

GPU computing use case for H.E.S.S./CTA

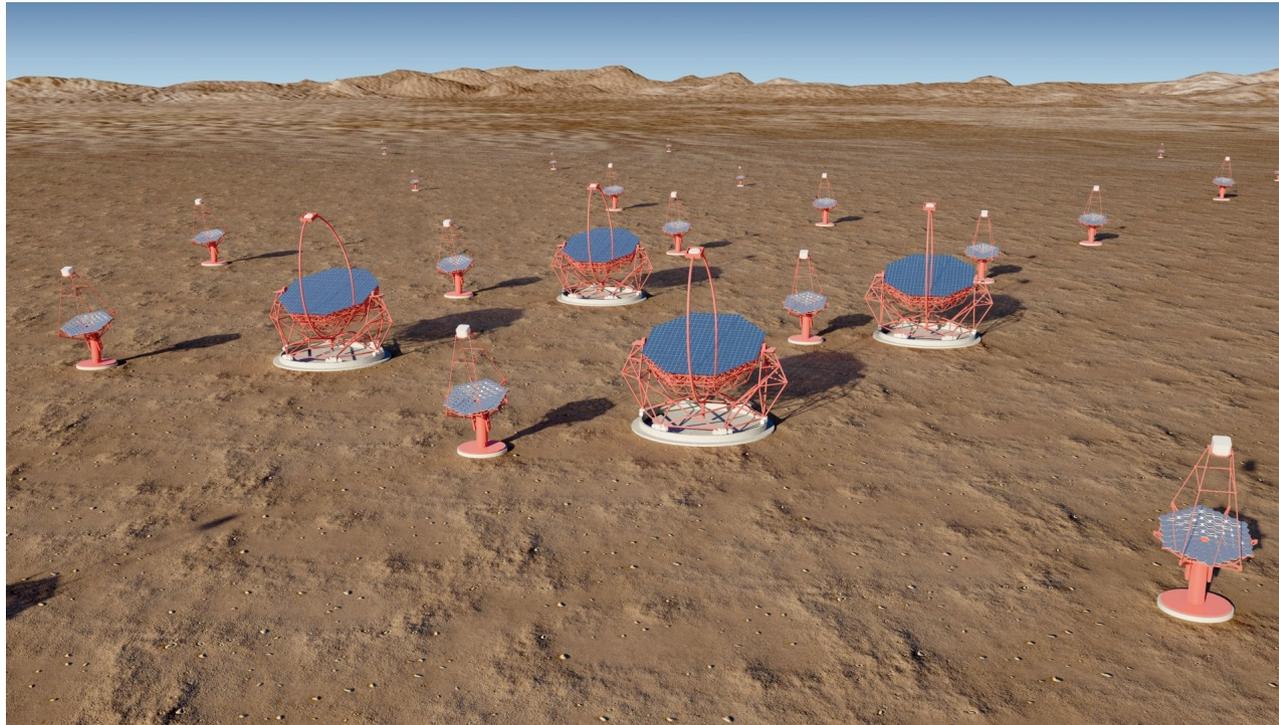
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Réunion GPU LIP6+LPNHE, 17/05/2016

The present: H.E.S.S. (High Energy Stereoscopic System)

