

# Characterization of image difference for transient detection

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# Outline

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- ▶ Alard-Lupton PSF-Matching subtraction
  - ▶ Basics
  - ▶ Results
  - ▶ Artifacts
- ▶ Image Subtraction on LSST Stack
  - ▶ Qualitative results
  - ▶ Classification of detections
  - ▶ Light curves
- ▶ Other methods: Zackay-Ofek optimal transient detection
- ▶ Conclusions
- ▶ Perspectives

# Introduction

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- ▶ **Transients:**

- ▶ Flux varying objects or moving objects
- ▶ Found by comparing flux intensities among several images

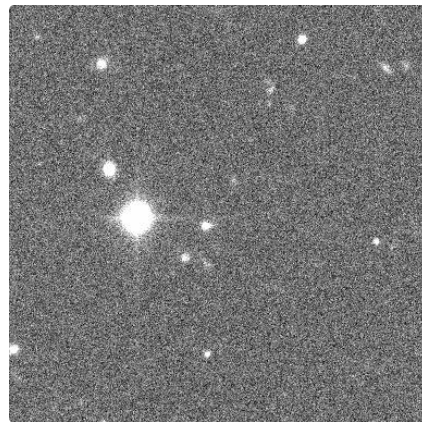
- ▶ **Flux variations:**

- ▶ Cosmic rays
- ▶ Plane misalignment
- ▶ Moving stars
- ▶ Instrumental defects
- ▶ Astrometric distortion

# Alard-Lupton<sup>1</sup> PSF-Matching subtraction: Basics

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- ▶ Two input images:
  - ▶  $T(x,y)$ : Template image, with the best PSF.
  - ▶  $S(x,y)$ : Science image. It usually has the variation of interest.
- ▶ Variable PSF is calculated on each image using a grid approximation.



$S(x,y)$



$T(x,y)$

<sup>1</sup> Alard, C. and Lupton, R. (1998). A Method for Optimal Image Subtraction. *Apj*, 503(1), pp.325-331.

# Alard-Lupton PSF-Matching subtraction: Basics

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- ▶ A Kernel convolution  $K(x,y;u,v)$  is calculated from the PSF of  $S(x,y)$  and  $T(x,y)$ .
- ▶ The index  $i$  runs over all permutations of  $n, p$  and  $q$ .  $\sigma_n$  is the width of the  $n$  component of the Gaussian and  $p$  and  $q$  are the degrees of the polynomial used to represent  $K$ .

$$K(u, v; x, y) = \sum_i a_i K_i(u, v)$$

$$K_i(u, v) = e^{-\frac{u^2+v^2}{2\sigma_n^2}} u^p v^q$$

# Alard-Lupton PSF-Matching subtraction: Basics

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- ▶  $T(x,y)$  is convolved with  $K(x,y,u,v)$  to PSF-match it to  $S(x,y)$ .
- ▶ The PSF-matched  $T(x,y)$  is finally subtracted to  $S(x,y)$  pixel to pixel.

$$S(x, y) - T(x, y) \otimes K(x, y, u, v) = D(x, y)$$

# Alard-Lupton PSF-Matching subtraction: Basics

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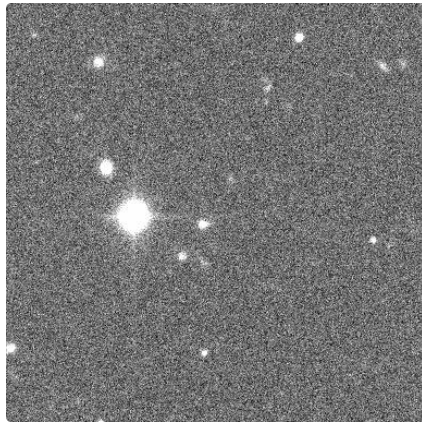
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- ▶  $D(x,y)$  is the Difference Image. Transients and variable objects are detected here.

