## On recent anomalies in flavor physics

Diego Guadagnoli
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The measured branching fraction is compatible with the previous measurement [3] and lies below SM expectations. For the $q^{2}$ region $1.0<q^{2}<6.0 \mathrm{GeV}^{2} / c^{4}$ the differential branching fraction of $\left(2.58_{-0.31}^{+0.33} \pm 0.08 \pm 0.19\right) \times 10^{-8} \mathrm{GeV}^{-2} c^{4}$ is more than $3 \sigma$ below the SM prediction of $(4.81 \pm 0.56) \times 10^{-8} \mathrm{GeV}^{-2} c^{4}[4,5,32]$.

```
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One of such "clean" observables is called $P_{5}^{\prime}$


- Tension seen in $P_{5}^{\prime}$ in [PRL 111, 191801 (2013)] confirmed [4.0, 6.0] and $[6.0,8.0] \mathrm{GeV}^{2} / c^{4}$ show deviations of $2.9 \sigma$ each
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- What cancels is the dependence on the large- $\mathrm{m}_{\mathrm{b}}$ form factors.
- Debate on the role of
- Subleading terms in $1 / m_{b}$
- cc loops and their resummation


## See:

Jäger \& Martin-Camalich, PRD 2016
Ciuchini et al., 1512.07157

## The $P_{5}^{\prime}$ anomaly: continued

The above said, this anomaly remains interesting (and more and more so)

- It occurs in the same kinematic range as $R_{K}$ namely $m_{\mu \mu}^{2} \in[1,6] \mathrm{GeV}^{2}$
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- Conclusion:

If it's new physics, it is expected to show up elsewhere in the $B \rightarrow K^{*} \mu \mu$ angular analysis.

Run II will tell for sure

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& B_{s} \rightarrow \mu \mu \\
& \text { (5) } \frac{B R\left(B_{s} \rightarrow \mu \mu\right)_{\mathrm{exp}}}{B R\left(B_{\mathrm{s}} \rightarrow \mu \mu\right)_{\mathrm{SM}}}=0.77 \pm 0.20
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- Exp error will go to: $\sim 10 \%$ by end of Run II
$\sim 5 \%$ w/ LHCb upgrade


## More discrepancies: <br> $b \rightarrow c$ decays

There are long-standing discrepancies in $b \rightarrow c$ transitions as well.

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- Yet: - Q1: Can we (easily) make sense of $\mathbf{0}$ to $\boldsymbol{\bullet}$ ?
- Q2: What are the most immediate signatures to expect?


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- Rotating $q$ and $\ell$ to the mass eigenbasis generates LFV interactions.


## Let's now turn to Q1:

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& \text { purely vector } \\
& \text { purely axial } \\
& \text { lepton current } \\
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- Note: $\quad C_{9}^{\mathrm{SM}}\left(m_{b}\right) \approx+4.2$

$$
C_{10}^{\mathrm{SM}}\left(m_{b}\right) \approx-4.4
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[Bobeth, Misiak, Urban, 99]
[Khodjamirian et al., 10]

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H_{\mathrm{SM}+\mathrm{NP}}(\bar{b} \rightarrow \bar{s} \mu \mu)=-\frac{4 G_{F}}{\sqrt{2}} V_{t b}^{*} V_{t s} \frac{\alpha_{\mathrm{em}}}{4 \pi}\left[\bar{b}_{L} \gamma^{\lambda} s_{L} \cdot\left(C_{9}^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu+C_{10}^{(u)} \bar{\mu} \gamma_{\lambda} \gamma_{5} \mu\right)\right] \text { purely vector } \quad \text { lepton current }
$$

i.e. in the SM

- Note: $C_{9}^{\mathrm{SM}}\left(m_{b}\right) \approx+4.2$

$$
C_{10}^{\mathrm{SM}}\left(m_{b}\right) \approx-4.4
$$

$$
\square C_{9}^{\mathrm{SM}}\left(m_{b}\right) \approx-C_{10}^{\mathrm{SM}}\left(m_{b}\right)
$$ also the lepton current has nearly $V$ - $A$ structure

## [Bobeth, Misiak, Urban, 99]

[Khodjamirian et al., 10]

## Let's now turn to Q1:

Can we (easily) make sense of data $\mathbf{( 1}$ to $\mathbf{5}$ ?

It is highly non-trivial that a simple consistent BSM picture exists to describe the above data $\mathbf{1}$ to $\boldsymbol{\bullet}$

