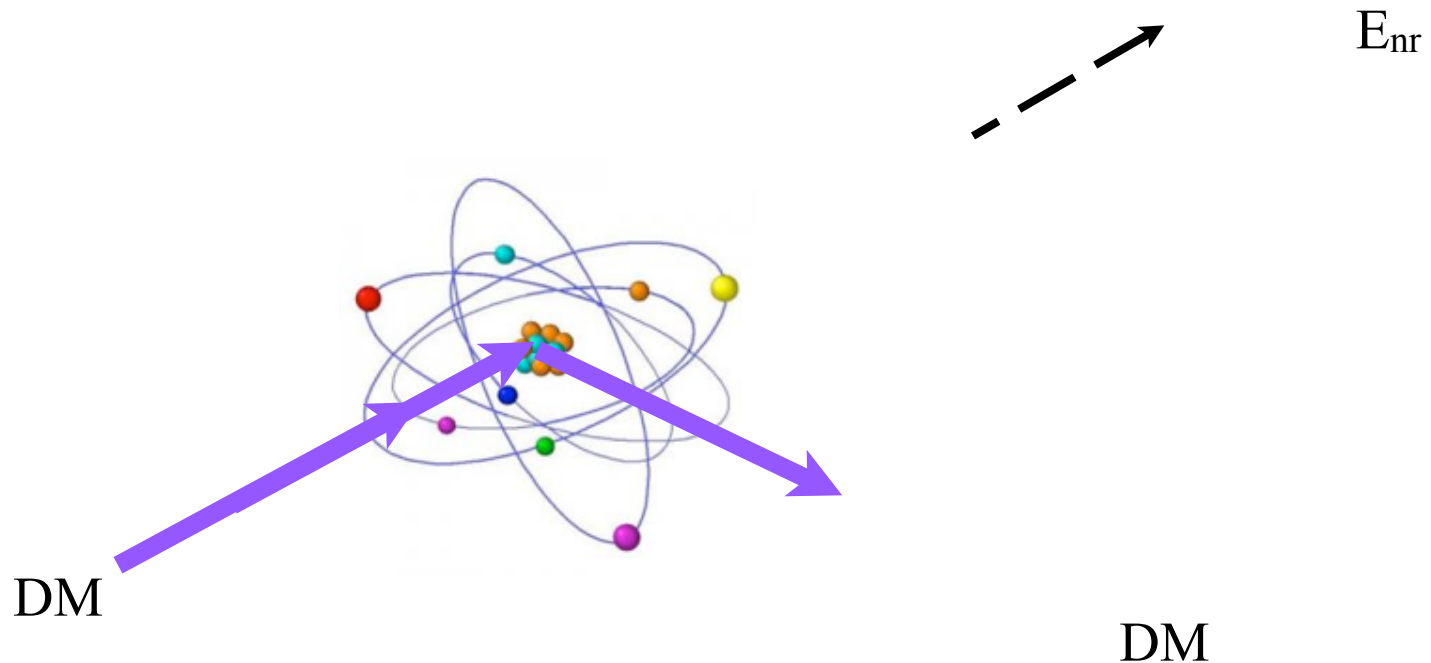


Status of the XENON Program: The fight against the background

J. Masbou, on behalf of the
XENON Collaboration

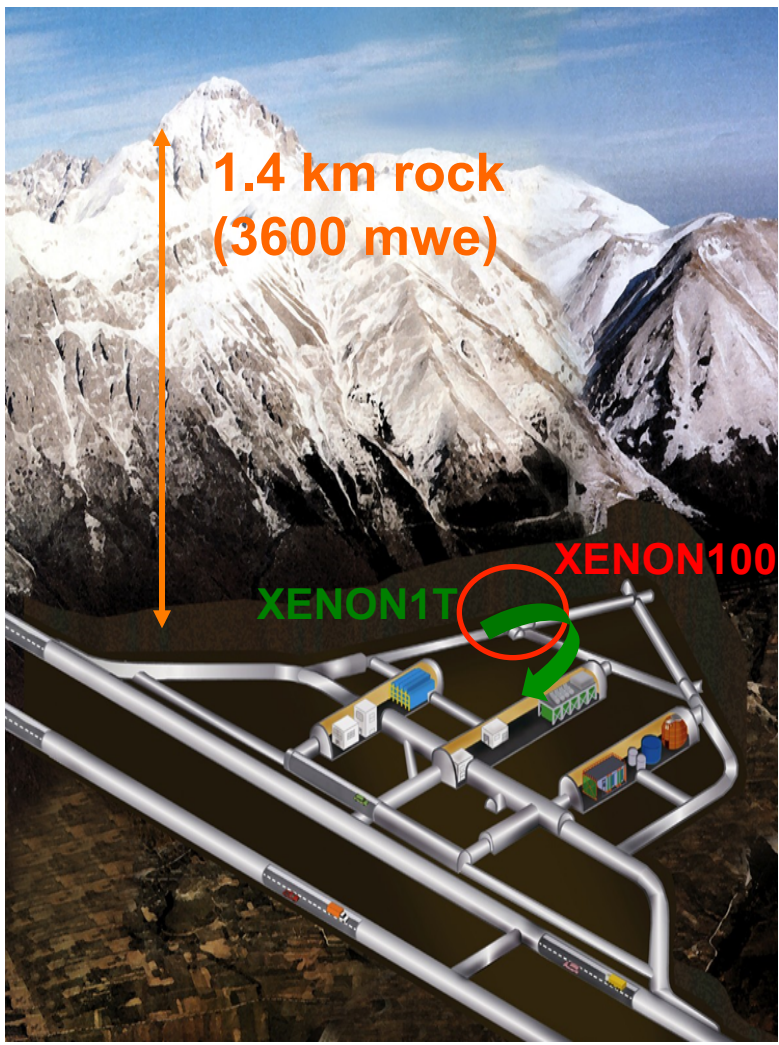
Our goal

Discover Dark Matter
with the most sensitive liquid xenon imaging detector
located in LNGS

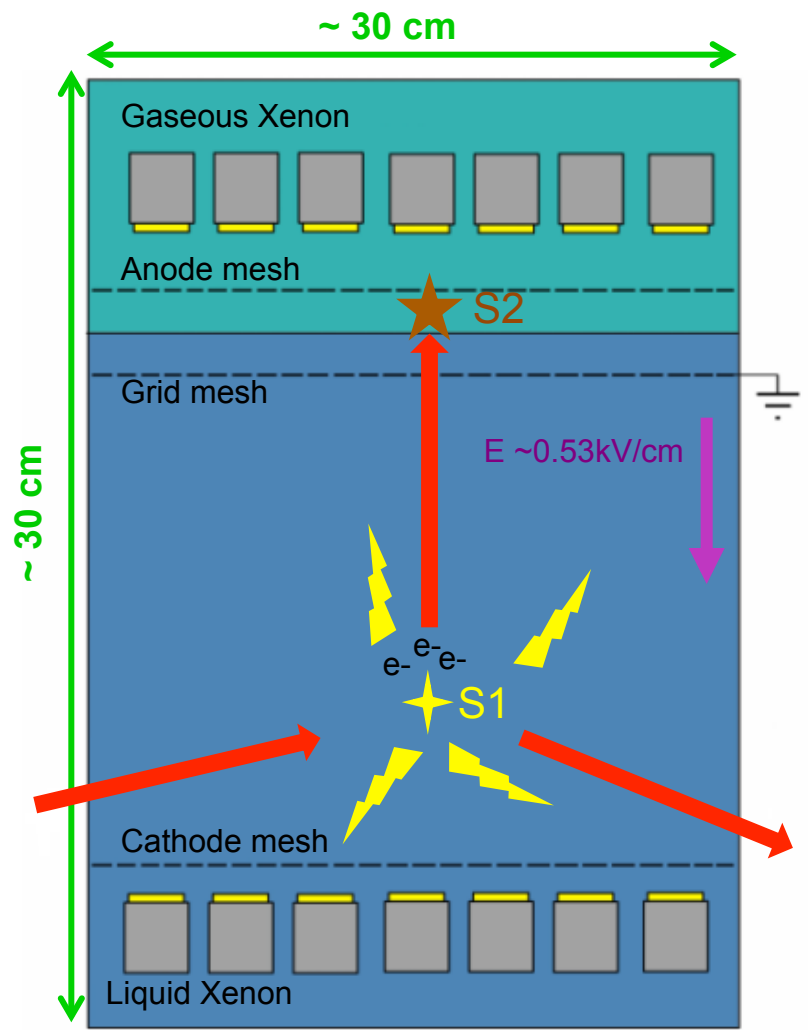


Location of the XENON experiment & Collaboration

21 Institutes
~150 members



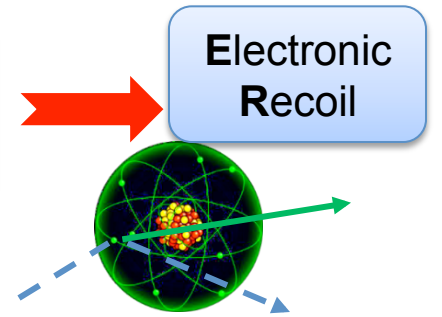
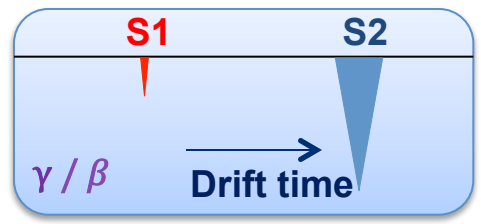
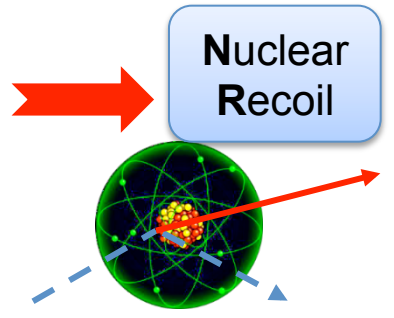
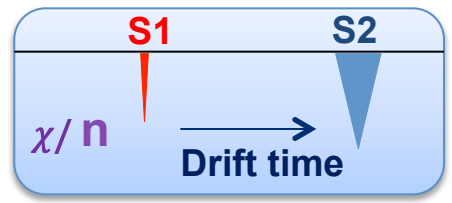
Two phase XENON TPC principle



S1:
 → Photon ($\lambda = 178 \text{ nm}$) from Scintillation process
 → Detected by PMTs (mainly bottom array)

S2:
 → Electrons drift
 → Extraction in gaseous phase
 → Proportional scintillation light

3D reconstruction :
 → X,Y from top array
 → Z from Drift time



E. Aprile et al. (XENON100), *Astropart. Phys.* 35, 573-590 (2012)

The XENON Dark Matter Program



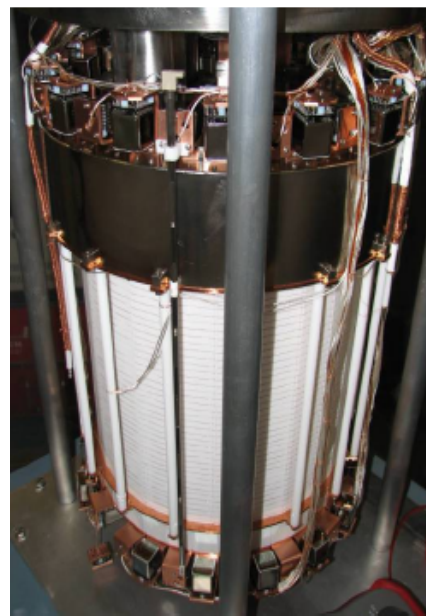
XENON10

Achieved (2007)

$\sigma_{\text{SI}} = 8.8 \cdot 10^{-44} \text{ cm}^2 @ 100 \text{ GeV}/c^2$
Phys.Rev.Lett. 100 (2008) 021303

Light DM:

$\sigma_{\text{SI}} = 7 \cdot 10^{-42} \text{ cm}^2 @ 7 \text{ GeV}/c^2$
Phys.Rev.Lett. 107 (2011) 051301



XENON100

Achieved (2012)

$\sigma_{\text{SI}} = 2.0 \cdot 10^{-45} \text{ cm}^2 @ 55 \text{ GeV}/c^2$
*E. Aprile et al. (XENON100),
Phys. Rev. Lett. 109 (2012)
arXiv:1207.5988*

**In operation
since 2009**



XENON1T

Projected (2017)

$\sigma_{\text{SI}} = \sim 2 \cdot 10^{-47} \text{ cm}^2$

**Science data by spring
2016**

Upgrade : XENONnT

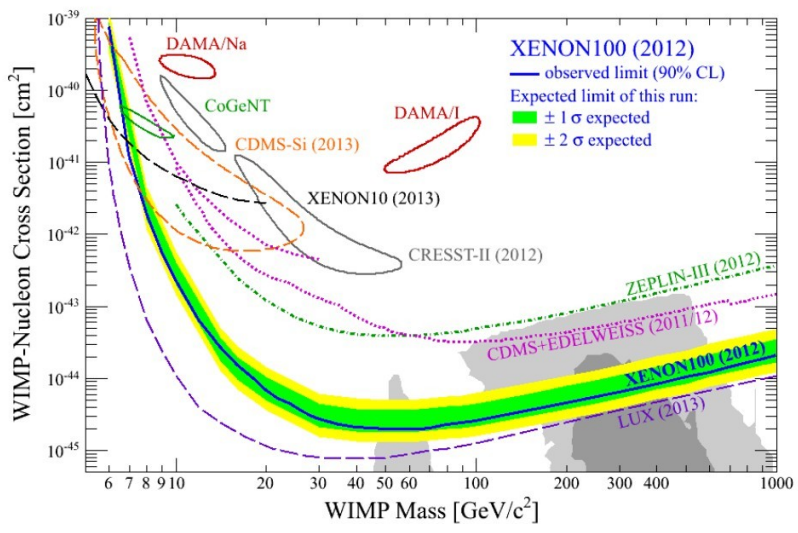
$\sigma_{\text{SI}} = \sim 2 \cdot 10^{-48} \text{ cm}^2$

XENON100 : Past Achievements

225 live days x 34 kg exposure

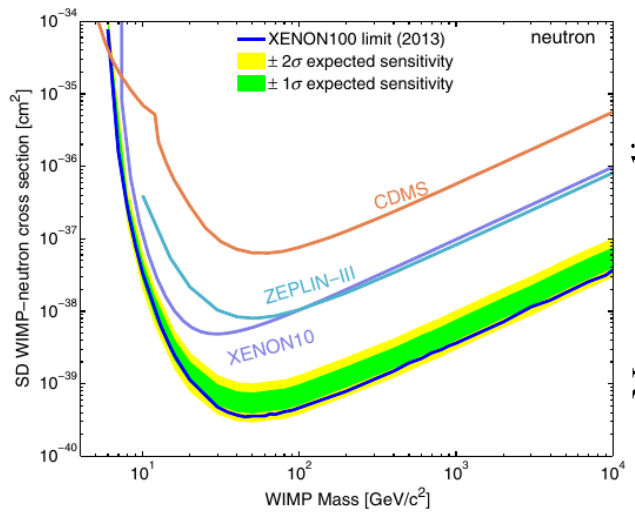
Profile likelihood analysis allows to set limits in the (m_χ, σ) parameter space

