

# HOW NEIL GERHEL'S CHANGED THE FIELD OF TRANSIENT ASTRONOMY

Paul O'Brien  
University of Leicester

(with thanks to the Swift team)



Neil Gehrels  
1952-2017

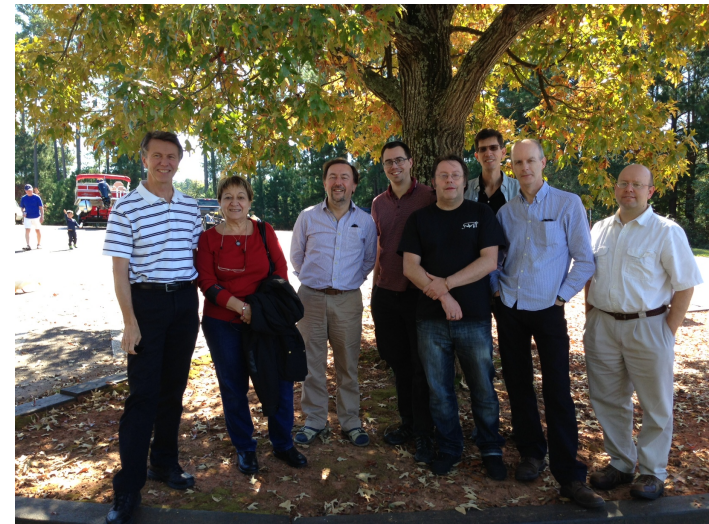
# NEIL GEHRELS



Music major (1972), climber, cyclist, volunteer and above all a great friend...



Neil worked on many projects, e.g. Voyager, balloon flights, CGRO, Integral, Swift, Fermi, WFIRST and more yet to come...





# THE BAND



# Adventures

Bicycle to Panama

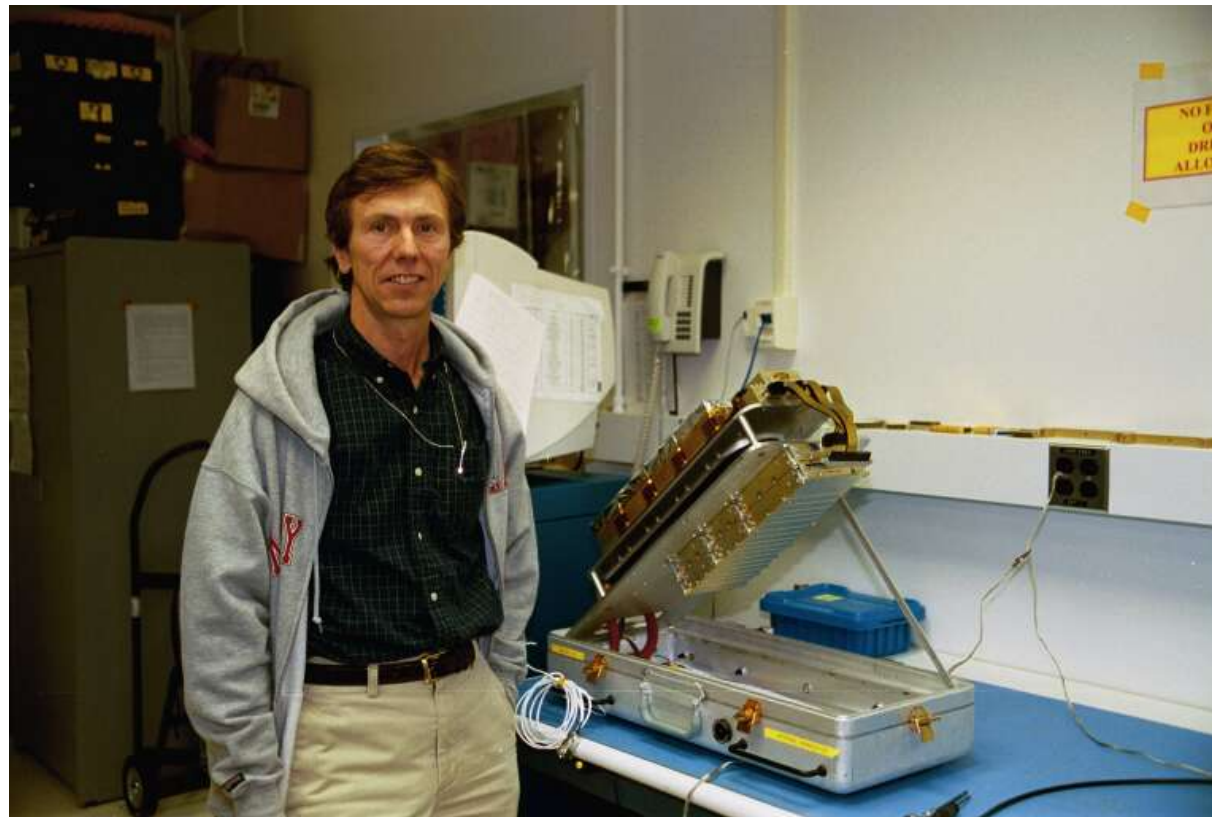


Hitchhike Round the World





So, what did we learn from Neil?



