

# Multipole modes within the finite amplitude method and application to the nuclear photo absorption cross section

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# Nuclear DFT

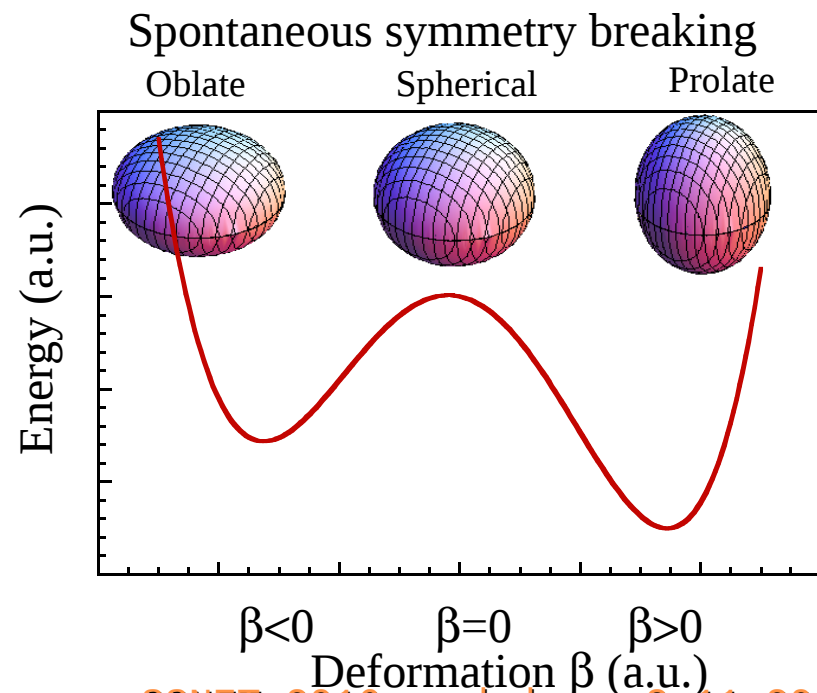
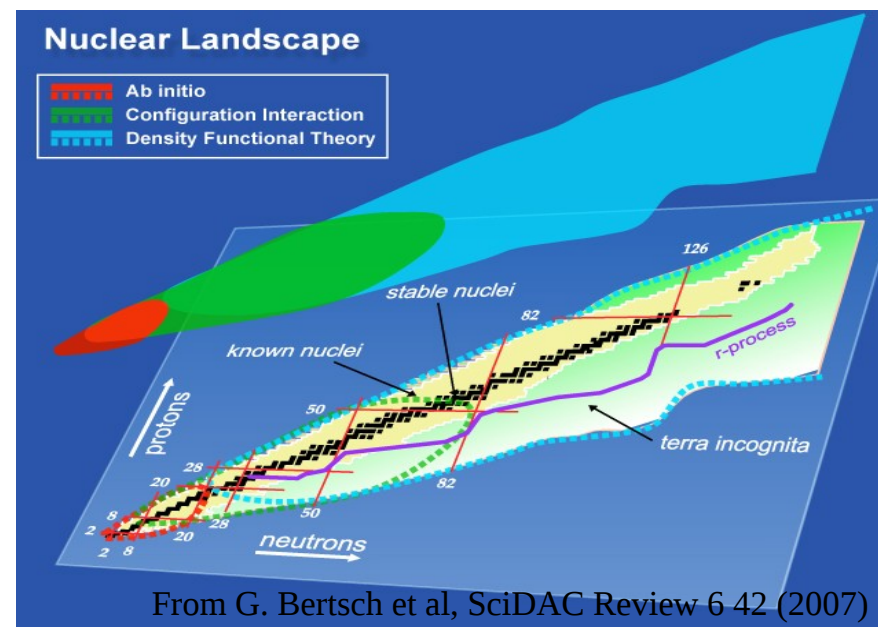
- The nuclear DFT is the only microscopic theory which can be applied throughout the entire nuclear chart
- Within the superfluid nuclear DFT, one needs to solve the Hartree-Fock-Bogoliubov (HFB) equation:

$$\begin{pmatrix} h - \lambda & -\Delta^* \\ \Delta & -h^* + \lambda \end{pmatrix} \begin{pmatrix} U_n \\ V_n \end{pmatrix} = E_n \begin{pmatrix} U_n \\ V_n \end{pmatrix}$$

- By solving these we obtain the quasiparticle energies  $E_n$  and the matrices  $U$  and  $V$  which determine the generalized Bogoliubov quasiparticle transformation:

$$\hat{b}_\alpha^\dagger = \sum_\beta (U_{\beta\alpha} \hat{c}_\beta^\dagger + V_{\beta\alpha} \hat{c}_\beta)$$

- Introduction correlations via spontaneous symmetry breaking



# Linear response

- To access the dynamical properties of the superfluid nuclei, in the framework of nuclear DFT, the linear response (that is, the QRPA) is one of the most often employed method
- Essentially, the QRPA corresponds to a small amplitude oscillations around the HFB ground state
- By diagonalizing the QRPA matrix,

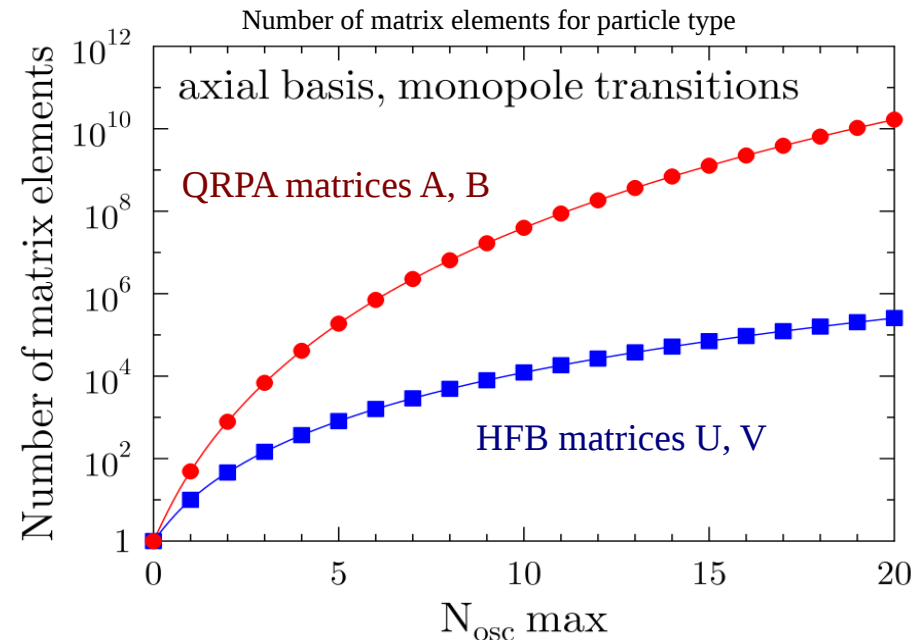
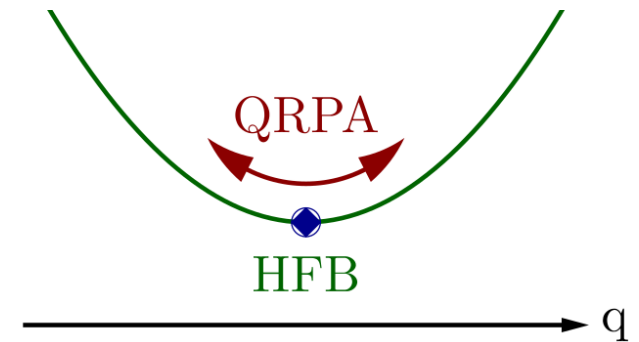
$$\begin{pmatrix} A & B \\ B^* & A^* \end{pmatrix} \begin{pmatrix} X \\ Y \end{pmatrix} = \omega \begin{pmatrix} X \\ -Y \end{pmatrix}$$

one can solve the eigenmodes of the QRPA phonon

- Computational cost of matrix QRPA (MQRPA) becomes huge when spherical symmetry is broken

⇒ Iterative QRPA method required!