Phys. Rev. Let. 109, 181301 (2012)

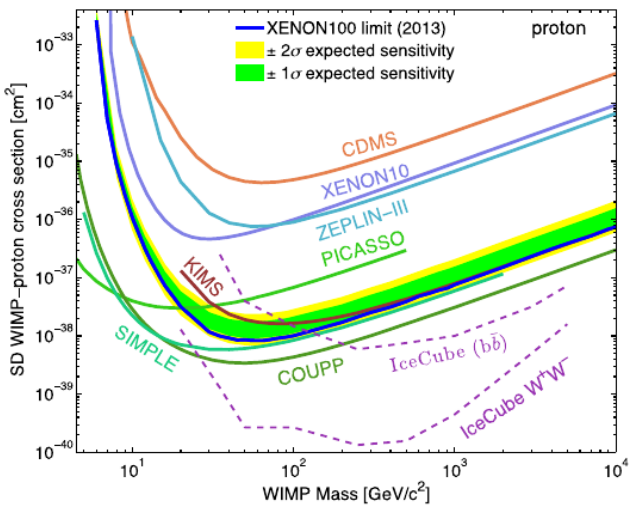


Spin-independent interaction

Phys. Rev. Let. 111, 021301 (2013)



Neutron coupling



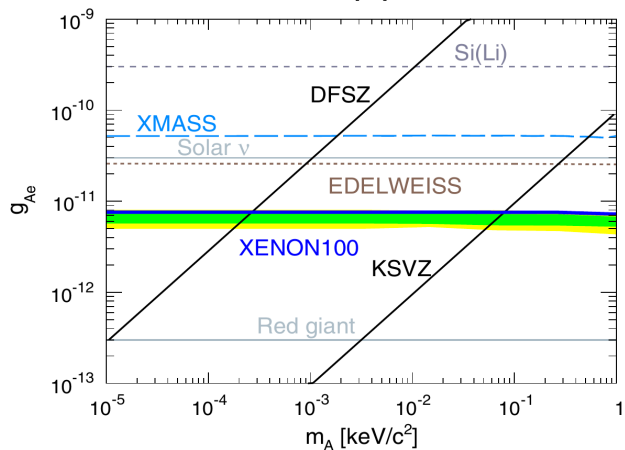
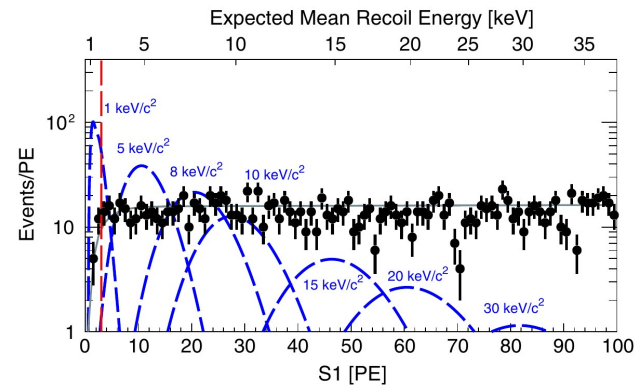
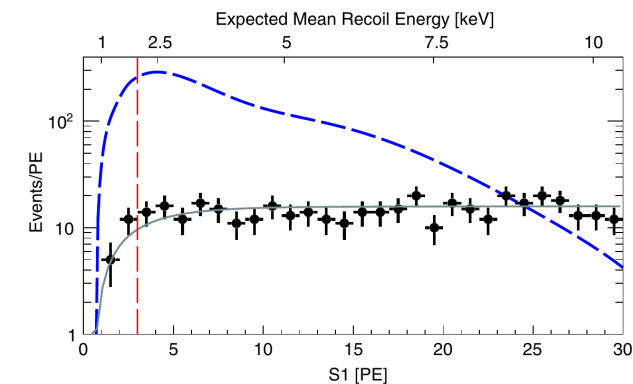
Proton coupling

Spin-dependent interaction

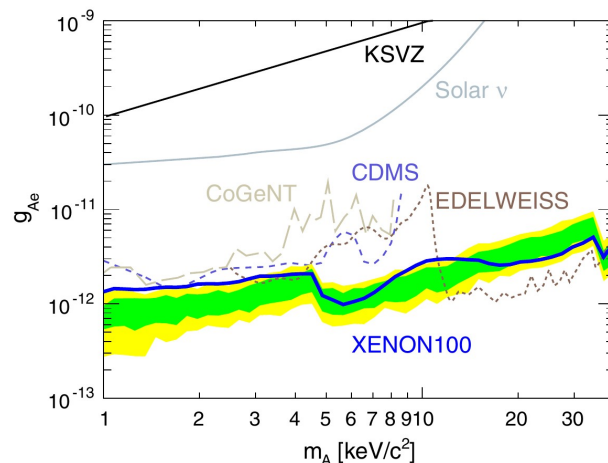
XENON100 : Past Achievements

225 live days x 34 kg exposure

- First axion results from the XENON100 experiment analyzing ER data
- Probing axion-electron coupling constant by exploiting the axioelectric effect in LXe



Solar Axions



Galactic ALPs (DM candidates)

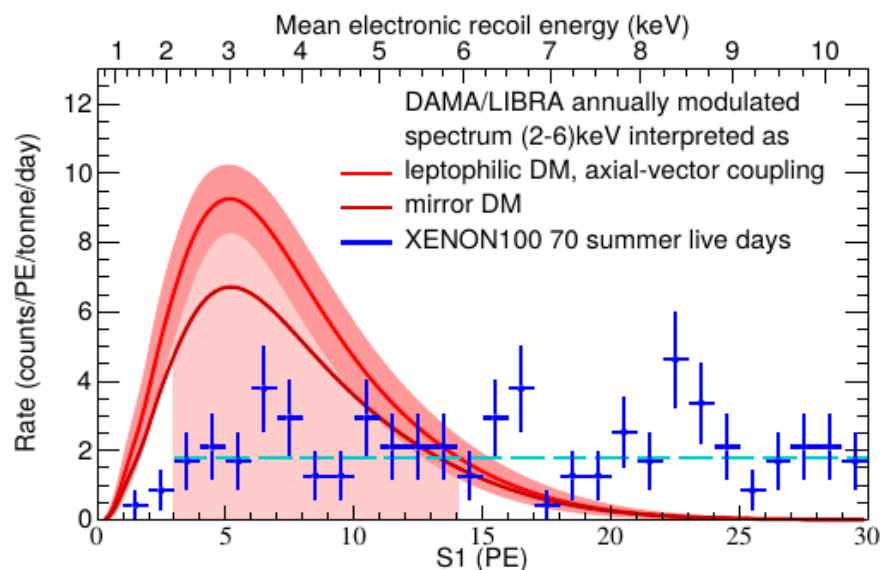
Phys. Rev. D 90, 062009
(2014)

XENON100 : Past Achievements

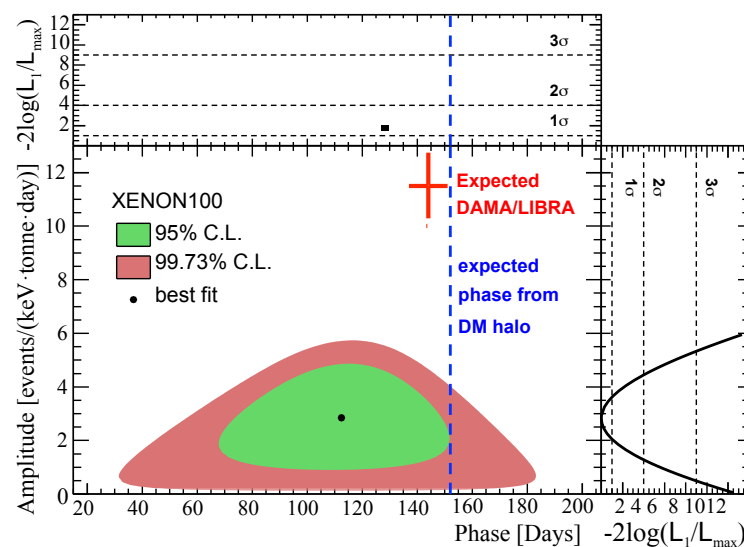
225 live days x 34 kg exposure

- Exclusion of several types of DM models as the cause of the annual modulation
- The DM interpretation of DAMA/ LIBRA annual modulation as being due to WIMPs electron scattering through axial vector coupling is disfavored at 4.8-sigma from a PL analysis

Science 349, 851 (2015)



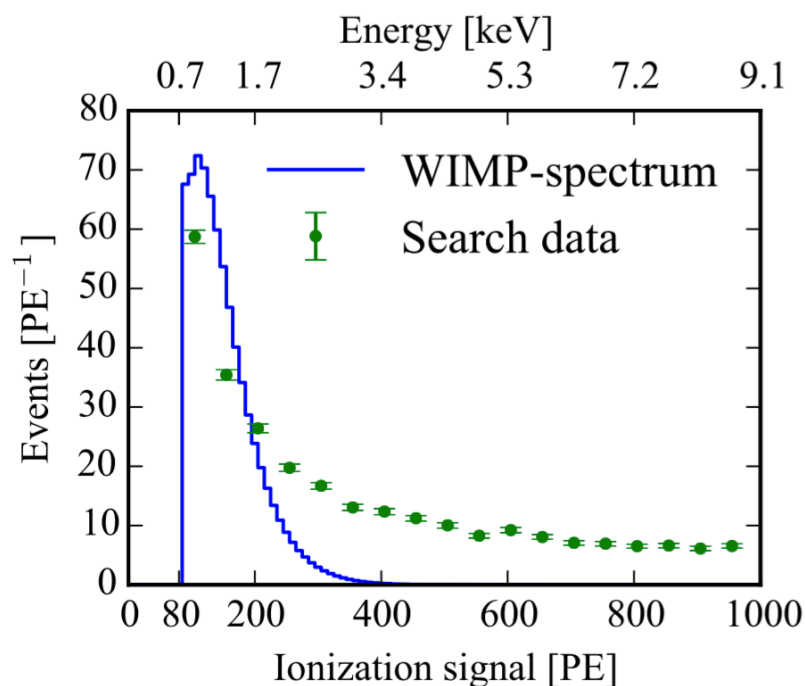
PRL 115, 091302 (2015)



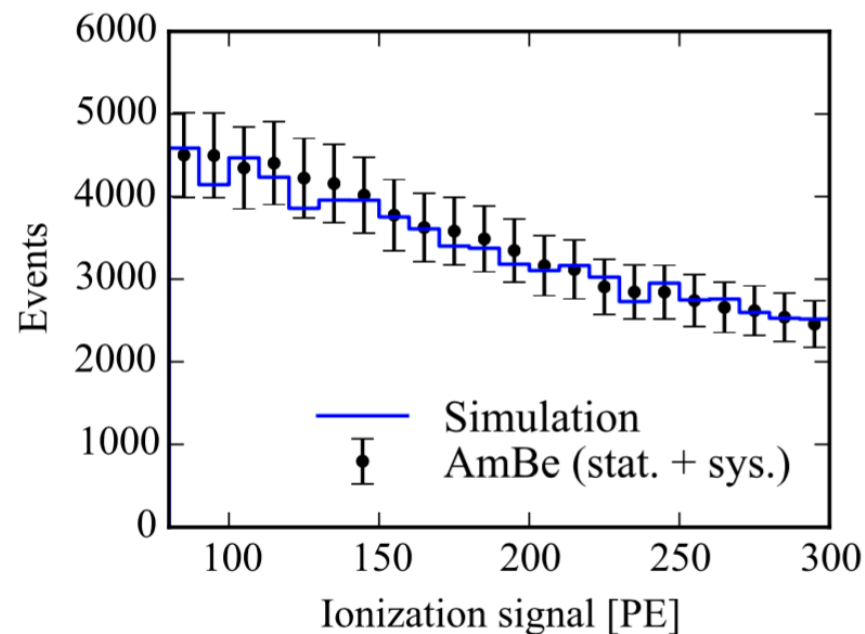
XENON100 : Low-mass dark matter

225 live days x 34 kg exposure

- No S1 signal → No 3D position reconstruction
- 0,7 keV threshold nuclear recoil

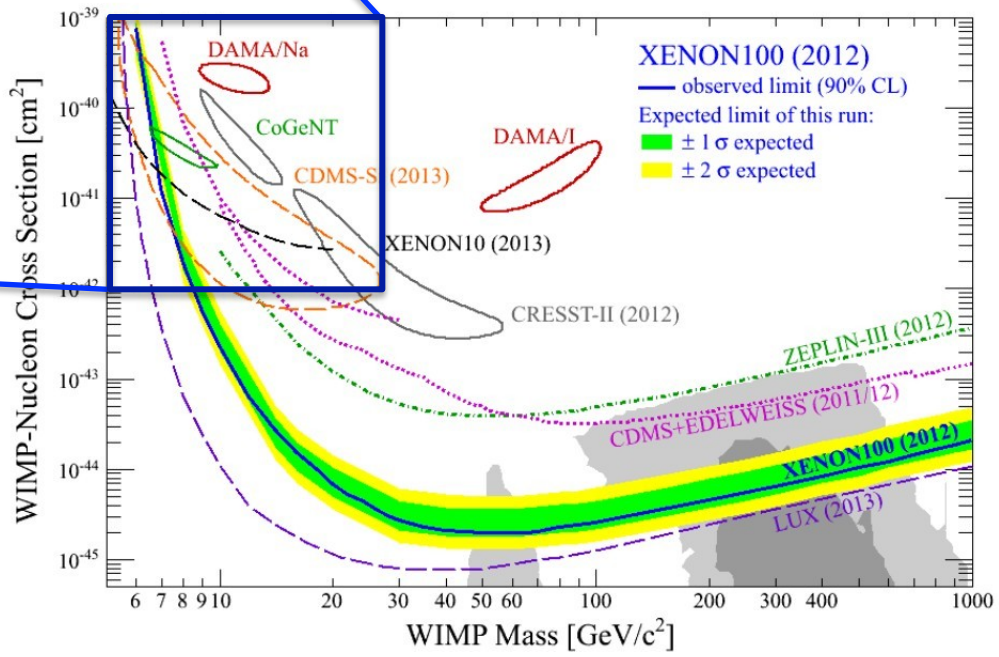
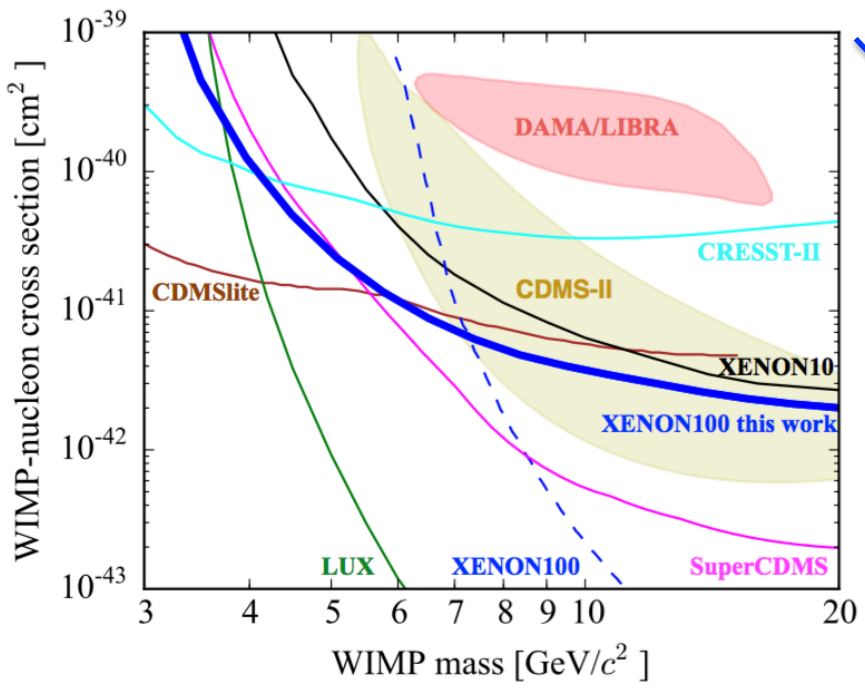