- Consider the following Hamiltonian

$$
\left.H_{\mathrm{SM}+\mathrm{NP}}(\bar{b} \rightarrow \bar{s} \mu \mu)=-\frac{4 G_{F}}{\sqrt{2}} V_{t b}^{*} V_{t s} \frac{\alpha_{\mathrm{em}}}{4 \pi}\left[\bar{b}_{L} \gamma^{\lambda} s_{L} \cdot\left(C_{9}^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu\right)+C_{10}^{(u)} \bar{\mu} \gamma_{\lambda} \gamma_{5} \mu\right)\right]
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i.e. in the SM also the lepton current has nearly $V-A$ structure

- Note: $\quad C_{9}^{\mathrm{SM}}\left(m_{b}\right) \approx+4.2$

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## [Bobeth, Misiak, Urban, 99]

[Khodjamirian et al., 10]

Let's assume the above $V-A$ structure to hold also beyond the SM, namely

$$
C_{9}^{(\ell)} \approx-C_{10}^{(\ell)} \quad \text { with } \quad C_{9,10}^{(\ell)}=C_{9,10}^{\mathrm{SM}}+C_{9,10}^{(\ell), \mathrm{NP}}
$$

Our main motivation is phenomenological: it fits the data.
However, there is more: see later
D. Guadagnoli, Flavor physics

## Model example:

Glashow, DG, Lane, PRL 2015

- Our model requirements are:

$$
\begin{aligned}
& -C_{9}^{(\ell)} \approx-C_{10}^{(\ell)} \quad(V-A \text { structure }) \\
& =\left|C_{9, \mathrm{NP}}^{(\mu)}\right| \gg\left|C_{9, \mathrm{NP}}^{(e)}\right| \quad \text { (LFNU) }
\end{aligned}
$$

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- This structure can be generated from a purely $3^{\text {rd }}$-generation interaction of the kind

$$
\begin{gathered}
H_{\mathrm{NP}}=G \bar{b}_{L}^{\prime} \gamma^{\lambda} b_{L}^{\prime} \bar{\tau}_{L}^{\prime} \gamma_{\lambda} \tau_{L}^{\prime} \\
\text { with } G=1 / \Lambda_{\mathrm{NP}}^{2} \ll G_{F}
\end{gathered}
$$

expected e.g. in partial-compositeness frameworks (see later)

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- Note: primed fields
- Fields are in the "gauge" basis (= primed)
- They need to be rotated to the mass eigenbasis

$$
\begin{aligned}
{b^{\prime}}_{L} \equiv\left(d_{L}^{\prime}\right)_{3}=\left(U_{L}^{d}\right)_{3 i} \underbrace{\substack{\text { mass } \\
\text { basis }}}_{\left(d_{L}\right)_{i}} \\
\left.\tau_{L}^{\prime} \equiv\left(\ell_{L}^{\prime}\right)_{3}=\left(U_{L}^{\ell}\right)_{3 i} \ell_{L}\right)_{i}
\end{aligned}
$$

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$$

expected e.g. in partial-compositeness frameworks (see later)

- Note: primed fields
- Fields are in the "gauge" basis (= primed)
- They need to be rotated to the mass eigenbasis
- This rotation induces LFNU and LFV effects


## Explaining $\mathbf{b} \rightarrow \mathbf{s}$ data

- Recalling our full Hamiltonian

$$
H_{\mathrm{SM}+\mathrm{NP}}(\bar{b} \rightarrow \bar{s} \mu \mu)=-\frac{4 G_{F}}{\sqrt{2}} V_{t b}^{*} V_{t s} \frac{\alpha_{\mathrm{em}}}{4 \pi}\left[\bar{b}_{L} \gamma^{\lambda} s_{L} \cdot\left(C_{9}^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu+C_{10}^{(\mu)} \bar{\mu} \gamma_{\lambda} \gamma_{5} \mu\right)\right]
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k_{S M}(S M \text { norm. factor) }
$$

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$$

the shift to the $C_{9}$ Wilson coeff. in the $\mu \mu$-channel becomes

$$
k_{\mathrm{SM}} C_{9}^{(u)}=k_{\mathrm{SM}} C_{9, \mathrm{SM}}+\frac{G}{2}\left(U_{L}^{d}\right)_{33}^{*}\left(U_{L}^{d}\right)_{32}\left|\left(U_{L}^{p}\right)_{32}\right|^{2}
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The NP contribution has opposite sign than the SM one if

$$
G\left(U_{L}^{d}\right)_{32}<0
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\end{aligned}
$$

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$$
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$$

- On the other hand, in the ee-channel

$$
