# Shapes and Symmetries in Nuclei Studied with GRETINA

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### **Outline**

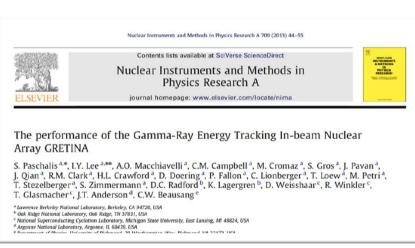
- GRETINA: A Brief Overview
- GRETINA Physics ⇒ Shapes and Symmetries
  - NSCL Campaign 1.0 2012 to 2013
    - "Bubble" Nucleus: 34Si
    - Rotational structure of <sup>32</sup>Mg
  - ANL Campaign 2014 to 2015
    - Octupole deformation in the Ba isotopes
- Summary

### **GRETINA**

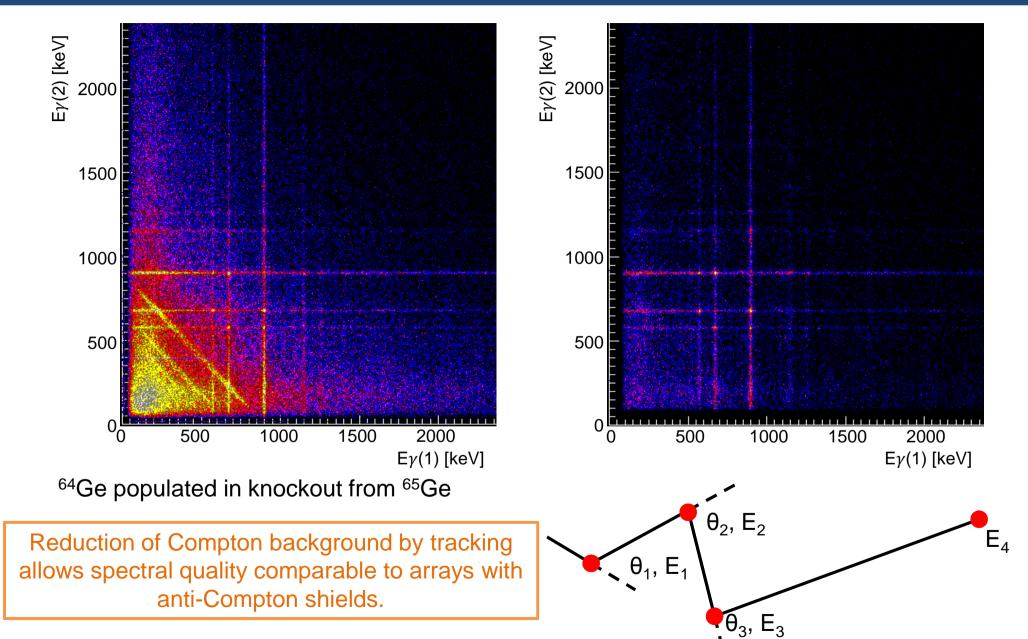
GRETINA, covering ¼ of the full solid angle, is the first realization of a γ-ray tracking array, the next generation of spectroscopic tools for the atomic nucleus.

- 28 x 36 fold segmented crystals housed in 7 (quad) modules
- Mechanical support structure, digital data acquisition and data processing software were all developed within the project
- Array represents significant advancements in all areas: detector technology,
   electronics and data
   processing

