



AGATA at GSI

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- Introduction
- Standard setup
- Specifities of the setup
- Selected experiments (techniques)
- Conclusion





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2012: 4 experiments



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2012: few nuclei investigated





2014: completed 2 experiments, 2 experiments performed



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Isomer data: Pb region, and one dedicated beam time



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Primary beam from SIS-18, 600 MeV/A to 1 GeV/A



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Production of relativistic beam



Relativistic fission or fragmentation



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Reaction products



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Selection and identification of the fragments: $B\rho - \Delta E - B\rho$ method



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Picture of the experimental area



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Exotic beam from the FRagment Separator (FRS)



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PreSPEC-AGATA setup



Fragmentation or relativistic coulomb-excitation



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γ rays detected with the AGATA and HECTOR+ detectors



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Reaction products detected in LYCCA



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Reaction products detected in LYCCA



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Beam size challenge



C. Domingo-Pardo, NIMA 694 (2012)

Issue with the beam spot size: hitting the detectors



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specificies of the setup expected efficiency versus reality



C. Domingo-Pardo, NIMA 694 (2012)

Expecting at least 25 crystals



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Up to 21 AGATA crystals (2014)



コントロント・エレート

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N. Lalović, NIMA 806, 258-266 (2016)

For 21 AGATA detectors at 1172 keV:

2.38% core common, 2.55% tracked, 3.3% calorimetric







Opening angle: Euroball cluster detector







Opening angle: AGATA segmented detector





Specificities of the setup Data-flow handling



NARVAL actors



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Specificities of the setup Data-flow handling



MBS DAQ system (FRS+LYCCA)



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Coupling of both data-flow



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Data-flow handling



Duplication of data: integrity checks



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Specificities of the setup Data-flow handling



Monitoring



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Selected experiments (techniques) In flight spectroscopy



IN FLIGHT DATA

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Courtesy: O. Wieland, R. Avigo

Wieland: Pygmy Dipole Resonance in $^{64}{\rm Fe}$ and the properties of neutron skin

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Study of nuclei toward the r-process



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Courtesy: O. Wieland, R. Avigo

Study of the dipole strength



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Selected experiments (techniques) 62-64





Access the dipole strength via relativistic coulomb excitation ($\beta \sim 0.7$)



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Observation of the 2^+ of 64 Fe: Essential for normalisation



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High energy structure observed with AGATA



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Courtesy: A. Boso, S. Milne, M. Bentley

RECCHIA/BENTLEY: TRANSITION RATES AND MIRROR ENERGY DIFFERENCES IN ISOBARIC MULTIPLETS

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Isospin triplet $A = 46^{Courtesy: A. Boso, S. Milne, M. Bentley}$



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First 2⁺ energy are similar in the $A = \overset{Courtesy: A. Boso, S. Milne, M. Bentley}{46 isospin triplet}$



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 $^{46}Cr^{46}V^{-46}Ti$



 $B(E2; 2^+
ightarrow 0^+)$ measurement with same conditions



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 $^{46}Cr^{46}V^{-46}Ti$



 $B(E2; 2^+
ightarrow 0^+)$ measurement with same conditions









Fragmentation of ${}^{58}Ni$ at 600 MeV/A



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 $\gamma\text{-ray}$ spectra obtained after safe coulomb excitation criteria



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 $^{46}Cr^{-46}V^{-46}Ti$



 $\gamma\text{-ray}$ spectra obtained after safe coulomb excitation criteria

With the triple target stack nice separation between the components



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Courtesy: A. Boso, S. Milne, M. Bentley

Estimation with a "plunger" type analysis



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Pietri: Shape evolution in Neutron-Rich $${\rm Zr}$$

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Hartree-Fock-Bogoliubov shape predictions

J.P. Delaroche, Phys. Rev. C 81:014303 (2010)



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J.P. Delaroche, Phys. Rev. C 81:014303 (2010)



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Image: A mathematical states and a mathem

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AGATE





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 $^{\text{Selected experiments (techniques)}}_{100-108} \text{Mo}$



D. Ralet, submitted to PRC Observed transition

in $^{100-108}Mo$



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D. Ralet, submitted to PRC Observed transition

in ^{100–108}Mo



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D. Ralet, submitted to PRC





100 - 108 Mo



D. Ralet, submitted to PRC

Systematics $B(E2; 2^+ \rightarrow 0^+)$





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 $^{\text{Selected experiments (techniques)}}_{100-108} \text{Mo}$



D. Ralet, submitted to PRC

Systematics $B(E2; 4^+ \rightarrow 2^+)$





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 $\frac{100-108}{100}$



D. Ralet, submitted to PRC



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Combining AGATA with high velocity fragments open great measurement opportunities

New tools and techniques have been tested and provided physics results

4 years after the first part of the AGATA@GSI campaign the first paper are being published





Thank to all the people that send me the material needed for this presentation:

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