k_{\mathrm{SM}} C_{9}^{(e)}=k_{\mathrm{SM}} C_{9, \mathrm{SM}}+\frac{G}{2}\left(U_{L}^{d}\right)_{33}^{*}\left(U_{L}^{d}\right)_{32}\left|\left(U_{L}^{\ell}\right)_{31}\right|^{2}
$$

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H_{\mathrm{SM}+\mathrm{NP}}(\bar{b} \rightarrow \bar{s} \mu \mu)=-\frac{4 G_{F}}{\sqrt{2}} V_{t b}^{*} V_{t s} \frac{\alpha_{\mathrm{em}}}{4 \pi}\left[\bar{b}_{L} \gamma^{\lambda} s_{L} \cdot\left(C_{9}^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu+C_{10}^{(\mu)} \bar{\mu} \gamma_{\lambda} \gamma_{5} \mu\right)\right]
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& =\beta_{\mathrm{SM}}+\quad+\quad \beta_{\mathrm{NP}}
\end{aligned}
$$

The NP contribution has opposite sign than the SM one if

$$
G\left(U_{L}^{d}\right)_{32}<0
$$

The NP contrib. in the eechannel is negligible, as

$$
\left|\left(U_{L}^{\ell}\right)_{31}\right|^{2} \ll\left|\left(U_{L}^{\ell}\right)_{32}\right|^{2}
$$

## Explaining $\mathbf{b} \rightarrow \mathbf{s}$ data

- The above setup implies:
$-\left|C_{9}\right| \simeq\left|C_{10}\right|$ also beyond the SM
- Corrections in the $\mu$ channel much larger than in the electron channel


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ㄷ

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R_{K} \approx \frac{\left|C_{9}^{(u)}\right|^{2}+\left|C_{10}^{(u)}\right|^{2}}{\left|C_{9}^{(e)}\right|^{2}+\left|C_{10}^{(e)}\right|^{2}} \simeq \frac{2\left|C_{10}^{\mathrm{SM}}+\delta C_{10}\right|^{2}}{2\left|C_{10}^{\mathrm{SM}}\right|^{2}}
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factors of 2:
equal contributions from $\left|C_{9}\right|^{2}$ and $\left|C_{10}\right|^{2}$

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$$

equal contributions from $\left|C_{9}\right|^{2}$ and $\left|C_{10}\right|^{2}$

- Note as well

$$
0.77 \pm 0.20=\frac{B R\left(B_{s} \rightarrow \mu \mu\right)_{\exp }}{B R\left(B_{s} \rightarrow \mu \mu\right)_{\mathrm{SM}}}=\frac{B R\left(B_{s} \rightarrow \mu \mu\right)_{\mathrm{SM}+\mathrm{NP}}}{B R\left(B_{s} \rightarrow \mu \mu\right)_{\mathrm{SM}}}=\frac{\left|C_{10}^{\mathrm{SM}}+\delta C_{10}\right|^{2}}{\left|C_{10}^{\mathrm{SM}}\right|^{2}}
$$

## Explaining $\mathbf{b} \rightarrow \mathbf{s}$ data

- The above setup implies:
$-\left|C_{9}\right| \simeq\left|C_{10}\right|$ also beyond the $S M$
- Corrections in the $\mu$ channel much larger than in the electron channel

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$$

implying (within our model) the correlations

$$
\frac{B R\left(B_{s} \rightarrow \mu \mu\right)_{\exp }}{B R\left(B_{s} \rightarrow \mu \mu\right)_{\mathrm{SM}}} \simeq R_{K} \simeq \frac{B R\left(B^{+} \rightarrow K^{+} \mu \mu\right)_{\exp }}{B R\left(B^{+} \rightarrow K^{+} \mu \mu\right)_{\mathrm{SM}}}
$$


D. Guadagnoli, Flavor physics

## Explaining $\mathbf{b} \rightarrow \mathbf{s}$ data

- So, the equal-size, opposite-sign corrections to $C_{9}{ }^{\text {SM }}$ and to $C_{10}{ }^{\text {SM }}$ (in the $\mu \mu$-channel only) would account for:
- $\quad R_{K}$ lower than 1
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- This pattern of corrections turns out to also accommodate the anomaly [LHCb, PRL 2013] in $B \rightarrow K^{*} \mu \mu$ angular data measured by $L H C b$, especially in $P_{5}{ }^{\prime}$


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- So, the equal-size, opposite-sign corrections to $C_{9}{ }^{S M}$ and to $C_{10}{ }^{\text {SM }}$ (in the $\mu \mu$-channel only) would account for:
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- This pattern of corrections turns out to also accommodate the anomaly [LHCb, PRL 2013] in $B \rightarrow K^{*} \mu \mu$ angular data measured by LHCb, especially in $P_{5}{ }^{\prime}$
- A fully quantitative test requires a global fit.

See in particular [Ghosh, Nardecchia, Renner, JHEP '14] and [Altmannshofer, Straub, EPJC '15]

