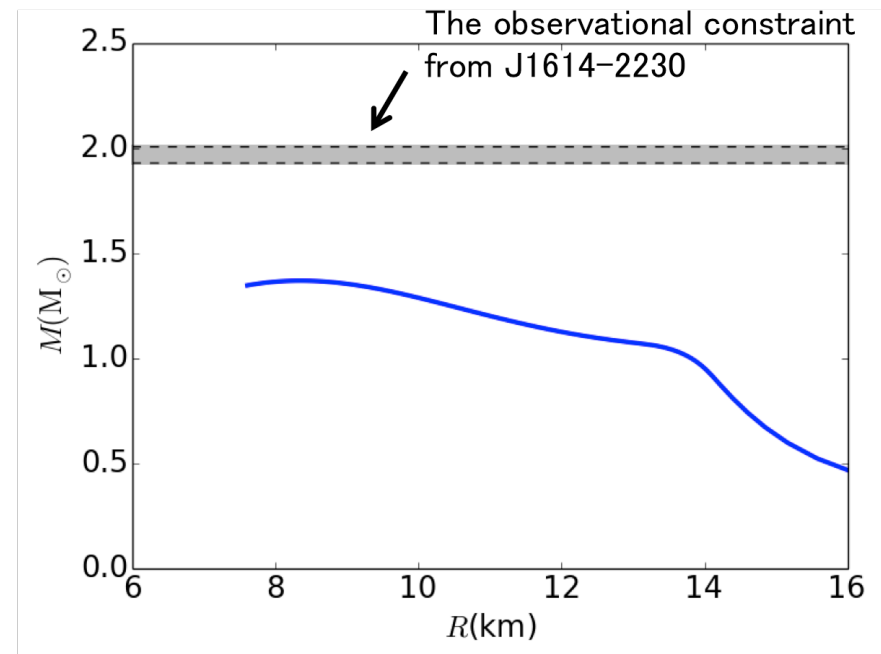
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# Introduction

- Tests of GR by NS-NS coalescence
- The existence of the  $2M_{\odot}$  NS[1]

The structure of NSs in  
alternatives to GR?  
 $2M_{\odot}$  can be explained?



MR relation with GS1 EOS

# Scalar-tensor theory

The scalar-tensor theory proposed in [2]

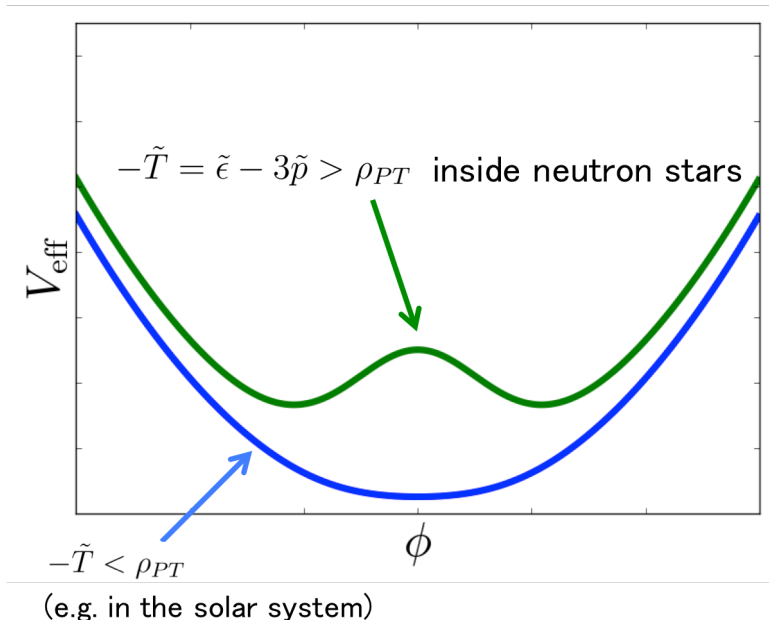
Include a massive scalar field,  $\phi$

$\phi$  = Dark matter

(Properties)

- The effective gravitational constant depends on  $\phi$ .
- The effective potential of  $\phi$  depends on the density of the surrounding matter.