QRPA oscillations around the HFB state (schematically)



# Finite amplitude method QRPA

FAM: T. Nakatsukasa, et. al., PRC 76, 024318 (2007)

1) Perform stationary HFB calculation

2) Introduce time-dependent q.p. operator as

$$\alpha_{\mu}(t) = (\alpha_{\mu} + \delta\alpha_{\mu}(t))e^{iE_{\mu}t}$$

3) Time-dependent HFB equation now reads

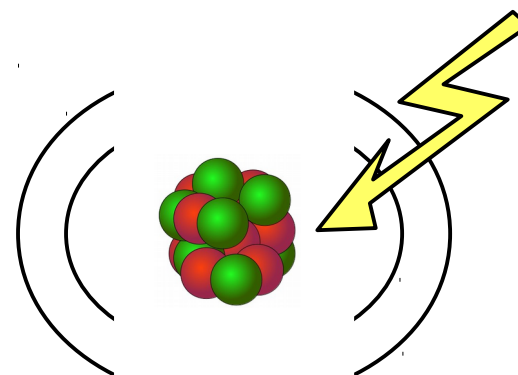
$$i\frac{d\delta\alpha_{\mu}(t)}{dt} = [H(t), \alpha_{\mu}(t)]$$

4) Define oscillating part as

$$\delta\alpha_{\mu}(t) = \eta \sum_{\nu} \alpha_{\nu}^{+} (X_{\nu\mu} e^{-i\omega t} + Y_{\nu\mu}^{*} e^{+i\omega t})$$

Here  $\eta$  is small, and hence the amplitude of oscillation is also small

5) Polarize system with an external field F



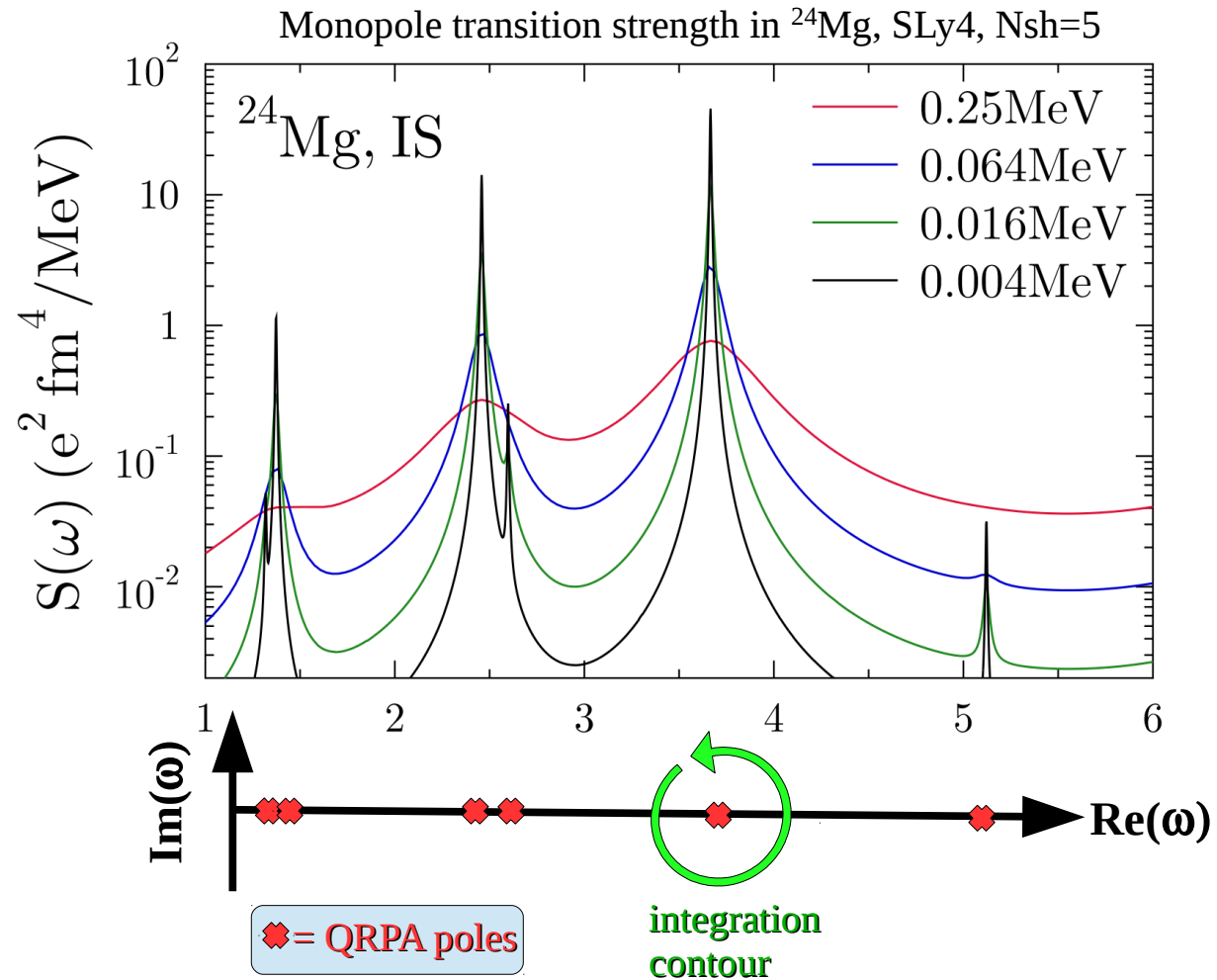
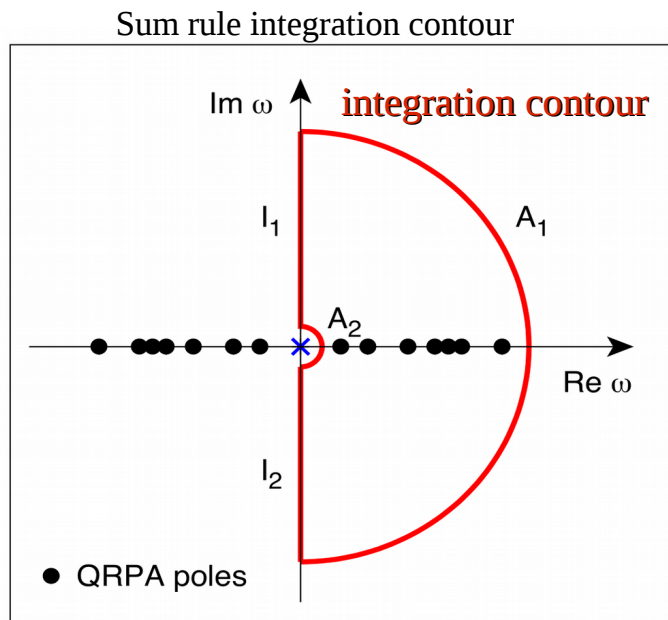
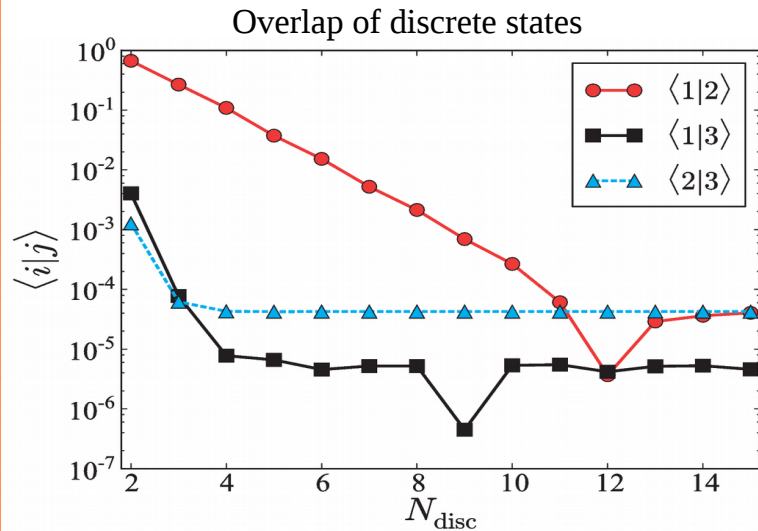
6) FAM equations then reads