```
new physics contributions to the Wilson coefficients. We find that the by far largest de-
crease in the \chi}\mp@subsup{\chi}{}{2}\mathrm{ can be obtained either by a negative new physics contribution to C}\mp@subsup{C}{9}{}\mathrm{ (with
C
(with C}\mp@subsup{C}{9}{\textrm{NP}}~-12%\times\mp@subsup{C}{9}{\textrm{SM}}\mathrm{ ). A positive NP contribution to Clo alone would also improve the
fit, although to a lesser extent.
[Altmannshofer, Straub, EPJC '15]
```


## LFV model signatures

As mentioned: if $R_{K}$ is signaling BSM LFNU, then expect BSM LFV as well

$$
\nabla \frac{B R\left(B^{+} \rightarrow K^{+} \mu e\right)}{B R\left(B^{+} \rightarrow K^{+} \mu \mu\right)}=\frac{\left|\delta C_{10}\right|^{2}}{\left|C_{10}^{S M}+\delta C_{10}\right|^{2}} \cdot \frac{\left|\left(U_{L}^{\ell}\right)_{31}\right|^{2}}{\left|\left(U_{L}^{\ell}\right)_{32}\right|^{2}} \cdot 2
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$$
\square B R\left(B^{+} \rightarrow K^{+} \mu e\right)<2.2 \times 10^{-8} \cdot \frac{\left|\left(U_{L}^{\ell}\right)_{31}\right|^{2}}{\left|\left(U_{L}^{\ell}\right)_{32}\right|^{2}}
$$

The current $B R(B+\rightarrow K+\mu e)$ limit yields the weak bound

$$
\left|\left(U_{L}^{t}\right)_{31} 1\left(U_{L}^{t}\right)_{32}\right|<3.7
$$

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$\checkmark \quad B R\left(B^{+} \rightarrow K^{+} \mu \tau\right) \quad$ would be even more promising, as it scales with $\left|\left(U_{L}^{\ell}\right)_{33} /\left(U_{L}^{\ell}\right)_{32}\right|^{2}$

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$\checkmark \quad$ An analogous argument holds for purely leptonic modes
$\checkmark \quad$ There is even an interesting signature outside $B$ physics: $K \rightarrow(\pi) \ell \ell^{\prime}$ It is measurable at NA62, that just started taking data

## More signatures



- Being defined above the EWSB scale, our assumed operator $G \bar{b}^{\prime}{ }_{L} \gamma^{\lambda} b^{\prime}{ }_{L} \bar{\tau}^{\prime}{ }_{L} \gamma_{\lambda} \tau^{\prime}{ }_{L}$ must actually be made invariant under $S U(3)_{c} \times S U(2)_{L} \times U(1)_{Y}$


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```
    Bhattacharya,
\[
\bar{b}_{L}^{\prime} \gamma^{\lambda} b^{\prime}{ }_{L} \bar{\tau}^{\prime}{ }_{L} \gamma_{\lambda} \tau^{\prime}{ }_{L}
\]
```



