AUTOMATIC LOOP CALCULATIONS OF NMSSM PROCESSES WITH SLOOPS

Guillaume CHALONS

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OVERVIEW OF THE MODEL AND MOTIVATIONS

NMSSM = Next-to Minimal Supersymmetric Standard Model :

$$W_{NMSSM}^{\mathbb{Z}_3} = W_{MSSM}^{\mu=0} + \lambda \hat{S} \hat{H}_u \hat{H}_d + \frac{\kappa}{3} \hat{S}^3$$

Important phenomenological consequences :

- Same Pros as MSSM: Hierarchy, Gauge Unification, Dark Matter...
- Solves elegantly the " μ problem" of the MSSM : $\mu_{eff} = \lambda \langle S \rangle$.
- Richer Higgs spectrum than the MSSM : 2 (1) CP-odd Higgs A_i , 3 (2) CP-even H_i^0 , 1 charged H^{\pm} .
- Enlarged neutralino sector $(\widetilde{B}^0, \widetilde{W}^0, \widetilde{H}^0_1, \widetilde{H}^0_2, \widetilde{S}^0) \rightarrow (\tilde{\chi}^0_1, \tilde{\chi}^0_2, \tilde{\chi}^0_3, \tilde{\chi}^0_4, \tilde{\chi}^0_5)$



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- ► Enlarged neutralino sector $(\widetilde{B}^0, \widetilde{W}^0, \widetilde{H}_1^0, \widetilde{H}_2^0, \widetilde{S}^0) \rightarrow (\widetilde{\chi}_1^0, \widetilde{\chi}_2^0, \widetilde{\chi}_3^0, \widetilde{\chi}_4^0, \widetilde{\chi}_5^0)$
- ${}^{\rm \mbox{\scriptsize sol}}$ Additional F-term + doublet-singlet mixing makes it easier to get $m_h^{\rm SM} \sim 125$ GeV
- Solution Evade LHC SUSY constraints by hiding $\not\!\!\!E_T$ (visible energy diluted)
- Solution Easier to get low mass DM (thanks to approx. U(1) R/PQ sym.)
- Successful EW baryogenesis (allows 1st order PT)
- rs . . .

- Like in the MSSM, the SM-like Higgs mass gets large rad. cor.
- Rad. cor. in Higgs production/decays can also be large
- To match planned exp. accuracy in Higgs measurements, inclusion of loop corrections mandatory



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AUTOMATIC TOOLS AVAILABLE ON THE MARKET

- <u>MMSSMTools</u>: Spectrum gen. + Decays + DM + Constraints (Das,Domingo,Ellwanger,Gunion,Hugonie,Jean-Louis,Teixeira)
- <u>SPheno and SARAH</u>: Spectrum generator + Decays + Constraints (Porod, Staub, Goodsell,Nickel)
- <u>NMSSMCALC</u>: Higgs masses + decays (Baglio, Gröber, Mühlleitner, Nhung, Rzehak, Spira, Streicher, Walz)
- Flexible SUSY: Spectrum Generator (Athron, Park, Stöckinger, Voigt)
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SLOOPS: Automatic generation of any process (mass, decay, cross-section) at 1L with EW & QCD corrections.

Orsay

AUTOMATIC TOOL FOR ONE-LOOP CALCULATIONS: SLOOPS



- Automatic derivation of the CT Feynman rules and computation of the CT's
- Models renormalized: SM, MSSM, NMSSM, Wino DM, xSM (w/ & w/o vs),
- Modularity between different renormalisation schemes.
- Non-linear gauge fixing.
- Checks: results UV,IR finite and gauge independent.

http://code.sloops.free.fr/

LPT Orsav

APPLICATION TO DM: GAMMA-RAY LINES

G.C, A. Semenov JHEP 1112 (2011) 055; G.C, M.J. Dolan, C. McCabe JCAP 1302 016

- $\chi\chi
 ightarrow \gamma\gamma$ is a smoking gun signature of DM and $E_\gamma \sim m_\chi$
- Interesting process to look at to test the good implementation of the framework at LO (no renormalisation needed).
- Main mechanism in the NMSSM



- $\blacktriangleright \langle \sigma v \rangle_{\gamma\gamma/Z^0}$ computed with SloopS extended to deal with the NMSSM & all contributions
- ► GI checked thanks to an extended NLG GF for the NMSSM GC,Semenov '11
- Modified version of LOOPTOOLS to handle vanishing Gram determinants at v = 0 Boudjema,Semenov,Temes '05



APPLYING FERMI-LAT CONSTRAINTS

- \blacktriangleright FERMI satellite has a dedicated search for $\gamma\text{-lines}$
- Concentrate on low-mass m_{χ} which is specific to NMSSM
- Extension of the work performed in (Vazquez et. al '10)
- ▶ Limits taken from *Vertongen*, *Weniger* '11 : extended range $m_{\chi} \in [1, 300]$ GeV.





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Routine of $\langle \sigma v \rangle_{\gamma\gamma}$ through SLOOPS built-in MicrOMEGAs (MSSM/NMSSM)



APPLICATION TO HIGGS: LOOP-INDUCED DECAYS

► Use of an effective $\mathcal{V}_{rad.}^{S}$ (GC, Domingo '12) to compute loop-induced $h_{i}^{0} \rightarrow \gamma \gamma / Z^{0}$ decays at the right kin. in a GI invariant way through the interface of SLOOPS with NMSSMTools (Effective potential now present in standard version of NMSSMTools)

Routines incorporated within MicrOMEGAs (private version)



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- ${}^{\scriptstyle
 m I\!S}$ No significant deviation from SM observed in $\mu_{\gamma\gamma}$
- $^{\rm IS}$ In the SM $\Gamma^H_{\gamma\gamma}/\Gamma^H_{\gamma Z}\sim 2/3$, is there room to observe a deviation from the SM of this ratio in the NMSSM ?



