

# Recent anomalies in flavour physics

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Electroweak Physics and the LHC**

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- **Introduction**

- Theoretical framework

- **Observables**

- Recent anomalies

- **Implications**

- Model independent global fits

- Implications for Wilson coefficients

- Assessment of the theoretical uncertainties

- Identifying the origin of the anomalies

- Model dependent interpretations

- **Conclusions**

## Indirect search for New Physics

- In the past, the objective of flavour physics experiments was focused on the tests of the Standard Model and the CKM paradigm  
→ This is now well established!
- Focus is now towards the new physics!  
And search for the indirect signs of new physics

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Main probes: Rare decays

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Main probes: **Rare decays**

## Why rare decays are interesting?

- Rare  $B$  decays occur at loop level
    - The SM contributions are very small and the NP contributions can have a comparable magnitude.
  - The theory ingredients are known at a very good accuracy!  
In particular: QCD corrections are known with a good precision!
  - Very promising experimental situation
    - Branching ratios and distributions can be measured precisely
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## Theoretical framework

### A multi-scale problem

- new physics:  $1/\Lambda_{\text{NP}}$
- electroweak interactions:  $1/M_W$
- hadronic effects:  $1/m_b$
- QCD interactions:  $1/\Lambda_{\text{QCD}}$

⇒ Effective field theory approach:

separation between low and high energies using Operator Product Expansion

- short distance: Wilson coefficients, computed perturbatively
- long distance: local operators

$$\mathcal{H}_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \left( \sum_{i=1 \dots 10, S, P} (C_i(\mu) \mathcal{O}_i(\mu) + C'_i(\mu) \mathcal{O}'_i(\mu)) \right)$$

New physics:

- Corrections to the Wilson coefficients:  $C_i \rightarrow C_i + \Delta C_i^{\text{NP}}$
- Additional operators:  $\sum_j C_j^{\text{NP}} \mathcal{O}_j^{\text{NP}}$



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# Operators

$$\mathcal{O}_1 = (\bar{s}\gamma_\mu T^a P_L c)(\bar{c}\gamma^\mu T^a P_L b)$$

$$\mathcal{O}_2 = (\bar{s}\gamma_\mu P_L c)(\bar{c}\gamma^\mu P_L b)$$

$$\mathcal{O}_3 = (\bar{s}\gamma_\mu P_L b)\sum_q(\bar{q}\gamma^\mu q)$$

$$\mathcal{O}_4 = (\bar{s}\gamma_\mu T^a P_L b)\sum_q(\bar{q}\gamma^\mu T^a q)$$

$$\mathcal{O}_5 = (\bar{s}\gamma_{\mu_1}\gamma_{\mu_2}\gamma_{\mu_3} P_L b)\sum_q(\bar{q}\gamma^{\mu_1}\gamma^{\mu_2}\gamma^{\mu_3} q)$$

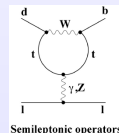
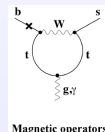
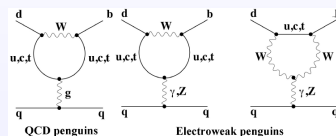
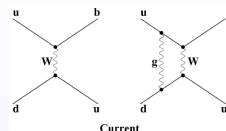
$$\mathcal{O}_6 = (\bar{s}\gamma_{\mu_1}\gamma_{\mu_2}\gamma_{\mu_3} T^a P_L b)\sum_q(\bar{q}\gamma^{\mu_1}\gamma^{\mu_2}\gamma^{\mu_3} T^a q)$$

$$\mathcal{O}_7 = \frac{e}{16\pi^2} \left[ \bar{s}\sigma^{\mu\nu} (m_s P_L + m_b P_R) b \right] F_{\mu\nu}$$

$$\mathcal{O}_8 = \frac{g}{16\pi^2} \left[ \bar{s}\sigma^{\mu\nu} (m_s P_L + m_b P_R) T^a b \right] G_{\mu\nu}^a$$

$$\mathcal{O}_9 = \frac{e^2}{(4\pi)^2} (\bar{s}\gamma^\mu b_L)(\bar{l}\gamma_\mu l)$$

$$\mathcal{O}_{10} = \frac{e^2}{(4\pi)^2} (\bar{s}\gamma^\mu b_L)(\bar{l}\gamma_\mu \gamma_5 l)$$



# Wilson coefficients

## Two main steps:

- Calculating  $C_i^{\text{eff}}(\mu)$  at scale  $\mu \sim M_W$  by requiring matching between the effective and full theories

$$C_i^{\text{eff}}(\mu) = C_i^{(0)\text{eff}}(\mu) + \frac{\alpha_s(\mu)}{4\pi} C_i^{(1)\text{eff}}(\mu) + \dots$$

- Evolving the  $C_i^{\text{eff}}(\mu)$  to scale  $\mu \sim m_b$  using the RGE:

$$\mu \frac{d}{d\mu} C_i^{\text{eff}}(\mu) = C_j^{\text{eff}}(\mu) \gamma_{ji}^{\text{eff}}(\mu)$$

driven by the anomalous dimension matrix  $\hat{\gamma}^{\text{eff}}(\mu)$ :

$$\hat{\gamma}^{\text{eff}}(\mu) = \frac{\alpha_s(\mu)}{4\pi} \hat{\gamma}^{(0)\text{eff}} + \frac{\alpha_s^2(\mu)}{(4\pi)^2} \hat{\gamma}^{(1)\text{eff}} + \dots$$

# Hadronic quantities

To compute the amplitudes:

$$\mathcal{A}(A \rightarrow B) = \langle B | \mathcal{H}_{\text{eff}} | A \rangle = \frac{G_F}{\sqrt{2}} \sum_i \lambda_i C_i(\mu) \langle B | \mathcal{O}_i | A \rangle(\mu)$$

$\langle B | \mathcal{O}_i | A \rangle$ : hadronic matrix element

How to compute matrix elements?

→ Model building, Lattice simulations, Light flavour symmetries,  
Heavy flavour symmetries, ...