# LESSON 1: PICK YOUR MOMENT



Swift: the right technology at  
the right time, with the right PI





# LESSON 1B: DON'T GIVE UP

November 2002

NASA Cancellation Review

Swift late and >20% over budget

Neil's approach: be honest, explain and have a plan

Result: extra funding and continue to launch



# WAITING, WAITING...

## Florida 2004 - Hurricane Alley

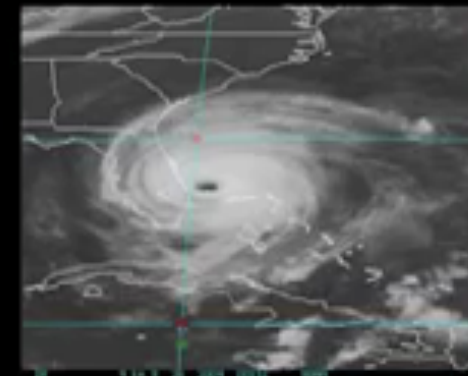
### Hurricanes:

- Charley Aug 13
- Frances Sept 4
- Ivan Sept 10
- Jeanne Sept 25

## Frances

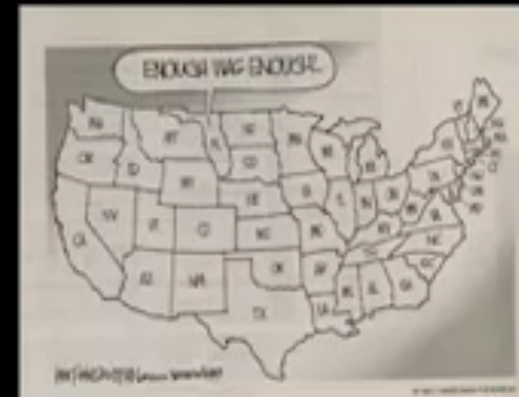


**Jeanne**



**Jeanne**

Vehicle  
Assembly  
Building  
damage





# HURRICANE FRANCES (SEPT 5)

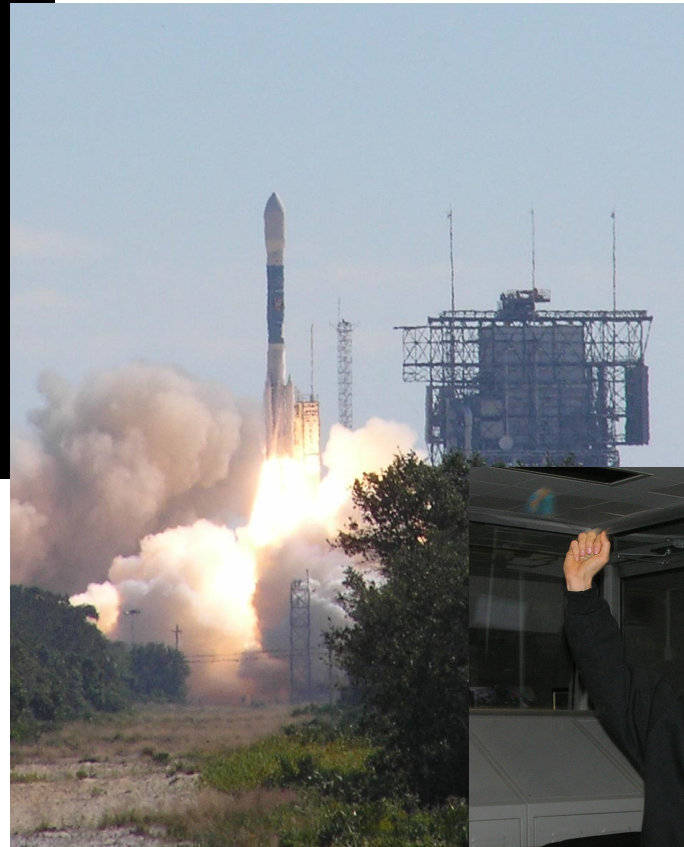


It's ok, the backup system worked

# DON'T DO THIS...



# DO THIS



20 November 2004

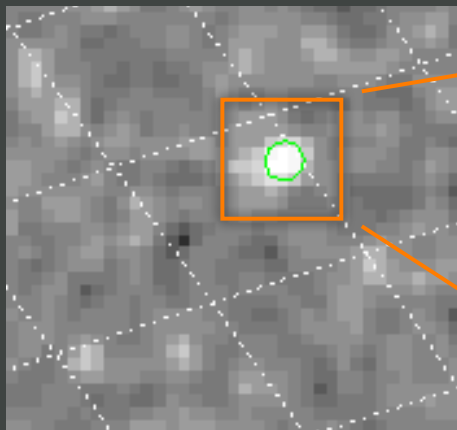




# LESSON 2: USE AUTOMATION

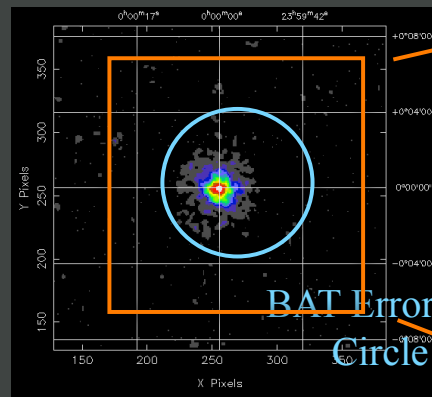
1. Burst Alert Telescope triggers on 'automated target', calculates position to  $\sim 1 - 3$  arc-minutes, transmits to ground via TDRSS and distributed by GCN.
2. Spacecraft autonomously slews to position in 1-2 minutes
3. X-ray Telescope:  $\sim 5$  arcsec prompt,  $\sim 2$  arcsec delayed position (distributed via GCN ASAP)
4. UV/Optical Telescope images field, transmits finding chart to ground
5. The Swift team analyzes the data in real time and sends out notice in  $\sim 5 - 20$  minutes.