$$\begin{aligned} (E_{\mu} + E_{\nu} - \omega) X_{\mu\nu}(\omega) + \delta H_{\mu\nu}^{20}(\omega) &= F_{\mu\nu}^{20} \\ (E_{\mu} + E_{\nu} + \omega) Y_{\mu\nu}(\omega) + \delta H_{\mu\nu}^{02}(\omega) &= F_{\mu\nu}^{02} \end{aligned}$$

7) Introduce a small imaginary width as  
 $\omega \rightarrow \omega + i\gamma$ .

Solve FAM eqs. iteratively for each  $\omega$ .

# Discrete low-lying states and sum rules



- FAM can be used to access discrete excited states, or to calculate sum rules
- Both methods are based on a contour integral in a complex plane

Discrete states: N. Hinohara, M. K., W. Nazarewicz, Phys. Rev. C 87, 064309 (2013)  
 Sum rules: N. Hinohara, M.K., W. Nazarewicz, E. Olsen, PRC 91, 044323 (2015)

# Multipole transitions with axial FAM-QRPA

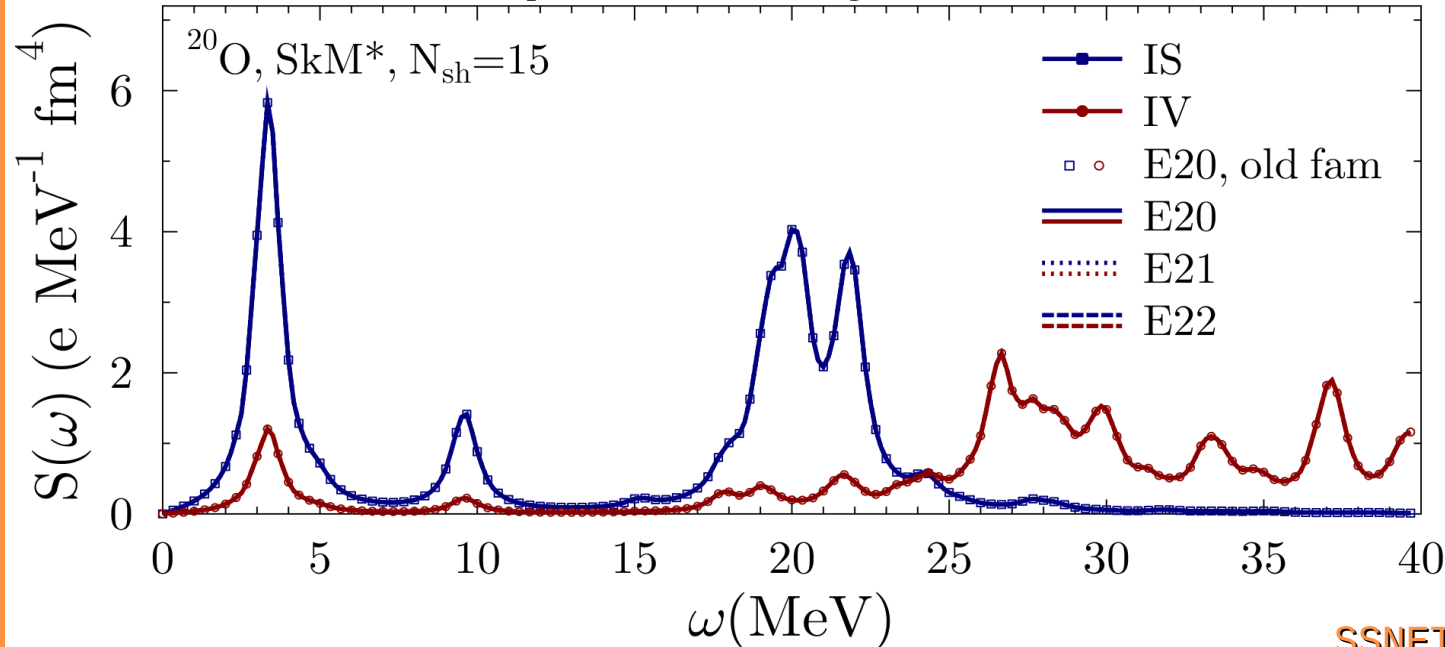
- For  $K \neq 0$  modes, the EM transition operator  $r^L Y_{LK}$  has a different block structure than  $h$  of HFB when using axial basis
- Need to explicitly linearize density dependent parts (expansion parameter  $\eta$  no longer needed)
- Implemented to HFBTHO
- Wigner-Eckart theorem in spherical nuclei allows to test the implementation
- No truncations in the q.p. space

Matrix structure in axial basis for  $K \neq 0$  modes

$$h_{\text{HFB}} = \begin{pmatrix} 1/2 & & & \\ & 3/2 & & \\ & & 5/2 & \\ & & & 7/2 \end{pmatrix}$$

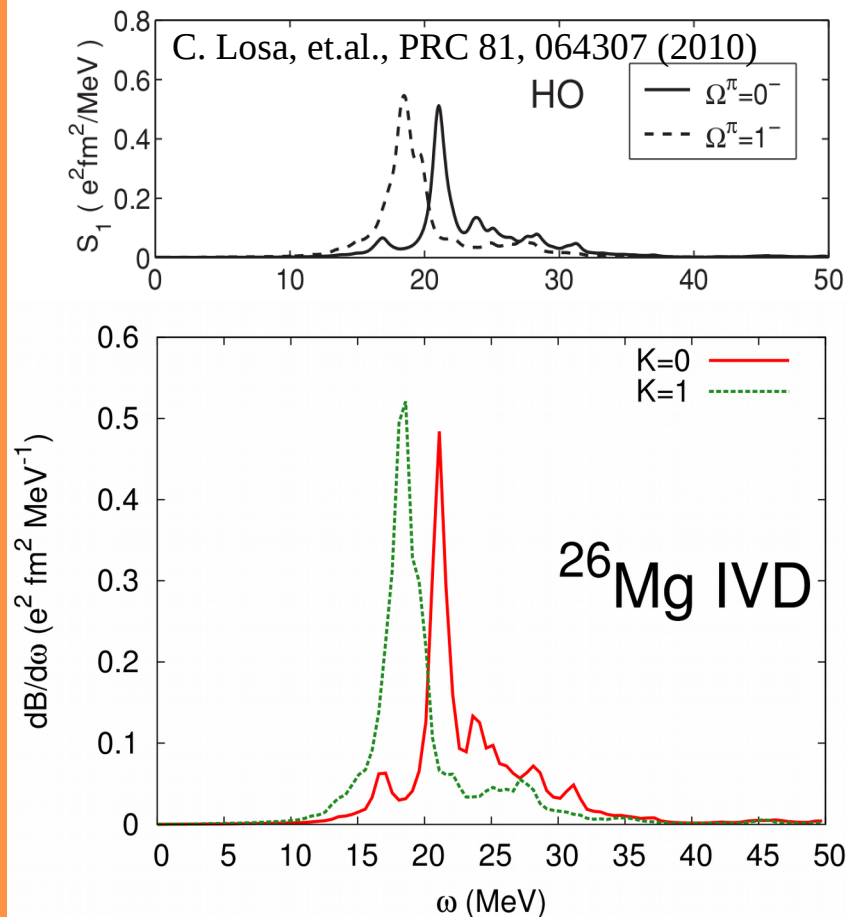
$$F^{20} = \begin{pmatrix} & & & \\ & & & \\ & & & \\ & & & \end{pmatrix}$$