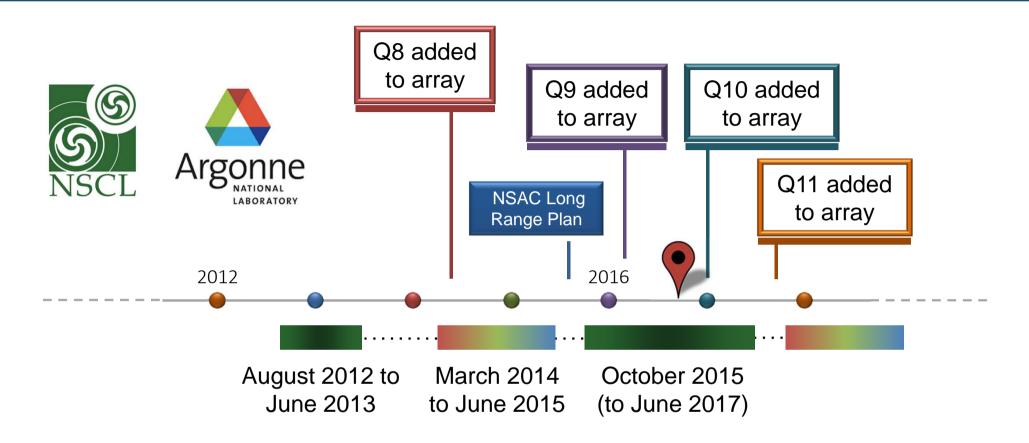




### Tracking: Compton Rejection

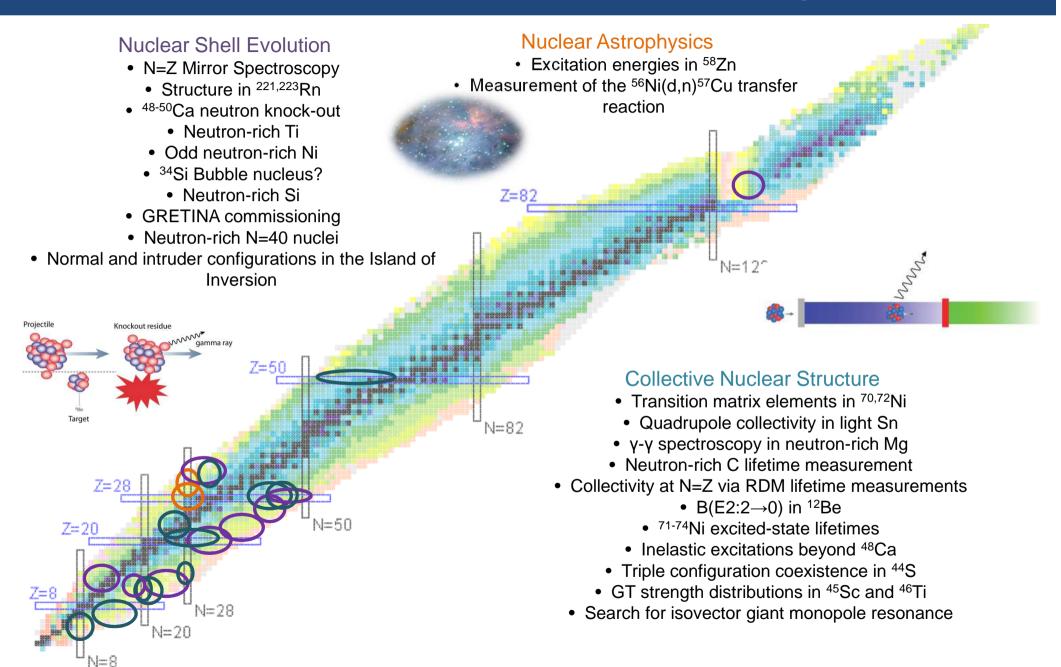


# GRETINA Physics Campaigns: Overview

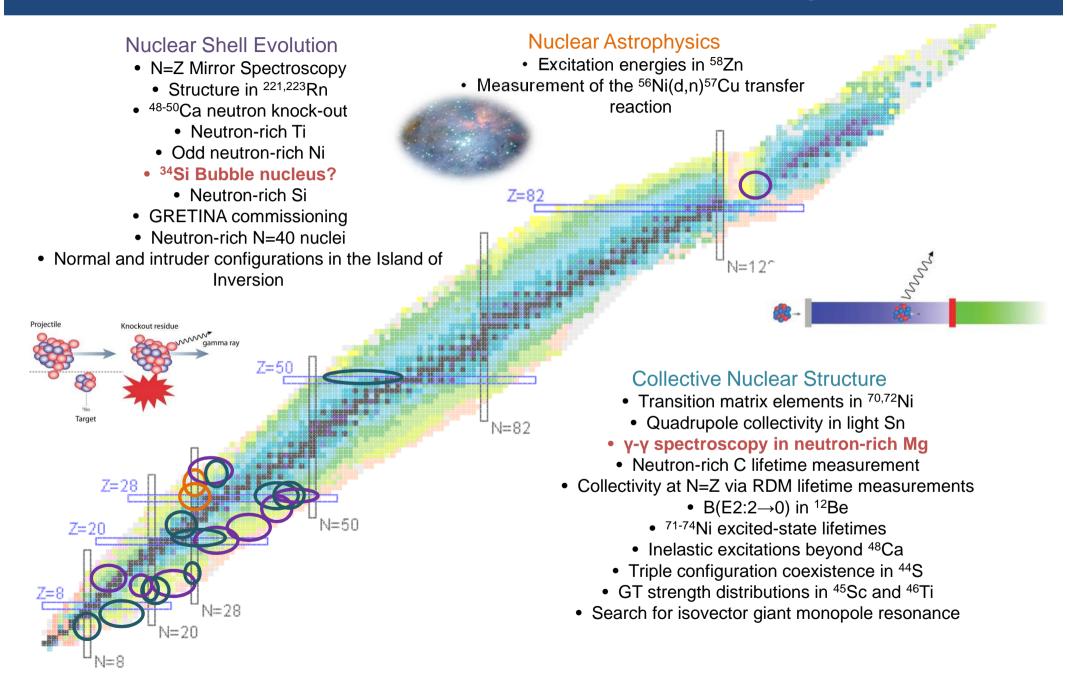


- First physics campaigns at NSCL and ANL (ATLAS/CARIBU) have already produced more than 25 physics papers
- Second NSCL campaign is ongoing, with 24 PAC approved experiments (~3600 hours)