**BAT Burst Image**



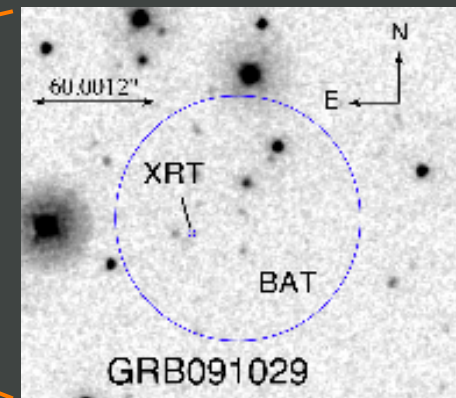
***T~30 sec***

**XRT Image**



***T~100 sec***

**UVOT Image**



***T~300 sec***

# GRB PRIORITY I: SHORT GRBS

Vol 437/6 October 2005/doi:10.1038/nature04142

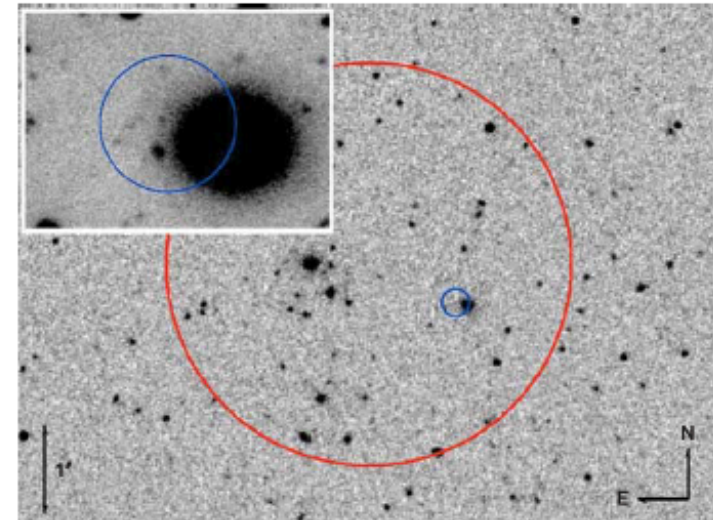
nature

## LETTERS

### A short $\gamma$ -ray burst apparently associated with an elliptical galaxy at redshift $z = 0.225$

N. Gehrels<sup>1</sup>, C. L. Sarazin<sup>2</sup>, P. T. O'Brien<sup>3</sup>, B. Zhang<sup>4</sup>, L. Barbier<sup>1</sup>, S. D. Barthelmy<sup>1</sup>, A. Blustin<sup>5</sup>, D. N. Burrows<sup>6</sup>, J. Cannizzo<sup>1,7</sup>, J. R. Cummings<sup>1,8</sup>, M. Goad<sup>3</sup>, S. T. Holland<sup>1,9</sup>, C. P. Hurkett<sup>3</sup>, J. A. Kennea<sup>6</sup>, A. Levan<sup>3</sup>, C. B. Markwardt<sup>1,10</sup>, K. O. Mason<sup>5</sup>, P. Meszaros<sup>6</sup>, M. Page<sup>5</sup>, D. M. Palmer<sup>1,1</sup>, E. Rol<sup>3</sup>, T. Sakamoto<sup>1,8</sup>, R. Willingale<sup>3</sup>, L. Angelini<sup>1,7</sup>, A. Beardmore<sup>3</sup>, P. T. Boyd<sup>1,7</sup>, A. Breeveld<sup>3</sup>, S. Campana<sup>12</sup>, M. M. Chester<sup>6</sup>, G. Chincarini<sup>1,2,13</sup>, L. R. Cominsky<sup>14</sup>, G. Cusumano<sup>15</sup>, M. de Pasquale<sup>5</sup>, E. E. Fenimore<sup>11</sup>, P. Giommi<sup>16</sup>, C. Gronwall<sup>6</sup>, D. Grupe<sup>6</sup>, J. E. Hill<sup>6</sup>, D. Hinshaw<sup>1,17</sup>, J. Hjorth<sup>18</sup>, D. Hullinger<sup>1,10</sup>, K. C. Hurley<sup>19</sup>, S. Klose<sup>20</sup>, S. Kobayashi<sup>6</sup>, C. Kouveliotou<sup>21</sup>, H. A. Krimm<sup>1,9</sup>, V. Mangano<sup>12</sup>, F. E. Marshall<sup>1</sup>, K. McGowan<sup>5</sup>, A. Moretti<sup>12</sup>, R. F. Mushotzky<sup>1</sup>, K. Nakazawa<sup>22</sup>, J. P. Norris<sup>1</sup>, J. A. Nousek<sup>6</sup>, J. P. Osborne<sup>3</sup>, K. Page<sup>3</sup>, A. M. Parsons<sup>1</sup>, S. Patel<sup>23</sup>, M. Perri<sup>16</sup>, T. Poole<sup>5</sup>, P. Romano<sup>12</sup>, P. W. A. Roming<sup>6</sup>, S. Rosen<sup>5</sup>, G. Sato<sup>22</sup>, P. Schady<sup>3</sup>, A. P. Smale<sup>24</sup>, J. Sollerman<sup>25</sup>, R. Starling<sup>26</sup>, M. Still<sup>1,9</sup>, M. Suzuki<sup>27</sup>, G. Tagliaferri<sup>12</sup>, T. Takahashi<sup>22</sup>, M. Tashiro<sup>27</sup>, J. Tueller<sup>1</sup>, A. A. Wells<sup>1</sup>, N. E. White<sup>1</sup> & R. A. M. J. Wijers<sup>26</sup>