# Symmetry breaking inside NSs



Gravity is weakened  
inside NSs

→ More massive NSs  
may be allowed.

$\eta$  :How much gravity is weakened

$\rho_{PT}$  :Critical density for symmetry breaking

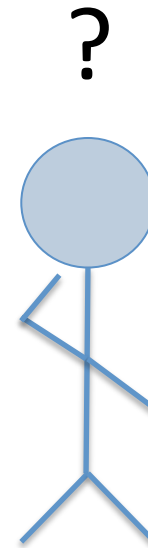
$\lambda_\phi$  :Compton wave length

# Structure of NS

$$\lambda_\phi \lesssim 10\text{km} \sim R_{\text{NS}}$$



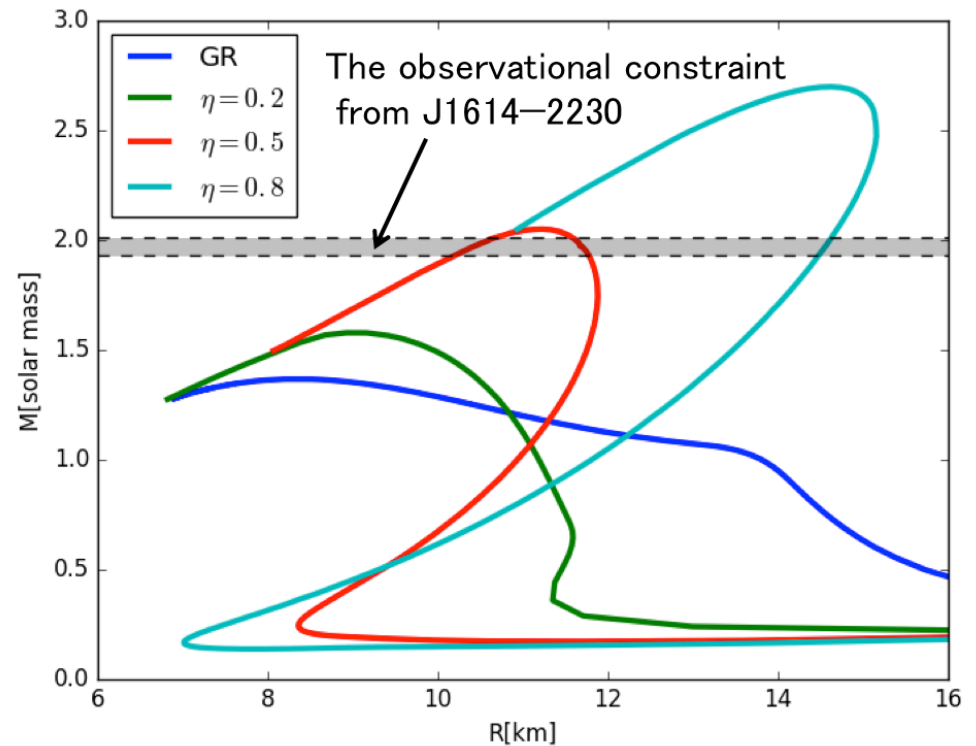
NS