Expected Spectrum for a 6 GeV WIMP



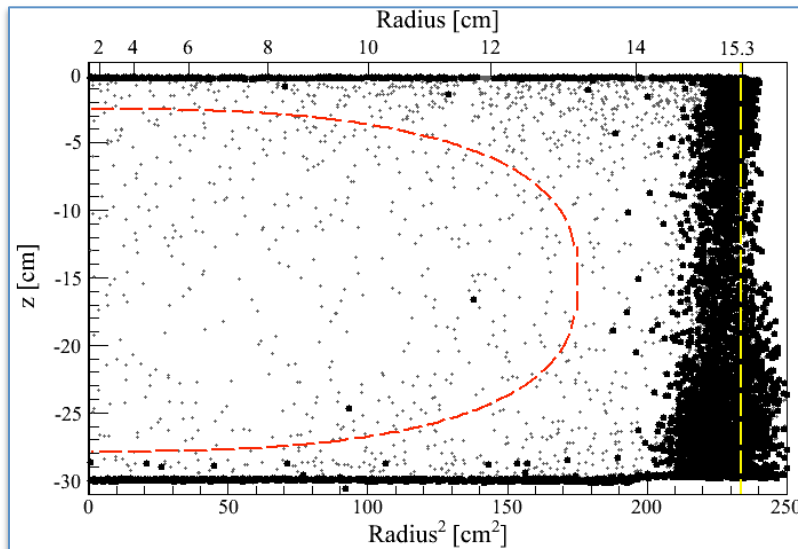
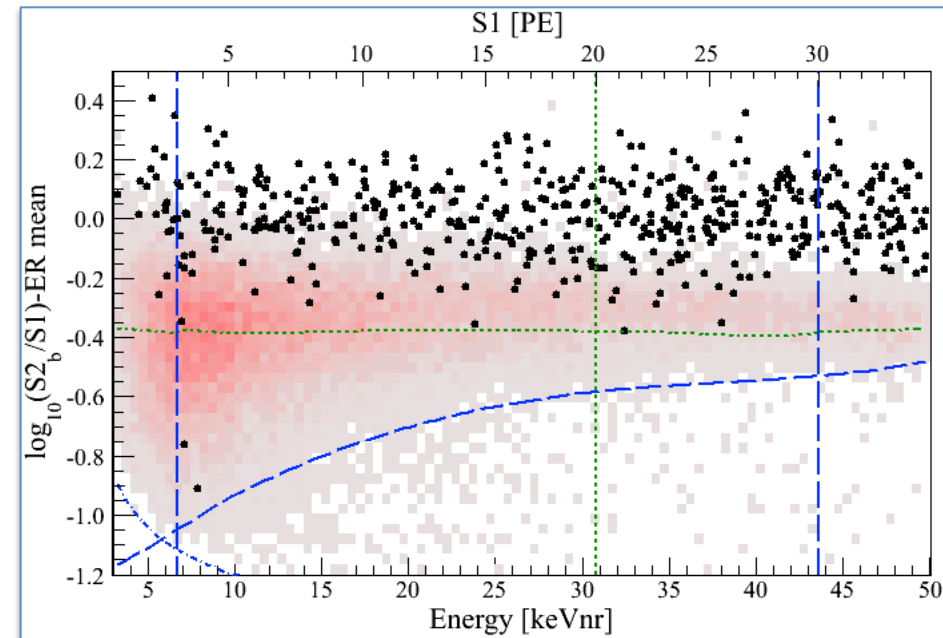
S2 spectrum of ²⁴¹AmBe

XENON100 : Low-mass dark matter



XENON100 : Blind analysis

- XENON100 did a blind analysis
- Event discrimination by S2/S1 separation
- Defined WIMP searching region:
 - S1 with benchmark region (3 - 30 pe)
 - S2 threshold cut (S2 > 150 pe)
 - 99.75 % ER rejection line

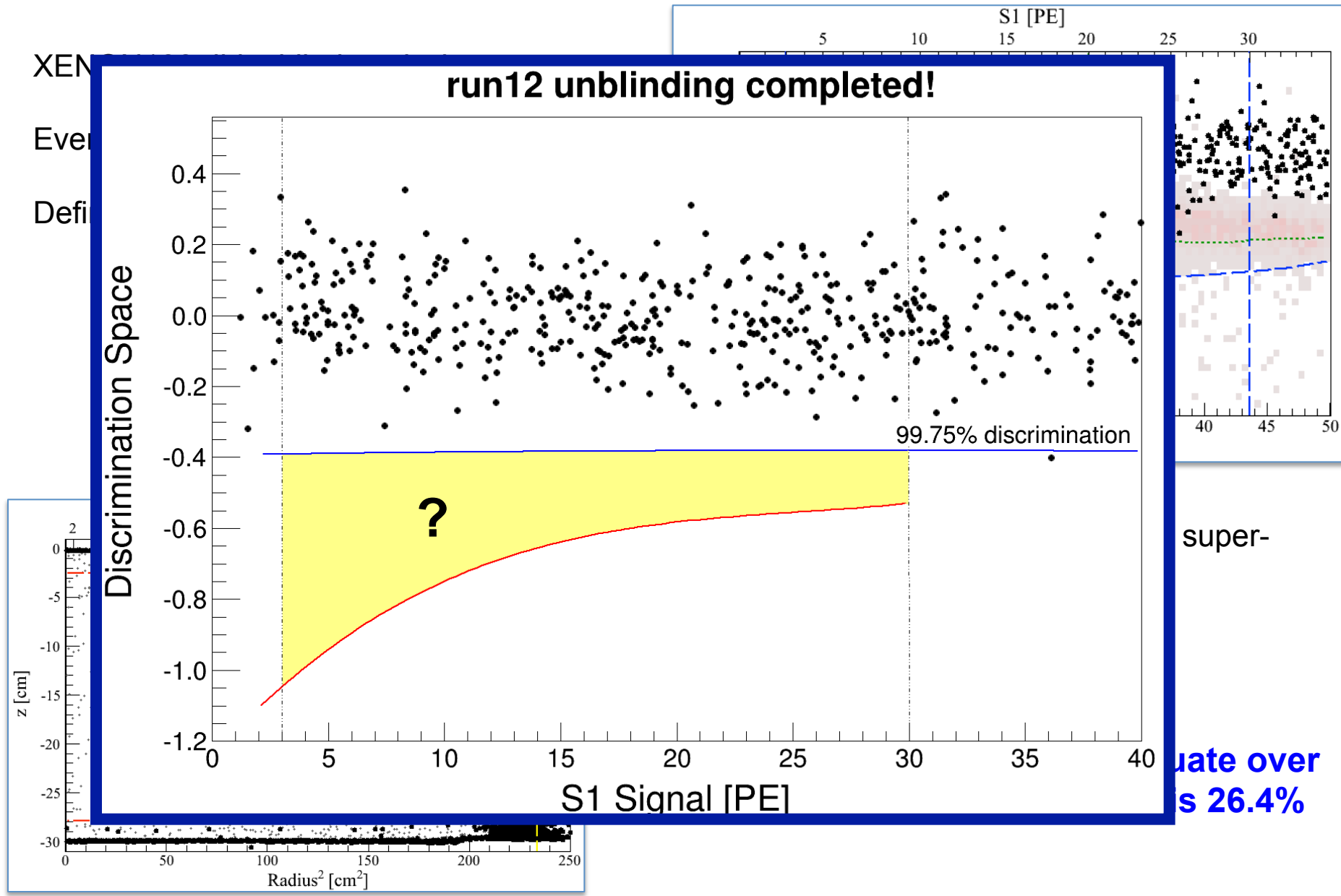


- Event rejection by defining a 34kg super-ellipse
- Double scatters excluded

Probability that 2 events fluctuate over the background expectation is 26.4%

XENON100 : New result is coming!

- XENON
- Event
- Defi

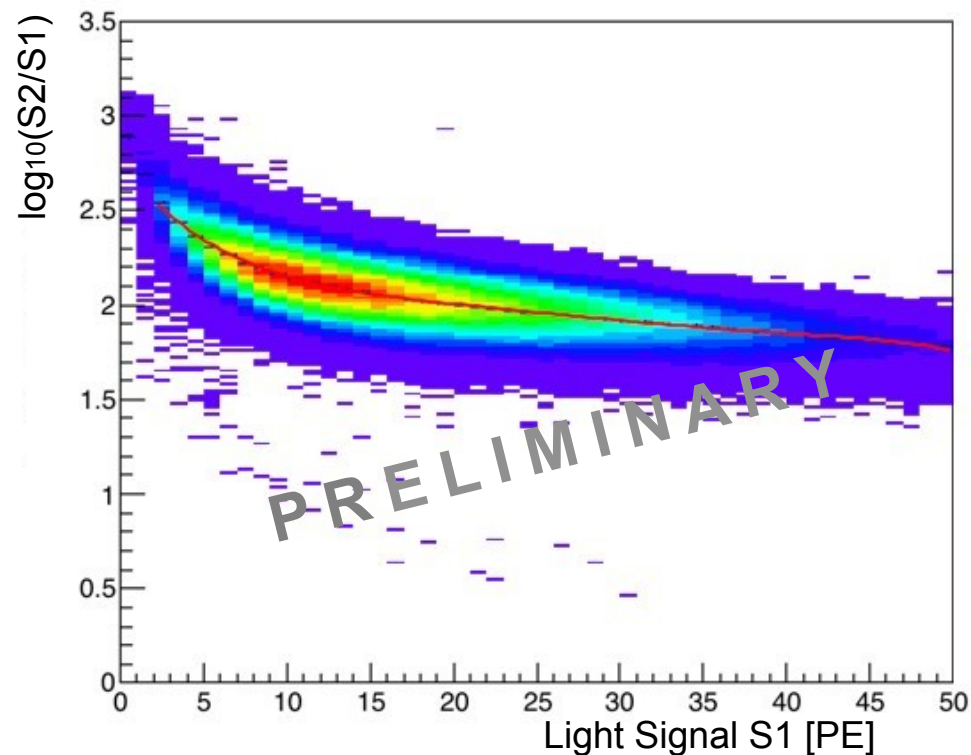
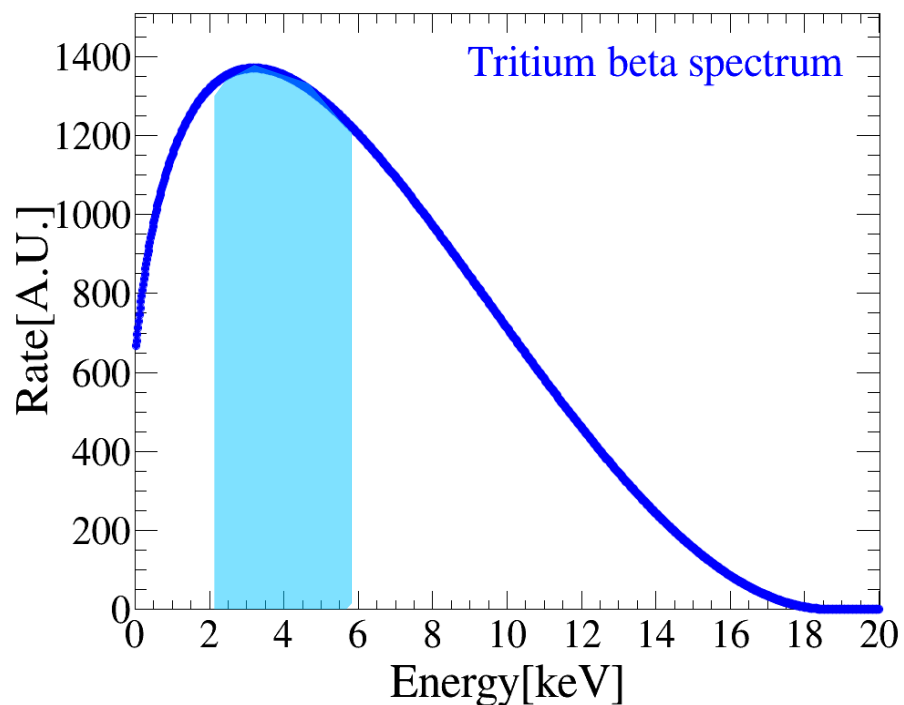


super-

ate over
s 26.4%

XENON100 : Low-mass dark matter

Intrinsic ER calibration with tritiated methane: pioneered by LUX [arXiv:1512.03133](https://arxiv.org/abs/1512.03133)