Gamma-ray bursts (GRBs) come in two classes<sup>1</sup>: long ( $>2$  s), soft-spectrum bursts and short, hard events. Most progress has been made on understanding the long GRBs, which are typically observed at high redshift ( $z \approx 1$ ) and found in subluminal star-forming host galaxies. They are likely to be produced in core-collapse explosions of massive stars<sup>2</sup>. In contrast, no short GRB had been accurately ( $<10''$ ) and rapidly (minutes) located. Here we report the detection of the X-ray afterglow from—and the localization of—the short burst GRB 050509B. Its position on the sky is near a luminous, non-star-forming elliptical galaxy at a redshift of 0.225, which is the location one would expect<sup>3,4</sup> if the origin of this GRB is through the merger of neutron-star or black-hole binaries. The X-ray afterglow was weak and faded below the detection limit within a few hours; no optical afterglow was detected to stringent limits, explaining the past difficulty in localizing short GRBs.

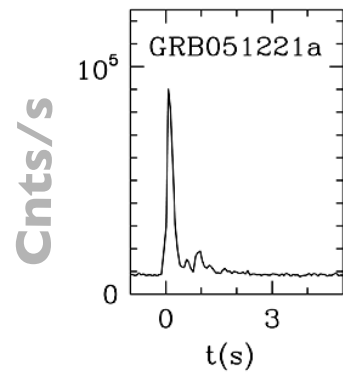


#### Clues for the merger origin of short GRBs:

- Early afterglow detections. None has a SN signature.
- In different types of host galaxies, including a few in elliptical/early-type galaxies, and most in star-forming galaxies
- Do not follow bright sights of hosts
- In regions of low star formation in star-forming galaxies
- Redshift distribution (a good fraction of low- $z$ ).
- Theoretical expectation should be “short”

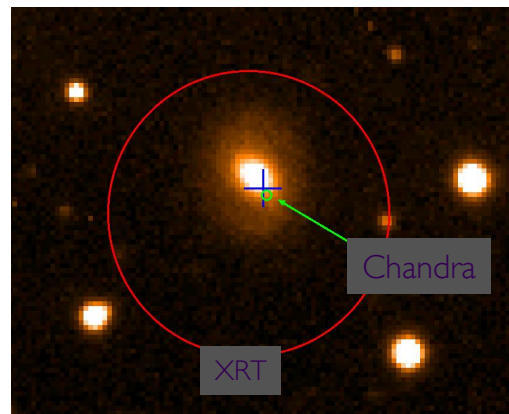


## Short GRB

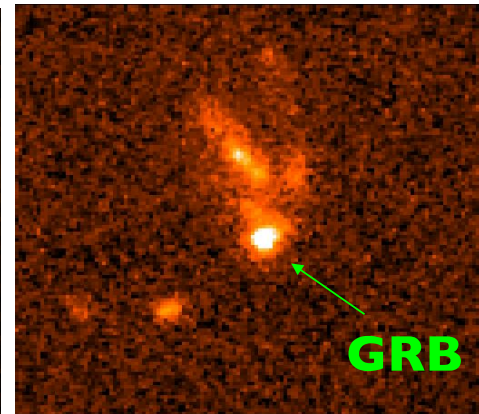


# Short vs Long GRBs

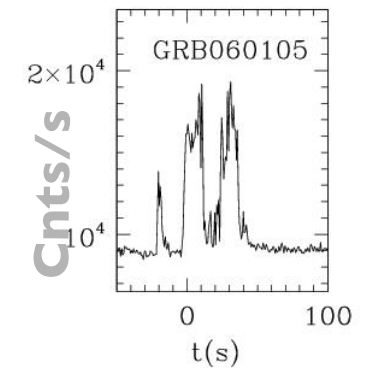
**GRB 050724 - *Swift***  
elliptical host



**GRB 990123 - *SAX***  
SF dwarf host



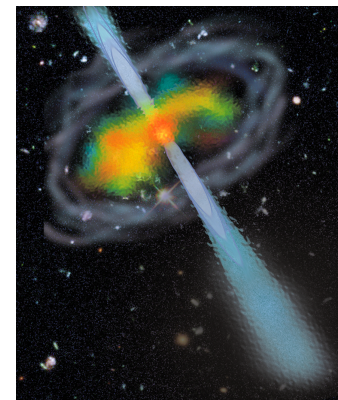
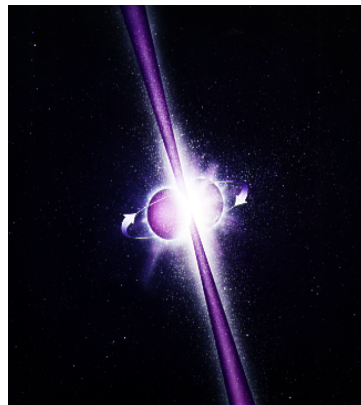
## Long GRB



In non-SF  
and SF galaxies

**No SNe  
detected**

Possible **merger**  
model – but what?