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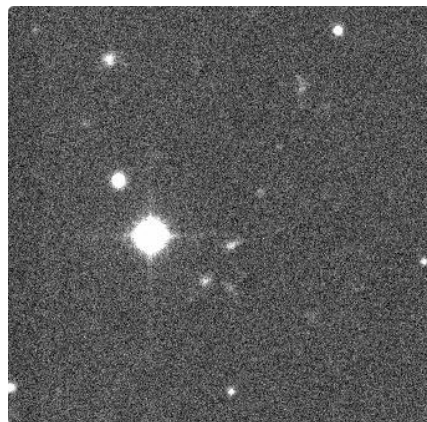
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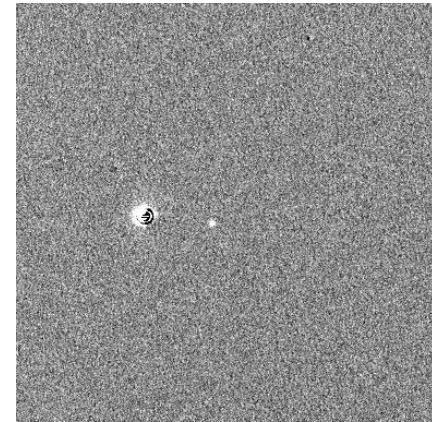
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$S(x,y)$



$T(x,y)$



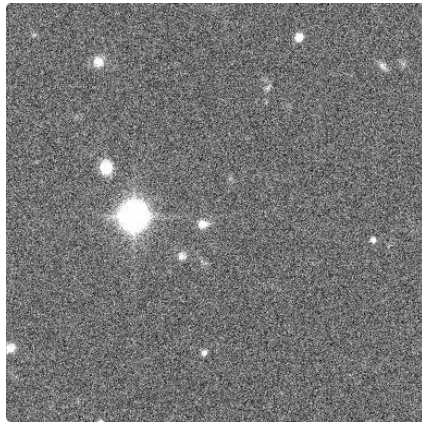
$D(x,y)$



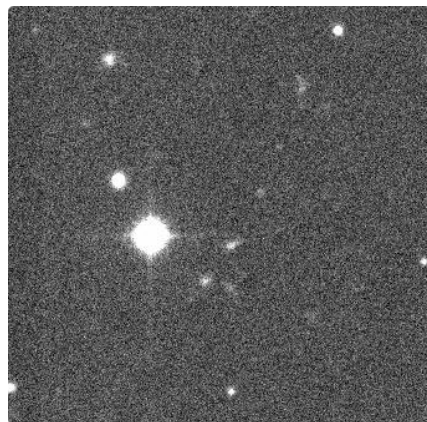
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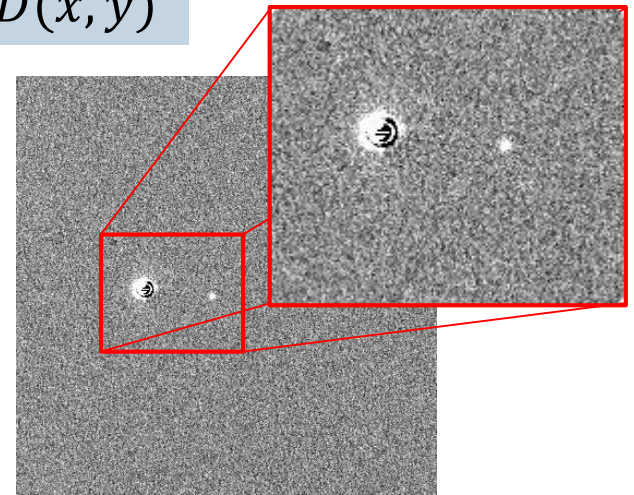
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$T(x,y)$



$D(x,y)$

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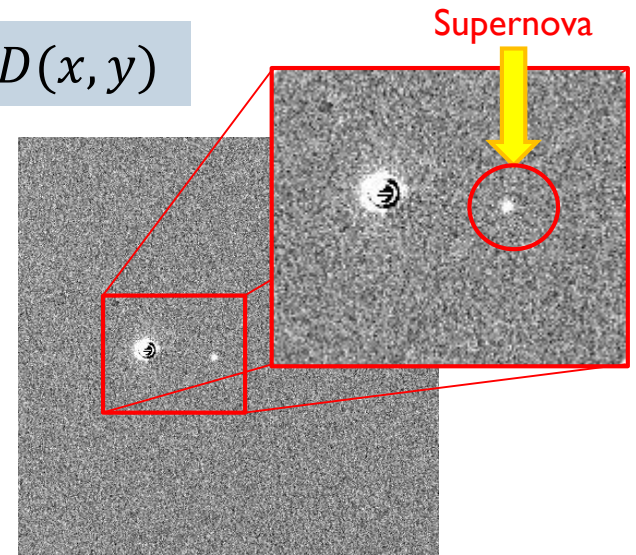
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$S(x,y)$