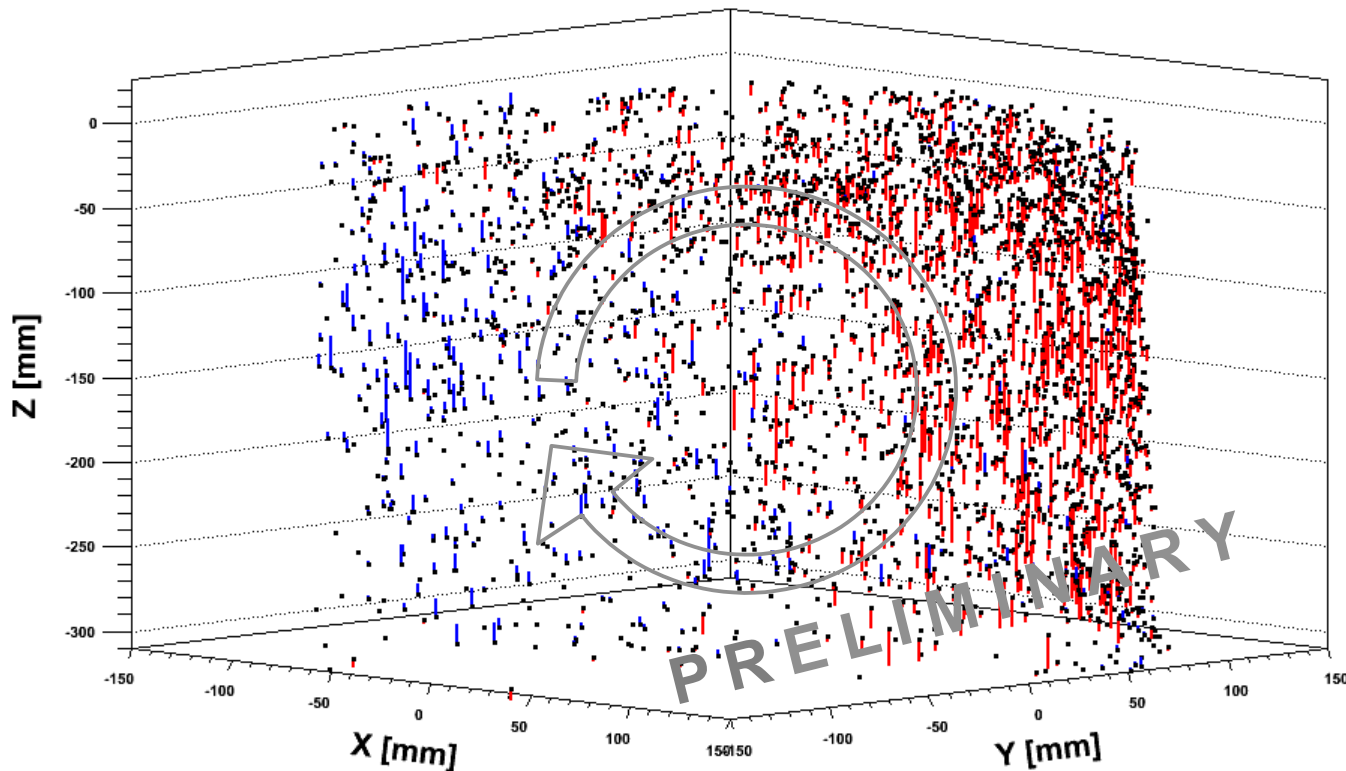
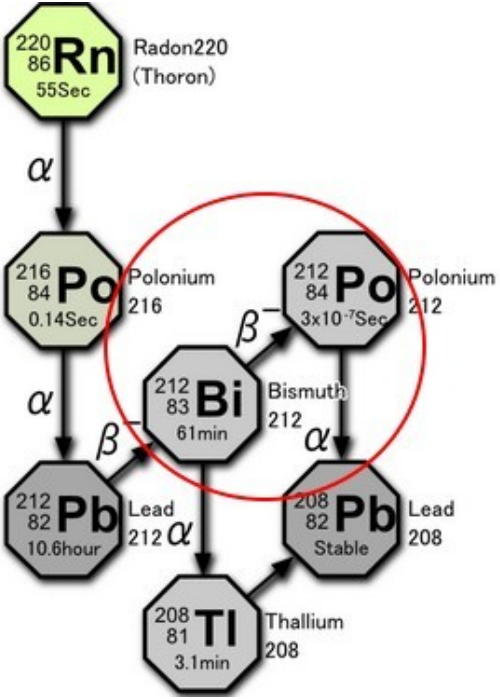
- high-stats calibration of XENON100 successfully performed twice
- ongoing: data analysis, removal strategies, application to XENON1T

XENON100 : Low-mass dark matter

- ^{228}Th source emanates ^{220}Rn
- ^{220}Rn is very short-lived (55 s)
- use beta-decays following ^{220}Rn to characterize low-E ER response
- prepare XENON1T use