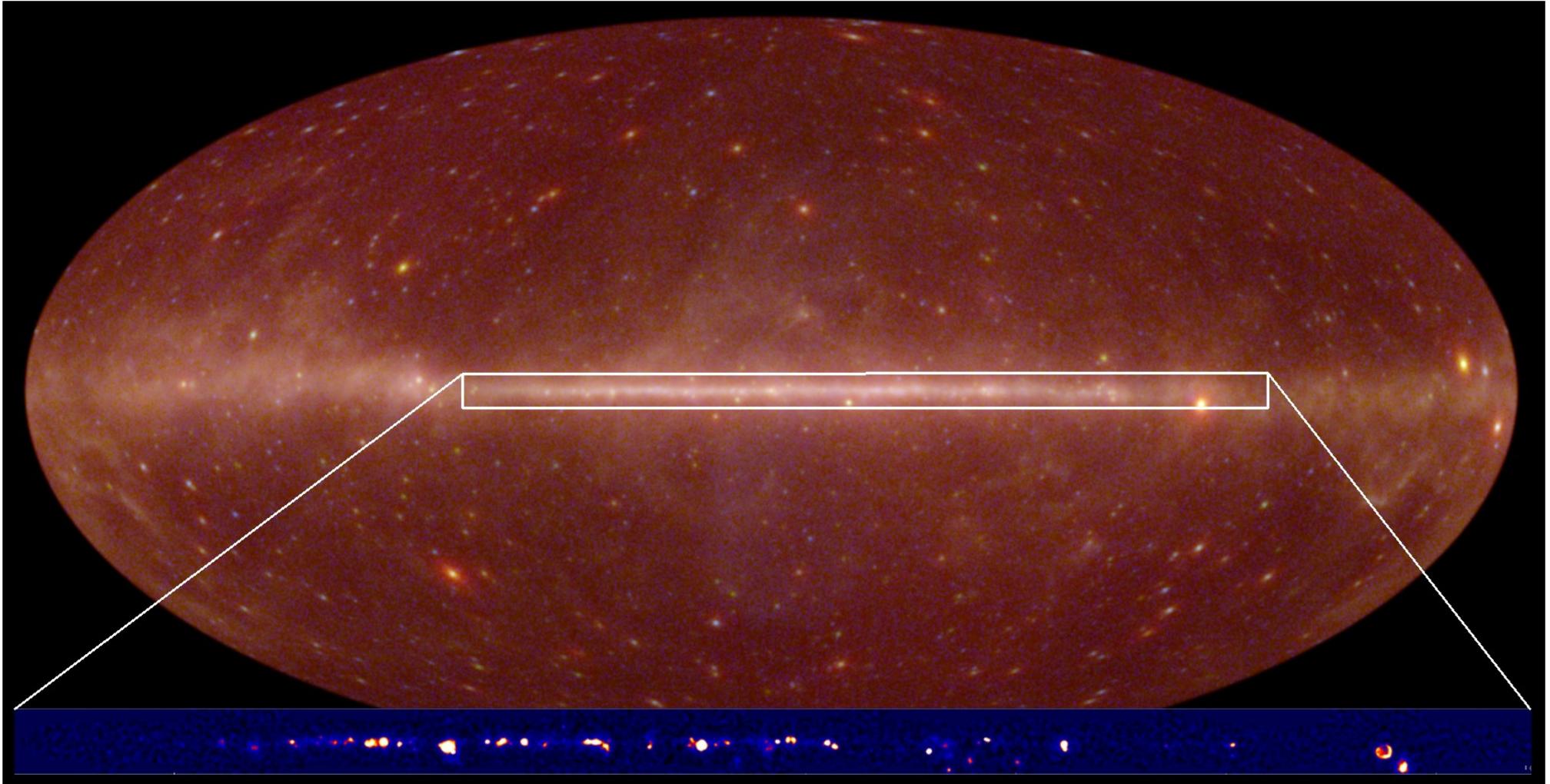


The future: CTA (Cherenkov Telescope Array)