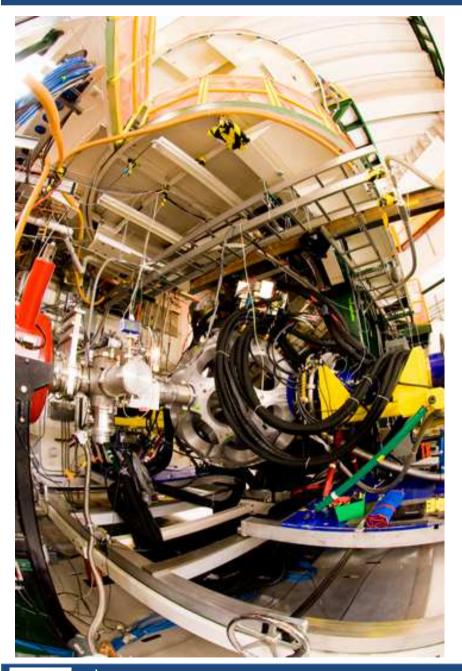
# NSCL GRETINA Physics Campaign 1.0

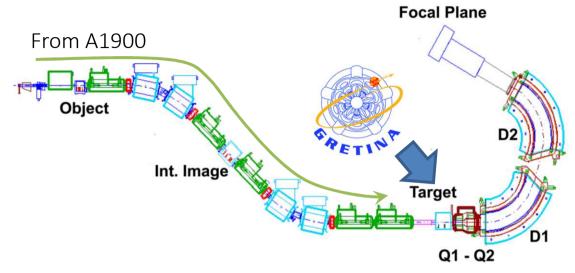


# NSCL GRETINA Physics Campaign 1.0



# First NSCL Physics Campaign (2012-2013)

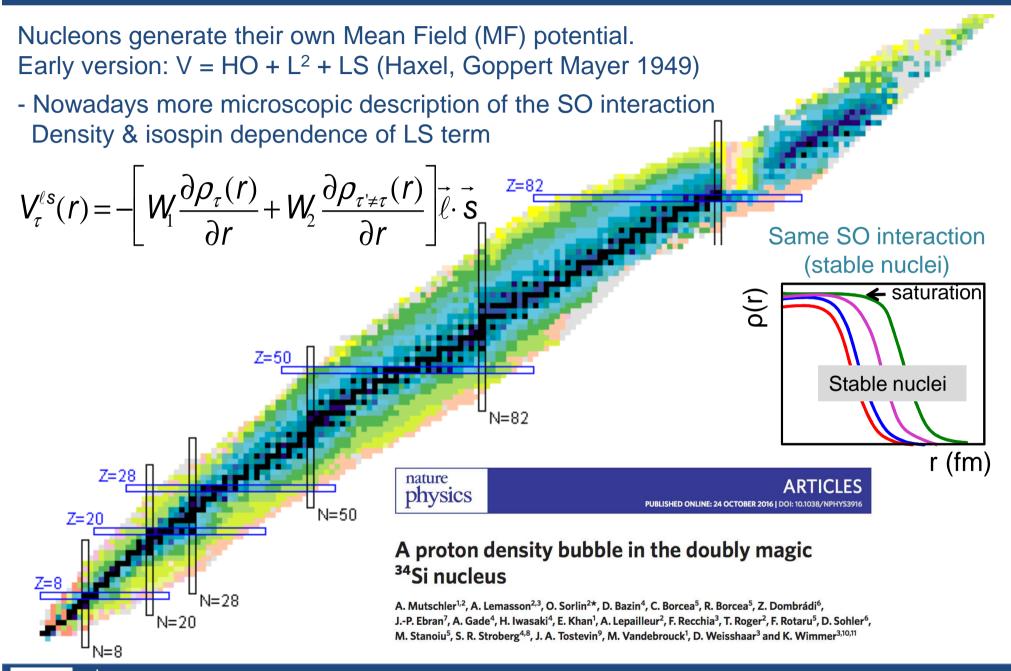




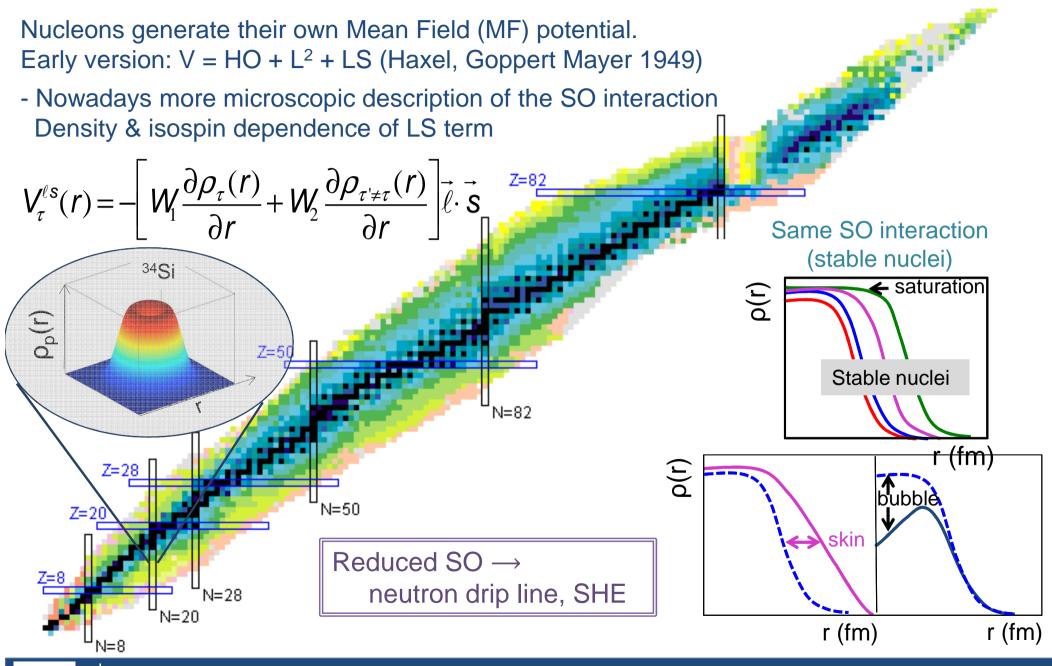
- First physics campaign coupled GRETINA with the S800 spectrograph for fast fragmentation beam (inbeam) spectroscopy
- Reaction channel selection is possible event-byevent in the S800

#### **Challenges:**

- High beam velocities (β ≈ 0.35) 2mm spatial resolution of GRETINA optimizes Doppler correction
- Low beam rates (as low as a few pps) efficiency of GRETINA, in singles and γγ is critical



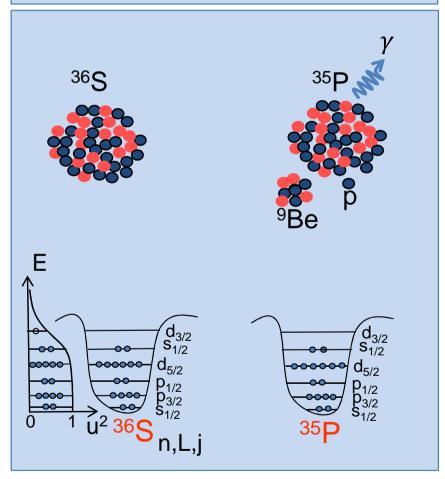
Nucleons generate their own Mean Field (MF) potential. Early version:  $V = HO + L^2 + LS$  (Haxel, Goppert Mayer 1949) - Nowadays more microscopic description of the SO interaction Density & isospin dependence of LS term  $V_{\tau}^{\ell s}(r) = - \left| W_{1} \frac{\partial \rho_{\tau}(r)}{\partial r} + W_{2} \frac{\partial \rho_{\tau' \neq \tau}(r)}{\partial r} \right| \vec{\ell} \cdot \vec{s}_{\square}$ Same SO interaction (stable nuclei) saturation Z = 50Stable nuclei N=82 <del>' (fm)</del> bubble N=50 Z = 20skin Reduced SO → neutron drip line, SHE N=28 N = 20r (fm) r (fm)

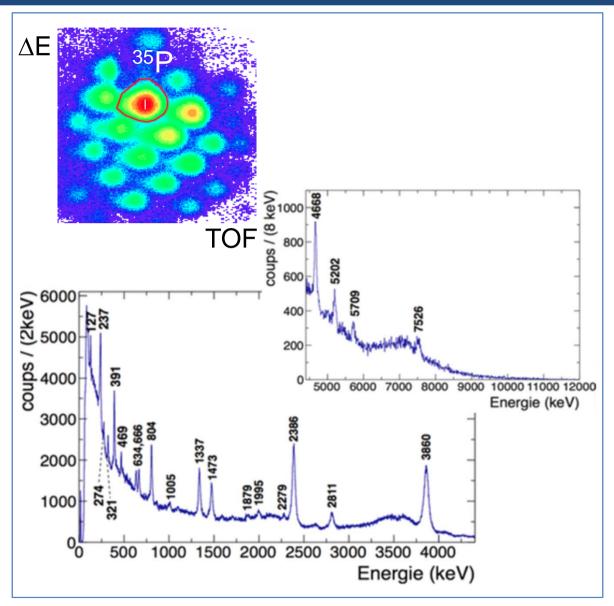


### Proton Densities in <sup>36</sup>S and <sup>34</sup>Si

Knock-out reactions at β≈0.4

$$\sigma_{\text{-1p}}(n,L) = C^2 S \left(j,n,L\right) \quad \sigma_{\text{sp}}(j,S_p) \ R_S \\ \text{occupancy} \qquad \text{reaction theory}$$



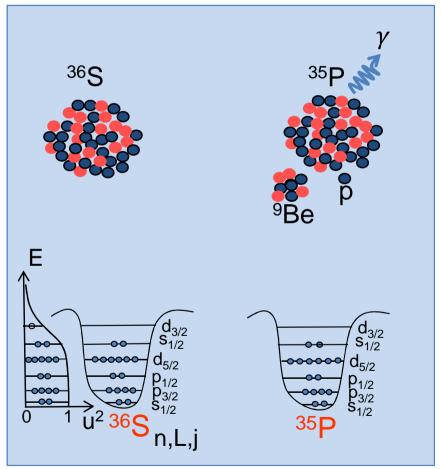


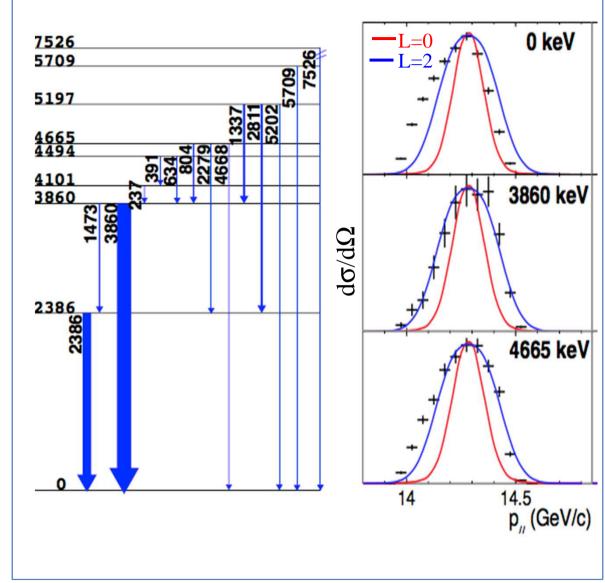
Protons knockout reaction gives composition of orbital occupancies. Large  $^{36}$ S(-1p) cross section for L=0 knock out  $\Rightarrow$  s<sub>1/2</sub> orbit is almost fully occupied in  $^{36}$ S