- ✓ Solar system experiments
- ✓ Binary pulsar observations

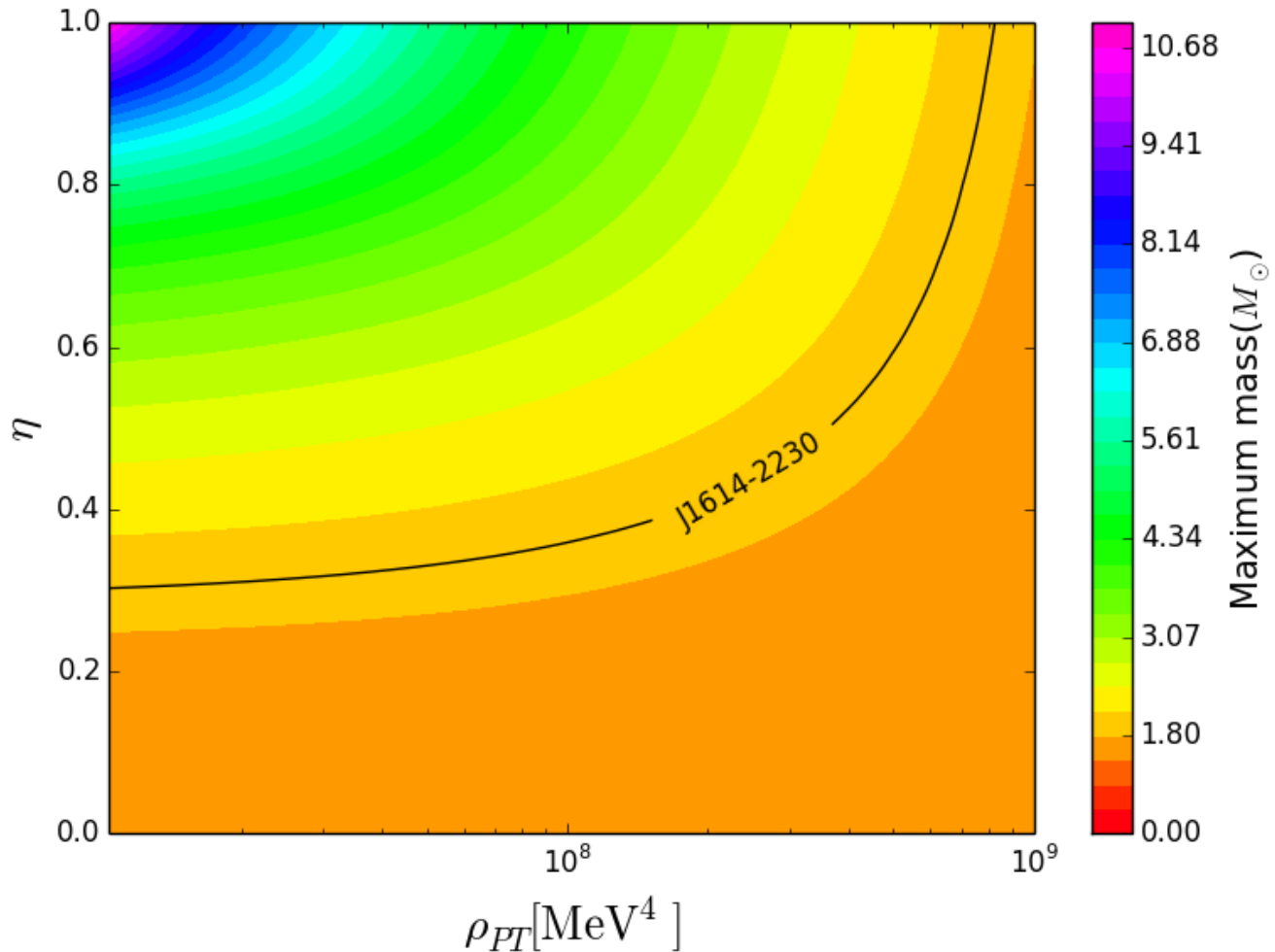
# Mass-Radius Relation

$2M_{\odot}$  is allowed!!



$$\rho_{\text{PT}} = 10^8 \text{MeV}^4, \quad \lambda_{\phi} = 10 \text{km}$$

# The maximum mass



$$\lambda_\phi / R_{\text{NS}} \ll 1$$

# Summary

- The existence of the  $2M_{\odot}$  NS is allowed in our model.

## Future work

- How to test our model by GWs.