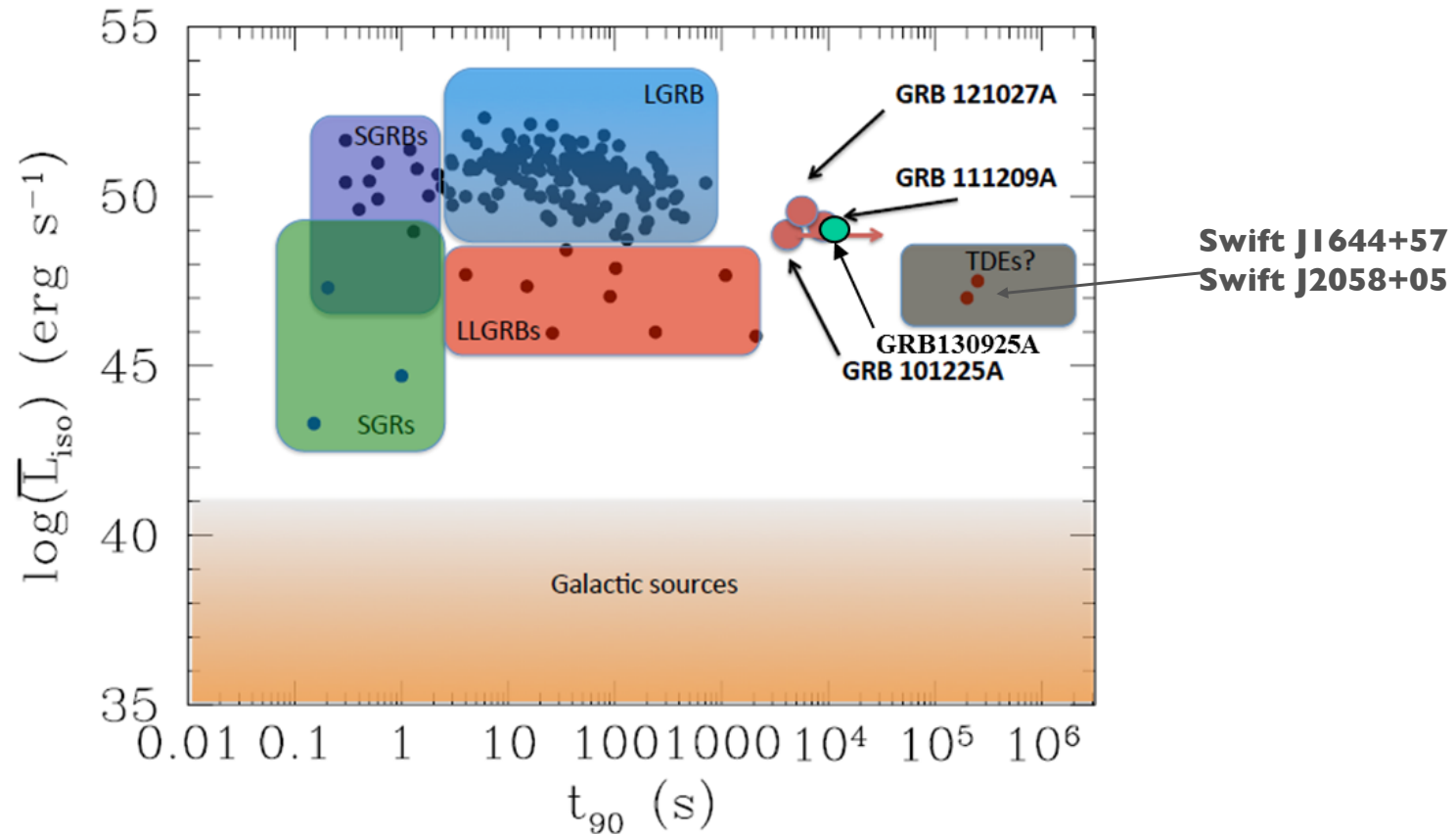
In SF  
galaxies

**Accompanied  
by SNe**

**Collapsar** model  
well supported

# THE TRANSIENT ZOO: NEW ANIMALS?

(LEVAN ET AL., 2014; EVANS ET AL. 2014)



Swift has found a small number of “ultra long” transients/GRBs, with  $T_{90} > 2000$  s

# LESSON 3: COMMUNITY INVOLVEMENT = NEW PRIORITIES

- Swift has a GI programme and a TOO web page which scientists use to submit requests for observations.
- For some programs (Neutrino, Fermi LAT GRB, LIGO/Virgo GW triggers), we have a 'backdoor' system that allows auto generated TOOs.

The screenshot shows a web browser window with the URL [swift.psu.edu](http://swift.psu.edu). The page header features the Penn State logo and the text "Mission Operations Center for Swift". A navigation bar includes links for Home, Mission, Observatory, Operations, and Additional Info. The main content area is titled "ToO Request - Source Information" and contains several form fields: "Source Name" (a text box), "Coordinates (J2000)" (with instructions to use decimal or HH MM SS.ss and links to HEASARC and UKSSDC calculators), "R.A." and "Declination" (text boxes), "Position Error (If Applicable)" (with a "90% Confidence Radius" field and a unit selector set to "arcminutes"), and "Type or Classification" (a list of radio button options: AGN, Be Binary System, Comet or Asteroid, Dwarf Nova, GRB, Nova, Pulsar, Supernova, and X-Ray Transient). A right-hand sidebar displays a welcome message for "Jamie Kennea", links for "My ToO Requests" and "Submit a ToO Request", a "Test Request" section with links for "Update Account Info", "Change Password", and "Log Out", and a "Tiled Observations" section with links for "Summary of Requests" and "Admin". At the bottom of the sidebar is a grid of logos for various institutions including Penn State, NASA, UCL, and others.