arXiv:1602.01138



Mapping of convection inside XENON100 with ^{220}Rn

XENON1T Systems



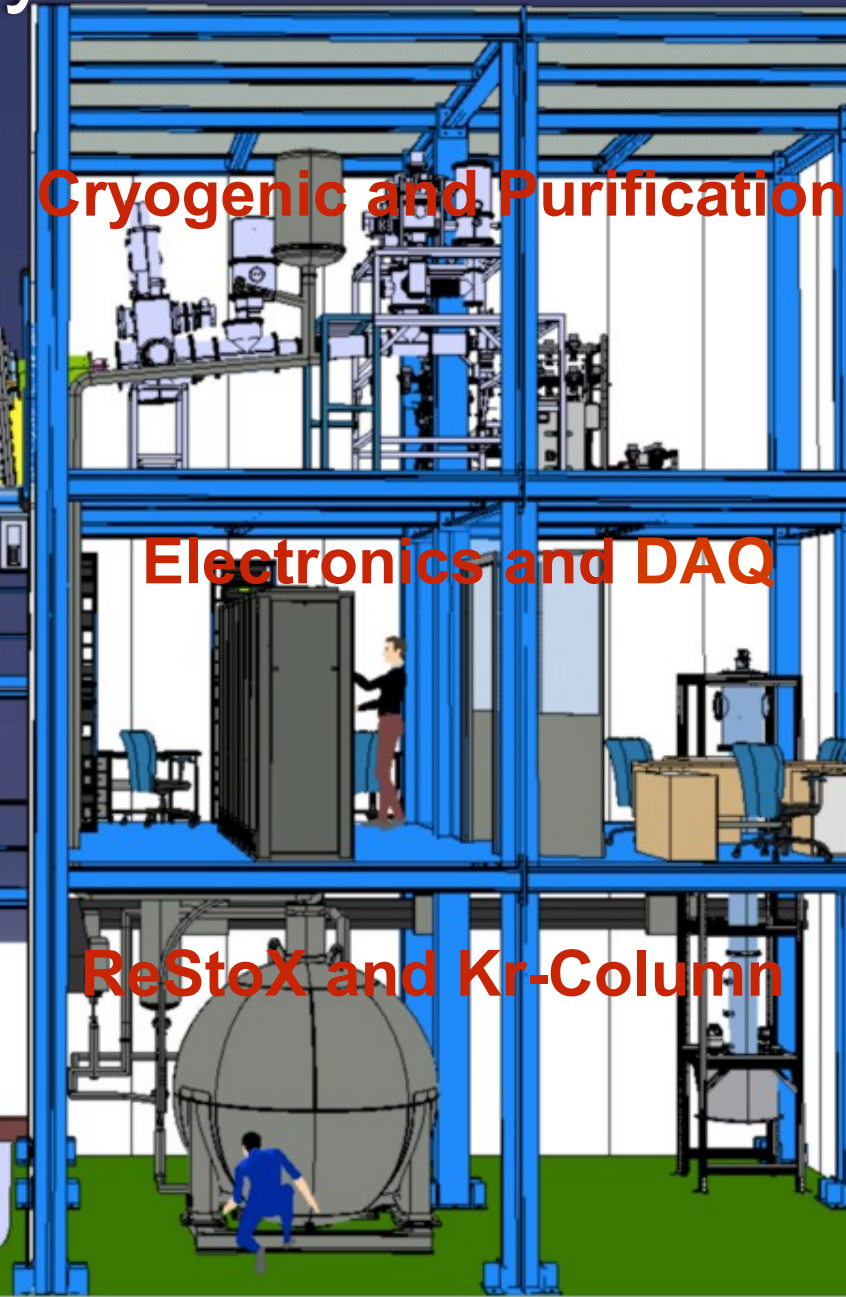
LXe Detector

Muon Veto Detector

Cryogenic and Purification

Electronics and DAQ

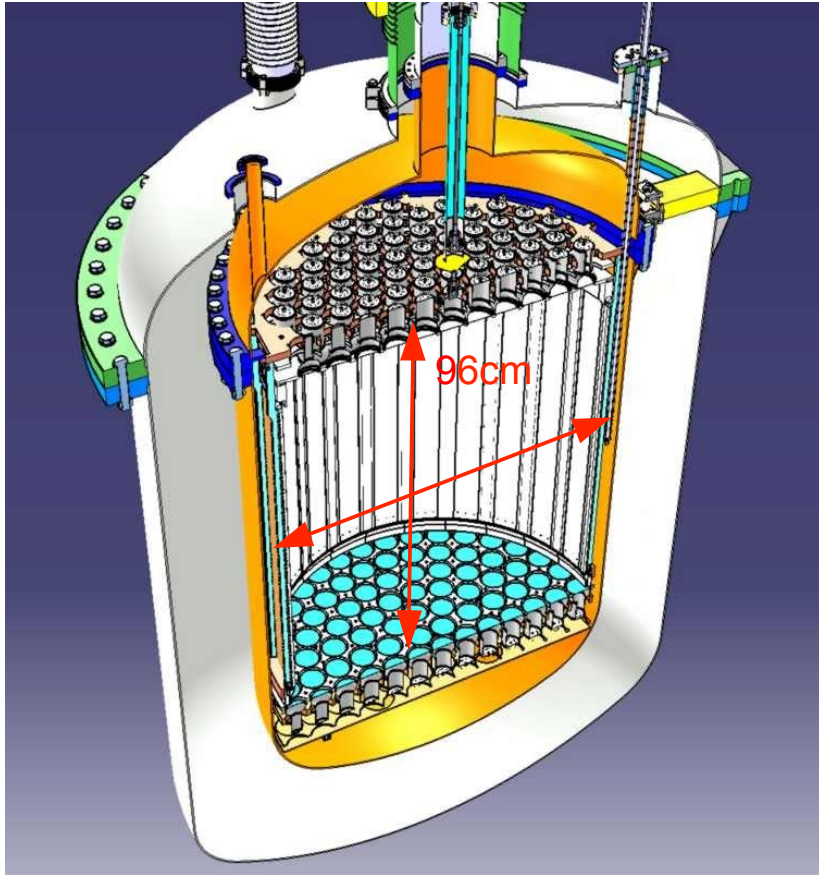
ReStoX and Kr-Column



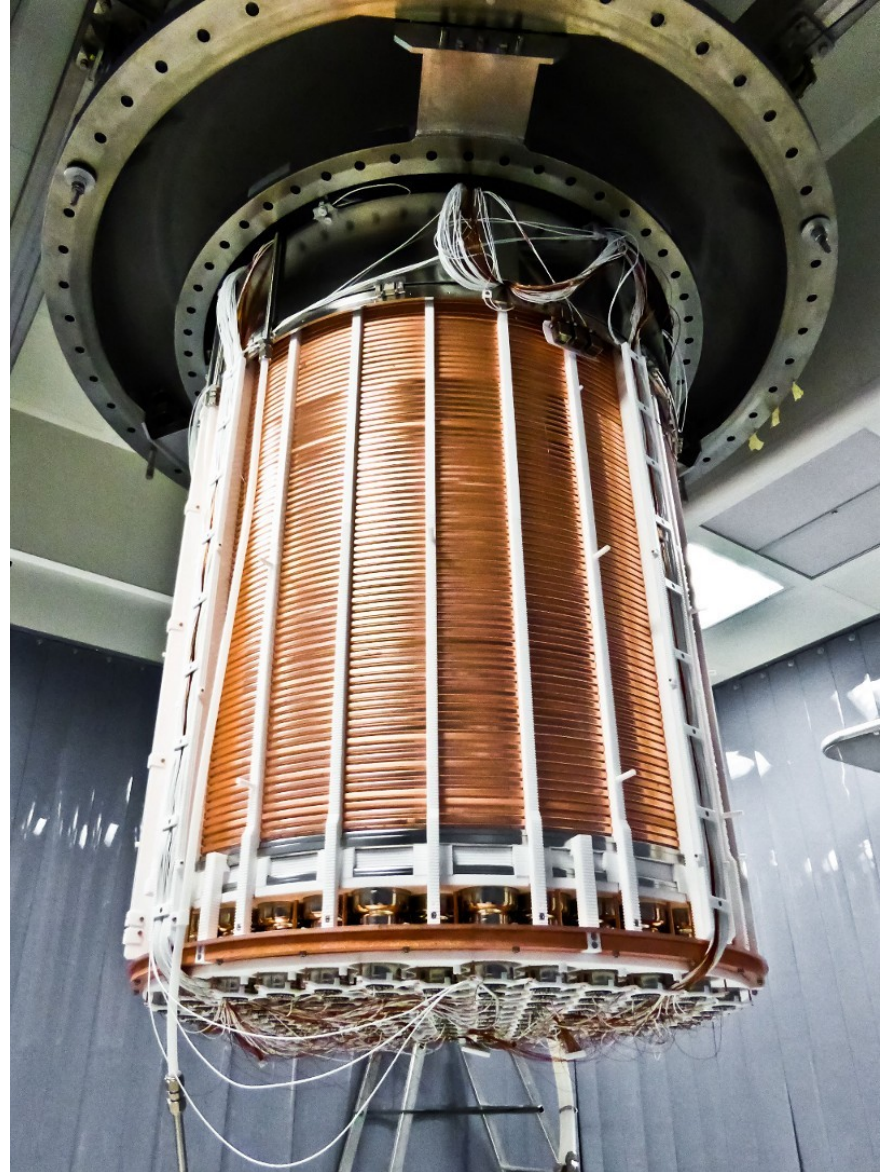




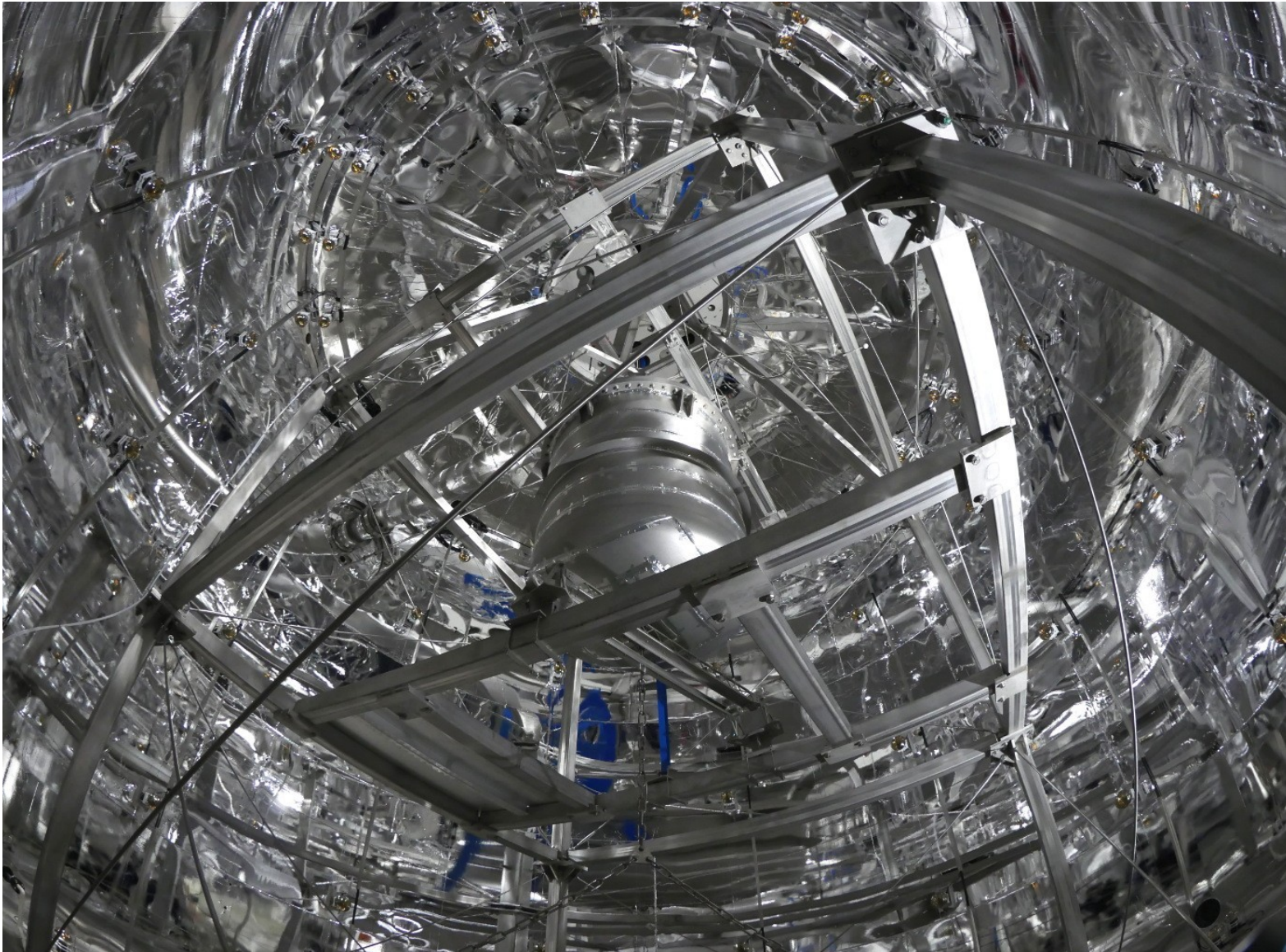
XENON1T: TPC



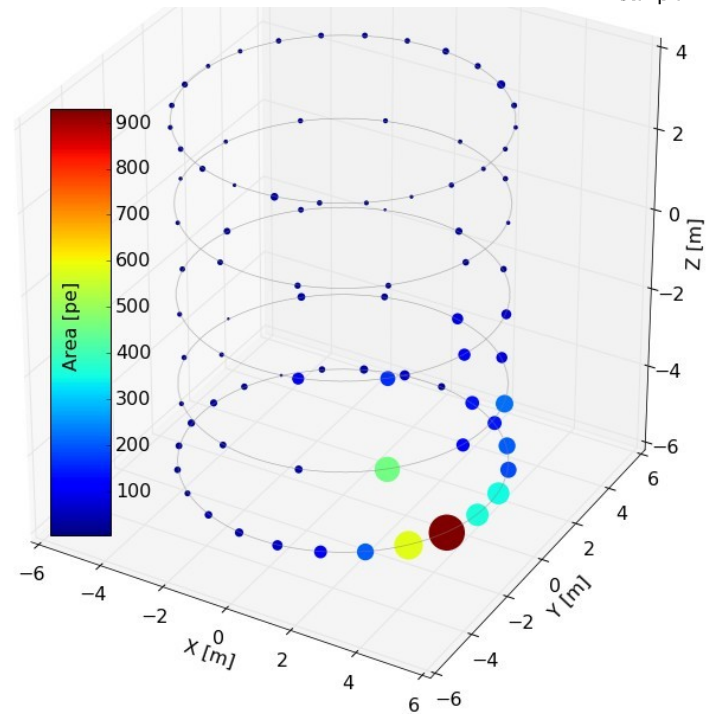
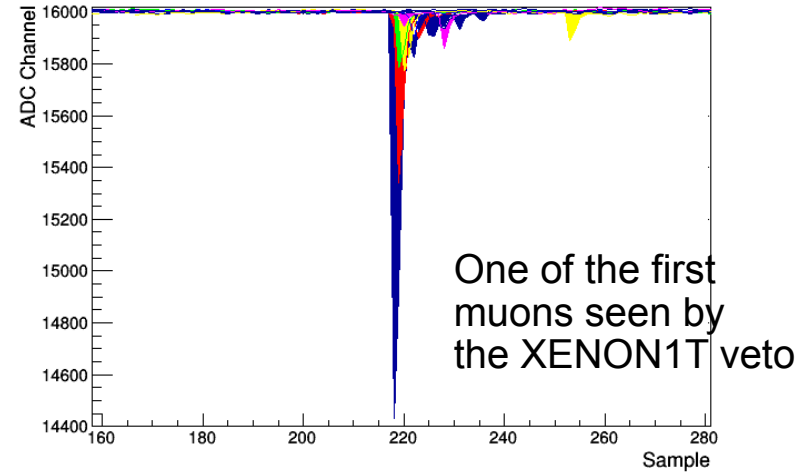
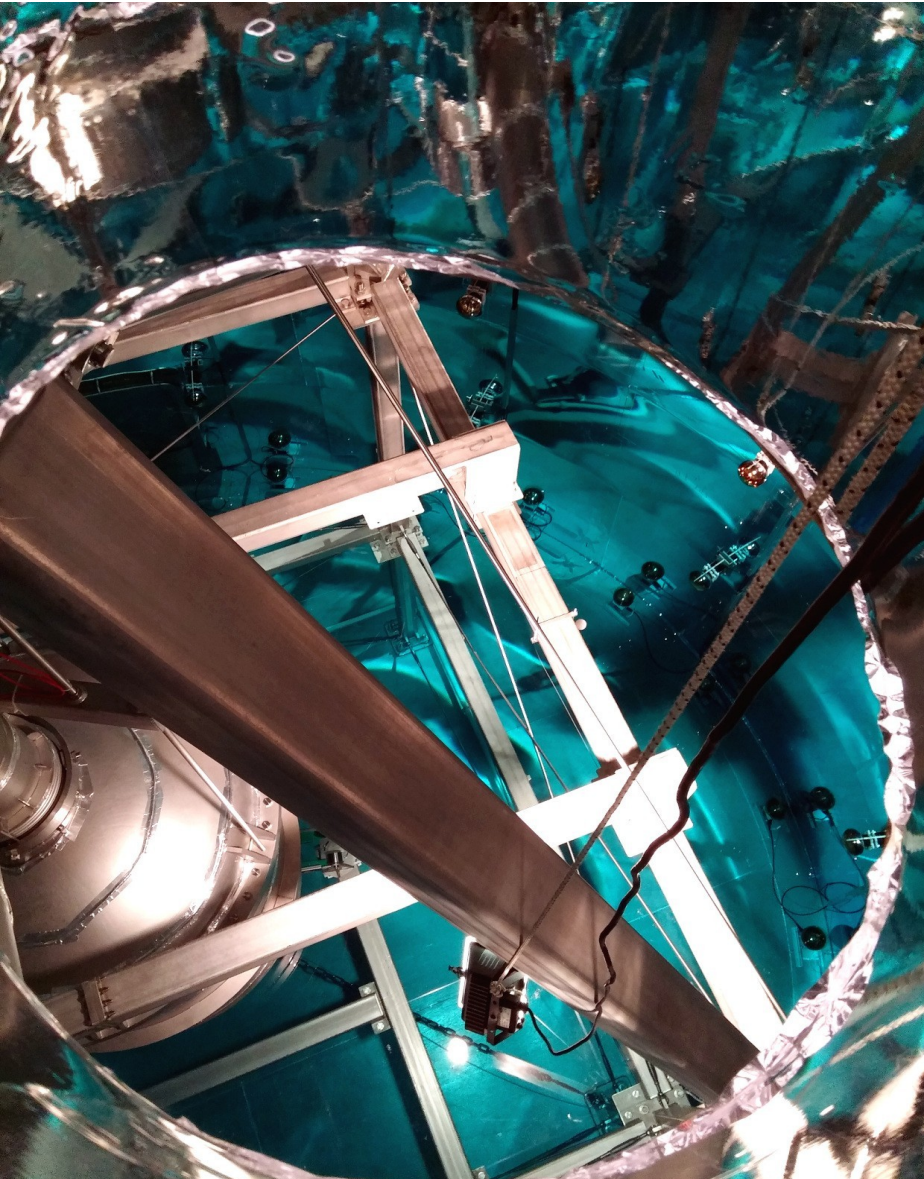
- 3.5 t liquid xenon in total
- **2.0t active target**
- ~1t after fiducialization
- 248+6 PMTs