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$100 \ {\rm GeV} <$	μ	< 500 ~GeV
$100~{\rm GeV} <$	M_2	$< 1000 \ GeV$
$0 {\rm GeV} <$	t_{eta}	< 20
0 <	λ,κ	< 0.7
$100 \ {\rm GeV} <$	A_{λ}	$< 1000 \ {\rm GeV}$
-1000 GeV <	A_{κ}	$< -100 { m ~GeV}$
-3000 GeV <	A_t	$< 3000 {\rm GeV}$
$400~{\rm GeV} <$	$m_{\tilde{Q},\tilde{U}_2}$	$< 2000 {\rm GeV}$

- ▶ $m_h^{
 m SM} \in [122, 128]$ GeV
- Various constraints from NMSSMTools
- Collider constraints on Higgs from HiggsBounds
- Higgs fits constraints from HiggsSignals



POSSIBLE DEVIATIONS IN $h_{1,2}^0 \rightarrow \gamma Z^0$?

If $h_1^0 \equiv H_{125} \rightarrow h_1^0 \rightarrow \gamma Z^0$ similar to SM expectations, look at $h_2^0 \equiv H_{125}$



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- ▶ Some enhancement for $a)S_{h23} < -0.5, b)S_{h23} > 0.7, c)S_{h23} \approx 0.4$
- h_2^0 significant doublet & singlet comp.
- μ_{VBF} mostly correlated with μ_{gg} .
- Large deviation could indicate a need for a lighter Higgs boson





SECTORS

- $\ensuremath{\,\cong}$ Fermion \rightarrow as in the SM
- $\ensuremath{\,^{\tiny \hbox{\tiny IM}}}$ Gauge \rightarrow as in the SM
- \blacksquare Sfermion \rightarrow as in the MSSM







G. Bélanger, V. Bizouard, F. Boudjema, GC, PRD93 (2016) 11, 115031 and ArXiv:1612.XXXX

All in all we have

$$\underbrace{g,g'}_{\text{SM}},\underbrace{v_u,v_d,s,\lambda,\kappa}_{\text{Higgs}\&\tilde{\chi}},A_{\lambda},A_{\kappa},m_{H_u}^2,m_{H_d}^2,m_{S}^2,M_1,M_2$$





All in all we have

$$\underbrace{\mathbf{g}, \mathbf{g}', \mathbf{v}}_{\text{gauge}}, \underbrace{\mathbf{t}_{\beta}, \lambda, \kappa, \mu}_{\mathbf{t}, \mathbf{A}_{\lambda}, \mathbf{A}_{\kappa}, m_{H_{u}}^{2}, m_{H_{d}}^{2}, m_{S}^{2}, M_{1}, M_{2}$$



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All in all we have

$$\underbrace{g,g',v}_{\text{gauge}}, \underbrace{t_{\beta}, \lambda, \kappa, \mu}_{H_{\beta}, \lambda, \kappa, \mu}, A_{\lambda}, A_{\kappa}, m_{H_{u}}^{2}, m_{H_{d}}^{2}, m_{5}^{2}, M_{1}, M_{2}$$

we trade some for physical parameters



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$$e, M_Z, s_W, \overbrace{t_\beta, \lambda, \kappa, \mu}^{\mathsf{Higgs} \& \tilde{\chi}} \underbrace{A_{\lambda}, A_{\kappa}, t_{h^0_u}, t_{h^0_d}, t_{h^0_s}}_{\mathsf{Higgs}}, \overbrace{M_1, M_2}^{\tilde{\chi}}$$

$$\ \text{Min. cond.} \implies \boxed{t_{h_i^0} \equiv 0}, \ i = u, d, s. \ \text{At 1L} \ \delta t_{h_i^0} = -t_{h_i^0}^{loop}$$

- Remains 8 counterterms to be determined .
- Equivalently, we need to find 8 exp inputs/definitions which are linked unambigously to the original 8 parameters
- \mathbb{I} \mathcal{W}_8 matrix system to invert with $\delta t_{\beta}, \delta \lambda, \delta \kappa, \delta \mu, \delta A_{\lambda}, \delta A_{\kappa}, \delta M_1, \delta M_2$ as variables

$$\overset{\mathcal{M}_{2},\mu}{\overset{\mathcal{W}_{2}}{\overset{\mathcal{W}_{2}}{\overset{\mathcal{U}}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}}{\overset{\mathcal{U}}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}{\overset{\mathcal{U}}}{\overset{\mathcal{U$$



$$\mathcal{W}_{8} = \underbrace{\mathcal{W}_{2}^{\chi^{\pm}}}_{\mathcal{W}_{2}^{\chi^{\pm}}} \oplus \underbrace{\mathcal{W}_{4}^{\chi^{0}}}_{\mathcal{W}_{4}^{\chi^{0}}} \oplus \underbrace{\mathcal{W}_{2}^{A^{0}}}_{\mathcal{W}_{2}^{\Lambda^{0}}} \rightarrow OS_{ijkl} \text{ (suited when only gaugino decays)}$$

$$\mathcal{W}_{8} = \underbrace{\mathcal{W}_{2}^{\chi^{\pm}}}_{\mathcal{W}_{2}^{\chi^{\pm}}} \oplus \underbrace{\mathcal{W}_{4}^{\chi^{0},A^{0},H^{\pm}(h^{0})}}_{\mathcal{H}_{3+3}} \rightarrow OS_{ijkA_{1}A_{2}H^{\pm}(h_{\alpha})}$$



$$\begin{array}{c} & & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ &$$



$$\begin{array}{c} & & & & \\ & & & \\ & & & \\ &$$



ON-SHELL SCHEMES WITH MASSES ONLY

$$\begin{array}{c} & & & & \\ & & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & & \\ & & & \\ & & & \\ & & & & \\ & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & &$$