<http://www.swift.psu.edu>

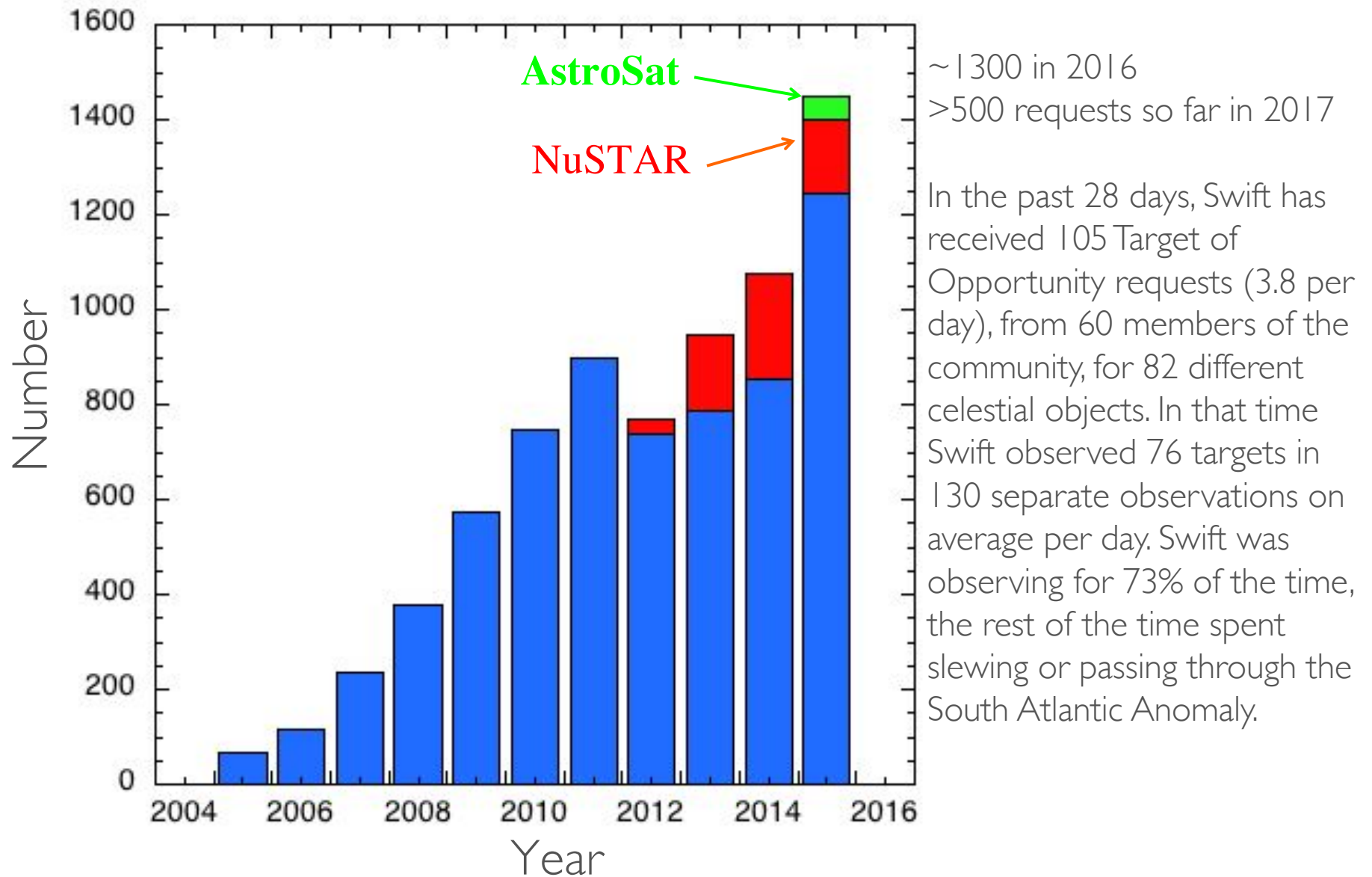


# Game-Changing Discoveries Every Year

- 2005: Short burst mystery solution. NS-NS mergers
- 2005: Flares & bright afterglows in GRBs
- 2008: Discovery of a SN shock break-out X-ray flash
- 2008: Naked-eye GRB from reverse shock jet physics
- 2009: Discovery of 2 GRBs with  $z > 8$
- 2005-2011: SFR and metallicity evolution to  $z > 5$
- 2010: Many galaxy mergers in hosts of absorbed AGN
- 2011: Tidal disruption super-flare of star eaten by BH
- 2012: Discovery of very young (2500 year old) SNR
- 2013: Anti-glitch *spin down* in magnetar 11E 2259+586
- 2013: Evidence for kilonova emission in a short GRB
- 2015: Extreme glitch in pulsar B0540-69
- 2015: V404 Cyg bright outburst for a micro-blazar
- 2015: Magnetar powered SN with ultra-long GRB
- 2015: ASASSN 14li tidal disruption event
- 2015: UV pulse in iPTF SN Ia, single degenerate explosion
- 2016: Starspots on Proxima Centauri
- 2017: NGC4151 monitoring changes disk reprocessing model

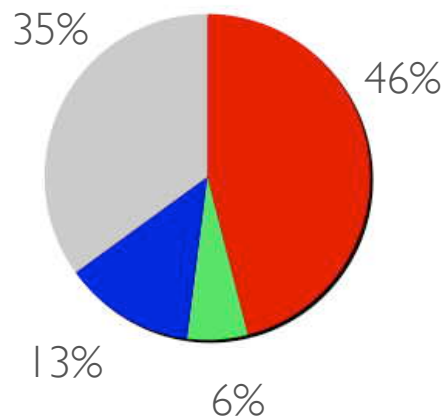
**Continuous breakthroughs with *Swift*:  
*many not GRB related***