Quadrupole transition strength in  $^{20}\text{O}$



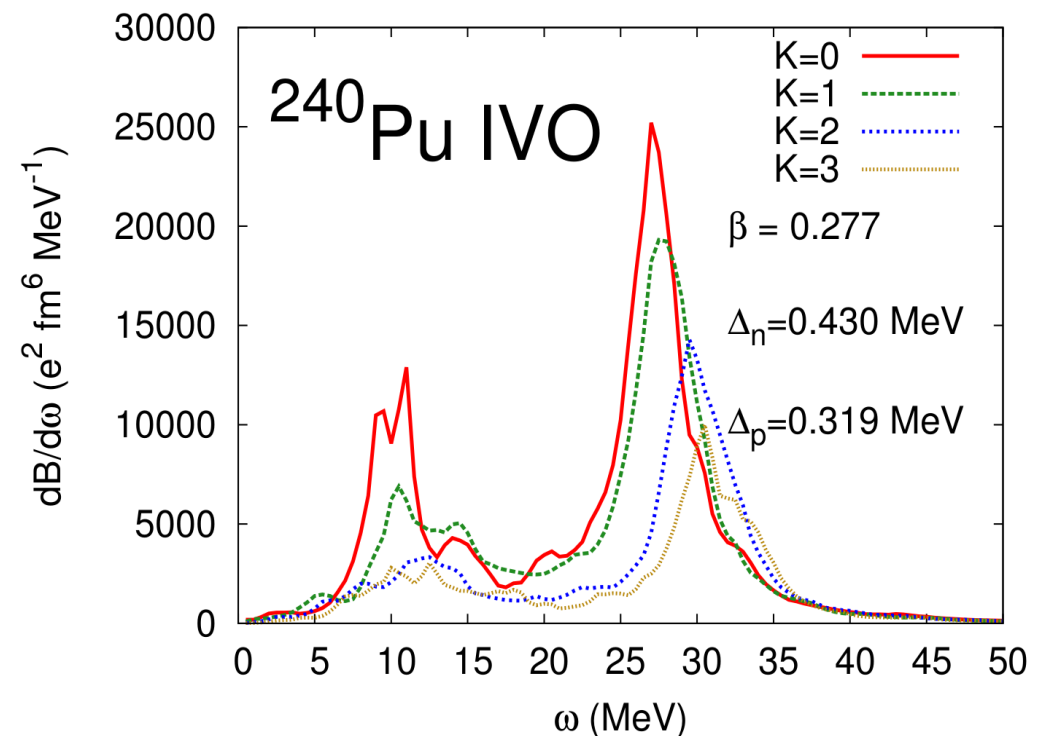
# Multipole transitions with axial FAM-QRPA

- Comparison to earlier MQRPA calculation agrees very well ( $^{26}\text{Mg}$ ,  $N_{\text{sh}}=15$ , SkM\*, isovector dipole mode)



See: M.K, N.Hinohara, W.Nazarewicz,  
Phys. Rev. C 92, 051302(R) (2015)

- Method feasible for actinide nuclei with  $N_{\text{sh}}=20$ , without any truncations
- Example of isovector octupole transition strength in  $^{240}\text{Pu}$



- Calculation of the strength function is trivially parallelizable

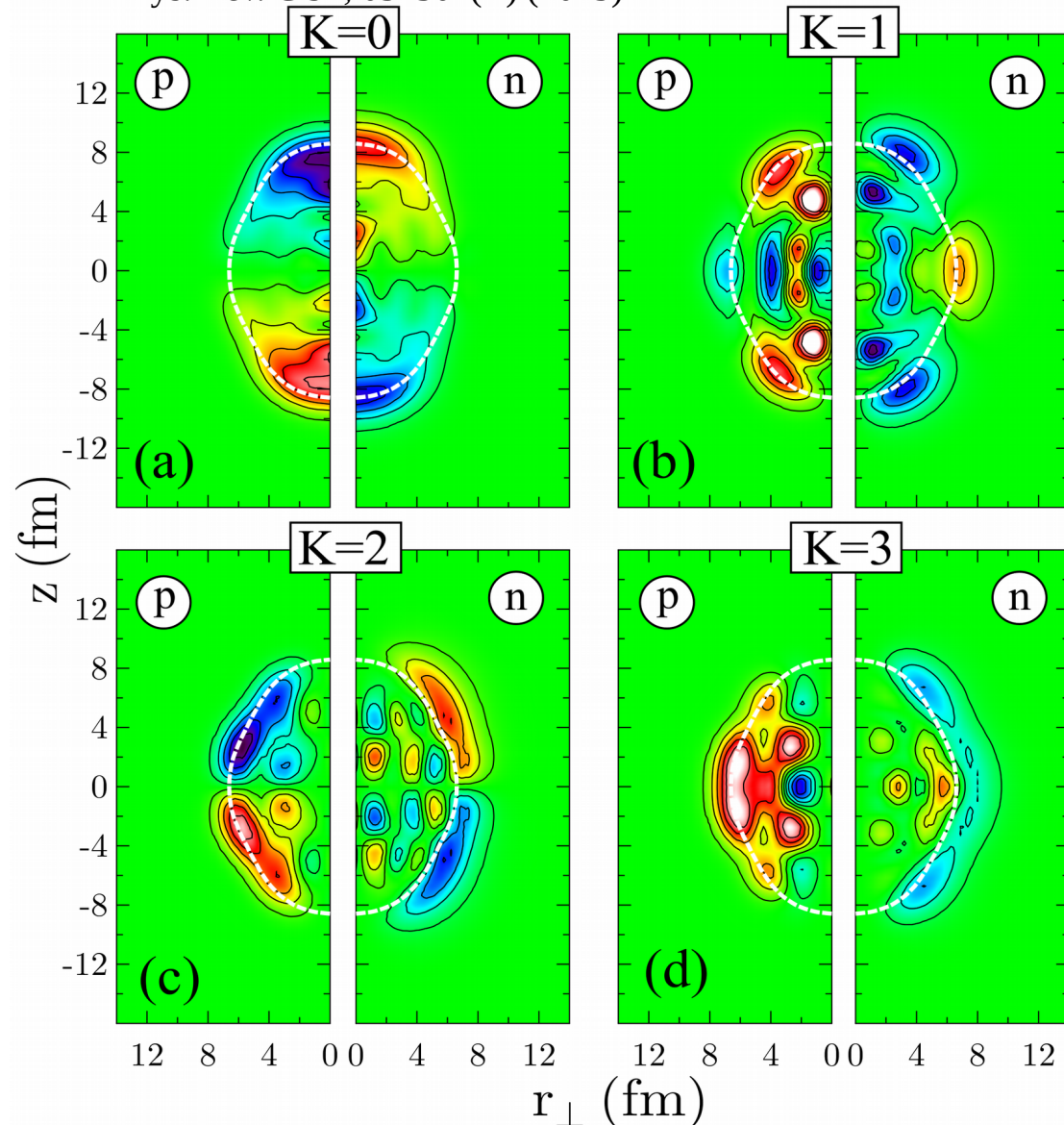
# Multipole transitions with axial FAM-QRPA