### Proton Densities in <sup>36</sup>S and <sup>34</sup>Si

Knock-out reactions at β≈0.4

$$\sigma_{\text{-1p}}(n,L) = C^2 S \left(j,n,L\right) \quad \sigma_{sp}(j,S_p) \ R_S \\ \text{occupancy} \qquad \text{reaction theory}$$





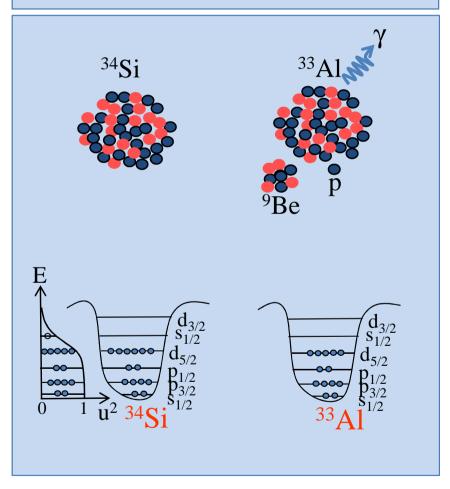


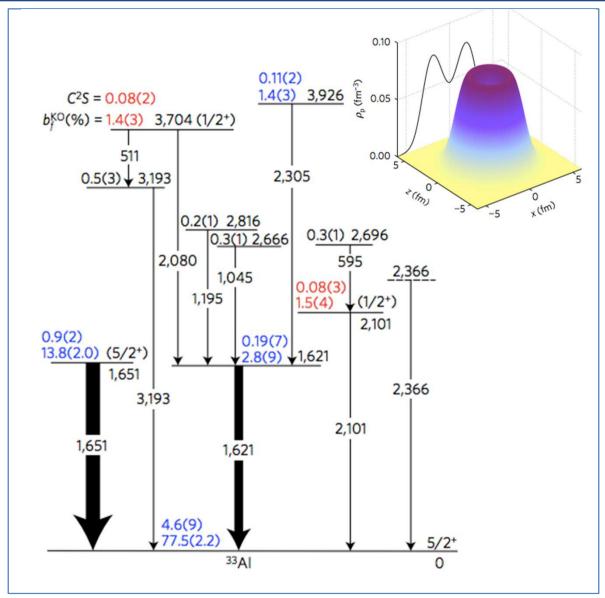
 $^{36}$ S – full  $d_{5/2}$  and  $s_{1/2}$  orbitals, minimum scattering to upper  $d_{3/2}$  orbital

### Proton Densities in <sup>36</sup>S and <sup>34</sup>Si

Knock-out reactions at β≈0.4

$$\sigma_{\text{-1p}}(n,L) = C^2 S \left(j,n,L\right) \quad \sigma_{\text{sp}}(j,S_p) \ R_S \\ \text{occupancy} \qquad \text{reaction theory}$$





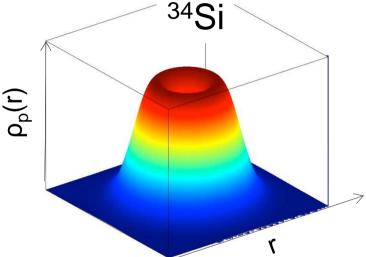


 $^{34}$ Si − ~ empty proton  $s_{1/2}$  orbitals, neutron  $s_{1/2}$  is full  $\Rightarrow$  proton bubble

Nucleons generate their own Mean Field (MF) potential. Early version:  $V = HO + L^2 + LS$  (Haxel, Goppert Mayer 1949)

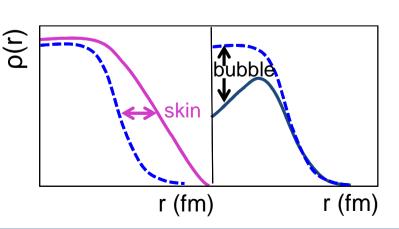
- Nowadays more microscopic description of the SO interaction Density & isospin dependence of LS term

Isospin dependence  $W_1 \approx W_2$  in RMF (no isospin dependence)

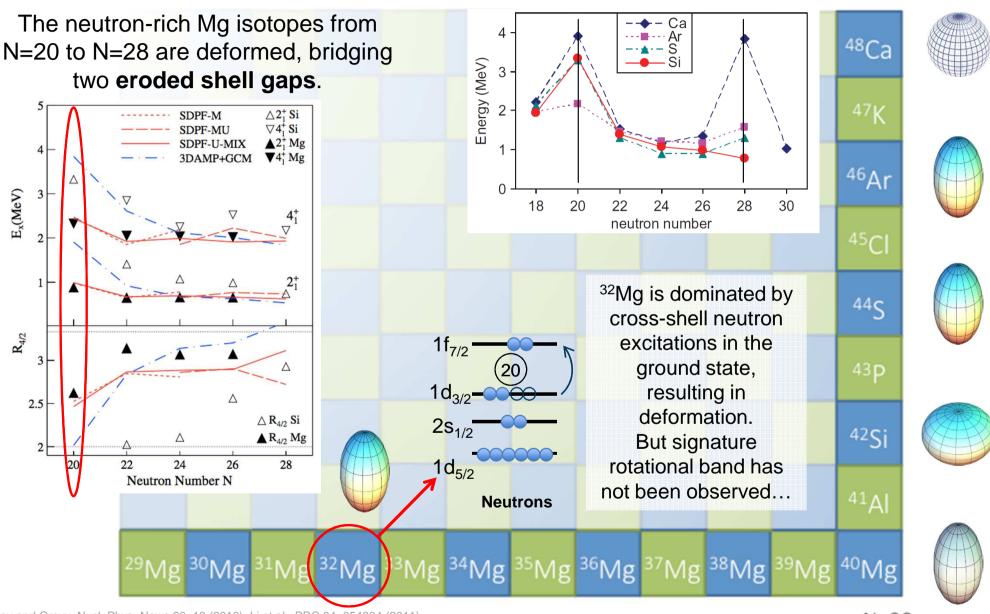


Density dependence

$$V_{\tau}^{\ell s}(r) = -\left[W_{1} \frac{\partial \rho_{\tau}(r)}{\partial r} + W_{2} \frac{\partial \rho_{\tau' \neq \tau}(r)}{\partial r}\right] \vec{\ell} \cdot \vec{s}$$



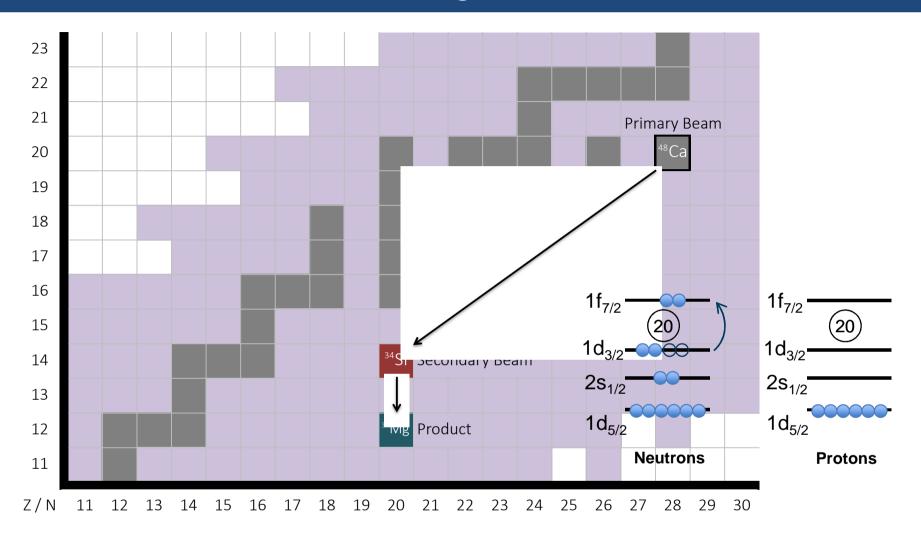
# Structure in Neutron-Rich Mg



Gaudefroy and Grevy, Nucl. Phys. News 20, 13 (2010); Li et al., PRC 84, 054304 (2011). Nowacki and Poves, PRC 79, 014310 (2009); Doornenbal et al., PRL 111, 212502 (2013)