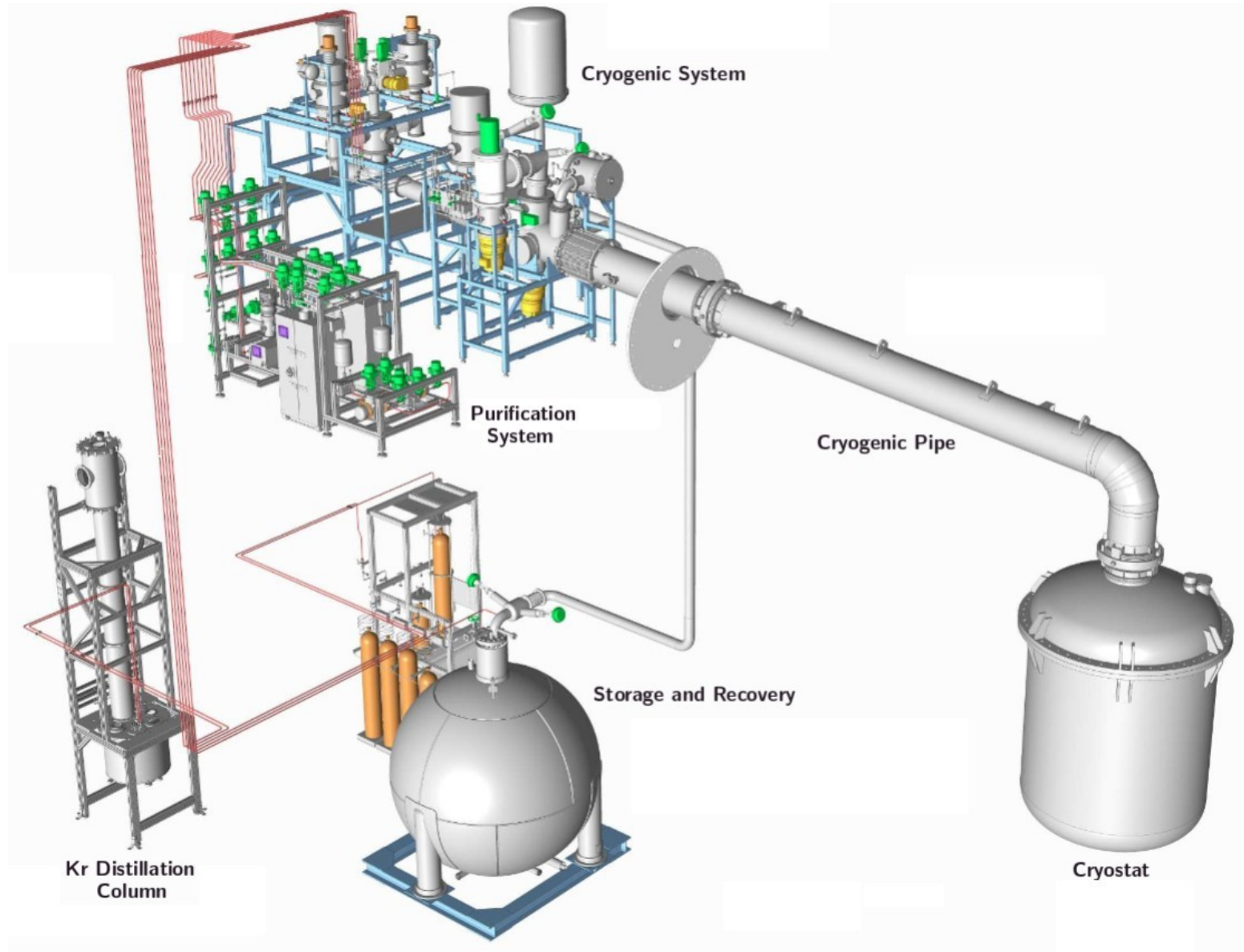
XENON1T: TPC in the Water Tank



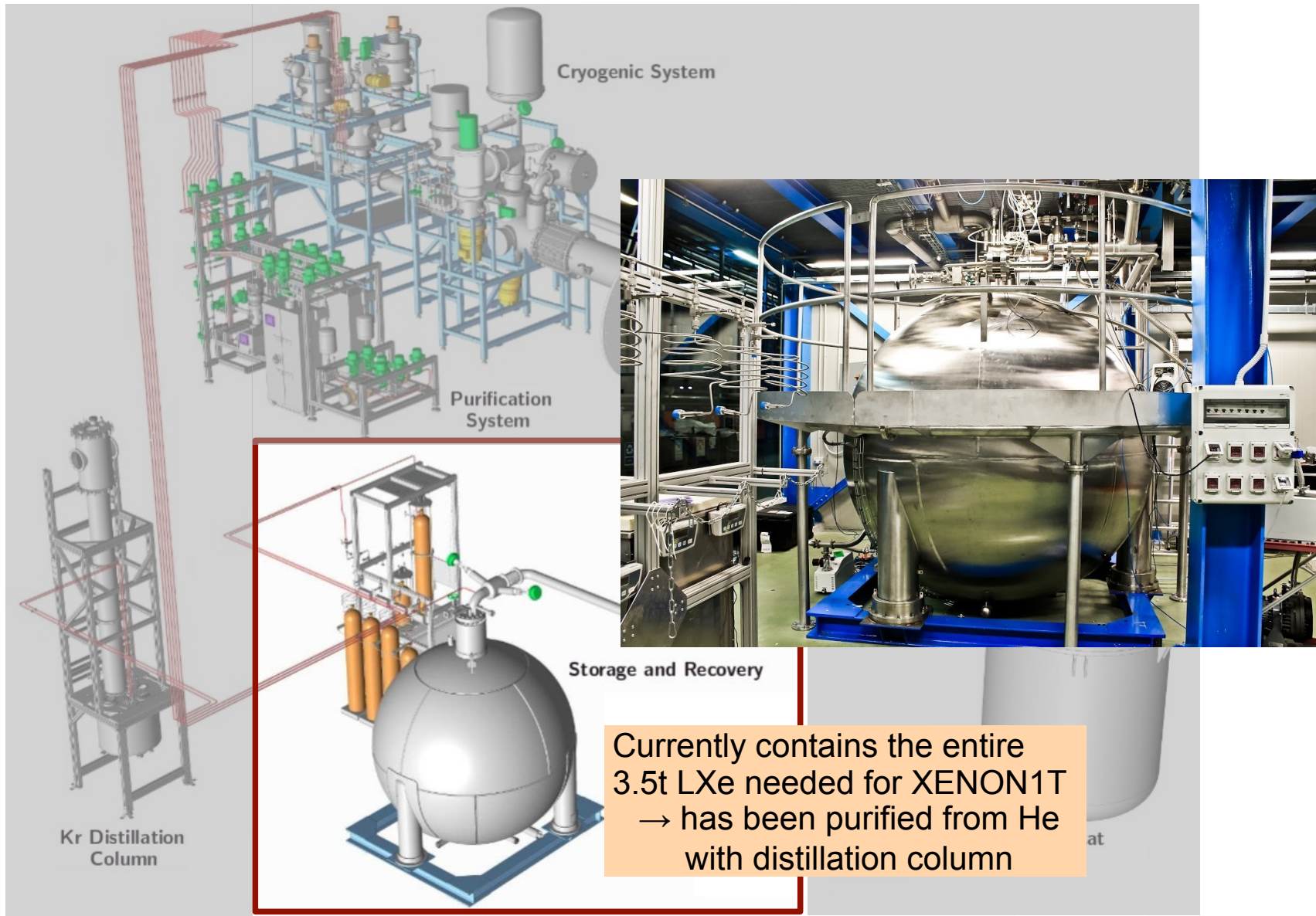
XENON1T: Muon Veto



XENON1T: Cryogenic System



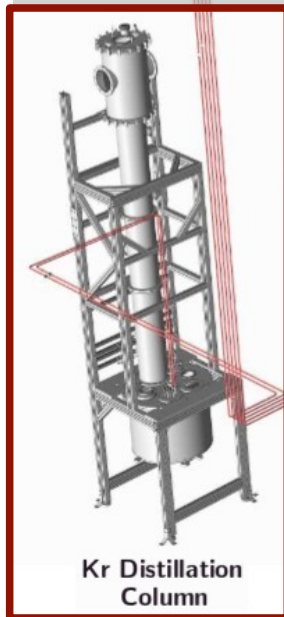
XENON1T: ReStoX



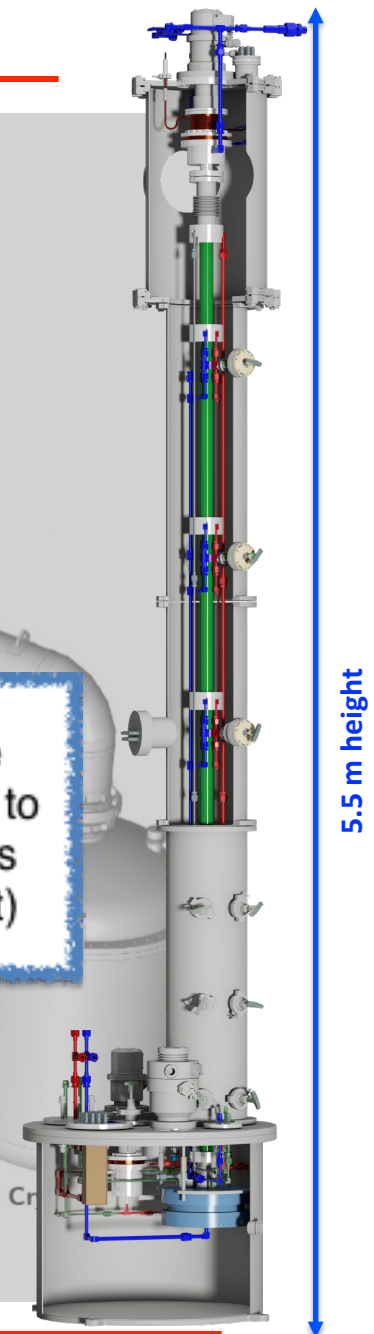
XENON1T: Kr Column

Separation factor: $10^4 - 10^5$
Kr removal : ${}^{\text{nat}}\text{Kr}/\text{Xe} < 0.026$ ppt
Factor 10 better than required

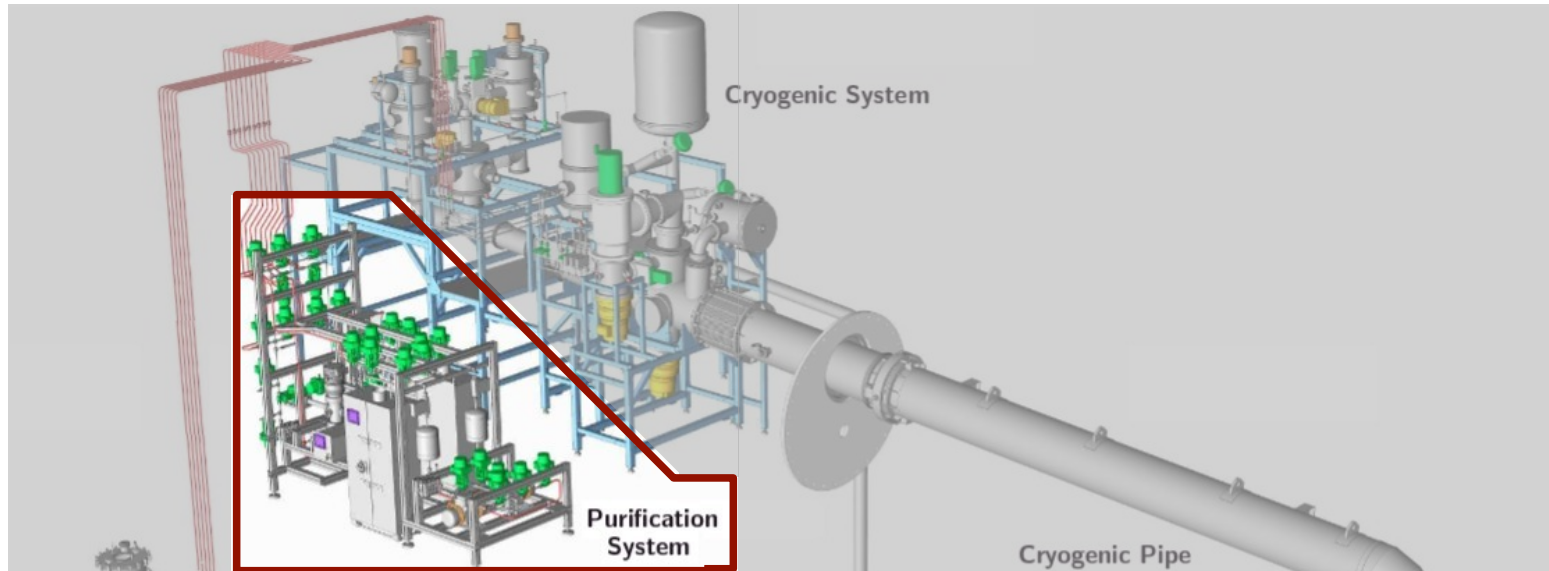
ready to remove
 ${}^{\text{nat}}\text{Kr}$ from Xe



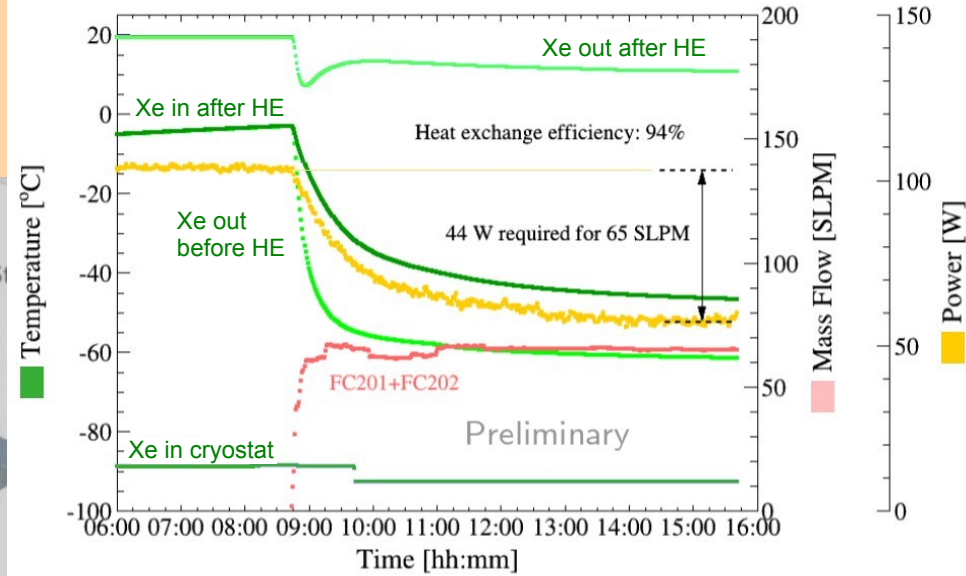
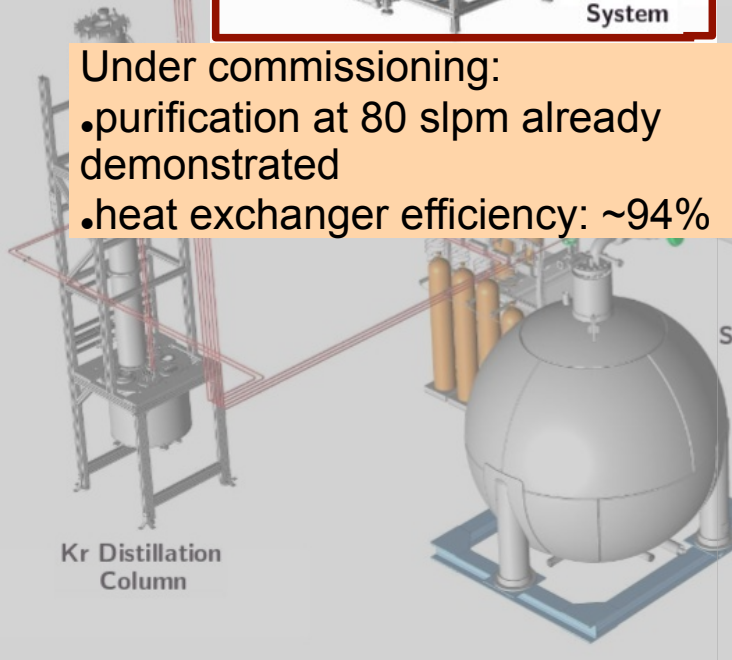
The ${}^{85}\text{Kr}$ concentration in the atmosphere is strongly related to nuclear proliferation (weapons testing, nuclear plant accident)