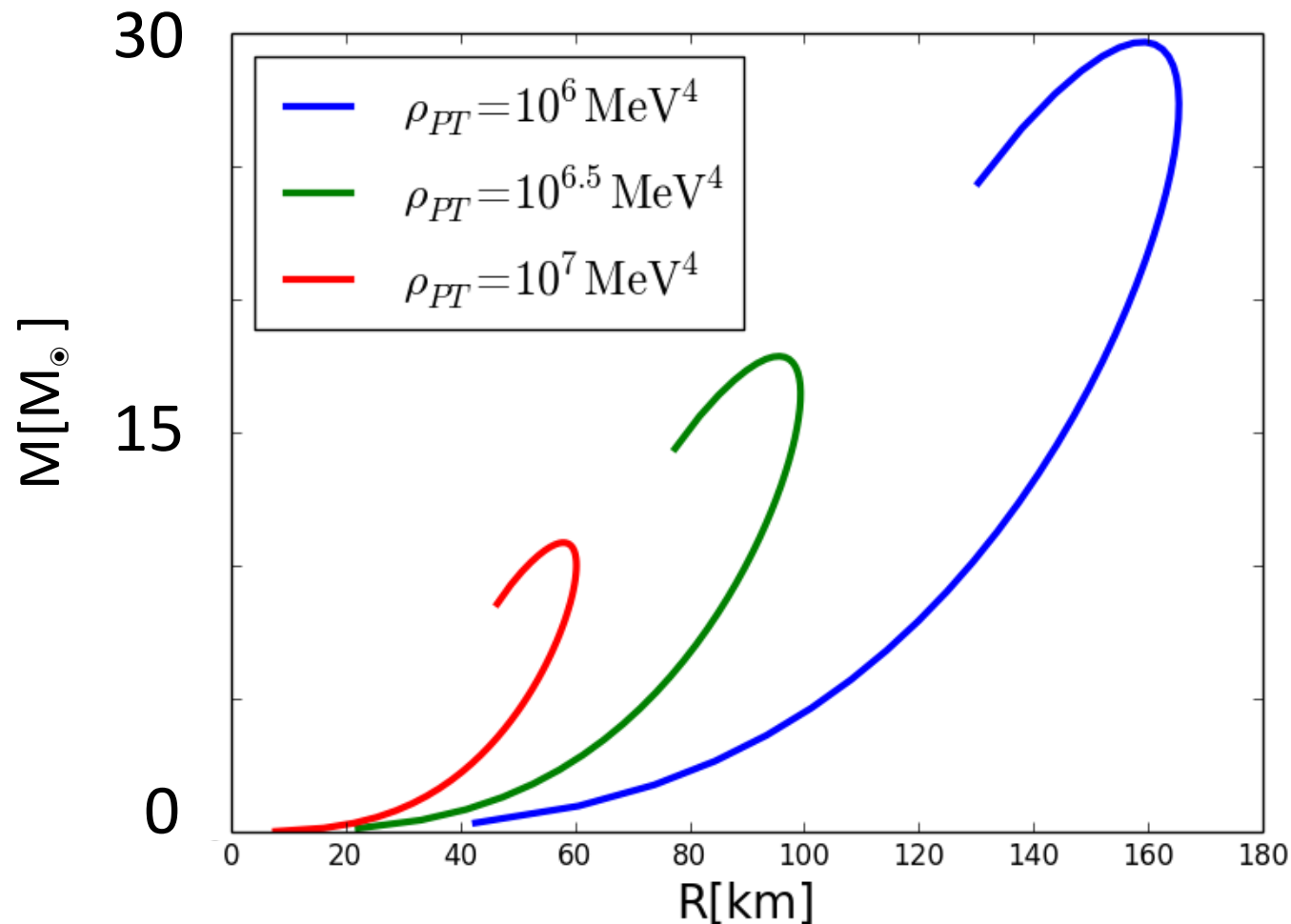
Detections of massive NSs?

Tidal deformability?



Bonus slide

# $30M_{\odot}$ neutron stars



.....but it seems to be inconsistent with the X-ray observations.

# Scalar force

The new force, scalar force, appears.