N=28

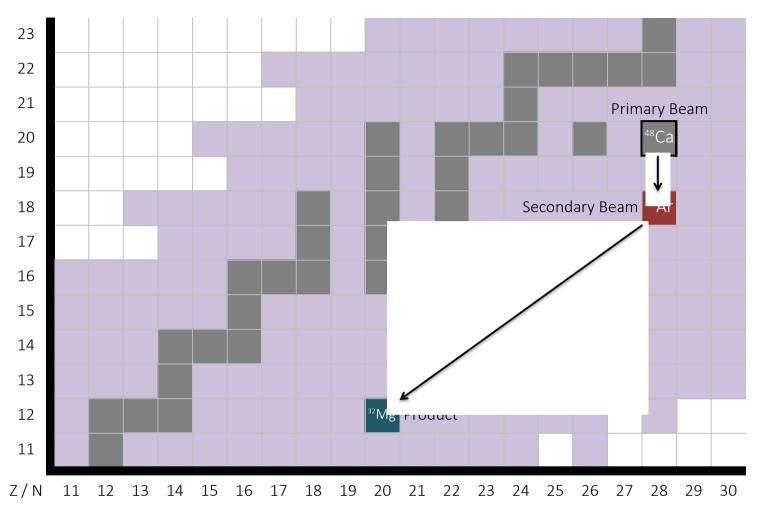
### Fast beams - fragmentation



#### "2-Step Reaction"

- Powerful, Selective: enables "direct reactions" (nucleon removal/pickup)
- Main approach today

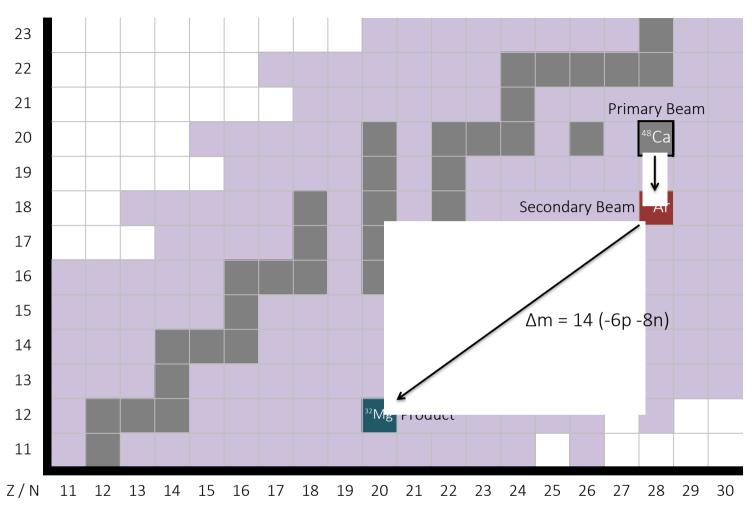
### Fast beams - fragmentation



#### "1-Step Reaction"

- Less selective enables "more statistical" population.
- Angular momentum imparted to a nucleus produced in a projectile fragmentation reaction is related to the mass difference ( $\Delta$ m) between the projectile and the final fragment nucleus

### Fast beams - fragmentation



#### "1-Step Reaction"

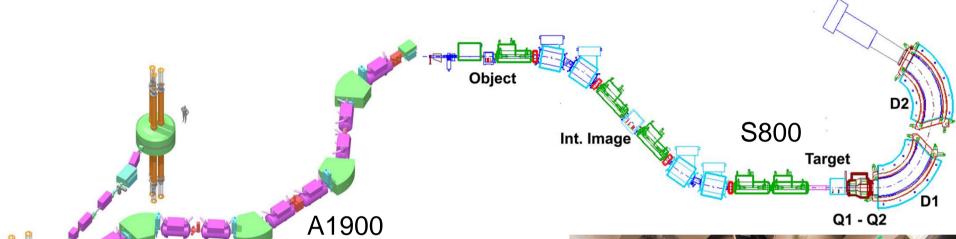
- Less selective enables "more statistical" population.
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# Experiment E11029: GRETINA@S800

GOAL: To see and to study the yrast rotational states (band)

With the intense beams available at NSCL and the availability of GRETINA, we have

have the tools in place to do this

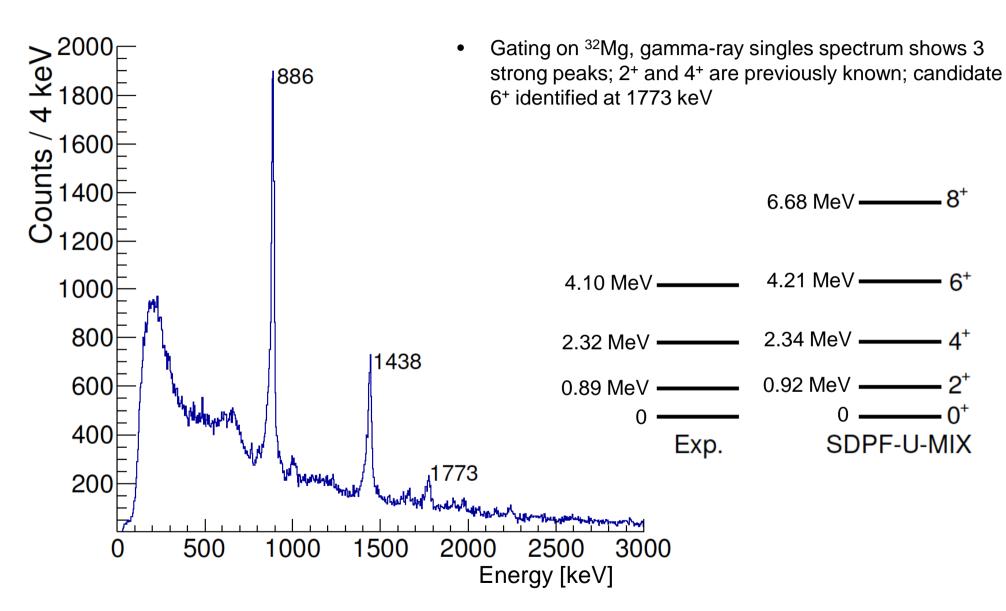


- Primary Beam <sup>48</sup>Ca; secondary beam <sup>46</sup>Ar
- Multi-nucleon knockout into <sup>32</sup>Mg
- GRETINA: All 7 modules at 90 degrees
  - Minimize light particle induced background