→ Describe hadronic matrix elements in terms of **hadronic quantities**

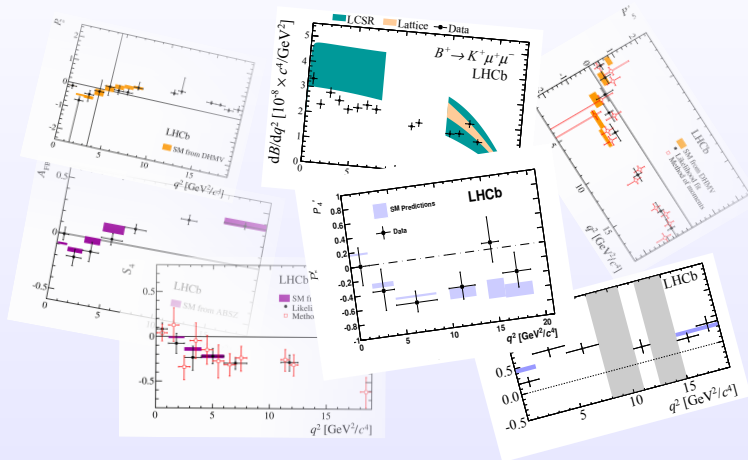
Two types of hadronic quantities:

- **Decay constants**: Probability amplitude of hadronising quark pair into a given hadron
- **Form factors**: Transition from a meson to another through flavour change

## LHCb measurements

Impressive effort in studying **exclusive**  $b \rightarrow s \ell \ell$  transitions at LHCb with the measurements of a large number of independent angular observables!

$B \rightarrow K \mu^+ \mu^-$ ,  $B \rightarrow K^+ e^+ e^-$ ,  $B \rightarrow K^* \mu^+ \mu^-$  ( $F_L$ ,  $A_{FB}$ ,  $S_i$ ,  $P_i$ ),  $B_s \rightarrow \phi \mu^+ \mu^-$ , ...

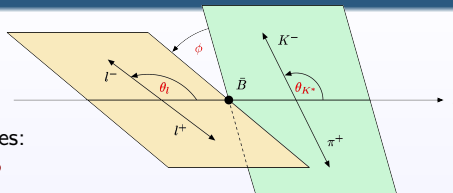


Deviations from the SM predictions in  $B \rightarrow K^* \mu^+ \mu^-$ ,  $B_s \rightarrow \phi \mu^+ \mu^-$  and  $R_K$ : “anomalies”

# $B \rightarrow K^* \mu^+ \mu^-$ – Angular distributions

## Angular distributions

The full angular distribution of the decay  $\bar{B}^0 \rightarrow \bar{K}^{*0} \ell^+ \ell^-$  ( $\bar{K}^{*0} \rightarrow K^- \pi^+$ ) is completely described by four independent kinematic variables:  $q^2$  (dilepton invariant mass squared),  $\theta_\ell$ ,  $\theta_{K^*}$ ,  $\phi$



Differential decay distribution:

$$\frac{d^4 \Gamma}{dq^2 d \cos \theta_\ell d \cos \theta_{K^*} d \phi} = \frac{9}{32\pi} J(q^2, \theta_\ell, \theta_{K^*}, \phi)$$

$$J(q^2, \theta_\ell, \theta_{K^*}, \phi) = \sum_i J_i(q^2) f_i(\theta_\ell, \theta_{K^*}, \phi)$$

↘ angular coefficients  $J_{1-9}$

↘ functions of the spin amplitudes  $A_0$ ,  $A_{\parallel}$ ,  $A_{\perp}$ ,  $A_t$ , and  $A_S$

Spin amplitudes: functions of Wilson coefficients and form factors

Main operators:

$$\mathcal{O}_9 = \frac{e^2}{(4\pi)^2} (\bar{s} \gamma^\mu b_L) (\bar{\ell} \gamma_\mu \ell), \quad \mathcal{O}_{10} = \frac{e^2}{(4\pi)^2} (\bar{s} \gamma^\mu b_L) (\bar{\ell} \gamma_\mu \gamma_5 \ell)$$

$$\mathcal{O}_S = \frac{e^2}{16\pi^2} (\bar{s}_L^\alpha b_R^\alpha) (\bar{\ell} \ell), \quad \mathcal{O}_P = \frac{e^2}{16\pi^2} (\bar{s}_L^\alpha b_R^\alpha) (\bar{\ell} \gamma_5 \ell)$$

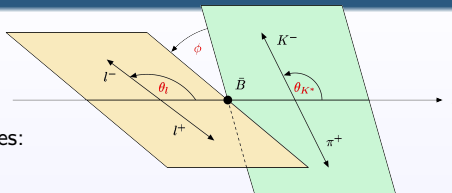
F. Kruger et al., Phys. Rev. D 61 (2000) 114028;

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## Differential decay distribution:

$$\frac{d^4\Gamma}{dq^2 d\cos\theta_\ell d\cos\theta_{K^*} d\phi} = \frac{9}{32\pi} J(q^2, \theta_\ell, \theta_{K^*}, \phi)$$

$$J(q^2, \theta_\ell, \theta_{K^*}, \phi) = \sum_i J_i(q^2) f_i(\theta_\ell, \theta_{K^*}, \phi)$$

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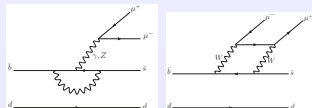
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# $B \rightarrow K^* \mu^+ \mu^-$ observables

**Optimised observables:** form factor uncertainties cancel at leading order

$$\langle P_1 \rangle_{\text{bin}} = \frac{1}{2} \frac{\int_{\text{bin}} dq^2 [J_3 + \bar{J}_3]}{\int_{\text{bin}} dq^2 [J_{2s} + \bar{J}_{2s}]}$$

$$\langle P_2 \rangle_{\text{bin}} = \frac{1}{8} \frac{\int_{\text{bin}} dq^2 [J_{6s} + \bar{J}_{6s}]}{\int_{\text{bin}} dq^2 [J_{2s} + \bar{J}_{2s}]}$$

$$\langle P'_4 \rangle_{\text{bin}} = \frac{1}{\mathcal{N}'_{\text{bin}}} \int_{\text{bin}} dq^2 [J_4 + \bar{J}_4]$$

$$\langle P'_5 \rangle_{\text{bin}} = \frac{1}{2\mathcal{N}'_{\text{bin}}} \int_{\text{bin}} dq^2 [J_5 + \bar{J}_5]$$

$$\langle P'_6 \rangle_{\text{bin}} = \frac{-1}{2\mathcal{N}'_{\text{bin}}} \int_{\text{bin}} dq^2 [J_7 + \bar{J}_7]$$

$$\langle P'_8 \rangle_{\text{bin}} = \frac{-1}{\mathcal{N}'_{\text{bin}}} \int_{\text{bin}} dq^2 [J_8 + \bar{J}_8]$$

with

$$\mathcal{N}'_{\text{bin}} = \sqrt{-\int_{\text{bin}} dq^2 [J_{2s} + \bar{J}_{2s}] \int_{\text{bin}} dq^2 [J_{2c} + \bar{J}_{2c}]}$$