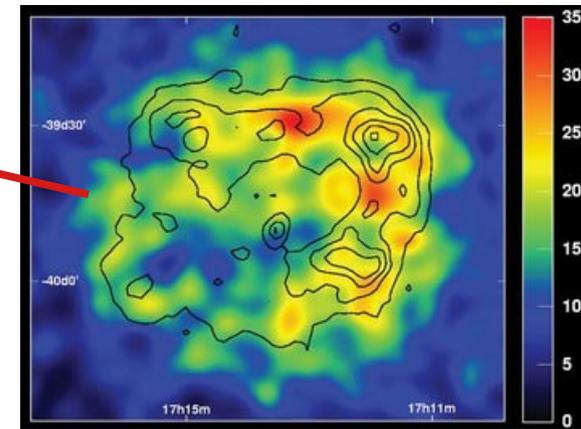
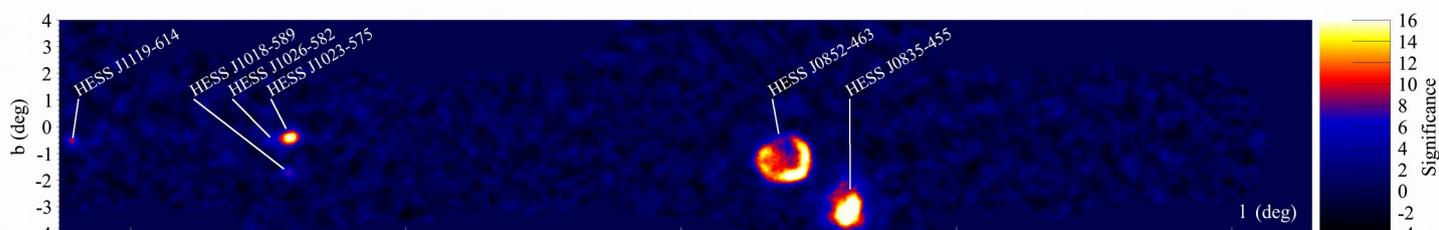
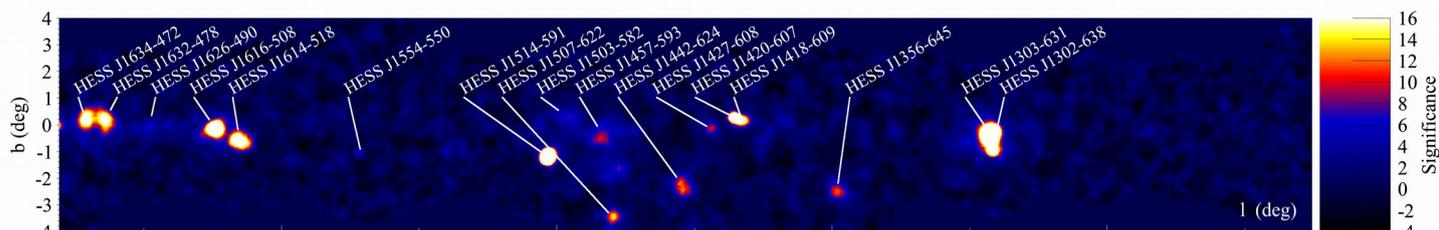
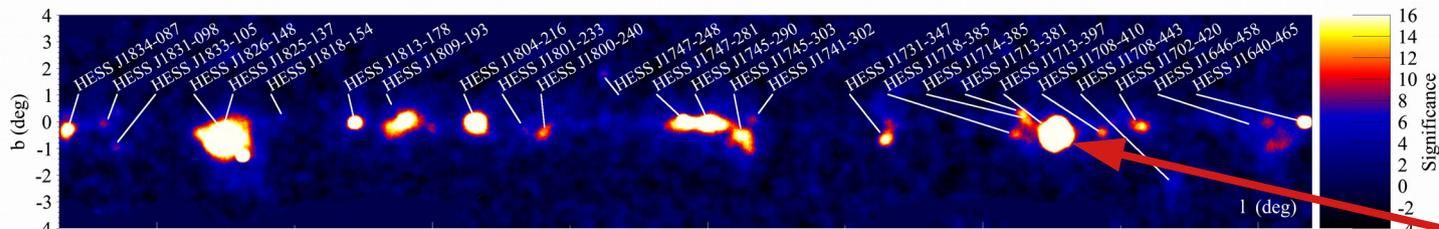
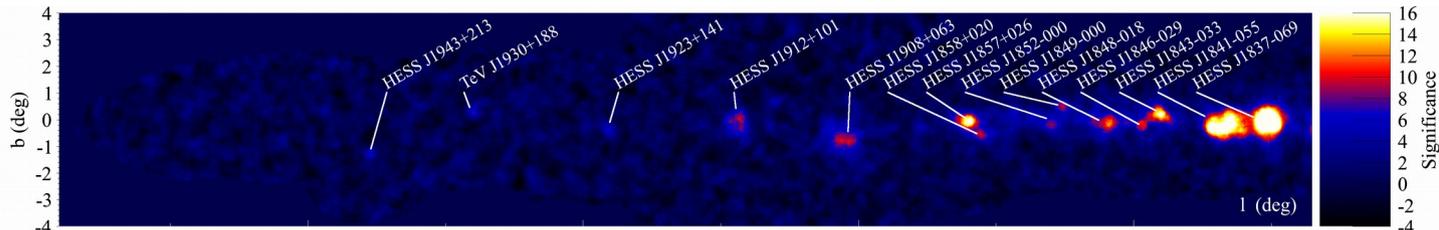