MIXED $\overline{\mathrm{DR}}$ -OS SCHEMES

$$\mathbb{I} \mathcal{W}_{8} = \mathcal{W}_{1, t_{\beta}} \oplus \widetilde{\mathcal{W}_{2}^{\chi^{\pm}}} \oplus \widetilde{\mathcal{W}_{3}^{\chi^{0}}} \oplus \widetilde{\mathcal{W}_{3}^{\chi^{0}}} \to t_{ijk} \text{ (suited when only gaugino decays)}$$

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$$\begin{array}{c} & & & & \\ & & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & & \\ & & & \\ & & & \\ & & & & \\ & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & &$$

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$$\mathfrak{W}_{8} = \mathcal{W}_{1,t_{\beta}} \oplus \widetilde{\mathcal{W}_{2}^{\chi^{\pm}}} \oplus \widetilde{\mathcal{W}_{2+3}^{\chi^{0},A^{0},H^{\pm}(h^{0})}$$

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Or simply go all $\overline{\mathrm{DR}}$

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APPLICATION TO HIGGS DECAYS

Point A($Q_{susy} = 1117.25$ GeV, $m_t = 173$ GeV, $m_{h_1^0} = 125.45$ GeV(1-loop OS))

<i>M</i> ₁	700	λ	0.1	A_{κ}	0	m _{õ3}	1740	$m_{\tilde{D},\tilde{U}_{1,2}}$	1000
<i>M</i> ₂	1000	κ	0.1	A_t	4000	$m_{\tilde{U}_3}$	800	$m_{\tilde{L}_3}$	1000
M ₃	1000	μ	120	Ab	1000	$m_{\tilde{D}_3}$	1000	m _{ĩ3}	1000
tβ	10	A_{λ}	150	A,	1000	m _{Õ1,2}	1000	$m_{\tilde{L},\tilde{l}_{1,2}}$	1000
$\lambda A_{\lambda} =$	15GeV,	A_t/A_λ	~ 27					,	

Point B($Q_{susy} = 753.55$ GeV, $m_t = 146.94$ GeV, $m_{h_s^0} = 124.44$ GeV(1-loop OS))

<i>M</i> ₁	120	λ	0.67	A_{κ}	0	m _{Õ3}	750	$m_{\tilde{D},\tilde{U}_{1,2}}$	1500	
<i>M</i> ₂	300	κ	0.2	A t	1000	$m_{\tilde{U}_3}$	750	$m_{\tilde{L}_3}$	1500	
M ₃	1500	μ	200	Ab	1000	$m_{\tilde{D}_3}$	1500	$m_{\tilde{l}_3}$	1500	
tβ	1.92	A_{λ}	405	A	1000	m _{Õ1,2}	1500	$m_{\tilde{L},\tilde{l}_{1,2}}$	1500	4
$\lambda A_{\lambda} =$	271GeV	A_t/A_t	$_{\lambda}\sim 2.5$	1		· · · · ·				

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APPLICATION TO HIGGS DECAYS

		Point A	Point B				
h ⁰ 1	h ⁰ d	1.1%	22.5%				
	hŪ	98.6%	67.4%				
	h _s 0	0.3%	10.1%				
h ₂ 0	h ⁰ d	0.1%	0.%				
	h ⁰ U	0.3%	12.5%				
	h_s^0	99.6%	87.5%				
h ⁰ 3	h ⁰ d	98.8%	77.5%				
	h ⁰ u	1.1%	19.7%				
	h _s 0	0.1%	2.8%				
A ⁰ 1	a_d^0	0%	1.8%				
	a ⁰ _U	0%	0.5%				
	a_s^0	100%	97.7%				
A_2^0	a_d^0	99.0%	76.9%				
	a ⁰ _U	1.0%	20.8%				
a _s ⁰ 0.0% 2.3%							
Point A: h_u , h_s , h_d , a_s , a_d							
Point B: h _U , h _S , h _d , a _S , a _d							

		Point A	Point B			
$\tilde{\chi}_1^0$	\tilde{B}^0	-	56.6%			
•	₩ ⁰	-	32.3%			
	\tilde{h}^{0}	98.4%	10.3%			
	\tilde{s}^0	0.77%	0.8%			
$\tilde{\chi}_{2}^{0}$	\tilde{B}^0	-	4.0%			
-	Ŵ ⁰	-	2.6%			
	\tilde{h}^{0}	99.5%	19.3%			
	\tilde{s}^0	-	74.0%			
$\tilde{\chi}_{3}^{0}$	₿ ⁰	-	10.1%			
U	Ŵ ⁰	-	-			
	\tilde{h}^{0}	0.9%	78.9%			
	\tilde{s}^0	99.1%	11.0%			
$\tilde{\chi}_{4}^{0}$	₿ ⁰	99.6%	18.1%			
-	Ŵ ⁰	-	12.3%			
	\tilde{h}^{0}	-	55.8%			
	\tilde{s}^0	-	13.7%			
$\tilde{\chi}_{5}^{0}$	₿ ⁰	-	11.2%			
5	₩ ⁰	99.3%	52.8%			
	\tilde{h}^{0}	0.69%	35.7%			
	\tilde{s}^0	-	0.4%			
Point A: <i>ĥ</i> , <i>ĥ</i> , <i>ŝ</i> , <i>ĥ</i> , <i>ŵ</i>						
Point B: ĎěĎů						