# TOO requests increasing

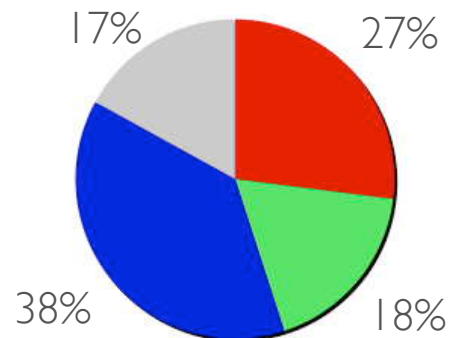


# Swift Operations & Science are Evolving

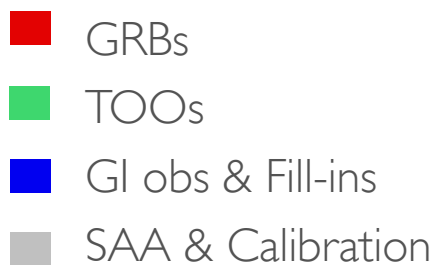
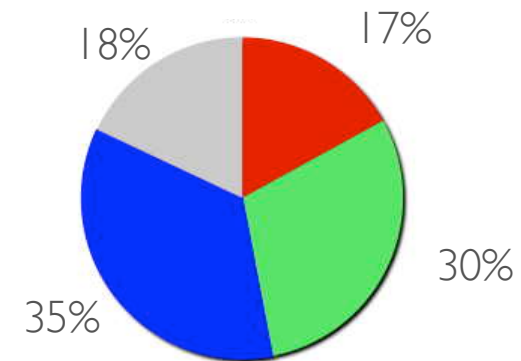
2005



2009



2014



Some TOOs in GI

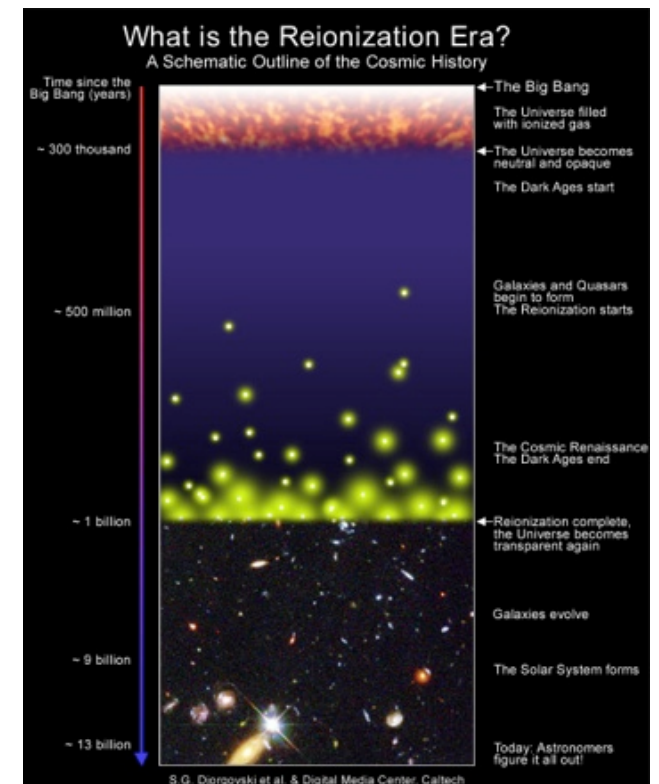
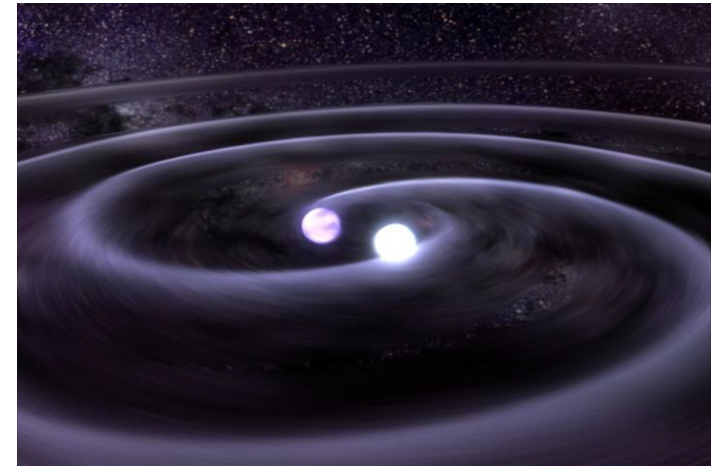
Fermi, NuSTAR, MAXI, INTEGRAL  
Spitzer, Kepler, Planck,  
Chandra, XMM, AstroSat  
VERITAS, MAGIC, HESS, HAWC,  
IceCube, ANTARES  
ALIGO-Virgo  
70 ground telescopes



# Current Prioritized Mission Objectives

- **Short burst progenitors**
  - Counterpart of ALIGO-Virgo detection
  - r-process kilonova emission
- **Epoch of reionization**
  - High signal-to-noise spectra of  $z > 7$  event
- **Supernovae & Cosmology**
  - UV light curves of SNe Ia in Hubble flow
- **Time-domain discovery**
  - New mission initiatives
  - Wide-field optical and radio surveys