- Worldwide effort: >1200 members, 200 institutes, 32 countries and counting...
- 2 sites: one per hemisphere (ongoing discussions: La Palma (Spain), Paranal (Chile))
- 3 types of telescopes: SST, MST, LST
- Broad energy range: ~ 10 GeV – >100 TeV
- 10 x sensitivity wrt. current instruments, ~ 5 x angular resolution

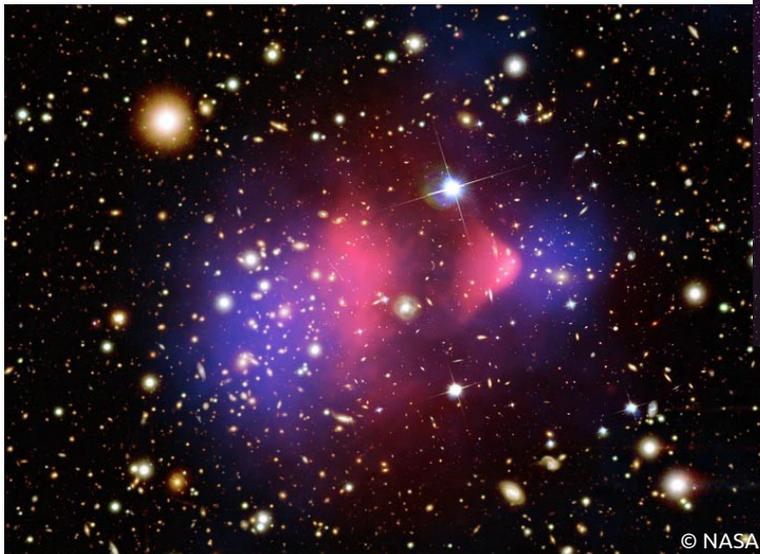
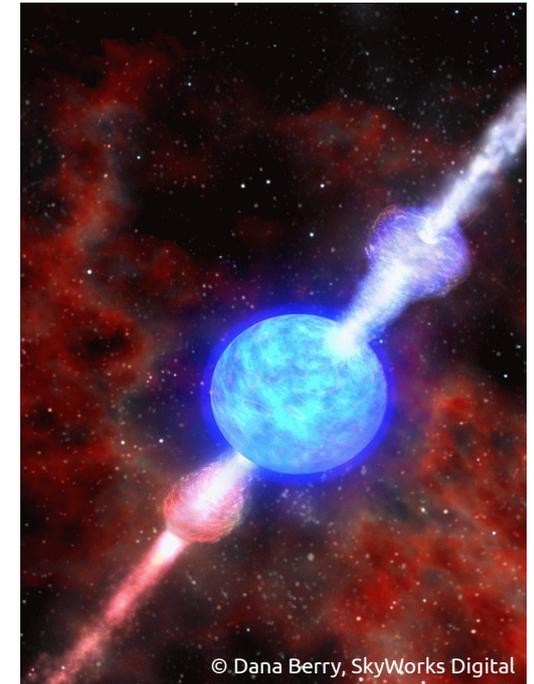
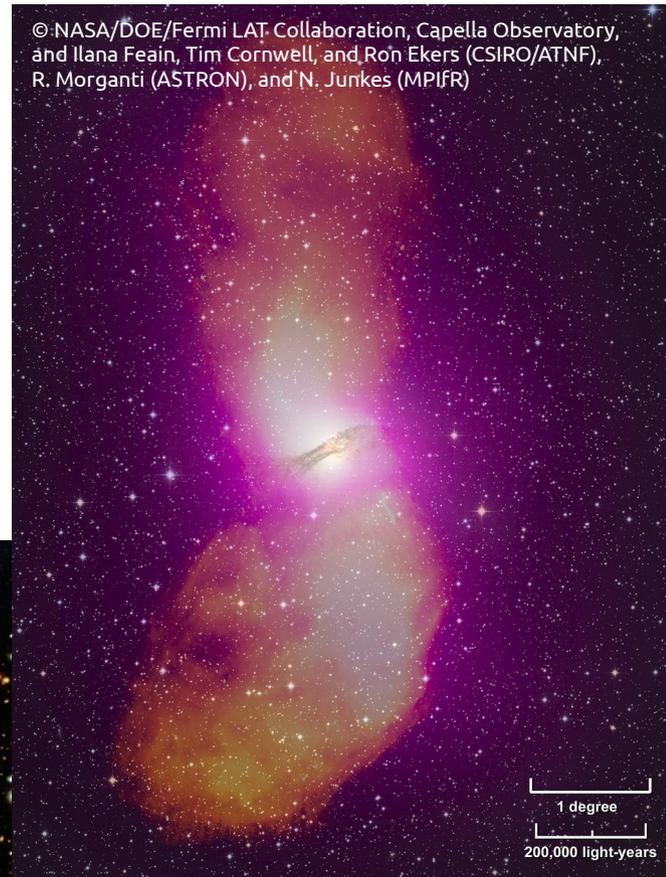
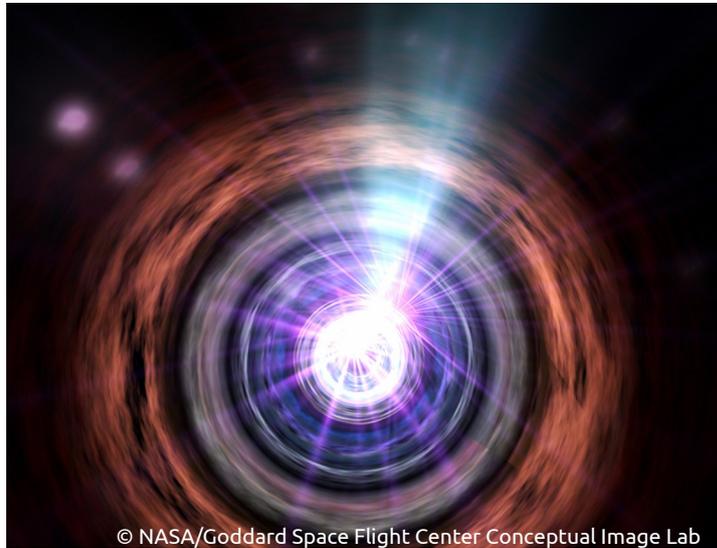
Very high energy astrophysics