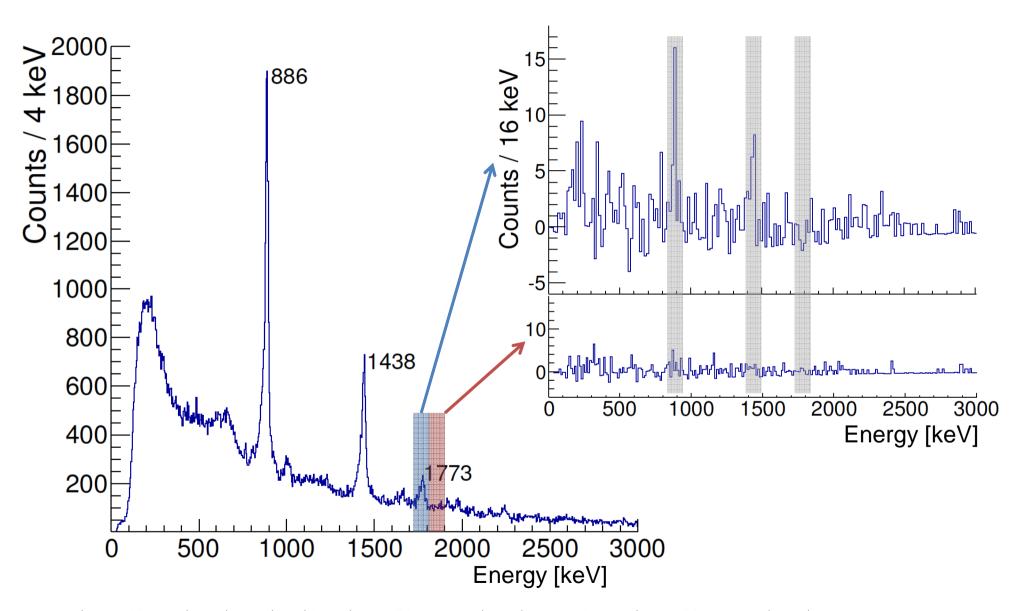
**Focal Plane** 

# Gamma-Ray Singles



C. Detraz et al., PRC 19, 164 (1979); S. Takeuchi et al., PRC 79, 054319 (2009); E. Caurier et al., PRC 90, 014302 (2014).

# Gamma-Ray Singles

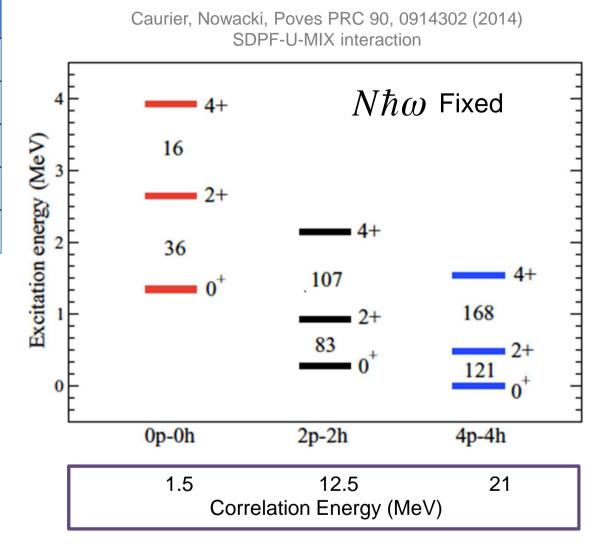


C. Detraz et al., PRC 19, 164 (1979); S. Takeuchi et al., PRC 79, 054319 (2009); E. Caurier et al., PRC 90, 014302 (2014).

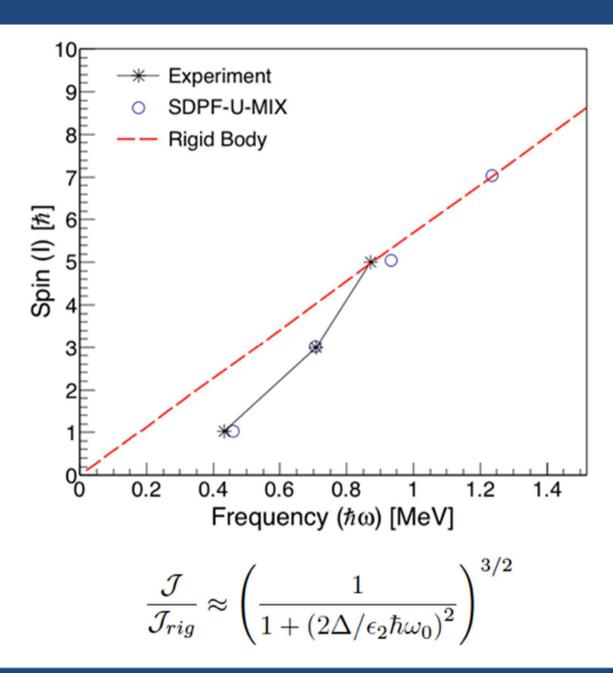
### <sup>32</sup>Mg Shell Model

| Spin | 2p-2h (%) | 4p-4h (%) | $Q_0$ |
|------|-----------|-----------|-------|
| 0    | 53        | 37        |       |
| 2    | 43        | 54        | 56.2  |
| 4    | 31        | 68        | 67.8  |
| 6    | 31        | 68        | 65.4  |
| 8    | 10        | 90        | 59.5  |

- <sup>32</sup>Mg mixed 2p-2h and 4p-4h
- "constant deformation"
- 4p-4h fixed configuration is an "ideal rotor" (superdeformed bands in <sup>36</sup>Ar and <sup>40</sup>Ca)