+ CP violating clean observables and other combinations

U. Egede et al., JHEP 0811 (2008) 032, JHEP 1010 (2010) 056

J. Matias et al., JHEP 1204 (2012) 104

S. Descotes-Genon et al., JHEP 1305 (2013) 137

**Or alternatively:**

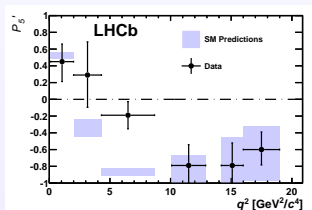
$$S_i = \frac{J_{i(s,c)} + \bar{J}_{i(s,c)}}{\frac{d\Gamma}{dq^2} + \frac{d\bar{\Gamma}}{dq^2}},$$

$$P'_{4,5,8} = \frac{S_{4,5,8}}{\sqrt{F_L(1 - F_L)}}$$

# The LHCb anomalies (1)

$B \rightarrow K^* \mu^+ \mu^-$  angular observables, in particular  $P'_5 / S_5$

- 2013 (1 fb<sup>-1</sup>): disagreement with the SM for  $P_2$  and  $P'_5$  (PRL 111, 191801 (2013))
- March 2015 (3 fb<sup>-1</sup>): confirmation of the deviations (LHCb-CONF-2015-002)
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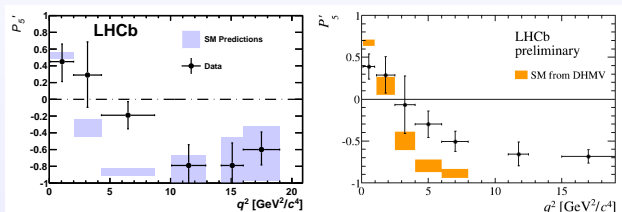


3.7 $\sigma$  deviation in the 3rd bin

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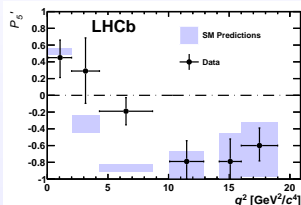
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2.9 $\sigma$  in the 4th and 5th bins  
(3.7 $\sigma$  combined)

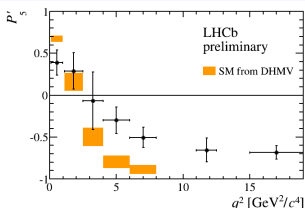
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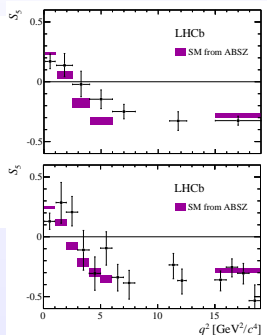
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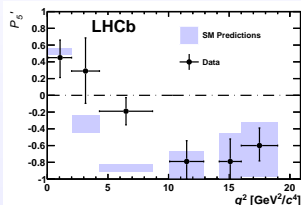


3.4 $\sigma$  combined fit (likelihood)

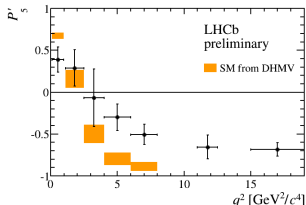
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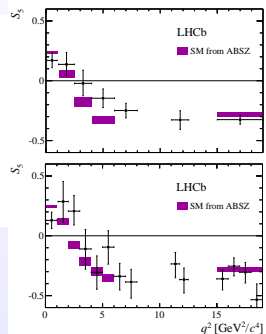
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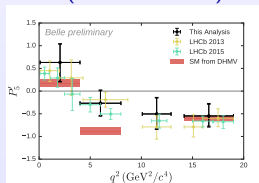


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3.4 $\sigma$  combined fit (likelihood)

Belle supports LHCb  
(arXiv:1604.04042)  
tension at 2.1 $\sigma$

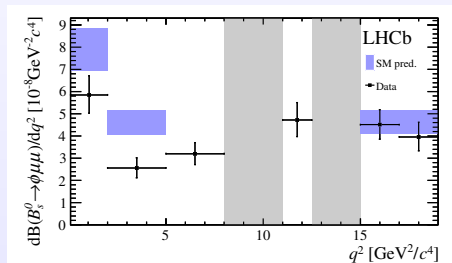


## The LHCb anomalies (2)

### $B_s \rightarrow \phi \mu^+ \mu^-$ branching fraction

- Same theoretical description as  $B \rightarrow K^* \mu^+ \mu^-$
- June 2015 ( $3 \text{ fb}^{-1}$ ): the differential branching fraction is found to be  $3.2\sigma$  below the SM predictions in the  $[1-6] \text{ GeV}^2$  bin

JHEP 1509 (2015) 179



## The LHCb anomalies (3)

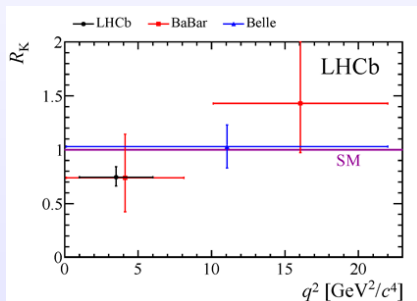
### Lepton universality in $B^+ \rightarrow K^+ \ell^+ \ell^-$

- June 2015 (3 fb<sup>-1</sup>): measurement of  $R_K$  in the [1-6] GeV<sup>2</sup> bin  
2.6 $\sigma$  tension in [1-6] GeV<sup>2</sup> bin