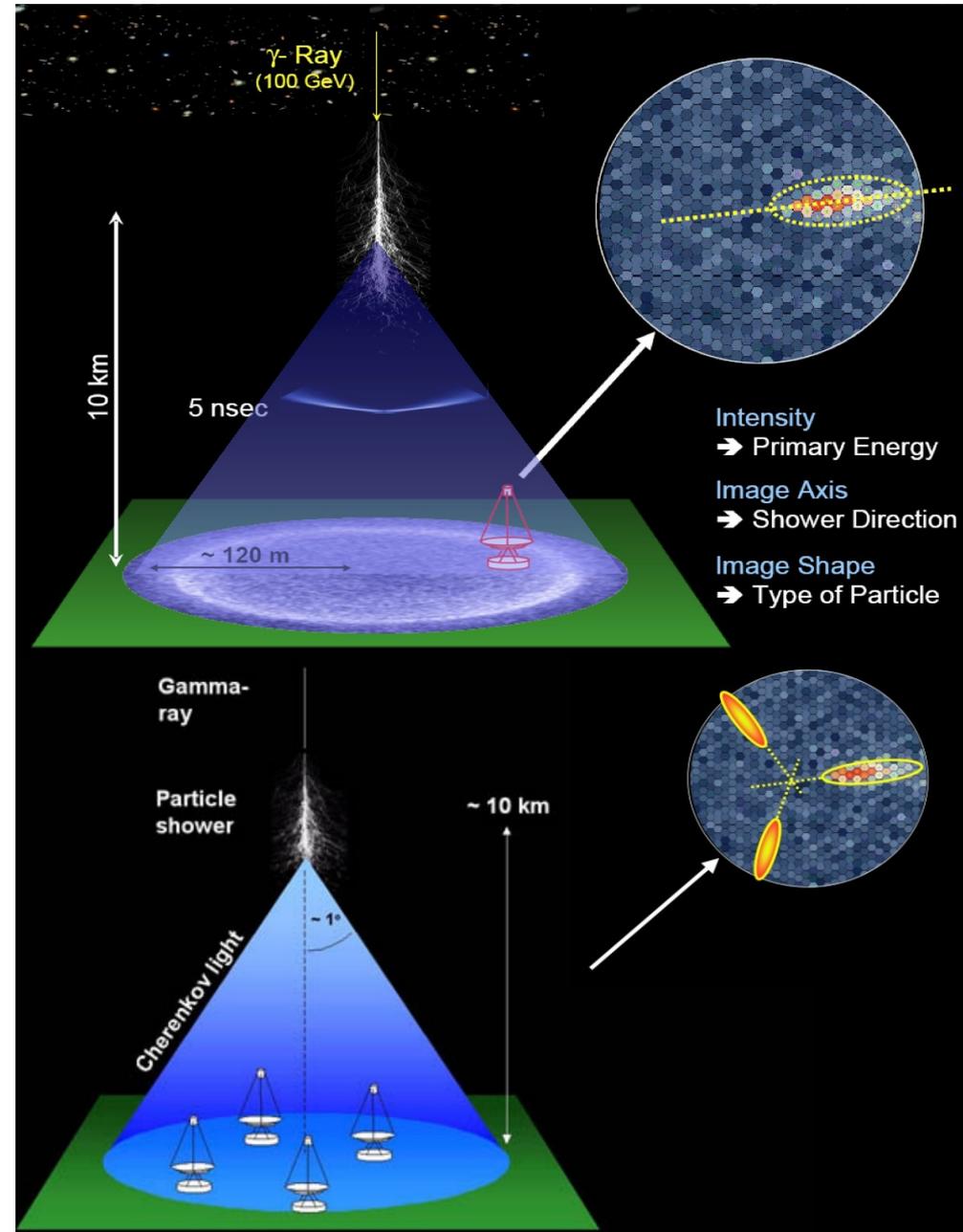
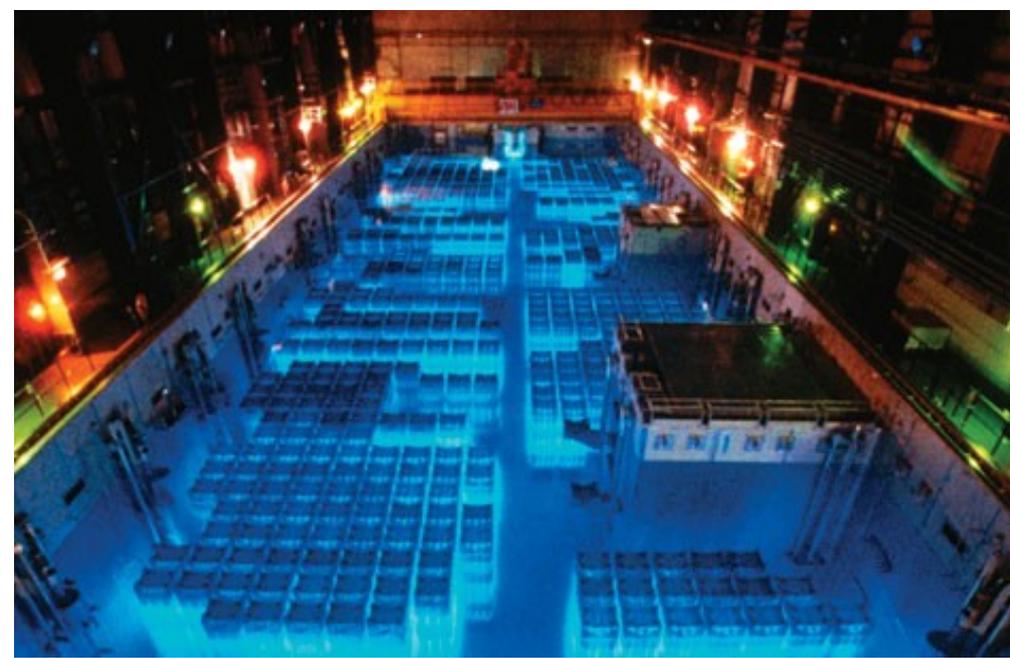
Very high energy astrophysics