$T(x,y)$

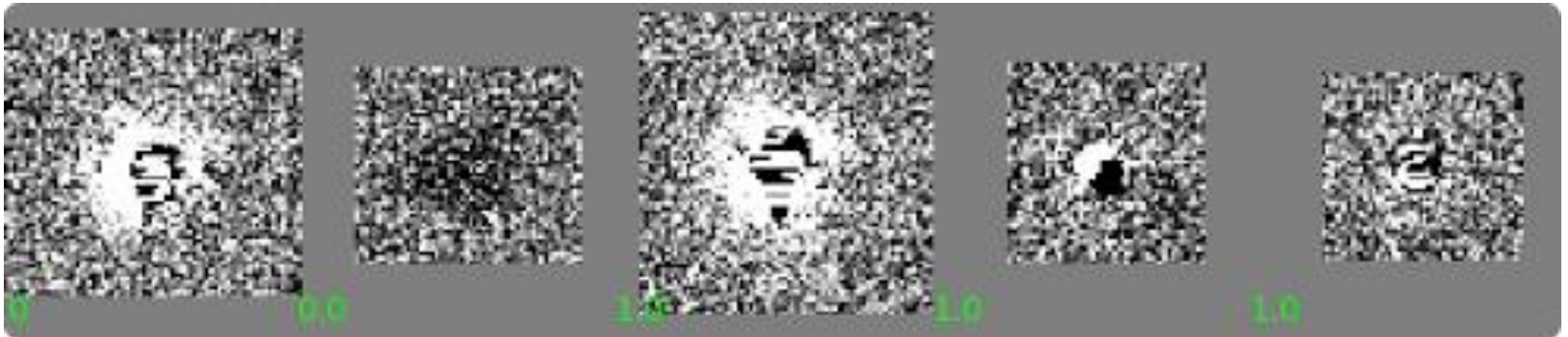


$D(x,y)$

# Alard-Lupton PSF-Matching subtraction: Artifacts

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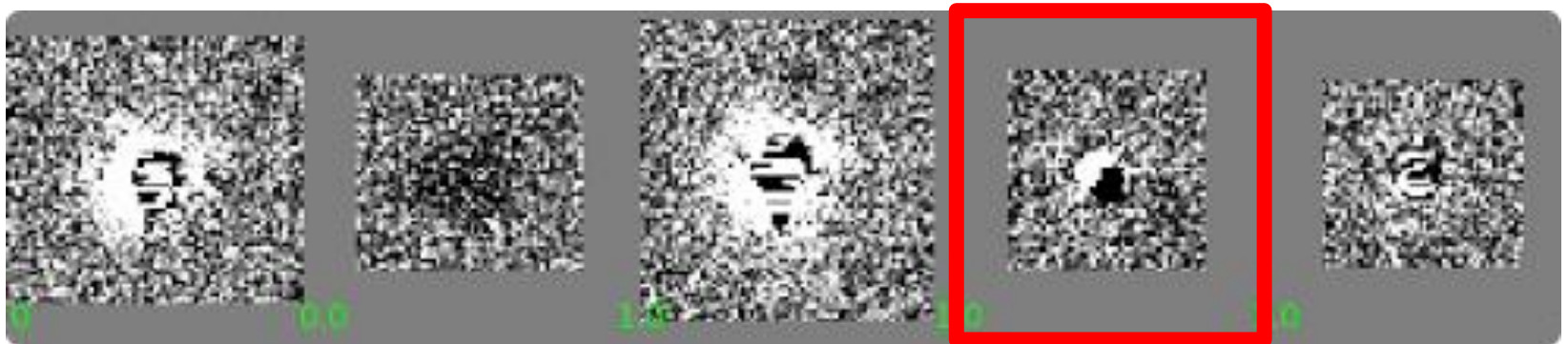
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# Alard-Lupton PSF-Matching subtraction: Artifacts

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  - ▶ Dipoles (negative and positive regions more or less balanced).