PRL 113, 151601 (2014)

$$R_K = BR(B^+ \rightarrow K^+ \mu^+ \mu^-) / BR(B^+ \rightarrow K^+ e^+ e^-)$$

$$R_K^{\text{SM}} \approx 1$$



$$R_K^{\text{exp}} = 0.745^{+0.090}_{-0.074}(\text{stat}) \pm 0.036(\text{syst})$$

## New Physics interpretation?

Many observables → **Global fits** of the latest LHCb data

Relevant Operators:

$$\mathcal{O}_7, \mathcal{O}_8, \mathcal{O}_{9\mu,e}^{(\prime)}, \mathcal{O}_{10\mu,e}^{(\prime)} \quad \text{and} \quad \mathcal{O}_{S-P} \propto (\bar{s}P_R b)(\bar{\mu}P_L \mu) \equiv \mathcal{O}_0^I$$

NP manifests itself in the shifts of the individual coefficients with respect to the SM values:

$$C_i(\mu) = C_i^{\text{SM}}(\mu) + \delta C_i$$

- Scans over the values of  $\delta C_i$
- Calculation of flavour observables
- Comparison with experimental results
- Constraints on the Wilson coefficients  $C_i$

Global fits using the latest LHCb results:

M. Ciuchini, M. Fedele, E. Franco, S. Mishima, A. Paul, L. Silvestrini, M. Valli, 1512.07157  
 T. Hurth, FM, S. Neshatpour, 1603.00865  
 B. Capdevila, S. Descotes-Genon, J. Matias, J. Virto, 1605.03156



## Global fits

Global fits of the observables obtained by minimisation of

$$\chi^2 = (\vec{O}^{\text{th}} - \vec{O}^{\text{exp}}) \cdot (\Sigma_{\text{th}} + \Sigma_{\text{exp}})^{-1} \cdot (\vec{O}^{\text{th}} - \vec{O}^{\text{exp}})$$

$(\Sigma_{\text{th}} + \Sigma_{\text{exp}})^{-1}$  is the inverse covariance matrix.

More than 100 observables relevant for leptonic and semileptonic decays:

T. Hurth, FM, S. Neshatpour, 1603.00865

- $\text{BR}(B \rightarrow X_s \gamma)$
- $\text{BR}(B \rightarrow X_d \gamma)$
- $\Delta_0(B \rightarrow K^* \gamma)$
- $\text{BR}^{\text{low}}(B \rightarrow X_s \mu^+ \mu^-)$
- $\text{BR}^{\text{high}}(B \rightarrow X_s \mu^+ \mu^-)$
- $\text{BR}^{\text{low}}(B \rightarrow X_s e^+ e^-)$
- $\text{BR}^{\text{high}}(B \rightarrow X_s e^+ e^-)$
- $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$
- $\text{BR}(B_d \rightarrow \mu^+ \mu^-)$
- $\text{BR}(B \rightarrow K^0 \mu^+ \mu^-)$
- $\text{BR}(B \rightarrow K^{*+} \mu^+ \mu^-)$
- $\text{BR}(B \rightarrow K^+ \mu^+ \mu^-)$
- $\text{BR}(B \rightarrow K^* e^+ e^-)$
- $R_K$
- $B \rightarrow K^{*0} \mu^+ \mu^-$ :  $\text{BR}, F_L, A_{FB}, S_3, S_4, S_5, S_7, S_8, S_9$   
in 8 low  $q^2$  and 4 high  $q^2$  bins
- $B_s \rightarrow \phi \mu^+ \mu^-$ :  $\text{BR}, F_L, S_3, S_4, S_7$   
in 3 low  $q^2$  and 2 high  $q^2$  bins

Computations performed using SuperIso public program

## Theoretical uncertainties and correlations (1)

- Monte Carlo analysis
- variation of the “standard” input parameters: masses, scales, CKM, ...
- decay constants taken from the latest lattice results
- use for the  $B_{(s)} \rightarrow V$  form factors of the lattice+LCSR combinations from 1503.05534, including correlations (Cholesky decomposition method)
- use for the  $B \rightarrow K$  form factors of the lattice+LCSR combinations from 1411.3161, including correlations
- for  $B_s \rightarrow \phi \mu^+ \mu^-$ , mixing effects taken into account

## Global fits

### Theoretical uncertainties and correlations (2)

- Two approaches for calculation of the exclusive decays:  
**full** form factors (7 independent FFs), **soft** form factors (2 independent FFs)
- full FF approach = soft FF approach +  $(1/m_b, \alpha_s)$  “**factorisable corrections**”
- Both approaches receive contributions from non-local 4 quark operators:  
“**non-factorisable corrections**”
- However, non-factorisable corrections only known at LO in  $(\Lambda/m_b, E_{K^*}/m_b)$
- Higher powers of  $(\Lambda/m_b, E_{K^*}/m_b)$  unknown = “**non-factorisation power corrections**”
- Evaluation of uncertainties from factorisable and non-factorisable power corrections:

$$A_k \rightarrow A_k \left( 1 + a_k \exp(i\phi_k) + \frac{q^2}{6 \text{ GeV}^2} b_k \exp(i\theta_k) \right)$$

Soft: parametrisation of both factorisable and non-factorisable power corrections

Full: parametrisation of only non-factorisable power corrections

$|a_k|$  between 10 to 60%,  $b_k \sim 2.5a_k$

Low recoil:  $b_k = 0$

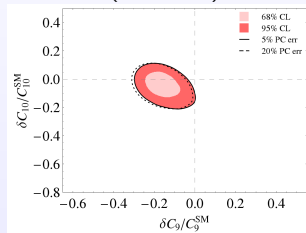
$\Rightarrow$  Computation of a (theory + exp) correlation matrix

## Fit results for two operators

### Fit results using full form factor approach:

- filled areas: 10% power correction errors
- solid line: 5% power correction errors
- dashed line: 20% power correction errors

$$(C_9 - C_{10})$$



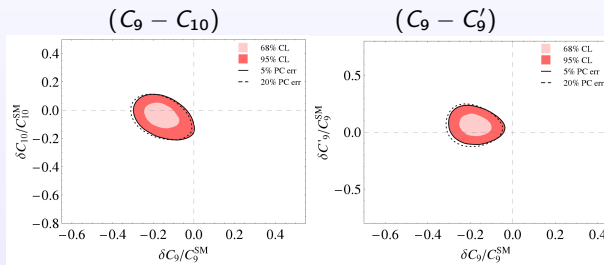
$$(C_9 - C_9')$$

$$(C_9^e - C_9^\mu)$$

## Fit results for two operators

### Fit results using full form factor approach:

- filled areas: 10% power correction errors
- solid line: 5% power correction errors
- dashed line: 20% power correction errors