Beware. B much more mixing, A quite pure

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AUTOMATIC LOOP CALCULATIONS OF NMSSM PROCESSES WITH SLOOPS

▶ singlets:
$$h_2^0, A_1^0, \tilde{\chi}_3^0, m_{h_2^0} = 240 \text{ GeV}, m_{h_3^0, A_2^0, H^{\pm}} \sim 570 \text{ GeV},$$

 $Q_{\text{SUSY}} = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}} = 1117 \text{ GeV}$

Decays	$t_{134A_1A_2}(Q_{\rm SUSY})$	$OS_{34h_2A_1A_2H^+}$	$\overline{\mathrm{DR}}(m_{parent})$	$\overline{\mathrm{DR}}(Q_{\mathrm{SUSY}})$



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$h_2^0 ightarrow A_1^0 A_1^0$	(128%)	(-12%)	(0.4%)	(-0.4%)

■ At LO (with
$$(A_{\kappa} = 0)$$
, $g_{h_2^0 A_1^0 A_1^0}$ stems from $\kappa^2 S^4$
 $\rightarrow g_{h_2^0 A_1^0 A_1^0} \propto \kappa^2 s \propto (\kappa s)^2 / s \propto \lambda / \mu(\kappa s)^2 \sim \lambda / m_{\widetilde{H}^{\pm}} \times m_{\widetilde{S}^0}^2$.



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 $\rightarrow g_{h_2^0 A_1^0 A_1^0} \propto \kappa^2 s \propto (\kappa s)^2 / s \propto \lambda / \mu(\kappa s)^2 \sim \lambda / m_{\widetilde{H}^{\pm}} \times m_{\widetilde{S}^0}^2$.

 $m_{\tilde{5}0}$ constrains $(\kappa s)^2$ and $m_{\tilde{H}^{\pm}}$ constrains μ well. Finite shift on λ is key. We have

$$\delta\lambda/\lambda|_{\rm fin.}^{\rm t_{134}}=$$
 62.26% and $\delta\lambda/\lambda|_{\rm fin.}^{\rm OS}=-7.88\%$

and loop correction is $\delta\Gamma/\Gamma \sim 2\delta\lambda/\lambda$ due to finite part of CT.



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 $Q_{\text{SUSY}} = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}} = 1117 \text{ GeV}$

Decays	$t_{134A_1A_2}(Q_{\mathrm{SUSY}})$	$OS_{34h_2A_1A_2H^+}$	$\overline{\mathrm{DR}}(m_{\scriptscriptstyle parent})$	$\overline{\mathrm{DR}}(Q_{\mathrm{SUSY}})$
$h_2^0 ightarrow A_1^0 A_1^0$	(128%)	(-12%)	(0.4%)	(-0.4%)

■ At LO (with (
$$A_{\kappa} = 0$$
), $g_{h_2^0 A_1^0 A_1^0}$ stems from $\kappa^2 S^4$
 $\rightarrow g_{h_2^0 A_1^0 A_1^0} \propto \kappa^2 s \propto (\kappa s)^2 / s \propto \lambda / \mu(\kappa s)^2 \sim \lambda / m_{\widetilde{H}^{\pm}} \times m_{\widetilde{S}^0}^2$.

 $m_{\tilde{5}^0}$ constrains $(\kappa s)^2$ and $m_{\tilde{H}^{\pm}}$ constrains μ well. Finite shift on λ is key. We have

$$\delta\lambda/\lambda|_{\rm fin.}^{\rm t_{134}}=$$
 62.26% and $\delta\lambda/\lambda|_{\rm fin.}^{\rm OS}=-7.88\%$

and loop correction is $\delta\Gamma/\Gamma \sim 2\delta\lambda/\lambda$ due to finite part of CT.

Small $\overline{\text{DR}}$ corrections \rightarrow pure virtual corrections negligible and κ, s do not run much (confirmed if one inspects the resp. $\beta_{\kappa,s}$ functions).

LPT Orsav

▶ singlets:
$$h_2^0, A_1^0, \tilde{\chi}_3^0, m_{h_2^0} = 240 \text{ GeV}, m_{h_3^0, A_2^0, H^{\pm}} \sim 570 \text{ GeV},$$

 $Q_{\text{SUSY}} = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}} = 1117 \text{ GeV}$

Decays	$t_{134A_1A_2}(Q_{\mathrm{SUSY}})$	$OS_{34h_2A_1A_2H^+}$	$\overline{\mathrm{DR}}(m_{\scriptscriptstyle parent})$	$\overline{\mathrm{DR}}(Q_{\mathrm{SUSY}})$
$h_2^0 ightarrow A_1^0 A_1^0$	(128%)	(-12%)	(0.4%)	(-0.4%)
$h_3^0 ightarrow { ilde \chi}_1^0 { ilde \chi}_3^0$	(122%)	(-3%)	(2%)	(0.3%)
$h_3^0 ightarrow ilde{\chi}_2^0 \overline{ ilde{\chi}_3^0}$	(126%)	(-35%)	(3%)	(1.1%)
$A_2^0 ightarrow ilde{\chi}_1^0 ilde{\chi}_3^0$	(130%)	(-31%)	(8%)	(6.2%)
$A^{ar 0}_2 o ilde\chi^{ar 0}_2 ilde\chi^{ar 0}_3$	(122%)	(-5%)	(-0.4%)	(-1.9%)
$H^+ ightarrow ilde{\chi}_1^+ ilde{\chi}_3^0$	(125%)	(-18%)	(3%)	(1.1%)

Same usual suspects, corrections in t_{134} accounted for by $\delta\lambda|_{\mathrm{fin.}}$. In $OS_{34h_2A_1A_2H^+}$ renormalisation of δt_β kicks in.