- The QRPA corresponds to a small amplitude oscillations around the static HFB density  $\rho_0$ .
- With isovector octupole operator, the induced transitions densities show a clear octupole-like shapes
- Low-lying iv. octupole modes in  $^{240}\text{Pu}$  show collectivity: The induced density covers large portion of the nuclear volume
- The angular dependency of the induced density is

$$\rho(\mathbf{r}) = \rho(r, z) \left( e^{+iK\phi} + e^{-iK\phi} \right) / 2$$

due to simplex-y symmetry.

Oscillating density  $\rho_f$  in  $^{240}\text{Pu}$  with isovector octupole operator at  $\omega=11\text{MeV}$  with SLy4 for different  $K$ -modes. From M.K, et.al., Phys. Rev. C 92, 051302(R) (2015)

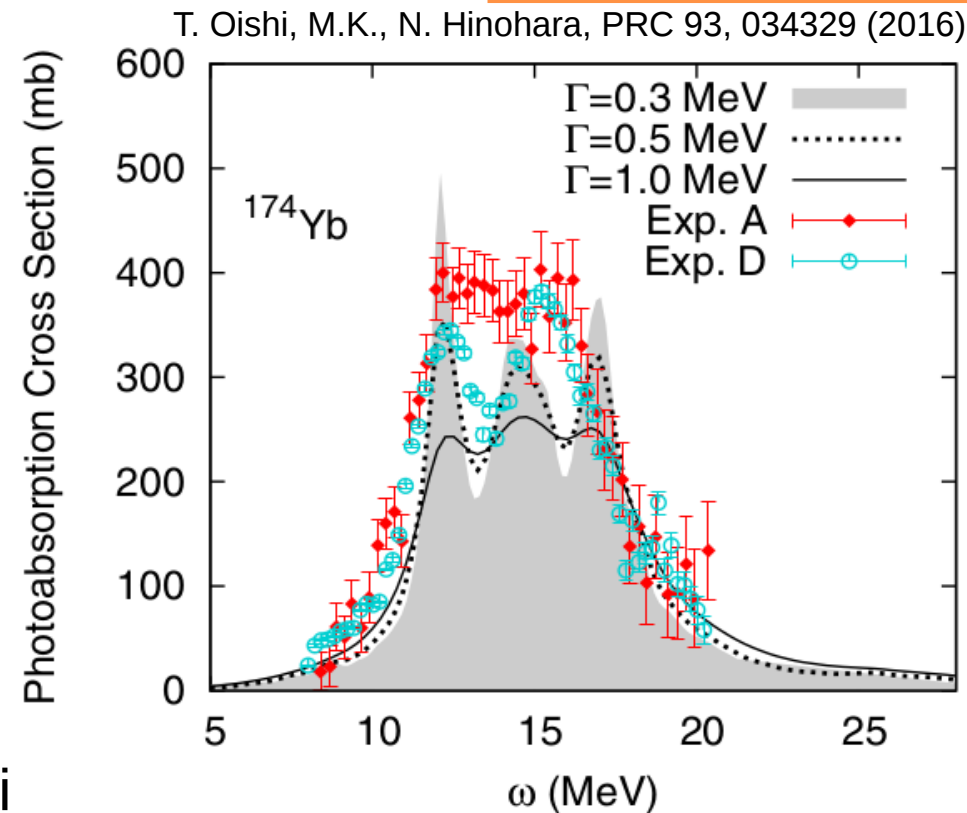


# Photo absorption in the heavy rare-earth nuclei

- The nuclear photo-absorption cross-section is connected to the electric isovector dipole transition

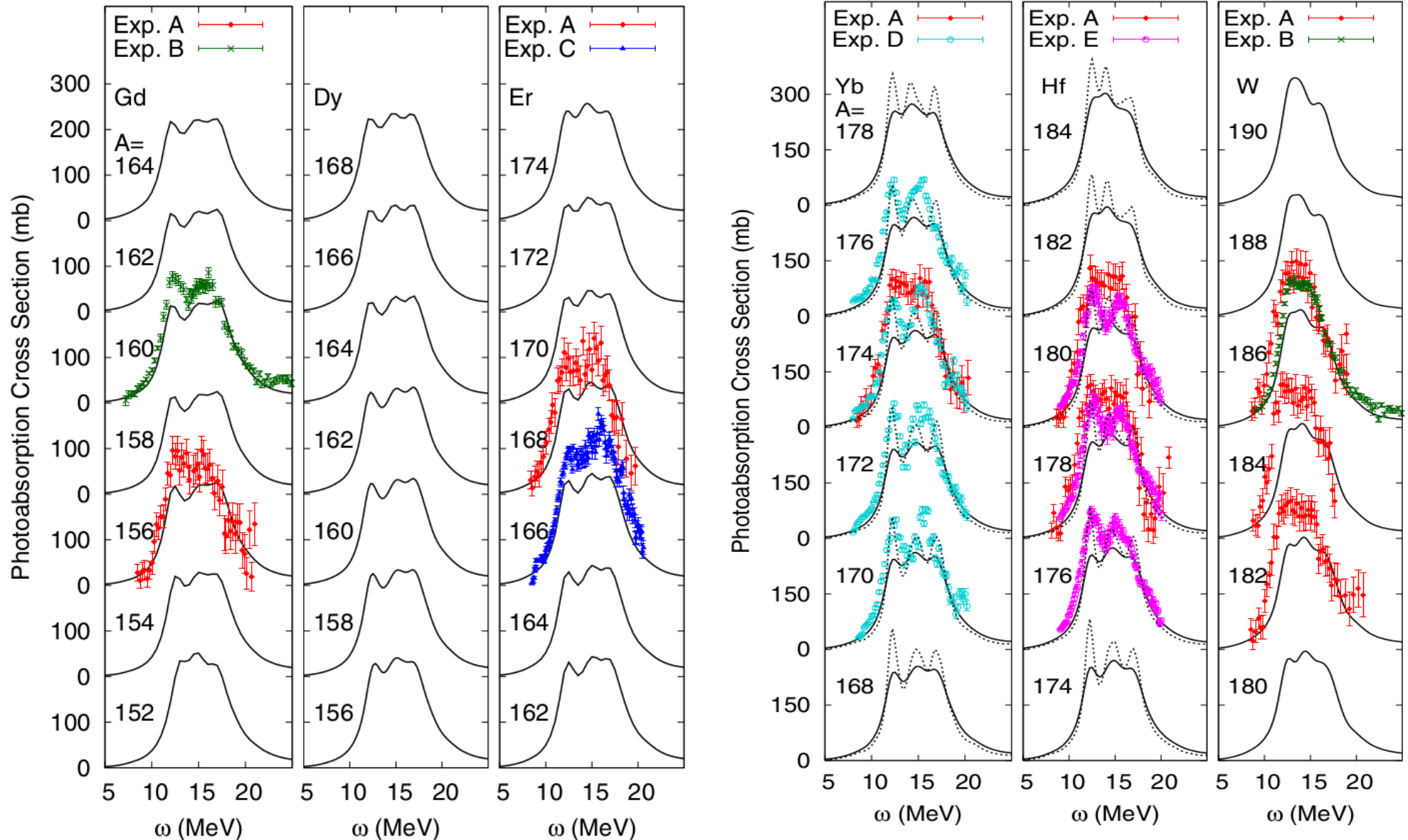
$$\sigma_{\text{abs}}(\omega) = \frac{4\pi\omega}{\hbar c} \sum_{K=0,\pm 1} \frac{dB(Q_{\text{IVD}}, \omega)}{d\omega}$$