# Alard-Lupton PSF-Matching subtraction: Artifacts

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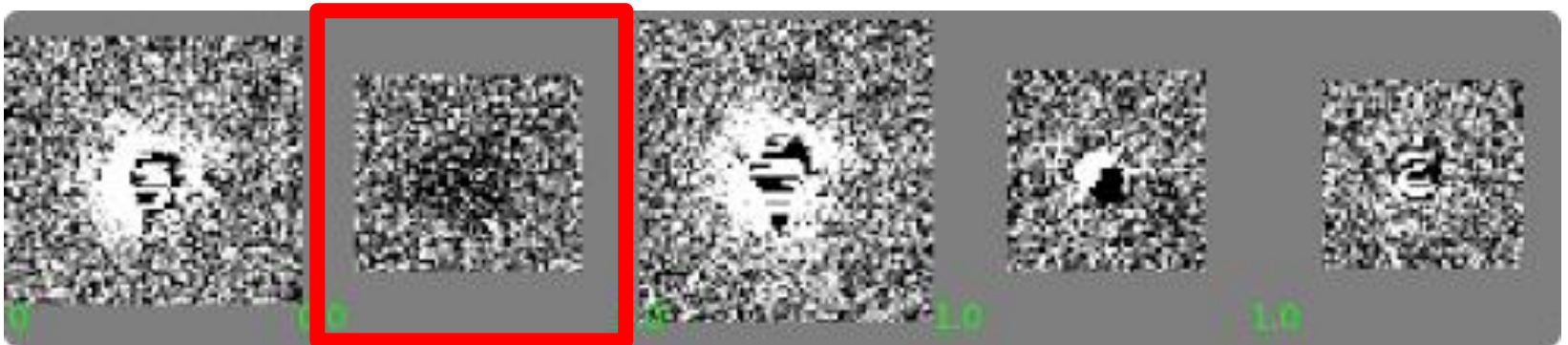
- ▶ Artifacts from Alard-Lupton method include:
  - ▶ Dipoles (negative and positive regions more or less balanced).
  - ▶ Fringes (kernel mismatch residuals).



# Alard-Lupton PSF-Matching subtraction: Artifacts

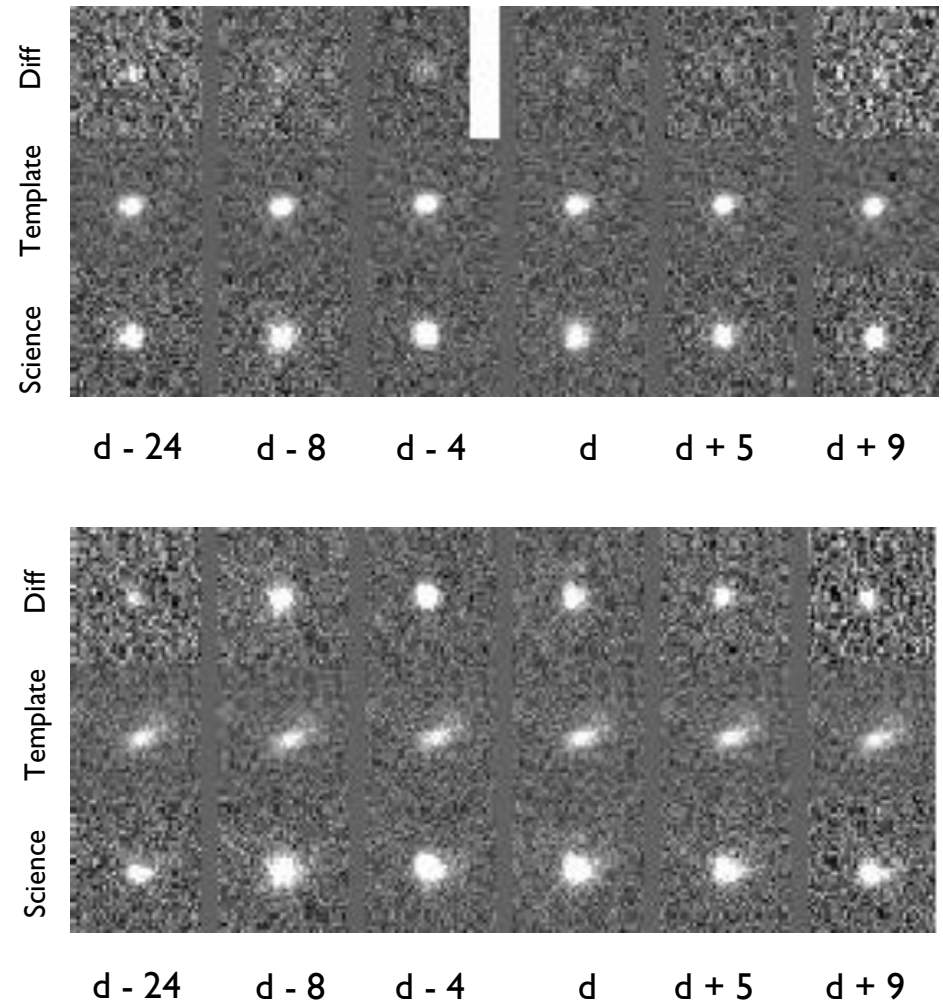
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- ▶ Artifacts from Alard-Lupton method include:
  - ▶ Dipoles (negative and positive regions more or less balanced).
  - ▶ Fringes (kernel mismatch residuals).
  - ▶ Noise.