Very high energy astrophysics



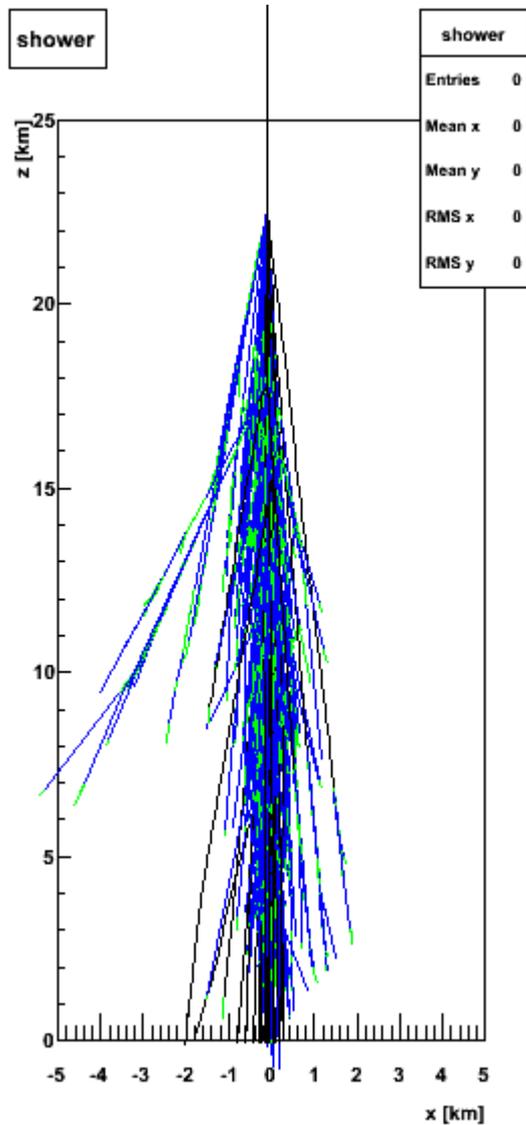
Imaging Atmospheric Cherenkov Telescopes