## More signatures

- Being defined above the EWSB scale, our assumed operator $G \bar{b}^{\prime}{ }_{L} \gamma^{\lambda} b^{\prime}{ }_{L} \bar{\tau}^{\prime}{ }_{L} \gamma_{\lambda} \tau^{\prime}{ }_{L}$ must actually be made invariant under $\operatorname{SU}(3)_{c} \times S U(2)_{L} \times U(1)_{Y}$

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See: 
    Bhattacharya, Dara, PLB 15
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- Thus, the generated structures are all of:
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Bhattacharya, Datta, London,
Shivashankara, PLB 15

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$b^{\prime} b^{\prime} \tau^{\prime} \tau^{\prime}$,
$t^{\prime} b^{\prime} \tau^{\prime} v_{\tau}^{\prime}$
- After rotation to the mass basis (unprimed), the last structure contributes to $\Gamma\left(b \rightarrow c \tau \bar{v}_{i}\right)$

$$
\square \text { Can explain BaBar }+ \text { Belle }+L H C b \text { deviations on } R\left(D^{(*)}\right)=\frac{B R\left(\bar{B} \rightarrow D^{(*)+} \tau^{-} \bar{v}_{\tau}\right)}{B R\left(\bar{B} \rightarrow D^{(*)+} \ell^{-} \bar{v}_{\ell}\right)}
$$

## Some models explaining $R_{k}$ and $R\left(D^{*}\right)$

- Introduce one single leptoquark scalar, transforming as (3, 1, $-1 / 3$ ) under $\operatorname{SU}(3)_{c} \times S U(2)_{L} \times U(1)_{Y}$
- One coupling does all the job: $\bar{Q}^{c}{ }_{L i} \lambda_{i j} i \tau_{2} L_{L j} \phi$


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- With $M_{\phi} \sim 1 \mathrm{TeV}$ and $O(1)$ generation-diagonal couplings, contributions are just the right size


## One model explaining all flavor

 anomalies and the diphoton resonance- New non-Abelian strongly interacting sector with $N_{\text {TC }}$ new "techni-fermions"

$$
\begin{aligned}
& \text { Buttazzo, Greljo, } \\
& \text { Isidori, Marzocca } \\
& \text { 1604.030^с }
\end{aligned}
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One of the pNGB is the $750-G e V$ state seen by Atlas \& CMS It couples to 2 gluons and decays to $2 \gamma$ via the anomaly

## One model explaining all flavor

 anomalies and the diphoton resonance: continued- There are also vector mesons, like QCD's rho.

$$
\begin{aligned}
& \text { Buttazzo, Greljo, } \\
& \text { Isidori, Marzocca } \\
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Their coupling to quarks and leptons nicely explains the flavor anomalies.

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- Vector mesons couple to techni-baryons, which in turn linearly mix with SM fermions.