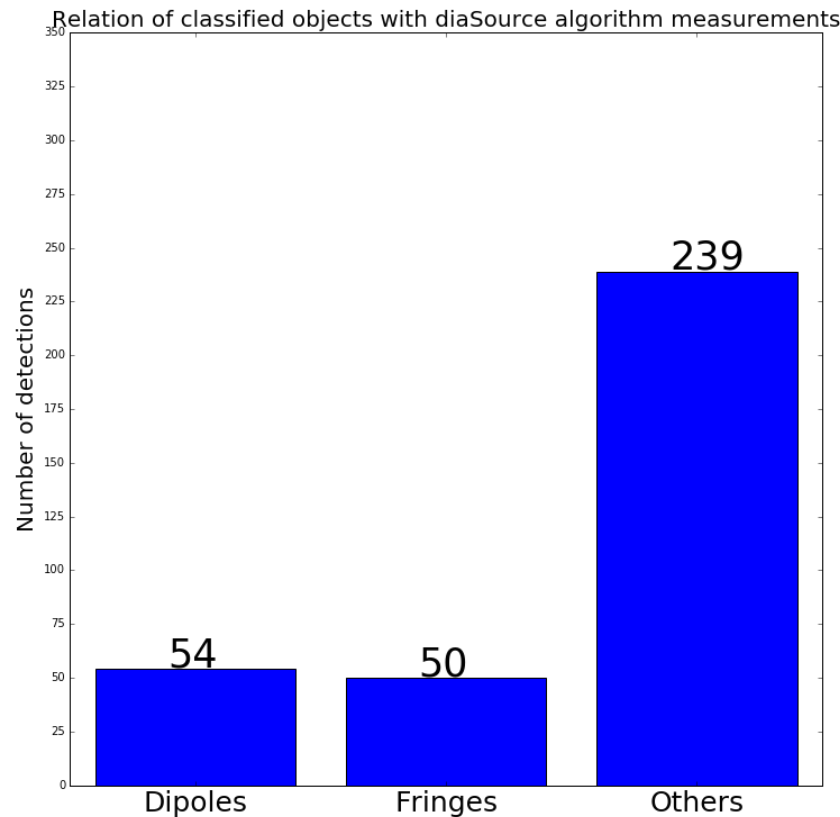
# Image Difference on Stack: Qualitative Results

- ▶ Image Difference used on CFHT images.
- ▶ Sources on Difference Images (DIASources) are detected and measured.
- ▶ Significant DIASources are selected as those with flux over  $5\sigma$ .