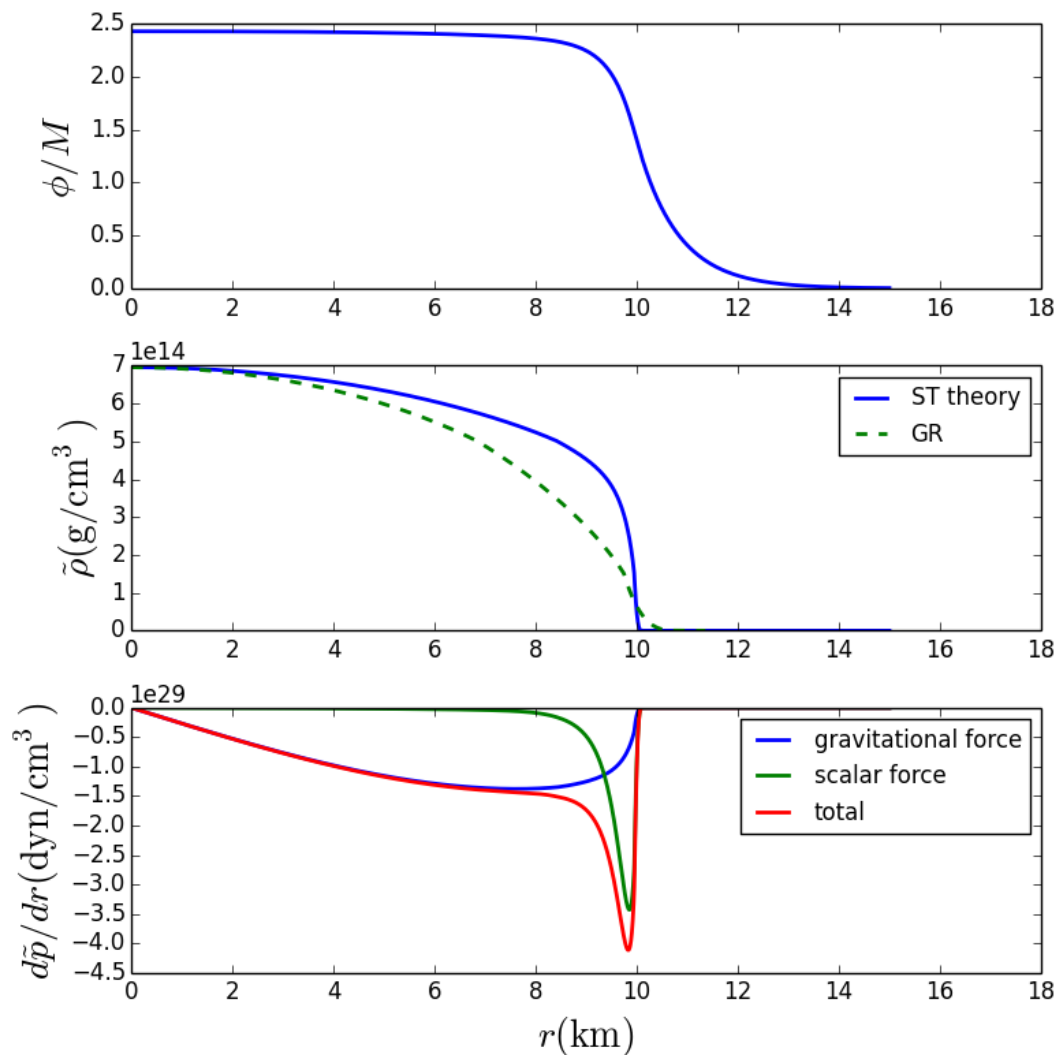
$$\tilde{p}' = -\frac{\tilde{\epsilon} + \tilde{p}}{2} \left( \underbrace{\nu'}_{\text{Gravitational force}} + \underbrace{2(\ln A(\phi))'}_{\text{Scalar force}} \right).$$