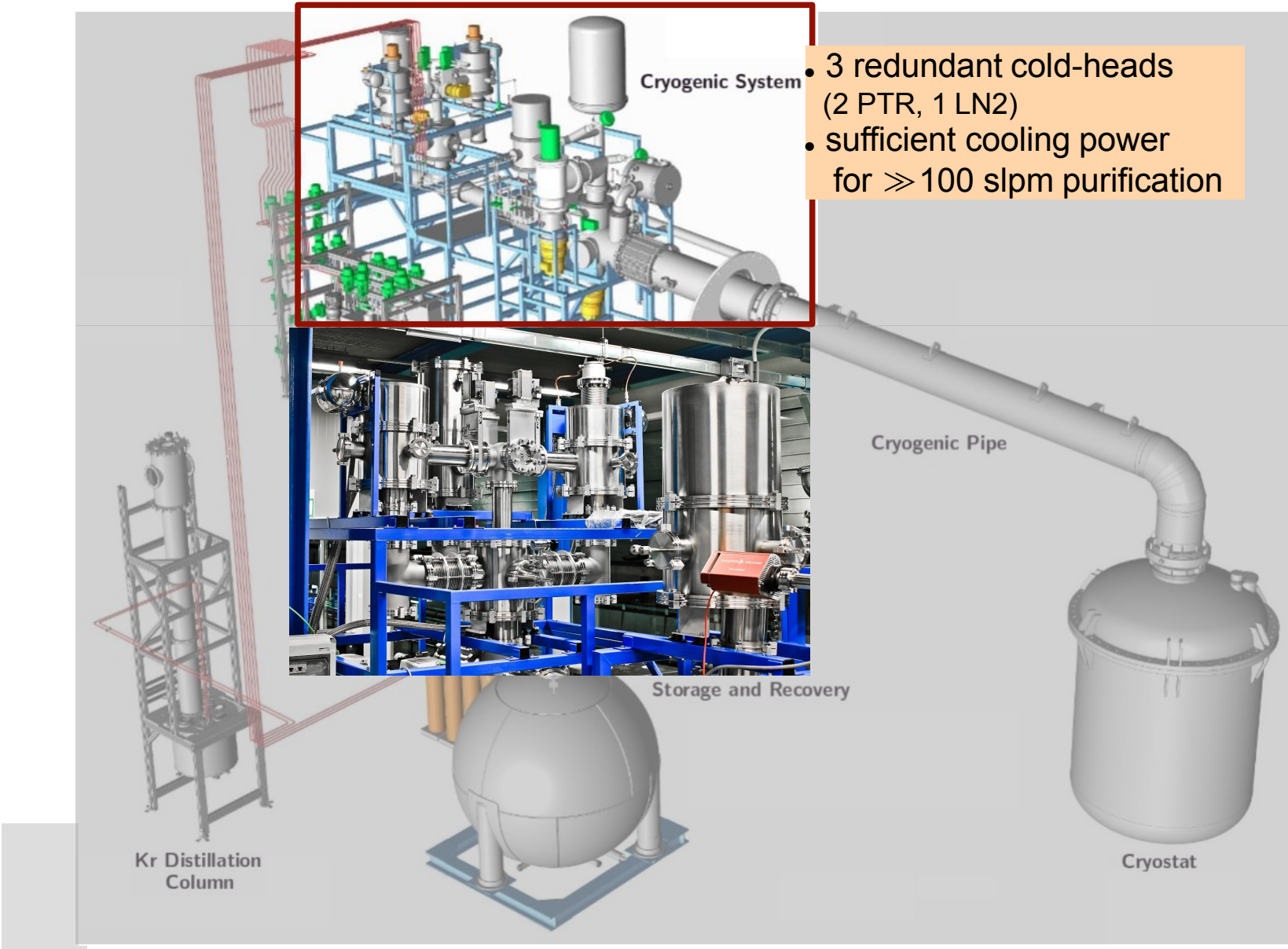
XENON1T: Purification System



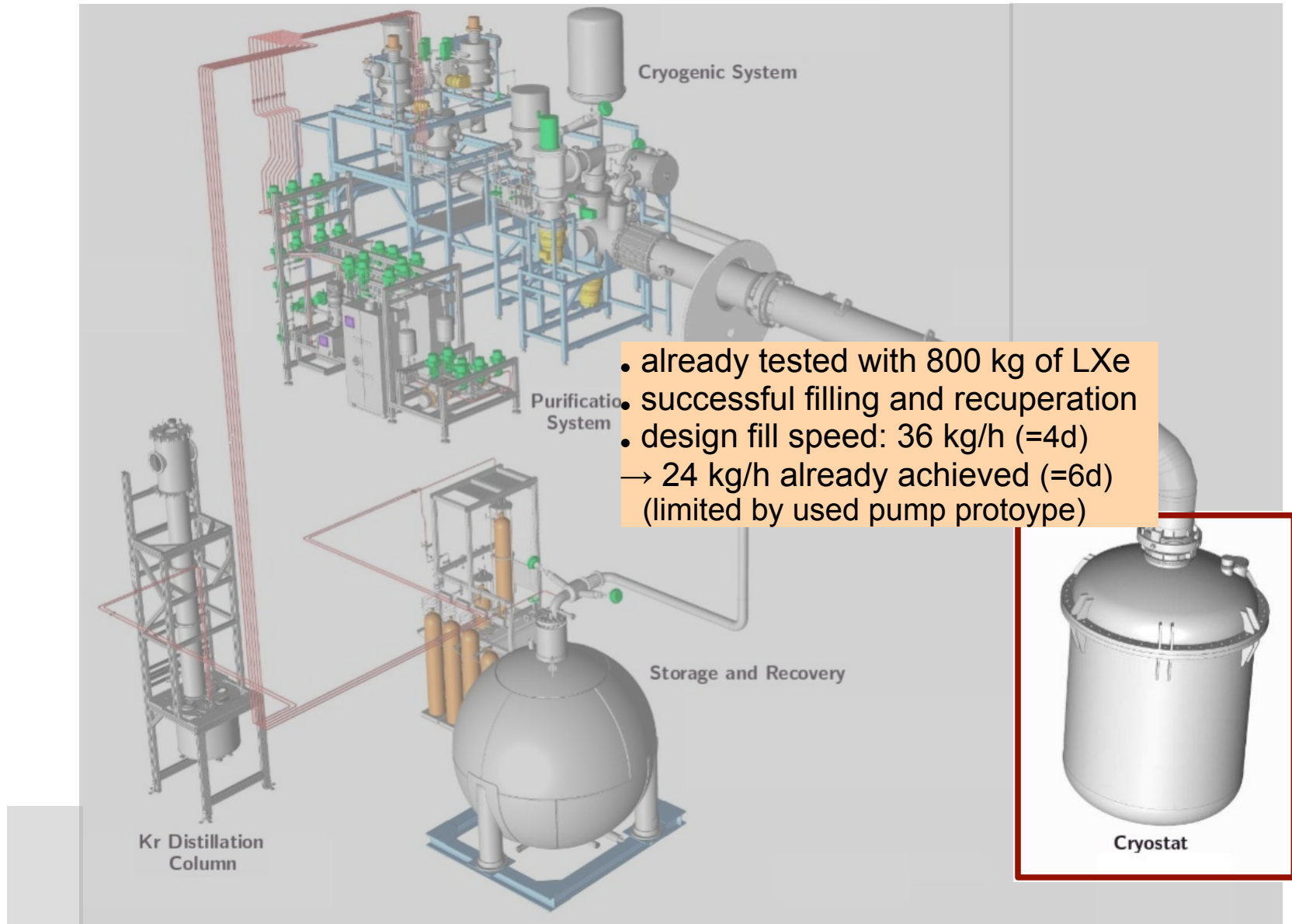
Under commissioning:
• purification at 80 slpm already demonstrated
• heat exchanger efficiency: ~94%



XENON1T: Cryogenic System



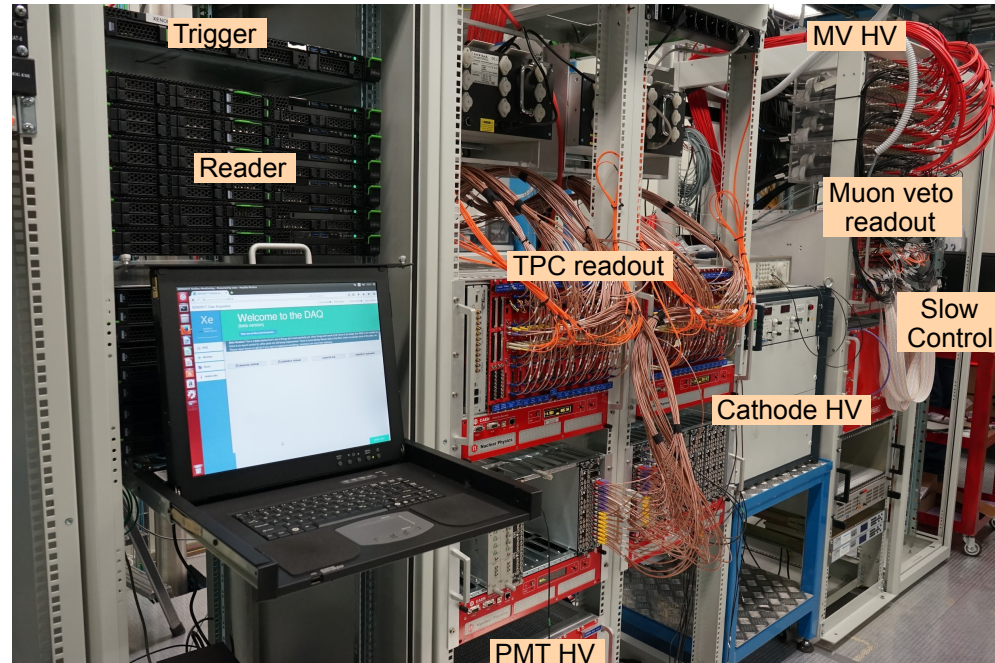
XENON1T: Cryostat



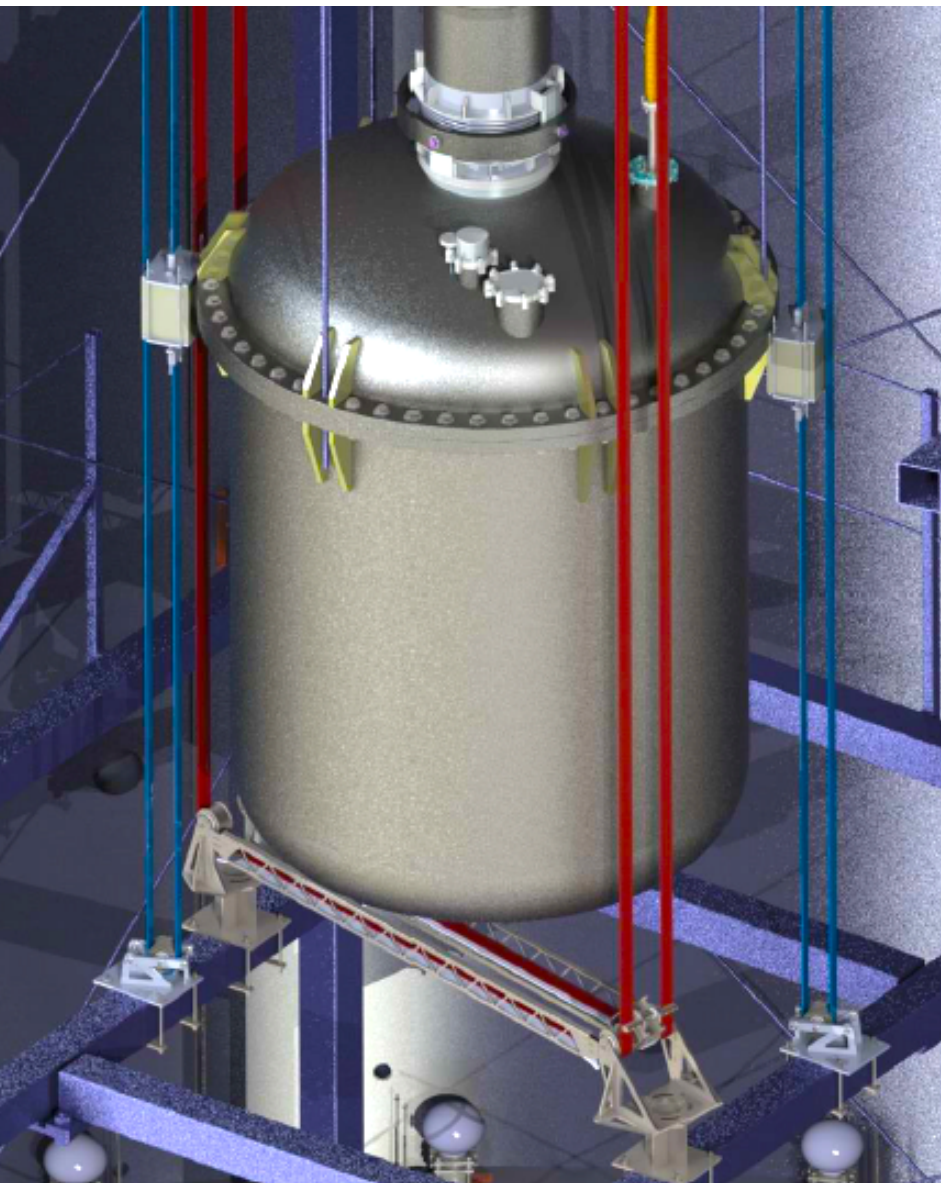
XENON1T: PMT, DAQ, Electronics



- all 254 PMTs operational
- DAQ electronics for TPC and muon veto installed in T-stabilized DAQ room
- detectors can be operated simultaneously (and time-synced) or independently
- PMT/DAQ commissioning ongoing



XENON1T: Calibration



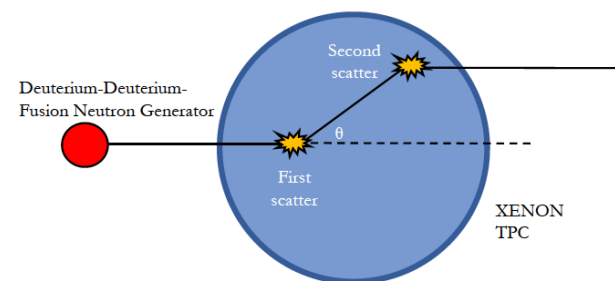
XENON1T is calibrated via:

LED: to periodically measure the gain of the PMT inside the TPC

INTERNAL CALIBRATION SOURCES: short-lived radioactive isotopes mixed to the xenon stream:

- ^{83m}Kr to calibrate the ER energy scale
- ^{220}Rn and TCH_3 for low-energy ER

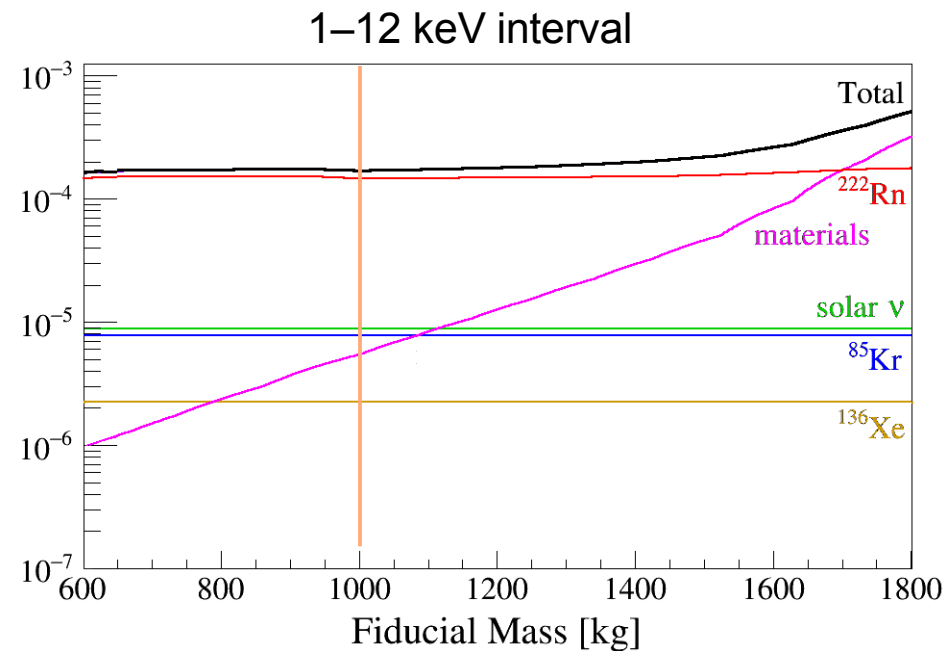
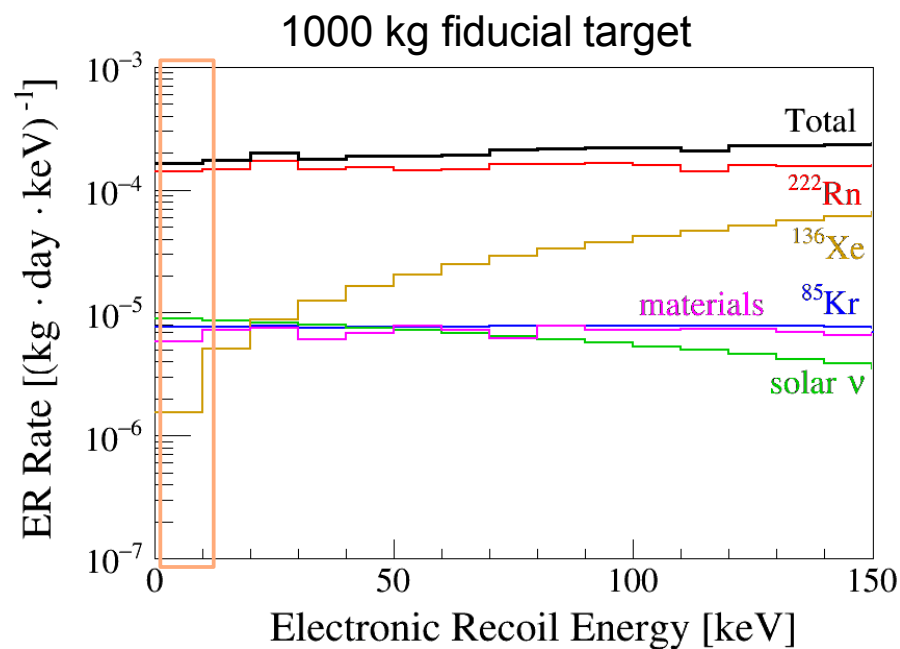
NEUTRON GENERATOR: the size of the XENON1T TPC allow to identifying double scatters of 2.5 MeV mono energetic neutrons produced by a D-D neutron generator



EXTERNAL CALIBRATION SOURCE to measure the purity of the LXe in the target and self shielding capability