# Image Difference on Stack: Classification of detections

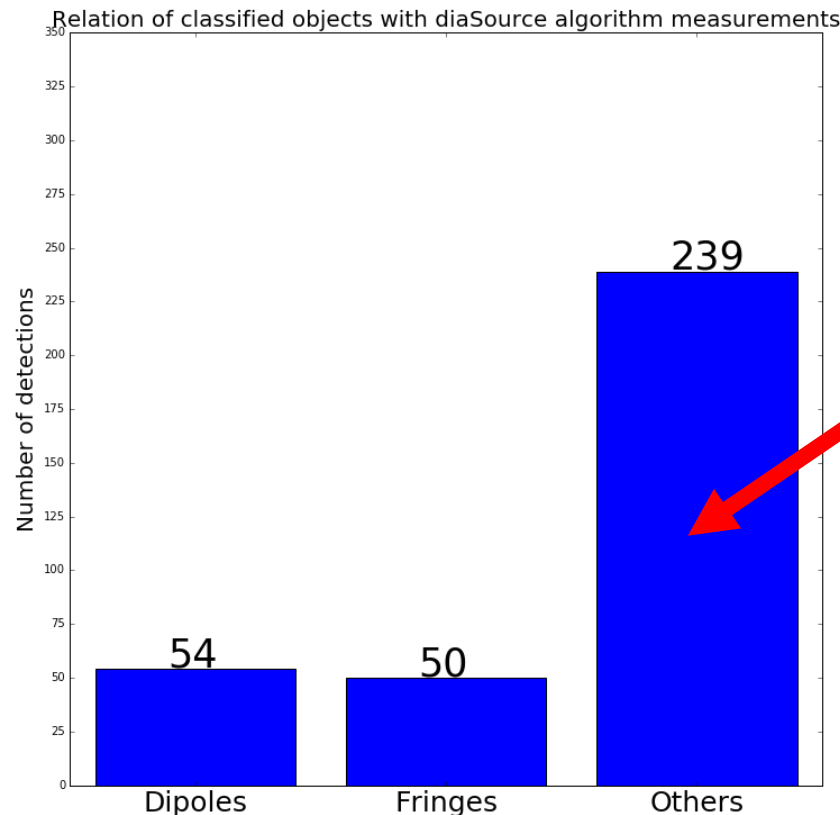
- ▶ In our first test set, we used 6 images on field D3 from the CCD 14. We had previously identified a known supernova present on all these images. There were around 2500 significative DIASources detected per square degree.





# Image Difference on Stack: Classification of detections

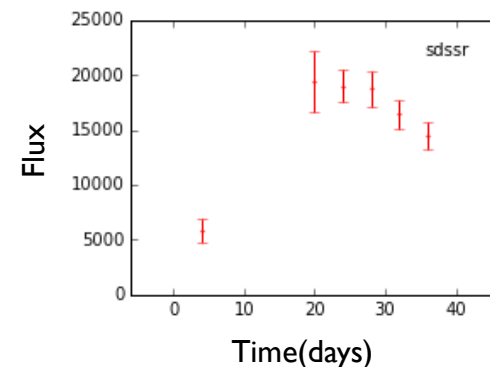
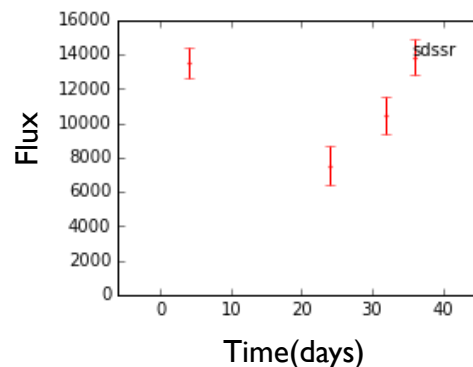
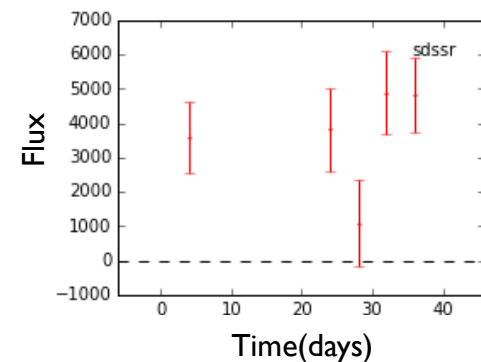
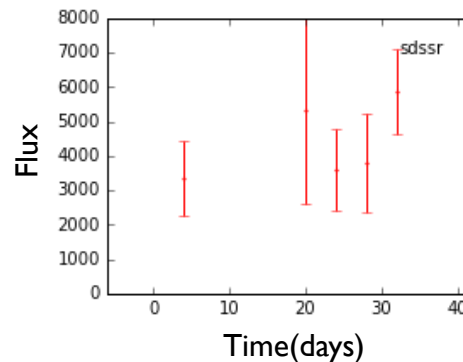
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Includes 6  
detections for  
the SN!