► singlets:
$$h_2^0, A_1^0, \tilde{\chi}_3^0, m_{h_2^0} = 240 \text{ GeV}, m_{h_3^0, A_2^0, H^{\pm}} \sim 570 \text{ GeV},$$

 $Q_{\text{SUSY}} = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}} = 1117 \text{ GeV}$

Decays	$t_{134A_1A_2}(Q_{\mathrm{SUSY}})$	$OS_{34h_2A_1A_2H^+}$	$\overline{\mathrm{DR}}(m_{\scriptscriptstyle parent})$	$\overline{\mathrm{DR}}(\mathcal{Q}_{\mathrm{SUSY}})$
$h_2^0 \rightarrow A_1^0 A_1^0$	(128%)	(-12%)	(0.4%)	(-0.4%)
$h_3^0 ightarrow ilde{\chi}_1^0 ilde{\chi}_3^0$	(122%)	(-3%)	(2%)	(0.3%)
$h_3^0 ightarrow ilde{\chi}_2^0 \overline{ ilde{\chi}_3^0}$	(126%)	(-35%)	(3%)	(1.1%)
$A_2^0 ightarrow ilde{\chi}_1^0 ilde{\chi}_3^0$	(130%)	(-31%)	(8%)	(6.2%)
$A^{ar 0}_2 o ilde\chi^{ar 0}_2 ilde\chi^{ar 0}_3$	(122%)	(-5%)	(-0.4%)	(-1.9%)
$H^+ ightarrow { ilde \chi}^+_1 { ilde \chi}^0_3$	(125%)	(-18%)	(3%)	(1.1%)
$h_3^0 ightarrow h_1^0 \overline{h_2^0}$	(116%)	(79%)	(52%)	(-1.7%)

Solution For $t_{134} \& OS_{34h_2A_1A_2H^+}$: same reasons as before



► singlets:
$$h_2^0, A_1^0, \tilde{\chi}_3^0, m_{h_2^0} = 240 \text{ GeV}, m_{h_3^0, A_2^0, H^{\pm}} \sim 570 \text{ GeV},$$

 $Q_{\text{SUSY}} = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}} = 1117 \text{ GeV}$

Decays	$t_{134A_1A_2}(Q_{\mathrm{SUSY}})$	$OS_{34h_2A_1A_2H^+}$	$\overline{\mathrm{DR}}(m_{\scriptscriptstyle parent})$	$\overline{\mathrm{DR}}(Q_{\mathrm{SUSY}})$
$h_2^0 \rightarrow A_1^0 A_1^0$	(128%)	(-12%)	(0.4%)	(-0.4%)
$h_3^0 ightarrow ilde{\chi}_1^0 ilde{\chi}_3^0$	(122%)	(-3%)	(2%)	(0.3%)
$h_3^0 \rightarrow ilde{\chi}_2^0 \overline{ ilde{\chi}_3^0}$	(126%)	(-35%)	(3%)	(1.1%)
$A_2^0 ightarrow ilde{\chi}_1^0 ilde{\chi}_3^0$	(130%)	(-31%)	(8%)	(6.2%)
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Solution For t_{134} & $OS_{34h_2A_1A_2H^+}$: same reasons as before

 $\label{eq:large} \hbox{${\rm I}$$ arge corrections $\overline{{\rm DR}}$? Pt A has small mixing: $g_{h_1h_2h_3} \simeq g_{h_uh_sh_d} \sim \lambda $ $ A_\lambda $ + 2\kappa\mu$. }$

$$16\pi^2 \frac{1}{A_\lambda} \frac{dA_\lambda}{A_\lambda} \sim 3h_t^2 \frac{A_t}{A_\lambda} \quad \text{recall} \quad \boxed{A_t/A_\lambda \sim 27!!}$$



► singlets:
$$h_2^0, A_1^0, \tilde{\chi}_3^0, m_{h_2^0} = 240 \text{ GeV}, m_{h_3^0, A_2^0, H^{\pm}} \sim 570 \text{ GeV},$$

 $Q_{\text{SUSY}} = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}} = 1117 \text{ GeV}$

Decays	$t_{134A_1A_2}(Q_{\mathrm{SUSY}})$	$OS_{34h_2A_1A_2H^+}$	$\overline{\mathrm{DR}}(m_{\scriptscriptstyle parent})$	$\overline{\mathrm{DR}}(Q_{\mathrm{SUSY}})$
$h_2^0 ightarrow A_1^0 A_1^0$	(128%)	(-12%)	(0.4%)	(-0.4%)
$h_3^0 \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_3^0$	(122%)	(-3%)	(2%)	(0.3%)
$h_3^0 ightarrow ilde{\chi}_2^0 ilde{\chi}_3^0$	(126%)	(-35%)	(3%)	(1.1%)
$A_2^0 \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_3^0$	(130%)	(-31%)	(8%)	(6.2%)
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Solution For t_{134} & $OS_{34h_2A_1A_2H^+}$: same reasons as before

 $\label{eq:large corrections $\overline{\mathrm{DR}}$? Pt A has small mixing: $g_{h_1h_2h_3} \simeq g_{h_uh_sh_d} \sim \lambda \middle| A_\lambda \middle| + 2\kappa\mu.$

$$16\pi^2 \frac{1}{A_\lambda} \frac{dA_\lambda}{A_\lambda} \sim 3h_t^2 \frac{A_t}{A_\lambda} \quad \text{recall} \left[\frac{A_t/A_\lambda \sim 27!!}{A_\lambda \sim 27!!} \right]$$