- The nuclear photo absorption cross section has also impact on the r-process dynamics
- We recently computed photo absorption cross sections in heavy rare-earth nuclei see T. Oishi, M.K., N. Hinohara, PRC 93, 034329 (2016)
- The results show that for heavier rare earths, the experimental cross section can not be reproduced fully, and there is a certain deficiency
- The position of the giant resonance is usually well reproduced



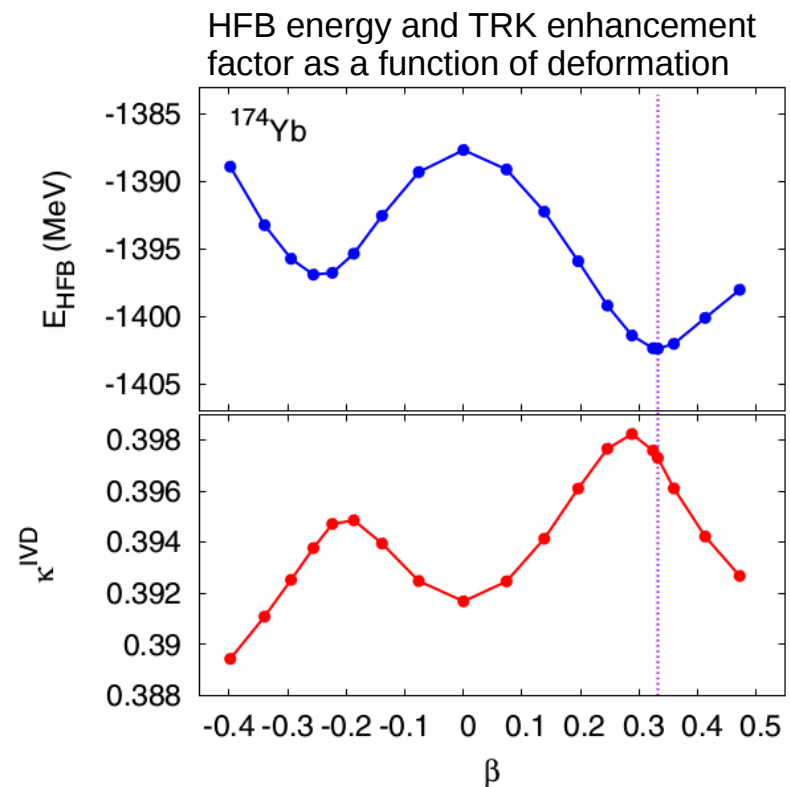
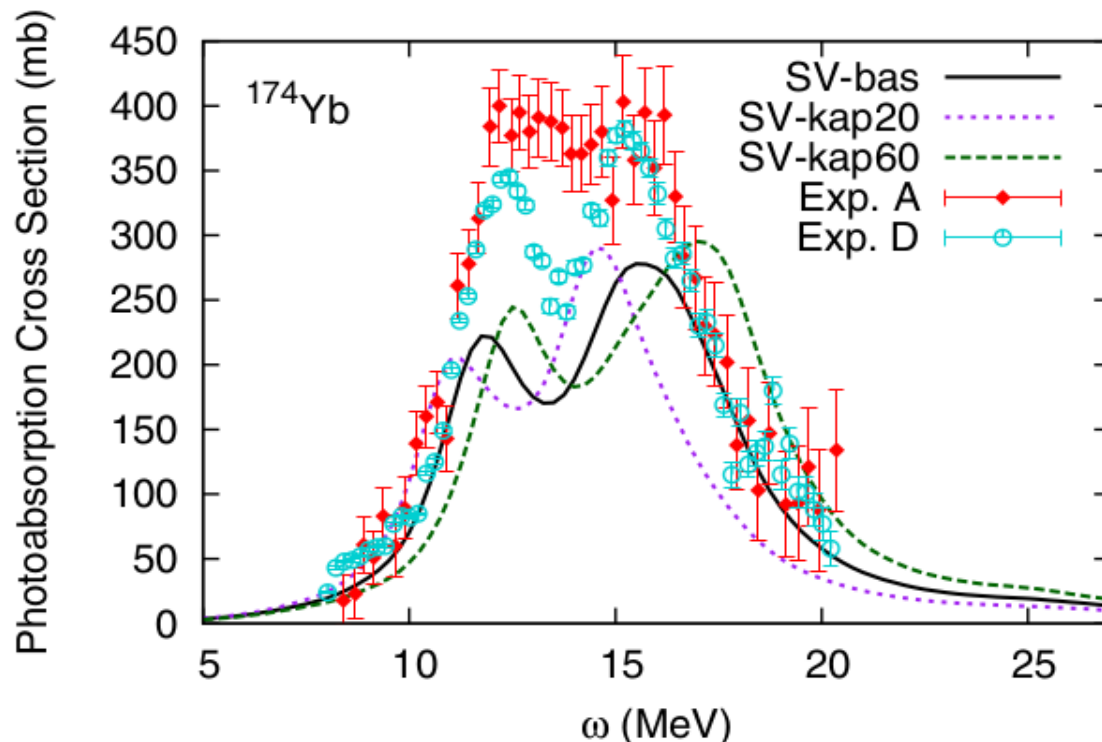
# Photo absorption in the heavy rare-earth nuclei

- T. Oishi, M.K., N. Hinohara, PRC 93, 034329 (2016)



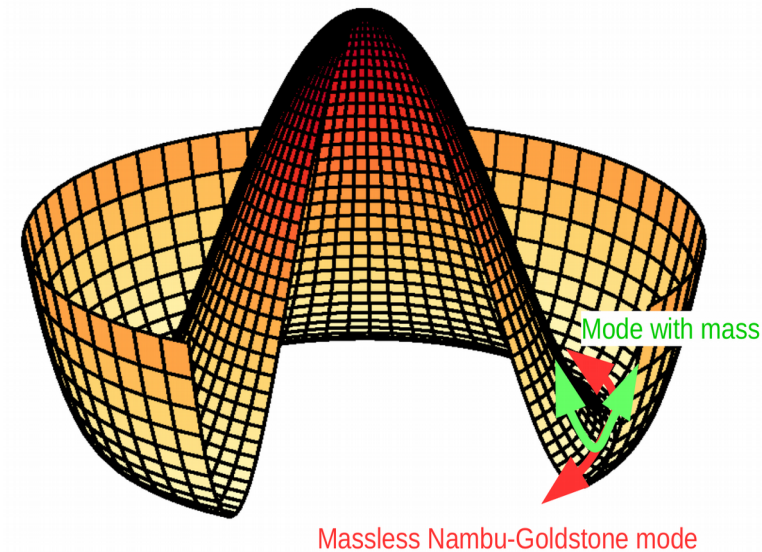
# Photo absorption in the heavy rare-earth nuclei