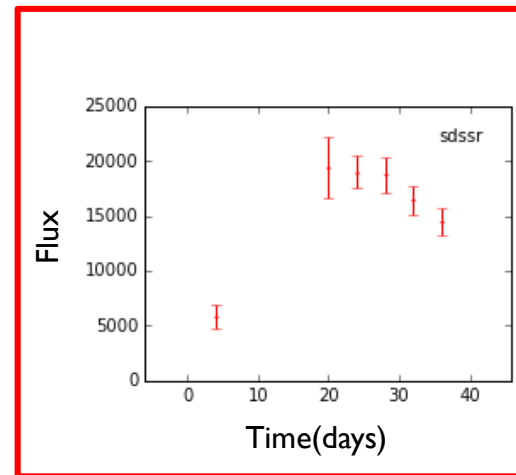
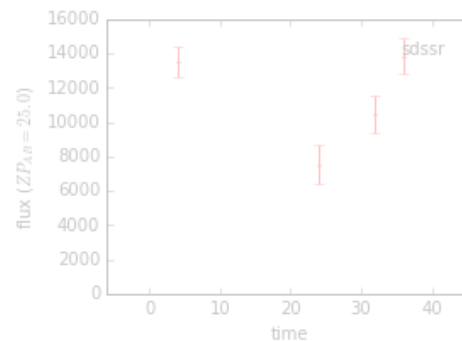
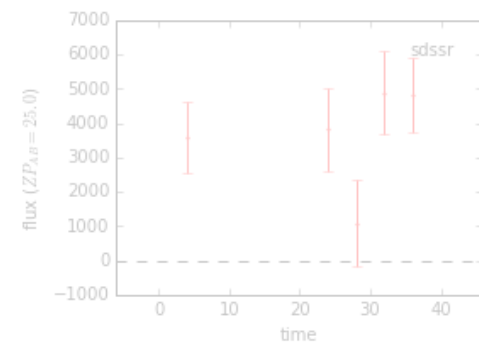
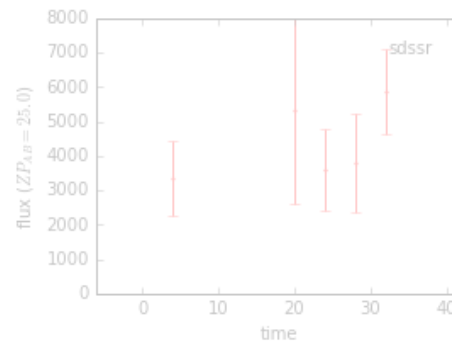
# Image Difference on Stack: Light curves

- ▶ Using Sncosmo library. We traced the light curves for all the detections.
- ▶ Light curves with less than 3 points were rejected.
- ▶ DIASources tagged as dipoles were dismissed.
- ▶ Candidates went from 343 to 8 (on the 6 images). The supernova is among these sources.