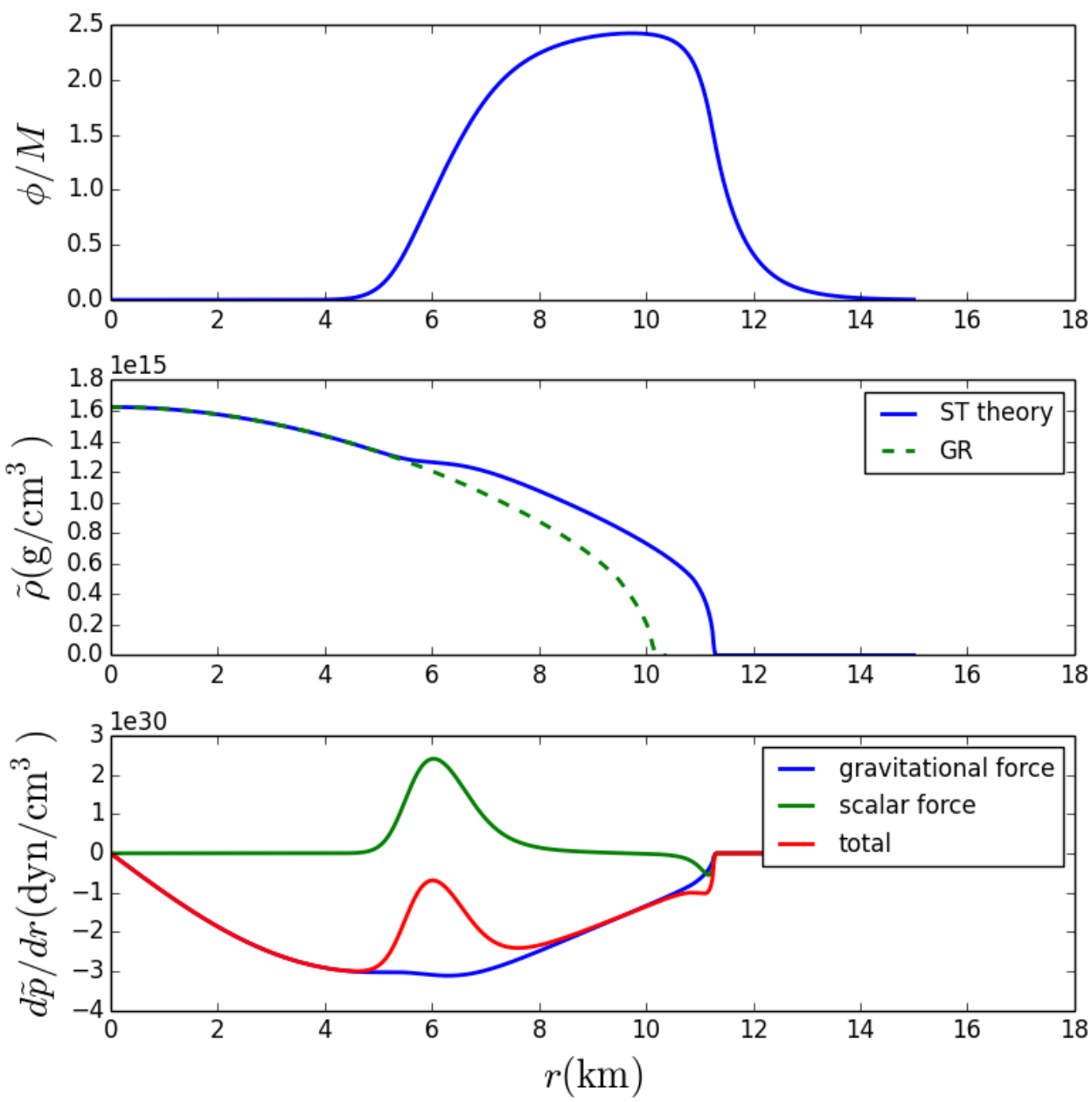
Gravitational force

Scalar force

$\phi$  decreases  $\rightarrow$  compresses the star

# Internal structure





# Allowed parameter region for DM