- By systematically varying the Thomas-Reiche-Kuhn (TRK) sum rule enhancement factor, the value of the energy weighted sum rule (EWSR) can be modified
- This, however, does not seem to improve situation when comparing to experimental data. A larger TRK factor just pushes the strength to higher energy, thus increasing the EWSR
- To improve the situation, future EDF parameter optimization schemes should incorporate data on various excitation modes

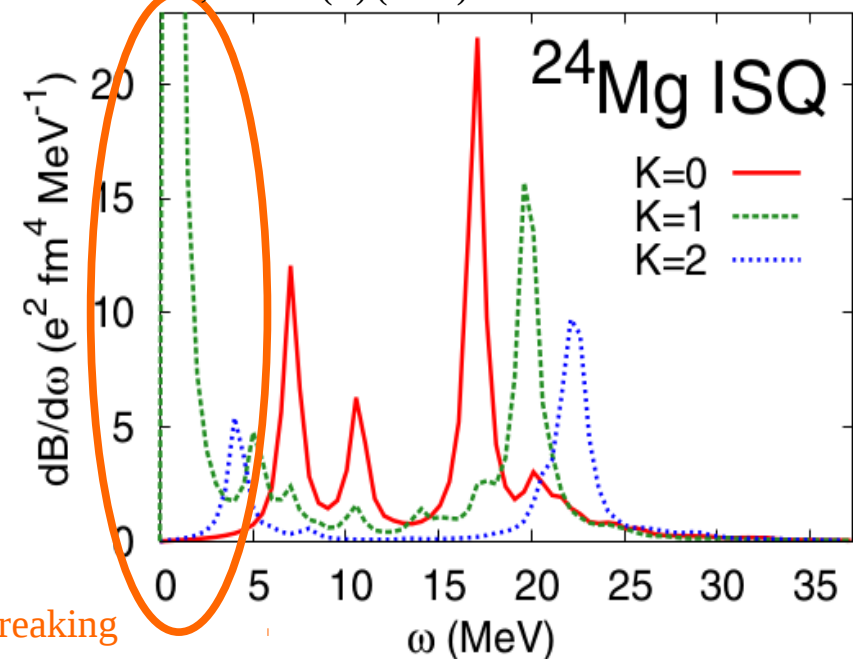


# Spurious modes

- If the underlying HFB solution breaks some symmetry, this causes a spurious QRPA mode to appear
- The spurious mode appears at zero energy
- This corresponds to the well-known Nambu-Goldstone mode
- The type of the spurious mode depends on which of the symmetry was broken
- For example, breaking of the rotational symmetry causes a spurious  $K^\pi=1^+$  mode to appear. This shows up when calculating isoscalar quadrupole modes
- Work towards removing spurious  $K^\pi=1^+$  mode is currently in progress



Isoscalar quadrupole transition strength. From M.K, N.Hinohara, W.Nazarewicz, Phys. Rev. C 92, 051302(R) (2015)



Spurious mode due to breaking of the spherical symmetry

## Summary and outlook

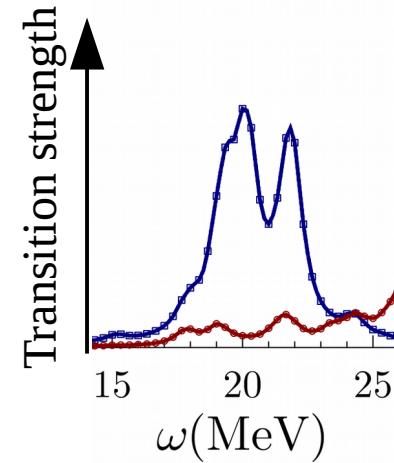
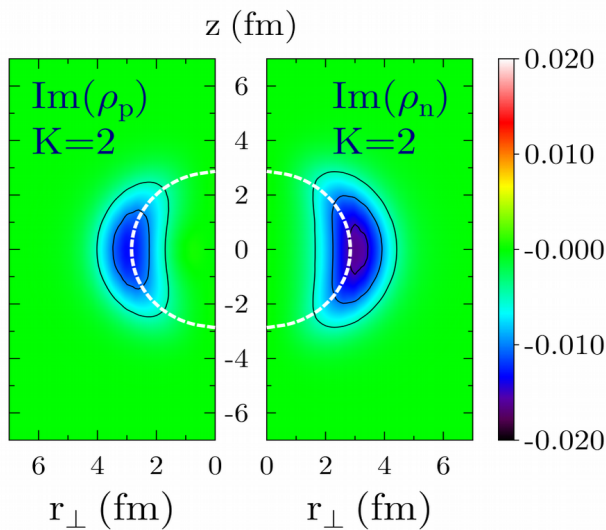
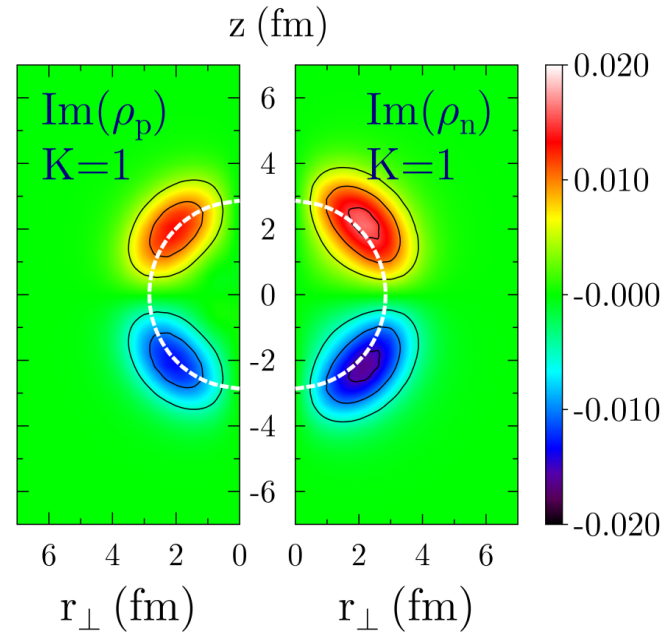
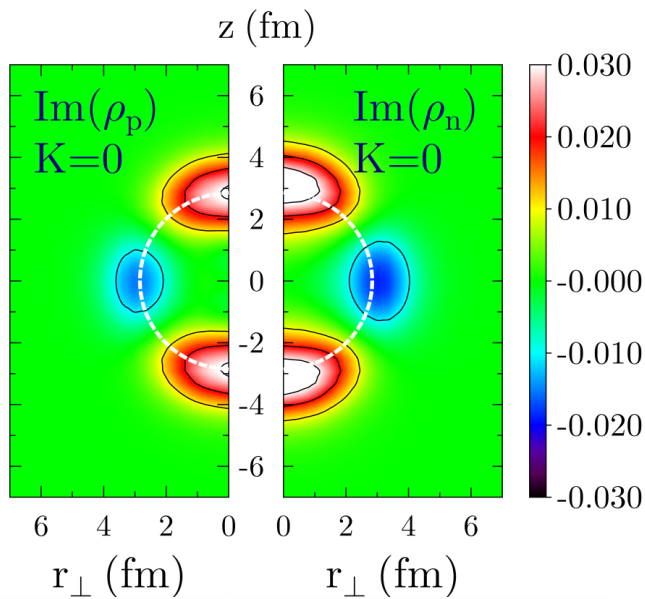
- FAM offers an computationally inexpensive way to solve the QRPA problem, by iterative means
- Many new FAM developments have been published recently:
  - Besides transition strength function, FAM can be used to compute discrete states and sum rules
  - New FAM module allows to compute arbitrary multipole mode with an axial code
  - FAM applied to photo absorption calculations in the heavy rare-earth region
- FAM can be trivially parallelized, allowing large scale surveys
- The newly developed FAM module allows systematic studies across significant portion of the nuclear chart (including actinides and superheavy nuclei)