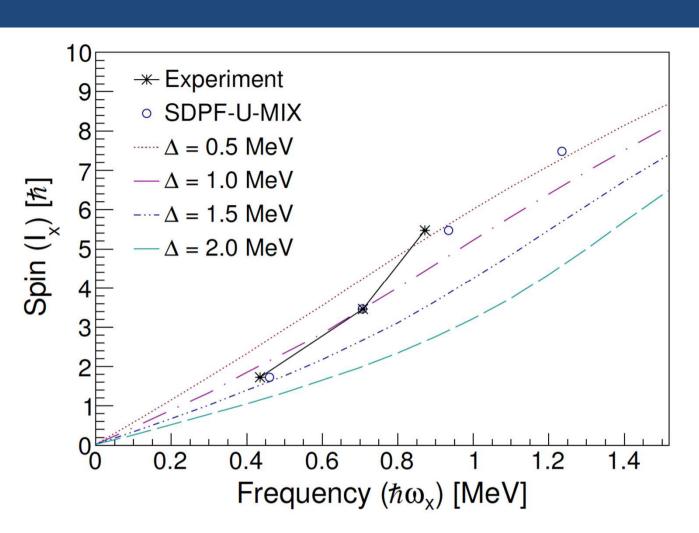


# Rotational band structure in <sup>32</sup>Mg



### Cranked Shell Model

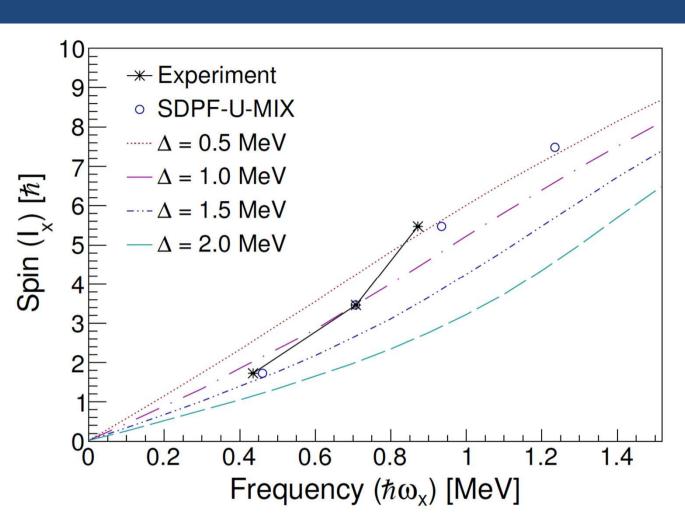
Cranked Shell Model calculations: motion of independent(quasi)particles in a rotating nucleus with a static deformed shape and pair field



$$h'_{q.p.} = h_{s.p.}(\epsilon) - \omega j_x - \Delta (P^+ + P^-) - \lambda \hat{N}$$

### Cranked Shell Model

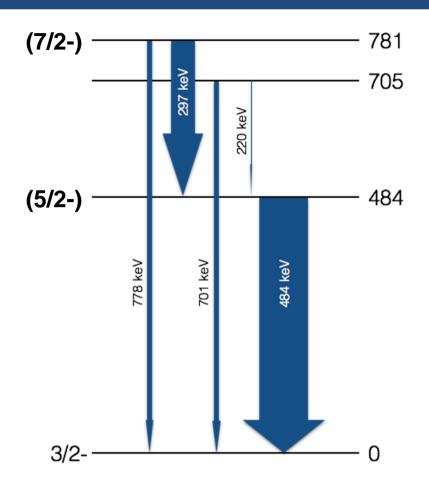
Cranked Shell Model calculations: motion of independent(quasi)particles in a rotating nucleus with a static deformed shape and pair field

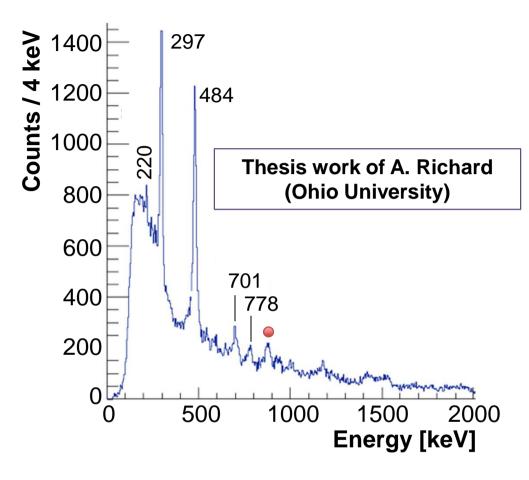


--> Alignment alone (with constant pair gap) does not reproduce upwards bend in the I-ω plot

--> To reproduce data, require change in pair gap by > 50%

### Rotational Band in <sup>33</sup>Mg





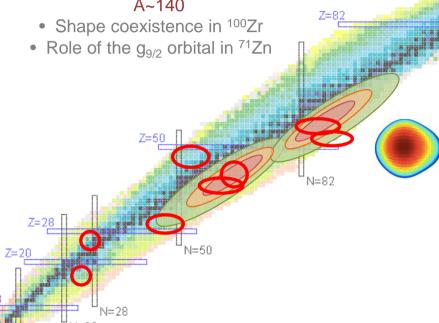
- <sup>33</sup>Mg populated in the same secondary fragmentation shows a rotational band welldescribed as a neutron (Nilsson 3/2[321]) coupled to the <sup>32</sup>Mg core
- Leading order estimates for E2/M1 intensities show remarkable agreement with experiment  $\lambda = \left[\frac{E_{\gamma}}{E_{\gamma'}}\right]^5 \frac{(I+1)(I-1+K)(I-1-K)/2K^2(2I+1)}{1+1.148[(g_K-g_R)/Q_0]^2(I+1)(I-1){E_{\gamma'}}^{-2}}$

# First ANL Physics Campaign (2014-2015)

#### **CARIBU Experiments**

- Shape coexistence in neutron-rich Ge isotopes <sup>72</sup>Ge and <sup>76</sup>Ge
  - Octupole strength in <sup>225</sup>Ra

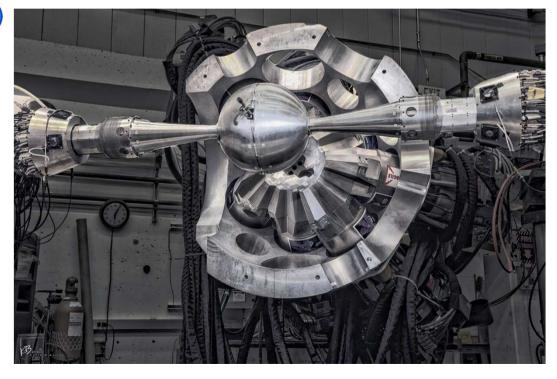
 Octupole strength in neutron-rich nuclei near A~140



ATLAS Stable Beam Experiments

- Comparison of the performance of GRETINA and Gammasphere
  - High-energy gamma-ray performance
  - The polarization sensitivity of GRETINA
  - Multi-particle-hole states in neutron-rich <sup>34</sup>P
- Understanding the low-energy enhancement of the photon-strength function

First physics campaign at ANL made use of the unique opportunities with CARIBU beams, and the combination of GRETINA with particle detectors such as CHICO-2 and the WashU Phoswich Wall.



### Octupole Deformation in 144Ba

PRL 116, 112503 (2016)

PHYSICAL REVIEW LETTERS

week ending 18 MARCH 2016

#### Direct Evidence of Octupole Deformation in Neutron-Rich 144Ba

B. Bucher, <sup>1,\*</sup> S. Zhu, <sup>2</sup> C. Y. Wu, <sup>1</sup> R. V. F. Janssens, <sup>2</sup> D. Cline, <sup>3</sup> A. B. Hayes, <sup>3</sup> M. Albers, <sup>2</sup> A. D. Ayangeakaa, <sup>2</sup> P. A. Butler, <sup>4</sup> C. M. Campbell, <sup>5</sup> M. P. Carpenter, <sup>2</sup> C. J. Chiara, <sup>2,6,†</sup> J. A. Clark, <sup>2</sup> H. L. Crawford, <sup>7,‡</sup> M. Cromaz, <sup>5</sup> H. M. David, <sup>2,§</sup> C. Dickerson, <sup>2</sup> E. T. Gregor, <sup>8,9</sup> J. Harker, <sup>2,6</sup> C. R. Hoffman, <sup>2</sup> B. P. Kay, <sup>2</sup> F. G. Kondev, <sup>2</sup> A. Korichi, <sup>2,10</sup> T. Lauritsen, <sup>2</sup> A. O. Macchiavelli, <sup>5</sup> R. C. Pardo, <sup>2</sup> A. Richard, <sup>7</sup> M. A. Riley, <sup>11</sup> G. Savard, <sup>2</sup> M. Scheck, <sup>8,9</sup> D. Seweryniak, <sup>2</sup> M. K. Smith, <sup>12</sup> R. Vondrasek, <sup>2</sup> and A. Wiens, <sup>5</sup>