 \blacksquare Correction in $\overline{\mathrm{DR}}$ driven by running of A_λ



► singlets:
$$h_2^0, A_1^0, \tilde{\chi}_3^0, m_{h_2^0} = 240 \text{ GeV}, m_{h_3^0, A_2^0, H^{\pm}} \sim 570 \text{ GeV},$$

 $Q_{\text{SUSY}} = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}} = 1117 \text{ GeV}$

Decays	$t_{134A_1A_2}(Q_{\mathrm{SUSY}})$	$OS_{34h_2A_1A_2H^+}$	$\overline{\mathrm{DR}}(m_{\scriptscriptstyle parent})$	$\overline{\mathrm{DR}}(Q_{\mathrm{SUSY}})$
$h_2^0 \rightarrow A_1^0 A_1^0$	(128%)	(-12%)	(0.4%)	(-0.4%)
$h_3^0 ightarrow ilde{\chi}_1^0 ilde{\chi}_3^0$	(122%)	(-3%)	(2%)	(0.3%)
$h_3^0 ightarrow ilde{\chi}_2^0 \overline{ ilde{\chi}_3^0}$	(126%)	(-35%)	(3%)	(1.1%)
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$A^{ar 0}_2 o ilde\chi^{ar 0}_2 ilde\chi^{ar 0}_3$	(122%)	(-5%)	(-0.4%)	(-1.9%)
$H^+ ightarrow { ilde \chi}^+_1 { ilde \chi}^0_3$	(125%)	(-18%)	(3%)	(1.1%)
$h_3^0 ightarrow h_1^0 \overline{h_2^0}$	(116%)	(79%)	(52%)	(-1.7%)

Solution For t_{134} & $OS_{34h_2A_1A_2H^+}$: same reasons as before

Solution Simpler Corrections $\overline{\mathrm{DR}}$? Pt A has small mixing: $g_{h_1h_2h_3} \simeq g_{h_uh_sh_d} \sim \lambda |A_{\lambda}| + 2\kappa\mu$.

$$16\pi^2 \frac{1}{A_\lambda} \frac{dA_\lambda}{A_\lambda} \sim 3h_t^2 \frac{A_t}{A_\lambda} \quad \text{recall} \boxed{A_t/A_\lambda \sim 27!!}$$

- \blacksquare Correction in $\overline{\mathrm{DR}}$ driven by running of A_λ
- An MSSM-like point with large A_t to reproduce correct m_H^{SM} entitled to large EW rad. cor., even in $\overline{\text{DR}}$, for decays driven by A_{λ} . For Pt A can be absorbed by setting $\overline{\mu} = Q_{\text{SUSY}}$. Q_{SUSY} always the right choice ?

> Mixing important $\lambda = 0.67$, no pure state, in principle better extraction of counterterms



- Mixing important $\lambda = 0.67$, no pure state, in principle better extraction of counterterms

Decays	SloopS	SloopS		
	$t_{123}(Q_{\rm SUSY})$	$OS_{12h_2A_1A_2H^+}$	$\overline{\mathrm{DR}}(m_{parent})$	$\overline{\mathrm{DR}}(Q_{\mathrm{SUSY}})$
$h_3^0 ightarrow ilde{\chi}_1^0 ilde{\chi}_2^0$	10.8%	(14%)	(5%)	(3%)
$h_3^0 \rightarrow A_1^0 Z^-$	(8.4%)	(3%)	(-3%)	(-8 %)
$h_3^0 \rightarrow h_2^0 h_1^0$	(-131.4%)	(-25%)	(-106%)	(-50%)
$h_3^{0} \rightarrow h_2^{\overline{0}} h_2^{\overline{0}}$	(41.8 %)	(6%)	(13%)	(-28%)
$A_2^0 ightarrow { ilde \chi}_1^+ { ilde \chi}_1^-$	(8.2 %)	(7%)	(2%)	(1%)
${\cal A}^0_2 o ilde{\chi}^0_1 ilde{\chi}^0_1$	(18.1%)	(32%)	(2%)	(2%)
$A_2^0 ightarrow Zh_2^0$	(-10.27 %)	(12%)	(-16%)	(-9%)
$A_2^{ar 0} ightarrow A_1^0 ar h_1^0$	(-40.9 %)	(-0.3%)	(-32%)	(-17%)
$H^+ ightarrow { ilde \chi}_1^+ { ilde \chi}_2^0$	(8.4%)	(6%)	(10%)	(8%)
$H^+ ightarrow W^+ h_2^0$	(-11%)	(11%)	(-18%)	(-10%)
$H^+ \rightarrow W^+ A_1^0$	(7.9%)	(2%)	(-3%)	(-9%)
$H^+ ightarrow { ilde \chi}_1^+ { ilde \chi}_1^0$	(12.5 %)	(21%)	(9%)	(9%)

- > Due to large mixing, dependence on parameters much more involved.
- > Still renormalisation of λ , t_{β} and running of A_{λ} (although smaller due to smaller A_t/A_{λ}) lead the corrections
- OS scheme gives reasonable corrections
- For $\overline{\text{DR}}$ even $\overline{\mu} = Q_{\text{SUSY}}$ does not absorb all the corrections most probably because A_{λ} is not the only driver of the decay



- ${\tt IS}$ Using <code>SloopS</code>, in principle up to 2 \rightarrow 3 processes at the 1L level can be evaluated
- Full on-shell renormalisation (all sectors) of the NMSSM at one-loop completed.
- \blacksquare Various schemes investigated \rightarrow large scheme dependence for some observables, depending on the scenario
- ^{ESF} Currently impossible to choose what is the best scheme for reconstructing parameters. As long as only predictions are concerned, $\overline{\rm DR}$ scheme sufficient but large pure EW corrections are possible in some scenarios (in particular when singlets are involved in MSSM-like points). Not always clear how to tame them by choosing appropriate $\bar{\mu}$
- Applications to astrophysics, colliders and cosmology (future)
- SloopS is not limited to SUSY model, any renormalisable model (ex: xSM, wino-DM model) can be implemented



BACKUP



GC, F. Domingo, PRD86 (2012) 115024; G. Bélanger, V. Bizouard, GC, PRD89 (2014) 9, 095023

• We know $m_h^{SM} \equiv m_H$ gets large rad. cor.