- To explain the flavor deviations, the mixing needs be hierarchical across generations (largest for the $3^{\text {rd }}$ one, as in partial compositeness)
- Integrating out the vector mesons then yields automatically (among the others) the effective operator

$$
H_{\mathrm{NP}}=G \bar{b}_{L}^{\prime} \gamma^{\lambda} b_{L}^{\prime} \bar{\tau}_{L}^{\prime} \gamma_{\lambda} \tau_{L}^{\prime}
$$

proposed in [Glashow, DG, Lane, PRL 15]

## Conclusions

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- Early to draw conclusions. But Run II will provide a definite answer
- Timely to propose further tests. One promising direction is that of LFV. Plenty of channels, many of which largely untested.
- Most promising theory direction (to me): Seeking for a correlated explanation of the diphoton excess and of the flavor anomalies.


## Spares

## Frequently made objection:

what about the SM? It has LFNU, but no LFV

Take the SM with zero v masses.

- Charged-lepton Yukawa couplings are LFNU, but they are diagonal in the mass eigenbasis (hence no LFV)

Or more generally, take the SM plus a minimal mechanism for $v$ masses.

- Physical LFV will appear in W couplings, but it's suppressed by powers of $\left(m_{v} / m_{w}\right)^{2}$

Bottom line: in the $S M+v$ there is LFNU, but LFV is nowhere to be seen (in decays)

- But nobody ordered that the reason (=tiny $m_{1}$ ) behind the above conclusion be at work also beyond the SM



## More quantitative LFV predictions

- More quantitative LFV predictions require knowledge of the $U_{L}{ }^{\ell}$


## Reminder:

$$
\left(U_{L}^{\ell}\right)^{\dagger} Y_{\ell} U_{R}^{\ell}=\hat{Y}_{\ell}
$$

- One approach:

DG, Lane, 1507.01412

- Appelquist-Bai-Piai ansatz:
the flavor-SU(3) rotations are not all independent. Choosing 3 to be the independent ones allows to predict one SM Yukawa in terms of the other two.
- One can thereby determine $Y_{t}$ in terms of $Y_{u}$ and $Y_{d}$
- But we don't know $Y_{u}$ and $Y_{d}$ entirely, so we take an (independently motivated) model for them, reproducing quark masses and the CKM matrix [Martin-Lane, PRD 2005].
- Another approach:

Boucenna, Valle, Vicente, PLB 2015

- One has $\left(U_{L}^{\ell}\right)^{\dagger} U_{L}^{\nu}=$ PMNS matrix
- Taking $U_{L}^{\nu}=1, U_{L}^{\ell}$ can be univocally predicted


## More quantitative LFV predictions

LFV predictions in one of the two scenarios of [DG, Lane]

|  | $B^{+} \rightarrow K^{+} \mu^{ \pm} \tau^{\mp}$ | $B^{+} \rightarrow K^{+} e^{ \pm} \tau^{\mp}$ | $B^{+} \rightarrow K^{+} e^{ \pm} \mu^{\mp}$ |
| :---: | :---: | :---: | :---: |
|  | $1.14 \times 10^{-8}$ | $3.84 \times 10^{-10}$ | $0.52 \times 10^{-9}$ |
| Exp: | $<4.8 \times 10^{-5}$ | $<3.0 \times 10^{-5}$ | $<9.1 \times 10^{-8}$ |


|  | $B_{s} \rightarrow \mu^{ \pm} \boldsymbol{r}^{\mp}$ | $B_{s} \rightarrow e^{ \pm} \tau^{\mp}$ | $B_{s} \rightarrow e^{ \pm} \mu^{\mp}$ |
| :---: | :---: | :---: | :---: |
| $1.37 \times 10^{-8}$ | $4.57 \times 10^{-10}$ | $1.73 \times 10^{-12}$ |  |
| Exp: | - | - | $<1.1 \times 10^{-8}$ |

All predictions are phase-space corrected.