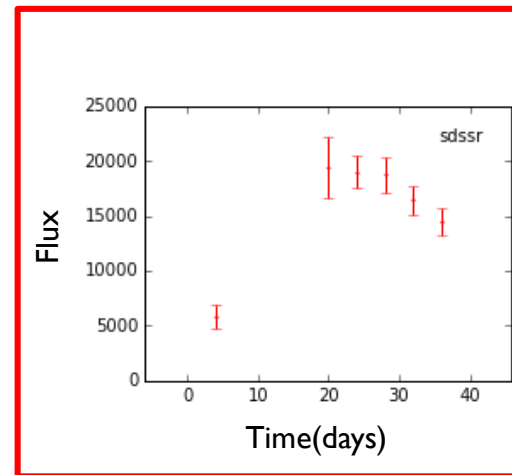
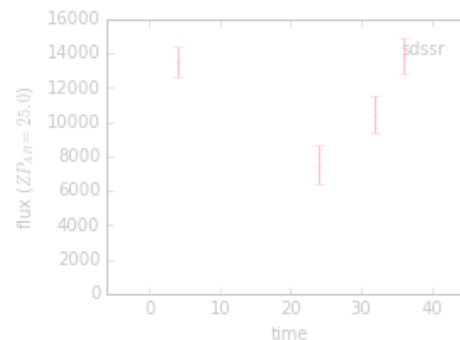
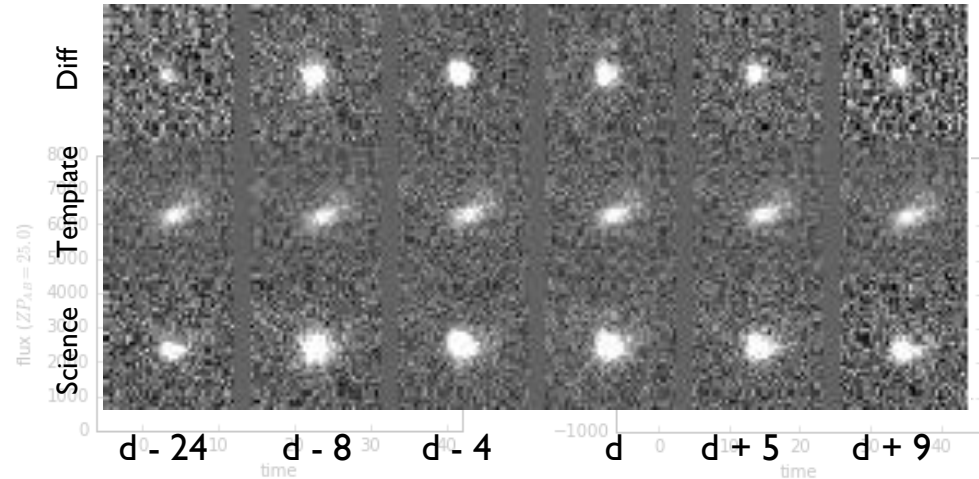
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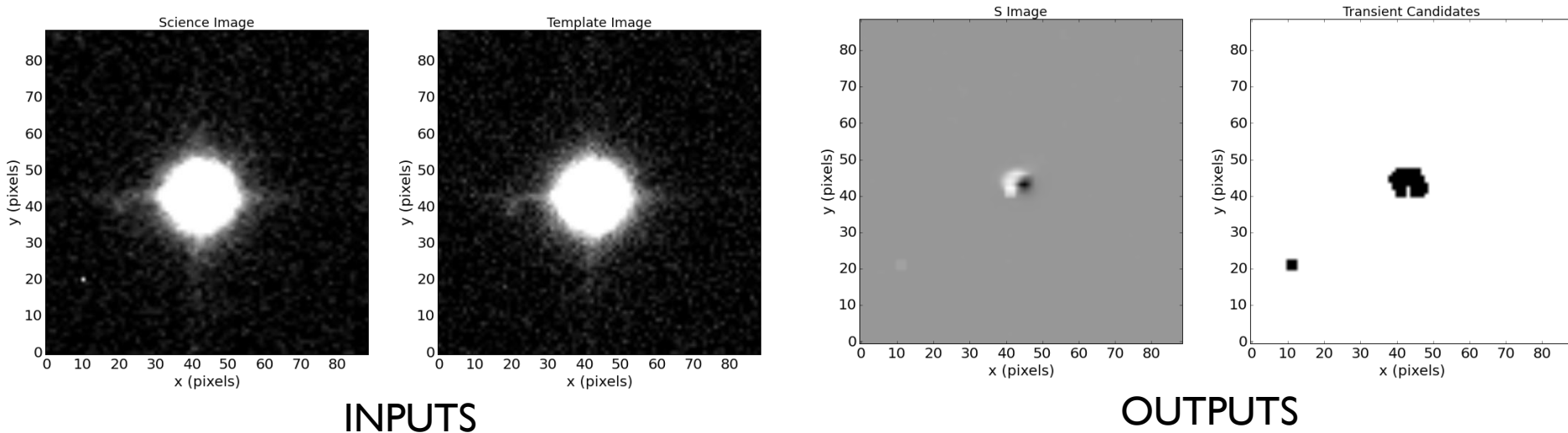
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# Other transient detection methods:

## Zackay-Ofek optimal transient detection<sup>2</sup>

- ▶ First implementation outside stack, using exclusively Python libraries.
- ▶ Zackay-Ofek method uses an optimal statistic approach for source detection (likelihood ratio test hypothesis: no source vs new source).



<sup>2</sup> Zackay, B., Ofek, E. and Gal-Yam, A. (2016). *Proper image subtraction - optimal transient detection, photometry and hypothesis testing*

# Conclusions

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- ▶ **On Alard-Lupton Image Subtraction method:**
  - ▶ Several transient objects were detected on the first test set and the Supernova was among them.
  - ▶ Characterization of sources detected on the difference image allows the creation of categories for further study.
  - ▶ It requires several tweaks to improve current results (including flexible testing).
- ▶ **On Zackay-Ofek method (viability and efficiency) are currently a focus of research.**
  - ▶ Benchmarking and comparison with Alard-Lupton algorithm are currently under investigation.

# Perspectives

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- ▶ On Alard-Lupton Image Subtraction method:
  - ▶ Improve selection criteria by using flux characteristics.
  - ▶ Implementation of automatic classification.
  - ▶ Study with more a priori transients.
- ▶ On Zackay-Ofek method:
  - ▶ Implementation on Stack.