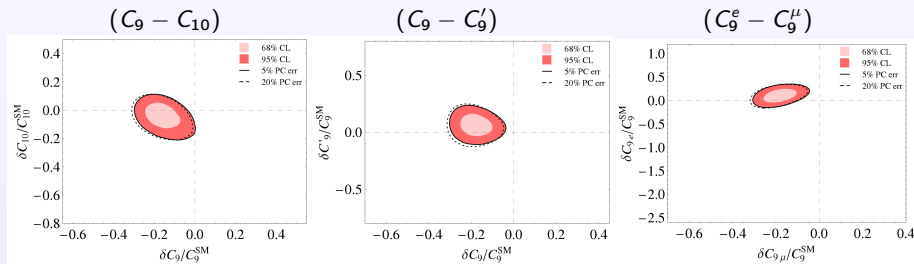


$$(C_9^e - C_9^\mu)$$

## Fit results for two operators

### Fit results using full form factor approach:

- filled areas: 10% power correction errors
- solid line: 5% power correction errors
- dashed line: 20% power correction errors



**About  $3\sigma$  deviations with the SM in all cases**

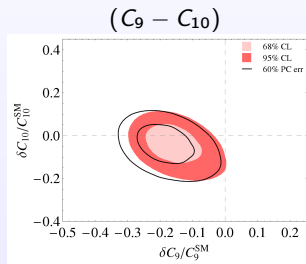
Power correction uncertainty between 5 and 20% does not change the picture.

Results using soft form factors are very similar

# Fit results for two operators: effect of power corrections

## Fits assuming different power correction uncertainties:

- 10% uncertainty (filled areas)
- 60% uncertainty (solid line)



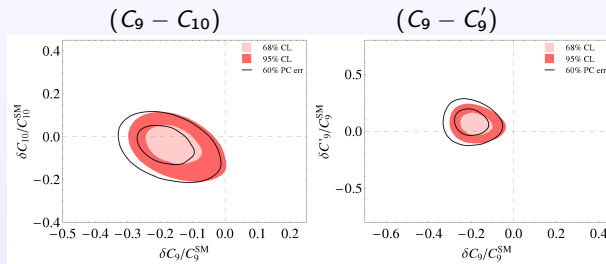
$$(C_9 - C'_9)$$

$$(C_9^e - C_9^\mu)$$

# Fit results for two operators: effect of power corrections

## Fits assuming different power correction uncertainties:

- 10% uncertainty (filled areas)
- 60% uncertainty (solid line)



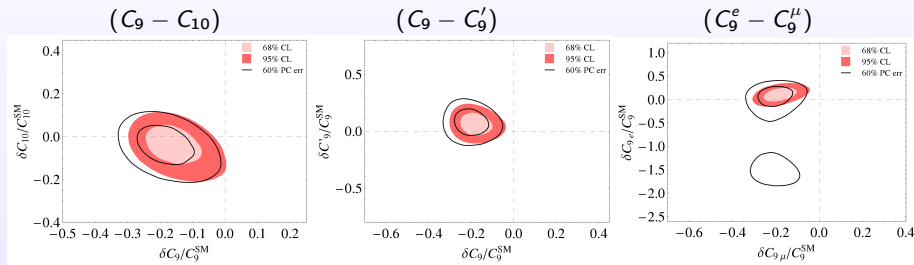
$$(C_9^e - C_9^\mu)$$



## Fit results for two operators: effect of power corrections

### Fits assuming different power correction uncertainties:

- 10% uncertainty (filled areas)
- 60% uncertainty (solid line)



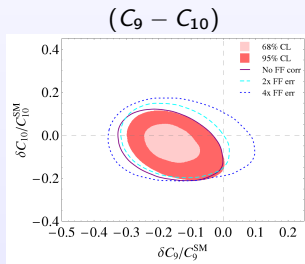
**Not a huge impact!**

60% power correction uncertainty leads to only 17-20% error at the observable level.

## Fit results for two operators: form factor dependence

### Fits with different assumptions for the form factor uncertainties:

- correlations ignored (solid line)
- normal form factor errors (filled areas)
- $2 \times$  form factor errors (dashed line)
- $4 \times$  form factor errors (dotted line)



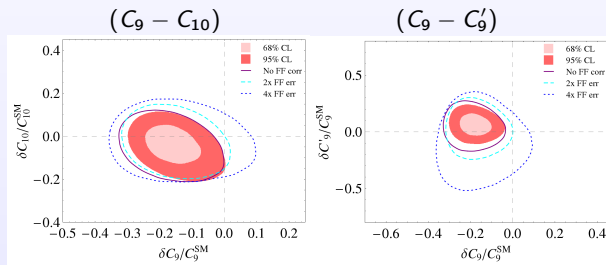
$$(C_9 - C'_9)$$

$$(C_9^e - C_9^\mu)$$

## Fit results for two operators: form factor dependence

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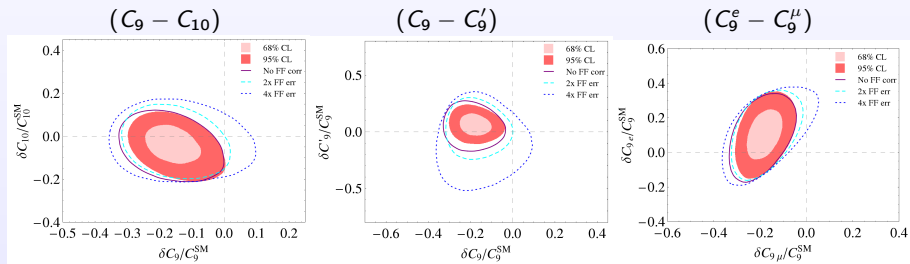


$$(C_9^e - C_9^\mu)$$

## Fit results for two operators: form factor dependence

### Fits with different assumptions for the form factor uncertainties:

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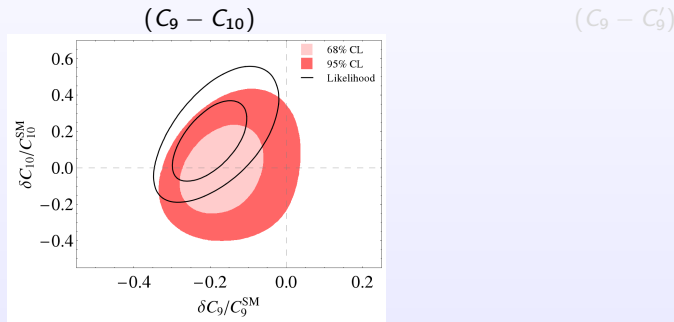
**The size of the form factor errors has a crucial role in constraining the allowed region!**