GC, F. Domingo, PRD86 (2012) 115024; G. Bélanger, V. Bizouard, GC, PRD89 (2014) 9, 095023

- We know $m_h^{SM} \equiv m_H$ gets large rad. cor.
- To get the right kinematics for its decays, better take $m_H^{\text{corr}} = \sqrt{s}$ on the external leg for comparison with exp.



GC, F. Domingo, PRD86 (2012) 115024; G. Bélanger, V. Bizouard, GC, PRD89 (2014) 9, 095023

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- To get the right kinematics for its decays, better take $m_H^{\text{corr}} = \sqrt{s}$ on the external leg for comparison with exp.
- Couplings of H to scalars/Goldstones also $\propto m_H \rightarrow$ one should ensure that $m_H^{\rm kin} \equiv m_H^{\rm coup}$ to maintain gauge invariance



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- Define an effective rad. Higgs potential



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- Couplings of *H* to scalars/Goldstones also $\propto m_H \rightarrow$ one should ensure that $m_H^{\text{kin}} \equiv m_H^{\text{coup}}$ to maintain gauge invariance
- Define an effective rad. Higgs potential

$$\begin{split} \mathcal{V}_{\rm rad.}^{S} &= m_{H_{u}}^{2} |H_{u}|^{2} + m_{H_{d}}^{2} |H_{d}|^{2} + \frac{\lambda_{1}}{2} |H_{d}|^{4} + \frac{\lambda_{2}}{2} |H_{u}|^{4} + \lambda_{3} |H_{u}|^{2} |H_{d}|^{2} + \lambda_{4} |H_{u} \cdot H_{d}|^{2} \\ &+ m_{5}^{2} |S|^{2} + \kappa^{2} |S|^{4} + \left[\frac{A_{5}}{3} S^{3} + h.c.\right] \\ &+ \lambda_{P}^{u} |S|^{2} |H_{u}|^{2} + \lambda_{P}^{d} |S|^{2} |H_{d}|^{2} + \left[A_{ud} S H_{u} \cdot H_{d} + \lambda_{P}^{M} S^{*2} H_{u} \cdot H_{d} + h.c.\right] \end{split}$$

The tree-level conditions resulting from the NMSSM read:

$$\begin{split} \lambda_1^0 &= \frac{g^2 + g'^2}{4} = \lambda_2^0 \quad ; \ \lambda_3^0 = \frac{g^2 - g'^2}{4} \quad ; \ \lambda_4^0 = \lambda^2 - \frac{g^2}{2} \quad ; \ \lambda_P^{u\,0} = \lambda^2 = \lambda_P^{d\,0} ; \\ \lambda_P^{M\,0} &= \lambda \kappa \qquad ; \ A_S^0 = \kappa A_\kappa \qquad ; \ A_{ud}^0 = \lambda A_\lambda \qquad ; \ \kappa^{0\,2} = \kappa^2 \end{split}$$



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- Define an effective rad. Higgs potential

$$\begin{split} \mathcal{V}_{\rm rad.}^{S} &= m_{H_{u}}^{2} |H_{u}|^{2} + m_{H_{d}}^{2} |H_{d}|^{2} + \frac{\lambda_{1}}{2} |H_{d}|^{4} + \frac{\lambda_{2}}{2} |H_{u}|^{4} + \lambda_{3} |H_{u}|^{2} |H_{d}|^{2} + \lambda_{4} |H_{u} \cdot H_{d}|^{2} \\ &+ m_{5}^{2} |S|^{2} + \kappa^{2} |S|^{4} + \left[\frac{A_{5}}{3} S^{3} + h.c.\right] \\ &+ \lambda_{P}^{u} |S|^{2} |H_{u}|^{2} + \lambda_{P}^{d} |S|^{2} |H_{d}|^{2} + \left[A_{ud} S H_{u} \cdot H_{d} + \lambda_{P}^{M} S^{*2} H_{u} \cdot H_{d} + h.c.\right] \end{split}$$

The tree-level conditions resulting from the NMSSM read:

$$\begin{split} \lambda_1^0 &= \frac{g^2 + g'^2}{4} = \lambda_2^0 \quad ; \ \lambda_3^0 &= \frac{g^2 - g'^2}{4} \quad ; \ \lambda_4^0 &= \lambda^2 - \frac{g^2}{2} \quad ; \ \lambda_P^{u0} &= \lambda^2 = \lambda_P^{d0} \; ; \\ \lambda_P^{M0} &= \lambda \kappa \qquad ; \ A_S^0 &= \kappa A_\kappa \qquad ; \ A_{ud}^0 &= \lambda A_\lambda \qquad ; \ \kappa^{02} &= \kappa^2 \end{split}$$