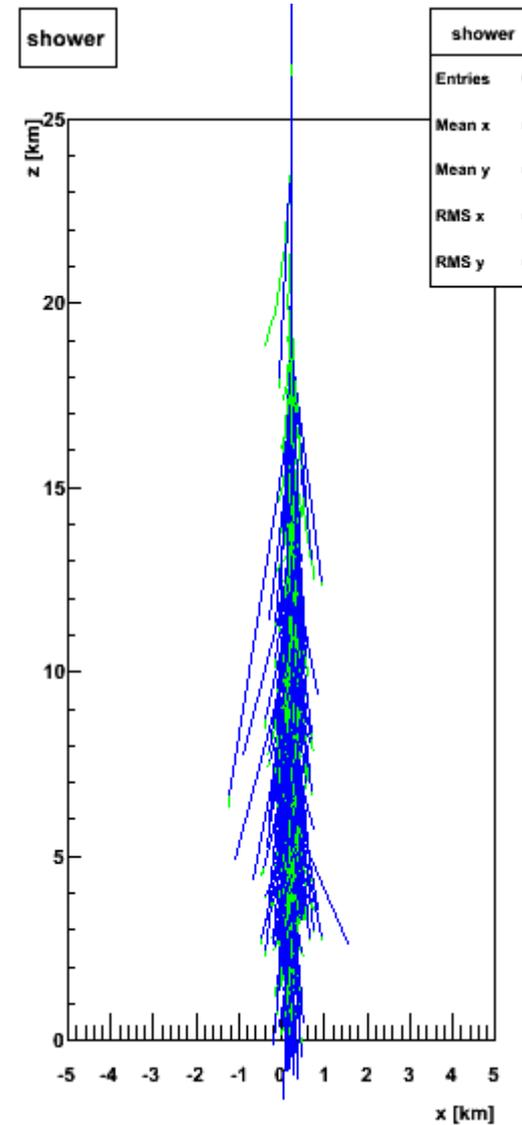
IACT data analysis

- Atmosphere **is part** of the detector → impossible to calibrate an IACT camera using a particle beam.
- ⇒ Energy scale calibration against extensive, massive MC simulations → EGI grid.
- Large background – mainly light cosmic ray nuclei (protons, helium) – over faint signal (γ -rays) to detect → event discrimination

Event discrimination



Proton 1 TeV



Gamma 1 TeV

Event discrimination

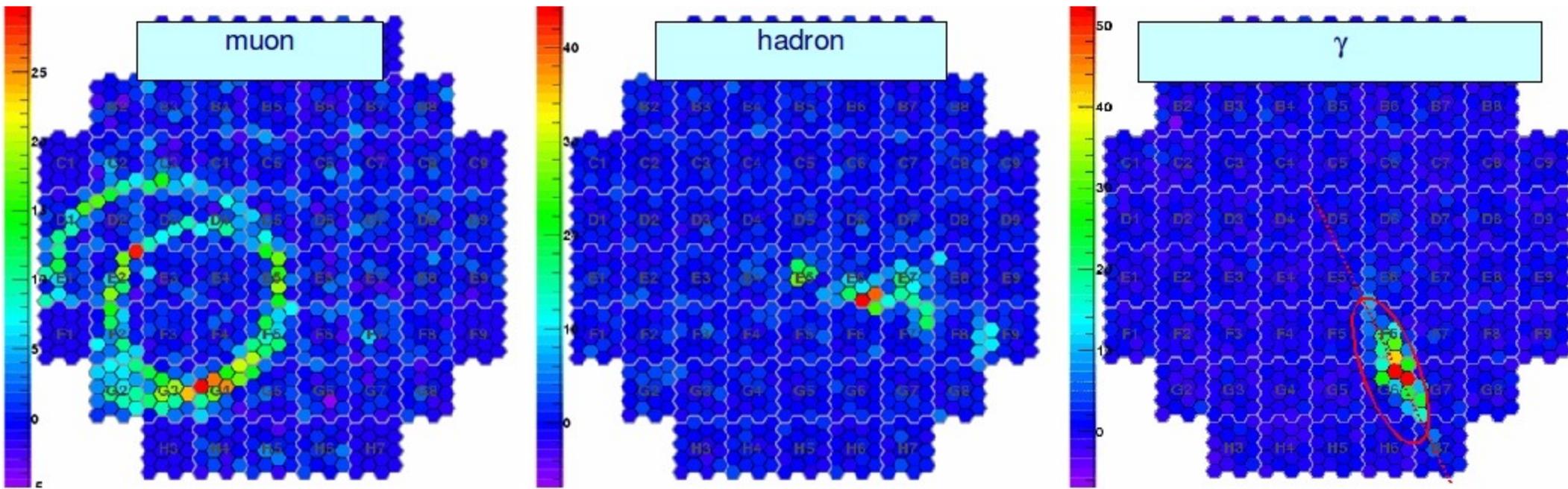


Image pattern recognition \rightarrow here could lie the power of GPU computing !

Calibration: pixel per pixel \rightarrow parallelizable