## Fit results for two operators: likelihood vs. method of moments

LHCb presented the  $B \rightarrow K^* \mu^+ \mu^-$  angular analysis with two different methods:

- likelihood fits: smaller uncertainties, but involves model-dependent assumptions
- method of moments: more robust, but larger uncertainties

How does the choice of method affect fits? Let's consider only  $B \rightarrow K^* \mu^+ \mu^-$  measurements.



likelihood fits: solid lines

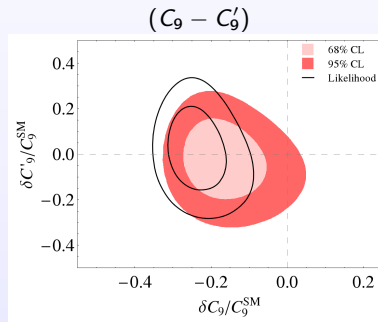
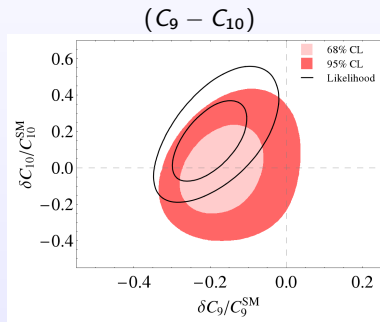
method of moments: filled areas

## Fit results for two operators: likelihood vs. method of moments

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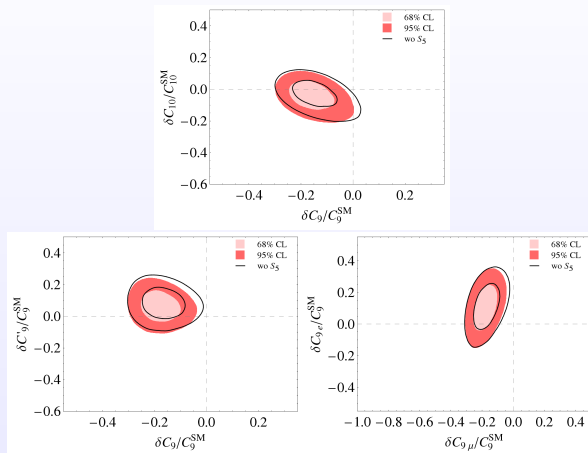
likelihood fits: solid lines

method of moments: filled areas

**Tension decreases using the method of moments results!**

## Role of $S_5$

Removing  $S_5$  from the fit:

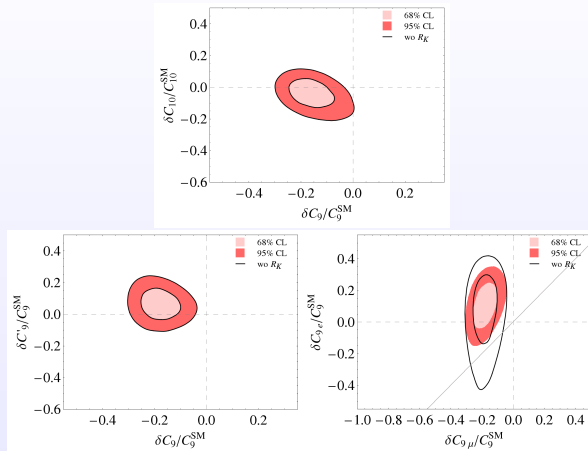


While the tension of  $C_9^{\text{SM}}$  and best fit point value of  $C_9$  is slightly reduced in the various two operator fits, still the tension exists at more than  $2\sigma$

→  $S_5$  is not the only observable which drives  $C_9$  to negative values!

## Role of $R_K$

### Removing $R_K$ from the fit:

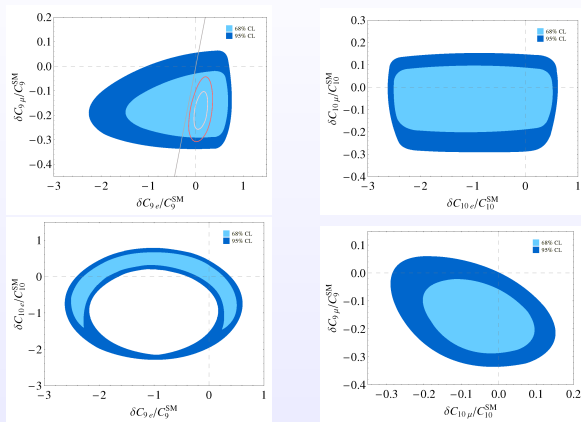


$R_K$  is the main measurement resulting in the best fit values for  $C_9^\mu$  and  $C_9^e$  which are in more than  $2\sigma$  tension with lepton-universality



# Fit results for four operators: $\{C_9^\mu, C_9^e, C_{10}^\mu, C_{10}^e\}$

No reason that only 2 Wilson coefficients receive contributions from new physics



**Larger ranges are allowed for the Wilson coefficients**

Considering 4 operator fits considerably relaxes the constraints on the Wilson coefficients leaving room for more diverse new physics contributions which are otherwise overlooked.

## Strategies to identify the origin of the observed anomalies (1)

### Unknown power corrections

- Significance of the anomalies depends on the assumptions on the power corrections
- Anomalies can be explained with 20-50% non-factorisable power corrections at the observable level in the critical bins (Ciuchini et al. 1512.07157)
- This corresponds to more than 100% error at the amplitude level (for  $S_3$ ,  $S_4$  and  $S_5$ )!
- Towards a calculations...