Reproduction of the corr. Higgs masses by a redefinition beyond LO such that $\lambda_i = \lambda_i^0 + \Delta \lambda_i$

$$\begin{cases} \Delta \lambda_{1} = \frac{1}{2v^{2}} \begin{bmatrix} \frac{m_{\mu}^{2}}{\log} S_{i1}^{2}}{\cos^{2} \beta} - \frac{m_{s0}^{2}}{\rho} P_{i1}^{\prime 2} \tan^{2} \beta - M_{Z}^{2} \end{bmatrix} \\ \Delta \lambda_{2} = \frac{1}{2v^{2}} \begin{bmatrix} \frac{m_{\mu}^{2}}{\log} S_{i2}^{\prime 2}}{\frac{1}{\sin^{2} \beta}} - \frac{m_{s0}^{2}}{\tan^{2} \beta} - M_{Z}^{2} \end{bmatrix} \\ \Delta \lambda_{3} = \frac{1}{2v^{2}} \begin{bmatrix} 2m_{H^{\pm}}^{2} + \frac{2m_{\mu}^{2}}{\sin^{2} \beta} - m_{s1}^{2} - M_{Z}^{2} \end{bmatrix} \\ \Delta \lambda_{4} = \frac{1}{2v^{2}} \begin{bmatrix} m_{s0}^{2} P_{i1}^{\prime 2} - m_{H^{\pm}}^{2} + M_{W}^{2} - \lambda^{2}v^{2} \end{bmatrix} \\ \Delta \lambda_{4} = \frac{1}{v^{2}} \begin{bmatrix} m_{s0}^{2} P_{i1}^{\prime 2} - m_{H^{\pm}}^{2} + M_{W}^{2} - \lambda^{2}v^{2} \end{bmatrix} \\ \Delta A_{4} = \frac{1}{3s} \begin{bmatrix} \frac{\sin 2\beta}{2s} m_{s0}^{2} P_{i1}^{\prime 2} - \frac{1}{v} m_{s0}^{2} P_{i1}^{\prime 2} P_{i2}^{\prime 2} \end{bmatrix} - \lambda A_{\lambda} \\ \Delta \lambda_{P}^{M} = \frac{1}{3s} \begin{bmatrix} \frac{\sin 2\beta}{2s} m_{s0}^{2} P_{i1}^{\prime 2} - \frac{1}{v} m_{s0}^{2} P_{i1}^{\prime 2} P_{i2}^{\prime 2} \end{bmatrix} - \lambda \kappa \\ \Delta A_{5} = \frac{1}{3s} \begin{bmatrix} \frac{v^{2} \sin^{2} 2\beta}{2s^{2}} m_{s0}^{2} P_{i1}^{\prime 2} - m_{s0}^{2} P_{i2}^{\prime 2} - \frac{v \sin 2\beta}{2s} m_{s0}^{2} P_{i1}^{\prime 2} P_{i2}^{\prime 2} \end{bmatrix} - \kappa A_{\kappa} \\ \Delta \kappa^{2} = \frac{1}{4s^{2}} \begin{bmatrix} m_{\mu0}^{2} S_{i3}^{2} + \frac{1}{3} m_{s0}^{2} P_{i2}^{\prime 2} - \frac{v^{2} \sin^{2} 2\beta}{3s^{2}} m_{s0}^{2} P_{i1}^{\prime 2} \end{bmatrix} - \kappa^{2} \\ \Delta \lambda_{P}^{\mu} = \frac{m_{\mu0}^{2} S_{i2} S_{i3}}{\frac{1}{2s V \cos \beta}} + \frac{1}{3s \tan \beta} \begin{bmatrix} \frac{\sin 2\beta}{s} m_{s0}^{2} P_{i1}^{\prime 2} - \frac{1}{2v} m_{s0}^{2} P_{i1}^{\prime 2} P_{i2}^{\prime 2} \end{bmatrix} - \lambda^{2} \\ \Delta \lambda_{P}^{\mu} = \frac{m_{\mu0}^{2} S_{i1} S_{i3}}{\frac{1}{2s V \cos \beta}} + \frac{\tan \beta}{3s} \begin{bmatrix} \frac{\sin 2\beta}{s} m_{s0}^{2} P_{i1}^{\prime 2} - \frac{1}{2v} m_{s0}^{2} P_{i1}^{\prime 2} P_{i2}^{\prime 2} \end{bmatrix} - \lambda^{2} \\ \Delta \lambda_{P}^{\mu} = \frac{m_{\mu0}^{2} S_{i1} S_{i3}}{\frac{1}{2s V \cos \beta}} + \frac{\tan \beta}{3s} \begin{bmatrix} \frac{\sin 2\beta}{s} m_{s0}^{2} P_{i1}^{\prime 2} - \frac{1}{2v} m_{s0}^{2} P_{i1}^{\prime 2} P_{i2}^{\prime 2} \end{bmatrix} - \lambda^{2} \\ \lambda_{P}^{\mu} = \frac{m_{\mu0}^{2} S_{i1} S_{i3}}{\frac{1}{2s V \cos \beta}} + \frac{\tan \beta}{3s} \begin{bmatrix} \frac{\sin 2\beta}{s} m_{s0}^{2} P_{i1}^{\prime 2} - \frac{1}{2v} m_{s0}^{2} P_{i1}^{\prime 2} P_{i2}^{\prime 2} \end{bmatrix} - \lambda^{2} \\ \lambda_{P}^{\mu} = \frac{m_{\mu0}^{2} S_{i1} S_{i3}}{\frac{1}{2s V \cos \beta}} + \frac{\tan \beta}{3s} \begin{bmatrix} \frac{\sin 2\beta}{s} m_{s0}^{2} P_{i1}^{\prime 2} - \frac{1}{2v} m_{s0}^{2} P_{i1}^{\prime 2} \end{bmatrix} \end{bmatrix}$$

 \blacktriangleright This form of $\mathcal{V}_{\rm rad.}^{S}$ is accurate at LL level. Implemented in SLOOPS

CHALONS Guillaume

AUTOMATIC LOOP CALCULATIONS OF NMSSM PROCESSES WITH SLOOPS

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