NASA SR proposal 2014: Swift came top, again

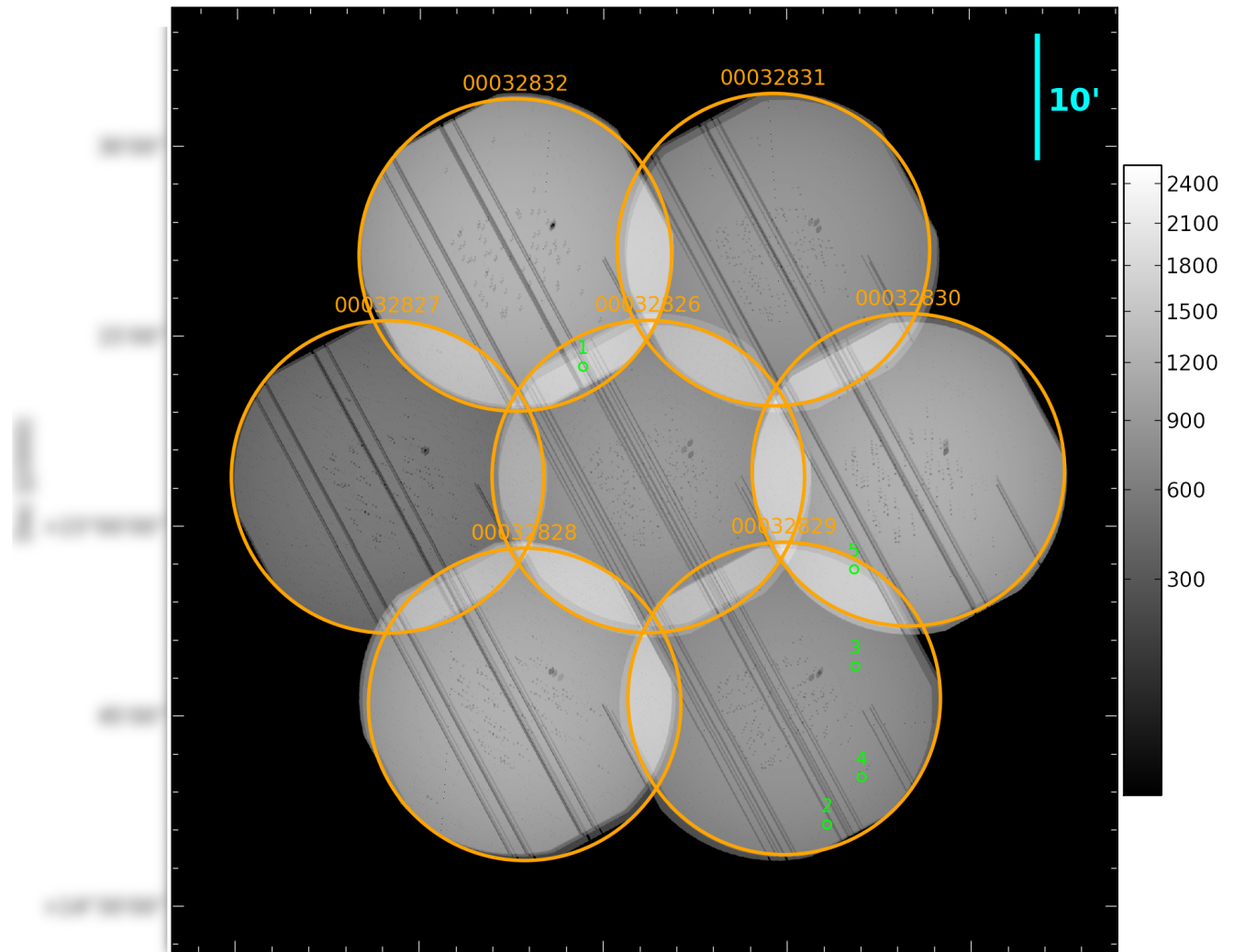


# TILING: EXAMPLE FOR ICECUBE

Swift response to a neutrino doublet from IceCube.

7-point tiling needed to cover the error region.

Multiple sources detected, but...



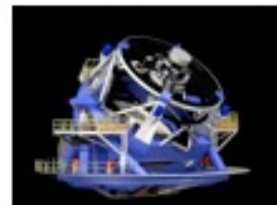
Coordinates redacted...

# OBSERVING LIGO/VIRGO TRIGGERS

- Trigger on compact binary coalescent events.
- Take BAYESTAR error region, convolve with Galaxy Catalog to target nearby galaxies.
- Upload PPST containing as many 60s exposures of the LIGO error regions for 48 hours to look for prompt emission.
- For following three days take as many 500s exposures of these galaxies again, in order to look for off-axis afterglows.
- Process validated and in use



# LESSON 4: THE NEXT MISSION



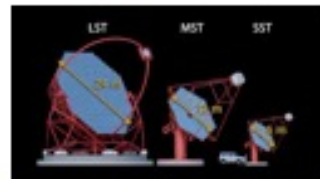
LSST



IceCube/ANTARES



LIGO-Virgo



CTA

## The Future



SVOM



GMT, TMT, E-ELT



JWST



SKA



ComPair  
AstroGam



Gamma  
400



Burst  
Cube



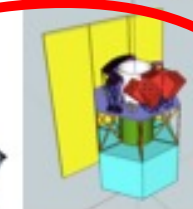
iWF-MAX



Einstein Probe



TAP



THESEUS

Want:  
Wide sky coverage,  
good localisation,  
on-board redshift

**‘Lobster’ X-ray  
technology**

Neil’s vision for the future (Huntsville, Nov. 2016)

# WE MISS YOU NEIL



Neil Gehrels, 1952-2017