*"Any reasonable calculation is better than a fit!"* – T. Hurth

- Problem: they are not calculable in QCD factorisation
- Alternative approaches exist based on light cone sum rule techniques (Khodjamirian et al. 1006.4945)  
→ the available partial calculation increases the tension in  $P'_5$

## Strategies to identify the origin of the observed anomalies (2)

### Cross-check with other $R_{\mu/e}$ ratios

- $R_K$  is theoretically very clean compared to the angular observables
- Its tension cannot be explained by power corrections
- Both tensions could be explained by new physics in  $C_9^\mu$

Cross-checks needed with other ratios. Our predictions (within the  $\{C_9^\mu, C_9^e\}$  set):

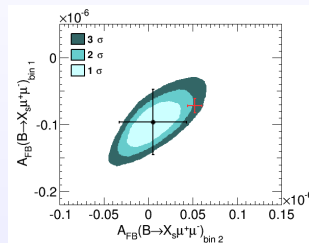
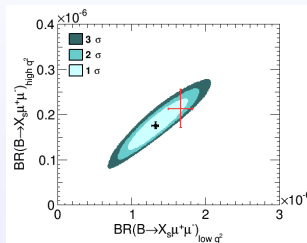
Observable	95% C.L. prediction
$\text{BR}(B \rightarrow X_s \mu^+ \mu^-) / \text{BR}(B \rightarrow X_s e^+ e^-)_{q^2 \in [1, 6] (\text{GeV})^2}$	[0.61, 0.93]
$\text{BR}(B \rightarrow X_s \mu^+ \mu^-) / \text{BR}(B \rightarrow X_s e^+ e^-)_{q^2 > 14.2 (\text{GeV})^2}$	[0.68, 1.13]
$\text{BR}(B^0 \rightarrow K^{*0} \mu^+ \mu^-) / \text{BR}(B^0 \rightarrow K^{*0} e^+ e^-)_{q^2 \in [1, 6] (\text{GeV})^2}$	[0.65, 0.96]
$\langle F_L(B^0 \rightarrow K^{*0} \mu^+ \mu^-) \rangle / \langle F_L(B^0 \rightarrow K^{*0} e^+ e^-) \rangle_{q^2 \in [1, 6] (\text{GeV})^2}$	[0.85, 0.96]
$\langle A_{FB}(B^0 \rightarrow K^{*0} \mu^+ \mu^-) \rangle / \langle A_{FB}(B^0 \rightarrow K^{*0} e^+ e^-) \rangle_{q^2 \in [4, 6] (\text{GeV})^2}$	[-0.21, 0.71]
$\langle S_5(B^0 \rightarrow K^{*0} \mu^+ \mu^-) \rangle / \langle S_5(B^0 \rightarrow K^{*0} e^+ e^-) \rangle_{q^2 \in [4, 6] (\text{GeV})^2}$	[0.53, 0.92]
$\text{BR}(B^0 \rightarrow K^{*0} \mu^+ \mu^-) / \text{BR}(B^0 \rightarrow K^{*0} e^+ e^-)_{q^2 \in [15, 19] (\text{GeV})^2}$	[0.58, 0.95]
$\langle F_L(B^0 \rightarrow K^{*0} \mu^+ \mu^-) \rangle / \langle F_L(B^0 \rightarrow K^{*0} e^+ e^-) \rangle_{q^2 \in [15, 19] (\text{GeV})^2}$	[0.998, 0.999]
$\langle A_{FB}(B^0 \rightarrow K^{*0} \mu^+ \mu^-) \rangle / \langle A_{FB}(B^0 \rightarrow K^{*0} e^+ e^-) \rangle_{q^2 \in [15, 19] (\text{GeV})^2}$	[0.87, 1.01]
$\langle S_5(B^0 \rightarrow K^{*0} \mu^+ \mu^-) \rangle / \langle S_5(B^0 \rightarrow K^{*0} e^+ e^-) \rangle_{q^2 \in [15, 19] (\text{GeV})^2}$	[0.87, 1.01]
$\text{BR}(B^+ \rightarrow K^+ \mu^+ \mu^-) / \text{BR}(B^+ \rightarrow K^+ e^+ e^-)_{q^2 \in [1, 6] (\text{GeV})^2}$	[0.58, 0.95]
$\text{BR}(B^+ \rightarrow K^+ \mu^+ \mu^-) / \text{BR}(B^+ \rightarrow K^+ e^+ e^-)_{q^2 \in [15, 22] (\text{GeV})^2}$	[0.58, 0.95]

## Strategies to identify the origin of the observed anomalies (3)

### Cross-check with inclusive modes

Inclusive decays are theoretically cleaner (see e.g. T. Huber, T. Hurth, E. Lunghi, JHEP 1506 (2015) 176)

At Belle-II, for inclusive  $b \rightarrow sll$ :



T. Hurth, FM, JHEP 1404 (2014) 097

T. Hurth, FM, S. Neshatpour, JHEP 1412 (2014) 053

Predictions based on our model-independent analysis

black cross: future measurements at Belle-II assuming the best fit solution

red cross: SM predictions

→ Belle-II will check the NP interpretation with theoretically clean modes

## New physics scenarios

Global fits: New physics is likely to appear in  $C_9$ :

$$O_9 = \frac{e^2}{(4\pi)^2} (\bar{s}\gamma^\mu b_L)(\bar{\ell}\gamma_\mu \ell)$$

It can also affect  $C'_9$  and  $C_{10}$  in a much lesser extent.

However, difficult to generate  $\delta C_9 \sim -1$  at loop level...

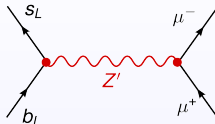
Need for tree level diagrams...

→ difficult in MSSM with MFV

### Mainstream scenarios:

- $Z'$  bosons
- leptoquarks
- composite models

## $Z'$ bosons



$Z'$  obvious candidate to generate the  $O_9$  operator

Needs:

- Flavour-changing couplings to left-handed quarks
- Vector-like couplings to leptons
- Flavour violation or non-universality in the lepton sector

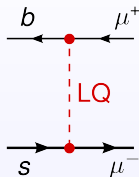
Strong constraints from  $B_s - \bar{B}_s$  mixing and LEP contact interactions.