### Future plans

- Assessment of the spurious rotational mode
- Implementation of M1 transitions (mostly done)

Backup slides

# Induced transition densities

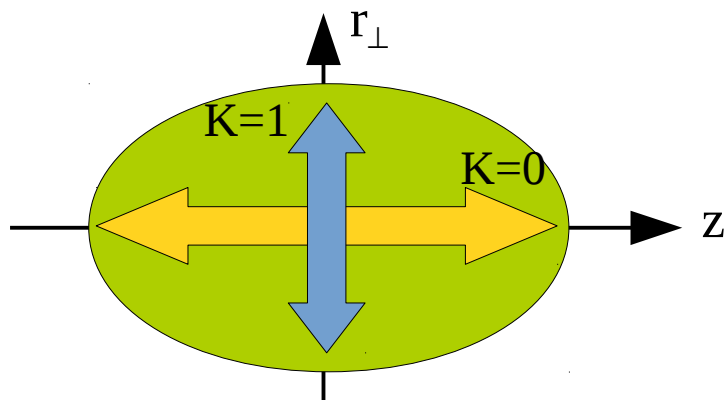


- The QRPA corresponds to a small amplitude oscillations around the static HFB density  $\rho_0$ .
- With FAM equations, we can isolate this oscillating part to induced density  $\rho_f$ .
- The shape of this oscillating density depends on the used transition operator, energy  $\omega$ , and underlying HFB solution
- Figs. show induced proton and neutron transition densities for  $^{20}\text{O}$  at energy of 20 MeV for isoscalar quadrupole transition

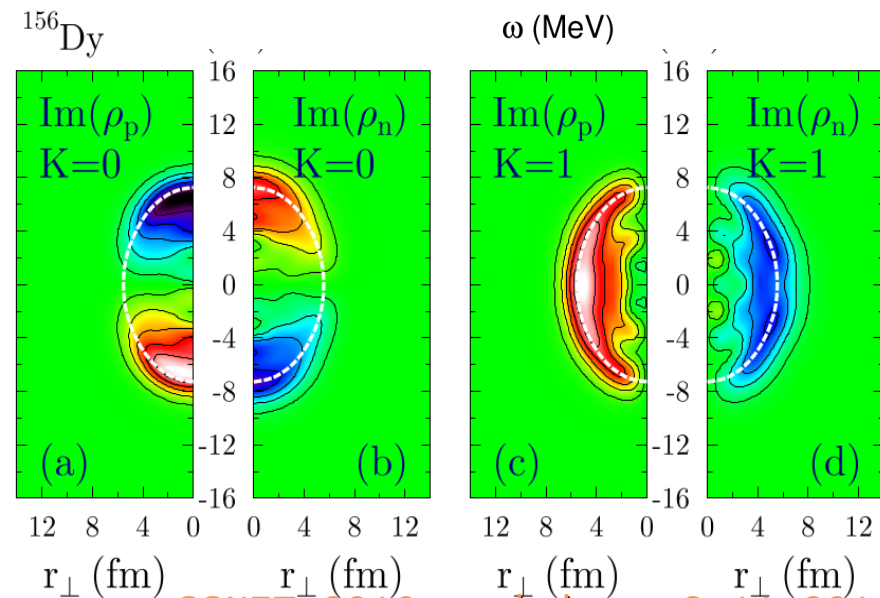
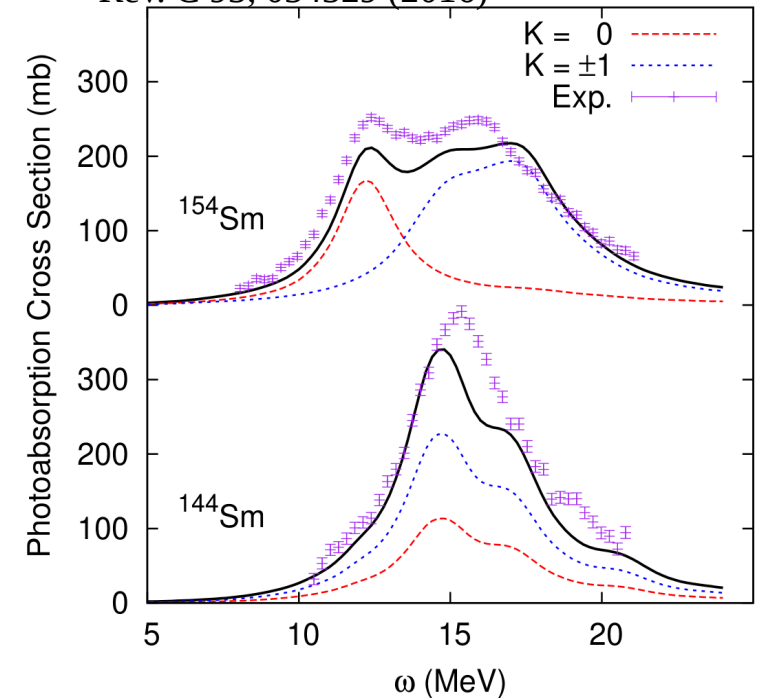
$$\rho(\mathbf{r}) = \rho(r, z) (e^{+iK\phi} + e^{-iK\phi}) / 2$$

# Deformation and splitting of the giant resonance

- Due to the deformation, the position of the giant resonance is different for different  $K$ -modes
- This happens for all  $L > 0$  multipole operators
- Semiclassically this could be explained such that the oscillating frequency is different to each direction. That is, e.g. with dipole operator, the  $K=1$  mode corresponds oscillation perpendicular to  $z$ -axis. With prolate deformation, this leads to higher  $\omega$  compared to  $K=0$  mode case, and conversely, other way around with oblate deformation.



Photoabsorption in Sm isotopes from FAM-QRPA with SkM\*. From T. Oishi, et.al, Phys. Rev. C 93, 034329 (2016)



## Removal of the spurious modes

- In practice, due to the finite size of the used basis, the spurious mode is usually located slightly above the zero energy
- It is possible to remove this mode within the FAM-QRPA framework. For the spurious center-of-mass mode, see e.g. T. Nakatsukasa, et.al., PRC 76, 024318 (2007) and N. Hinohara, PRC 92, 034321 (2015)
- The idea is to remove spurious contribution from calculated transition density as

$$\delta \rho_{\text{phys}}(\omega) = \delta \rho_{\text{cal}}(\omega) - \lambda_P \delta \rho_P - \lambda_R \delta \rho_R$$

where  $\lambda_R$  and  $\lambda_P$  are determined from canonicity condition  $[R, P] = i$

- This technique allows also to access pairing rotational mode, see: N. Hinohara, W. Nazarewicz, Phys Rev. Lett. 116, 152502 (2016)
- Currently we are working with removal of the spurious rotational mode

FAM-QRPA strength function for the response of the center of mass operator for  $^{26}\text{Mg}$  with SLy4. The calculation was done with  $N_{\text{sh}}$  oscillator shells. From N. Hinohara, Phys. Rev. C 92, 034321 (2015).