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<sup>2</sup>Argonne National Laboratory, Argonne, Illinois 60439, USA

<sup>3</sup>University of Rochester, Rochester, New York 14627, USA

<sup>4</sup>University of Liverpool, Liverpool L69 7ZE, United Kingdom

<sup>5</sup>Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA

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<sup>7</sup>Ohio University, Athens, Ohio 45701, USA

<sup>8</sup>University of the West of Scotland, Paisley PA1 2BE, United Kingdom

<sup>9</sup>SUPA, Scottish Universities Physics Alliance, Glasgow G12 8QQ, United Kingdom

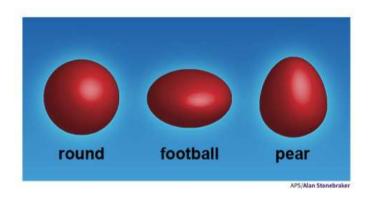
<sup>10</sup>CSNSM, IN2P3-CNRS, bâtiment 104-108, F-91405 Orsay Campus, France

<sup>11</sup>Florida State University, Tallahassee, Florida 32306, USA

<sup>12</sup>University of Notre Dame, Notre Dame, Indiana 46556, USA



- Understand octupole collectivity and effects on nuclear structure
- Important in the regions with  $Z,N \sim 34, 56, 88, 134,$  where  $\Delta J,\Delta I = 3$  orbitals lie near the Fermi surface



# Octupole Collectivity in 144,146Ba

Octupole deformation predicted with  $\Delta J_1\Delta I = 3$  orbitals near the Fermi surface

Qualitative agreement with theory (GCM), but quantitatively higher

Yield fit 144Ba

9-E1 decay

310

400

320

500

 $10^{3}$ 

 $10^{2}$ 

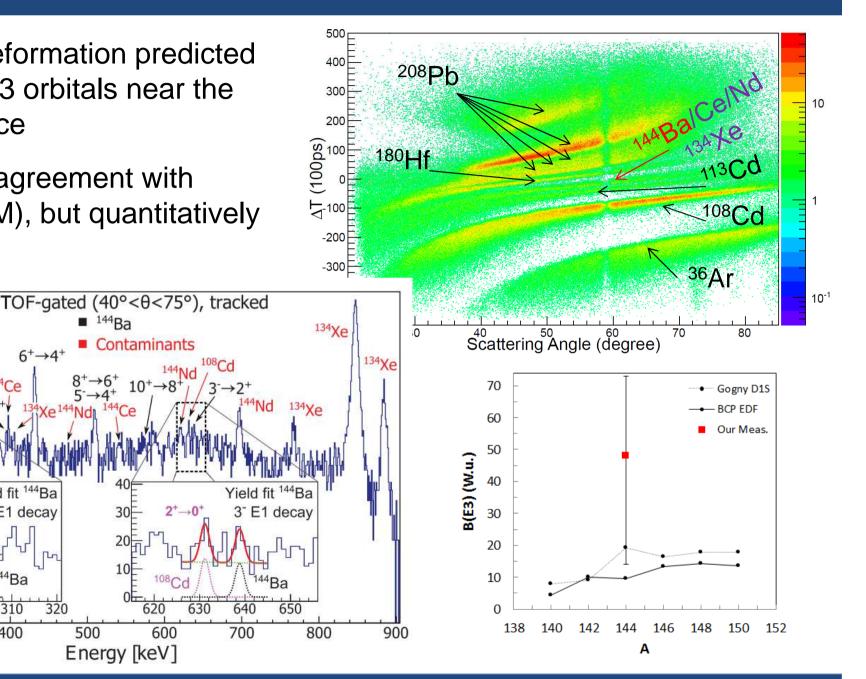
10<sup>-1</sup>

200

300

Counts

 $2^{+} \rightarrow 0^{+}$ 



Energy [keV]

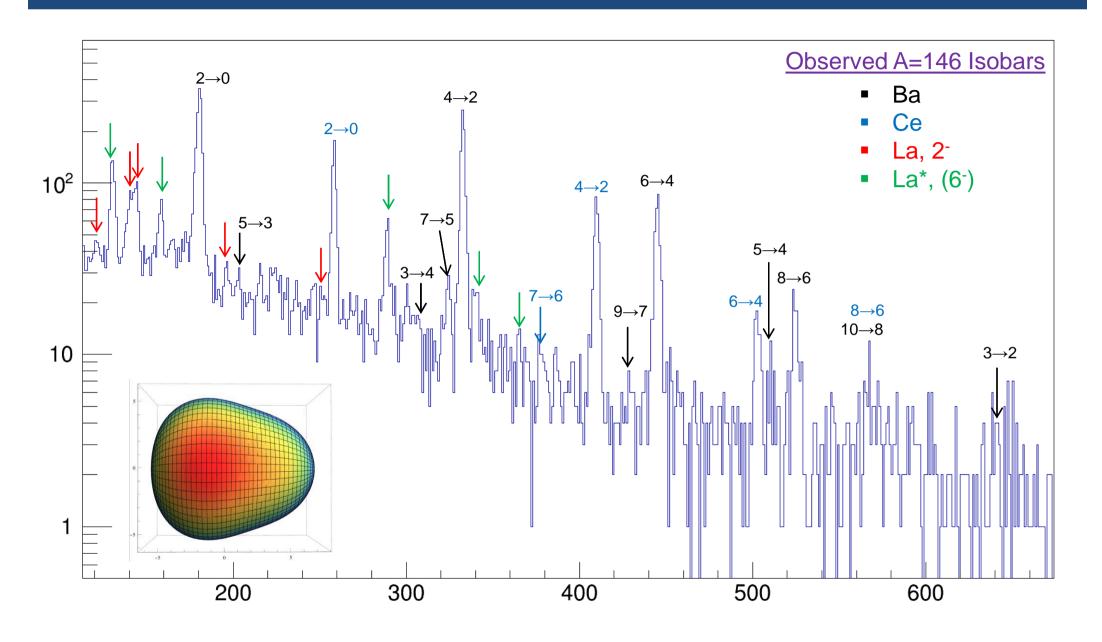
600

640

700

Contaminants

# Octupole Collectivity in 144,146Ba



# Summary

- GRETINA has proven an invaluable tool for nuclear structure studies, including sensitive tests for exotic nuclear shapes
- NSCL Physics Campaign 1.0:
  - <sup>34</sup>Si: "Bubble" nucleus offers unique laboratory for studying spin-orbit interaction
  - 32,33Mg: Rotational structure identified is in good agreement with rotational models; first suggestion of reduced pairing with increasing angular momentum
- ANL Campaign:
  - 144,146Ba Coulex consistent with octupole deformation



### Acknowledgements

#### **GRETINA Lead Laboratory - LBNL**

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Mario Cromaz
Rod Clark
Heather Crawford
Paul Fallon

http://gretina.lbl.gov/
http://greta.lbl.gov/



















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# Merci!