Anomalies consistent with a  $Z'$  of 1 to 10 TeV

Can appear in many models, like 331 models, gauge  $L_\mu - L_\tau$  models, ...

See e.g. Altmannshofer et al. 1308.1501, Gauld et al. 1308.1959, Buras et al. 1309.2466, Gauld et al. 1310.1082, Buras et al. 1311.6729, Altmannshofer et al. 1403.1269, Buras et al. 1409.4557, Glashow et al. 1411.0565, Crivellin et al. 1501.00993, Altmannshofer et al. 1411.3161, Crivellin et al. 1503.03477, Niehoff et al. 1503.03865, Crivellin et al. 1505.02026, Celis et al. 1505.03079, ...

# Leptoquarks



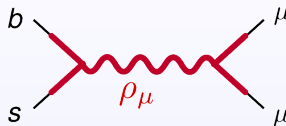
- t-channel diagrams
- Different possible representations, can be scalar (spin 0) or vector (spin 1)
- Cannot alter only  $C_9$ , but both  $C_9$  and  $C_{10}$  ( $= -C_9$ )
- Cannot be lepton flavour universal and conserve lepton number simultaneously

Model can be tested with  $R_{K^{(*)}}$  measurements and searches for  $b \rightarrow s\mu^{\pm}e^{\mp}$  and  $\mu \rightarrow e\gamma$

Possible scenario: two leptoquarks coupling to one lepton type only.

See e.g. Hiller et al. 1408.1627, Biswas et al. 1409.0882, Buras et al. 1409.4557, Sahoo et al. 1501.05193, Hiller et al. 1411.4773, Becirevic et al. 1503.09024, Alonso et al. 1505.05164, ...

## Composite models



- Neutral resonance  $\rho_\mu$  coupling to the muons via composite elementary mixing
- requires some compositeness for the muons
- can allow for lepton flavour violating couplings
- constrained by the LEP  $Z$ -width measurements and  $B_s - \bar{B}_s$  mixing

In addition, such models may explain the excesses observed in  $WW$ ,  $WZ$ ,  $Wh$  and  $\ell^+\ell^-$  resonance searches by ATLAS and CMS

See e.g. Gripaos et al. 1412.1791, Niehoff, et al. 1503.03865, Niehoff et al. 1508.00569, Carmona et al. 1510.07658, ...



# Conclusion

- Latest LHCb results, based on the  $3 \text{ fb}^{-1}$  data set still show some tensions with the SM predictions
- Model independent fits point to  $C_9^{NP} \sim -1$ , and new physics in muonic  $C_9^\mu$  is preferred
- In two operator fits there is more than  $2\sigma$  tension for  $\delta C_9^e = \delta C_9^\mu$
- The fit results do not depend very much on whether one uses soft or full form factor approach
- Factorisable power corrections have small effects at observable level
- The cross check with other not-yet-measured ratios (e.g.  $R_{K^*}$ ) and the inclusive measurements would be of importance to identify the origin of the anomalies

## Conclusion

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With the 750 GeV being gone, these are at the moment the only significant tensions with the SM at the LHC, so stay tuned! 🤔👉

## Backup

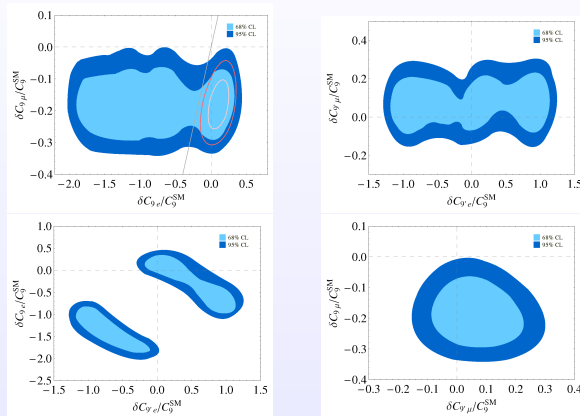
Backup

# Global fit results assuming new physics in one operator only

	b.f. value	$\chi^2_{\min}$	Pull <sub>SM</sub>	68% C.L.	95% C.L.
$\delta C_9/C_9^{\text{SM}}$	-0.18	123.8	$3.0\sigma$	$[-0.25, -0.09]$	$[-0.30, -0.03]$
$\delta C_9'/C_9^{\text{SM}}$	+0.03	131.9	$1.0\sigma$	$[-0.05, +0.12]$	$[-0.11, +0.18]$
$\delta C_{10}/C_{10}^{\text{SM}}$	-0.12	129.2	$1.9\sigma$	$[-0.23, -0.02]$	$[-0.31, +0.04]$
$\delta C_9^\mu/C_9^{\text{SM}}$	-0.21	115.5	$4.2\sigma$	$[-0.27, -0.13]$	$[-0.32, -0.08]$
$\delta C_9^e/C_9^{\text{SM}}$	+0.25	124.3	$2.9\sigma$	$[+0.11, +0.36]$	$[+0.03, +0.46]$

# Fit results for four operators: $\{C_9^\mu, C_9^{'\mu}, C_9^e, C_9^{'e}\}$

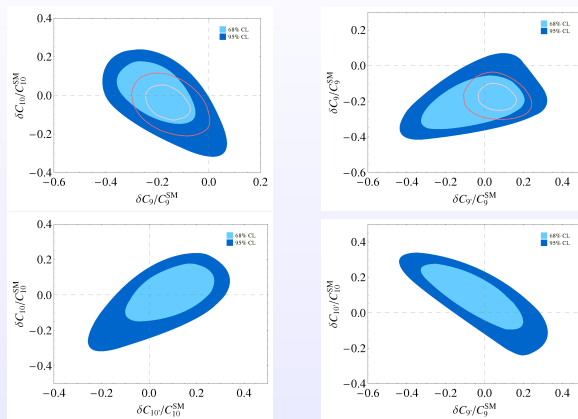
No reason that only 2 Wilson coefficients receive contributions from new physics



**Larger ranges are allowed for the Wilson coefficients**

# Fit results for four operators: $\{C_9, C'_9, C_{10}, C'_{10}\}$

No reason that only 2 Wilson coefficients receive contributions from new physics



Larger ranges are allowed for the Wilson coefficients