XENON1T: Electronic Recoil

arXiv:1512.07501, accepted by JCAP



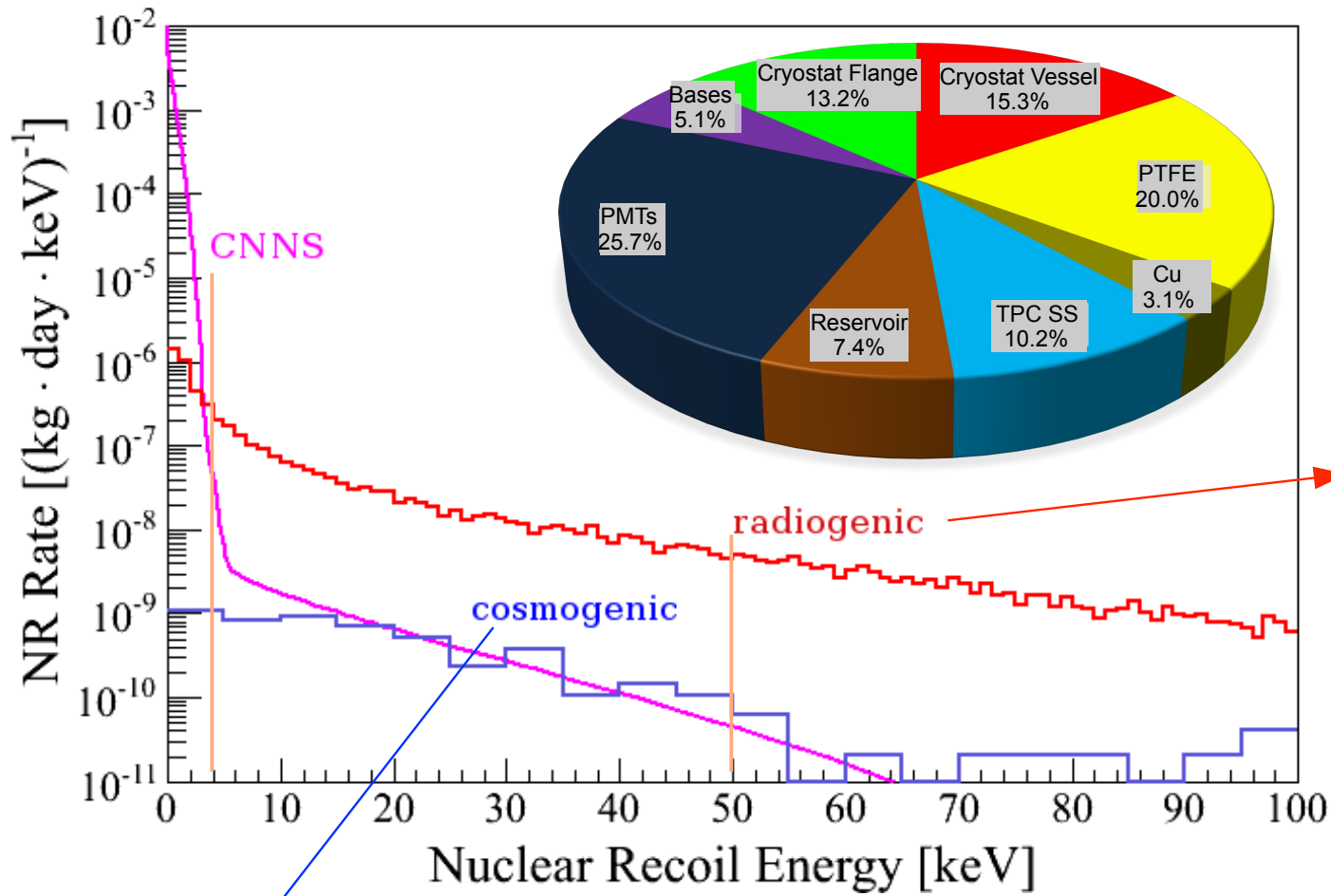
Assumed contamination:

²²²Rn: 10 μ Bq/kg (measured $19 \pm 4 \rightarrow 8$ expected after purification)

^{nat}Kr: 0.2 ppt

¹³⁶Xe: 8.9% natural abundance

XENON1T: Nuclear Recoil



material screening, e.g. EPJ C 75, 546 (2015)

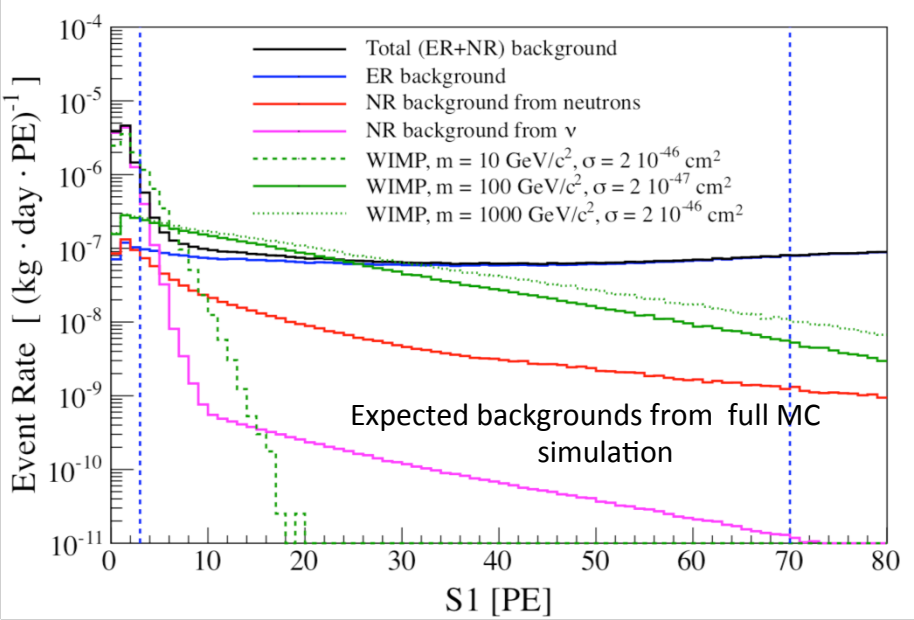
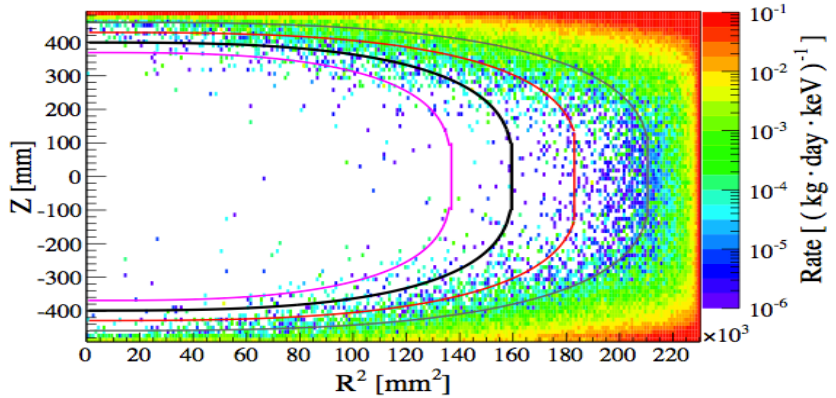
Muon veto design and performance: XENON1T, JINST 9, P11006 (2014)

Monte Carlo Simulation

Reproduce via software the performance of the XENON1T detector, and predict the sensitivity of the experiment

arXiv:1512.07501 [physics.ins-det]

Position of the ER background from the materials, negligible inside the 1 ton fiducial volume



1 ton fiducial volume S1 in [3,70] PE
 ER discrimination 99.75%
 NR acceptance 40%



Method:

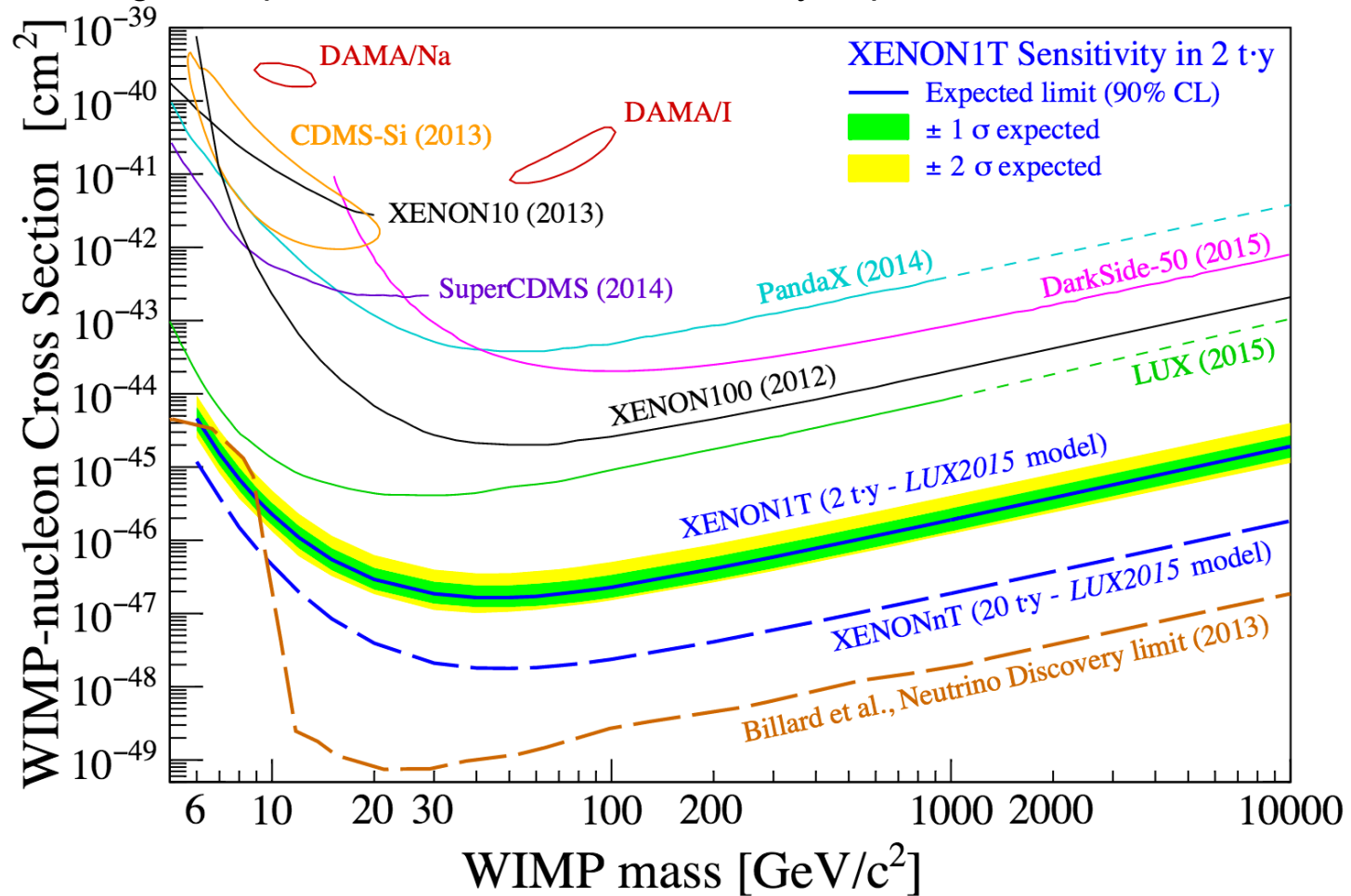
- Input from screening campaign by all detector components
- Monte Carlo simulation with GEANT4
- Statistical treatment

Source	Bkg (evts/ton/year)
ER (materials + intrinsic + solar ν)	1.6
NR from radiogenic neutrons	0.22
NR from ν coherent scattering	0.23
Total	2.05

XENON1T: Sensitivity

arXiv:1512.07501, accepted by JCAP

based on background predictions shown before, 2 t×y exposure:



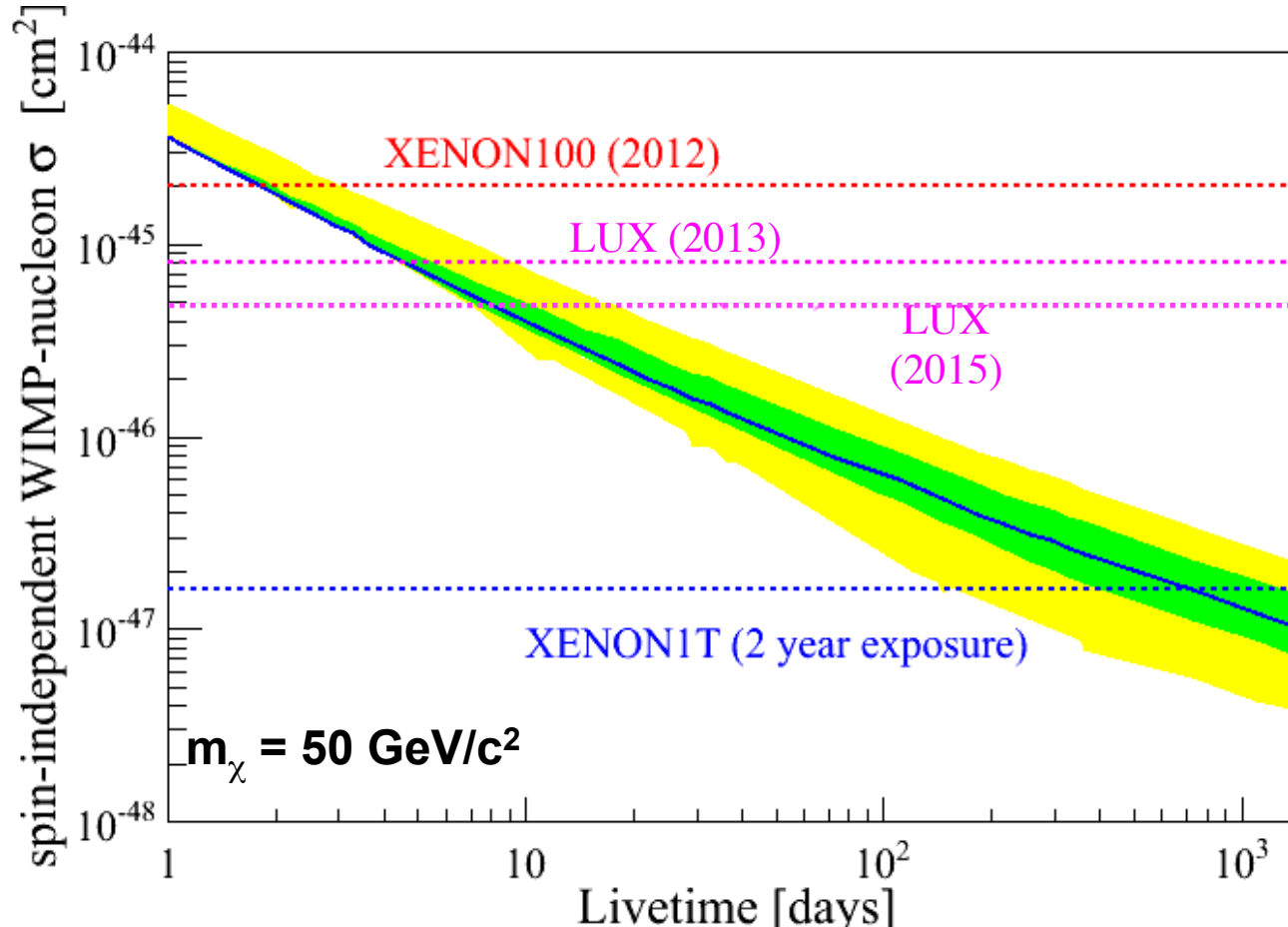
assumptions: energy interval: 4 – 50 keV,

ER rejection as XENON100: 99.5% @ 50% NR acc.
→ expected LY is 2x higher than in XENON100!

XENON1T: Sensitivity Vs Time

arXiv:1512.07501, accepted by JCAP

based on background predictions shown before, 2 t×y exposure:

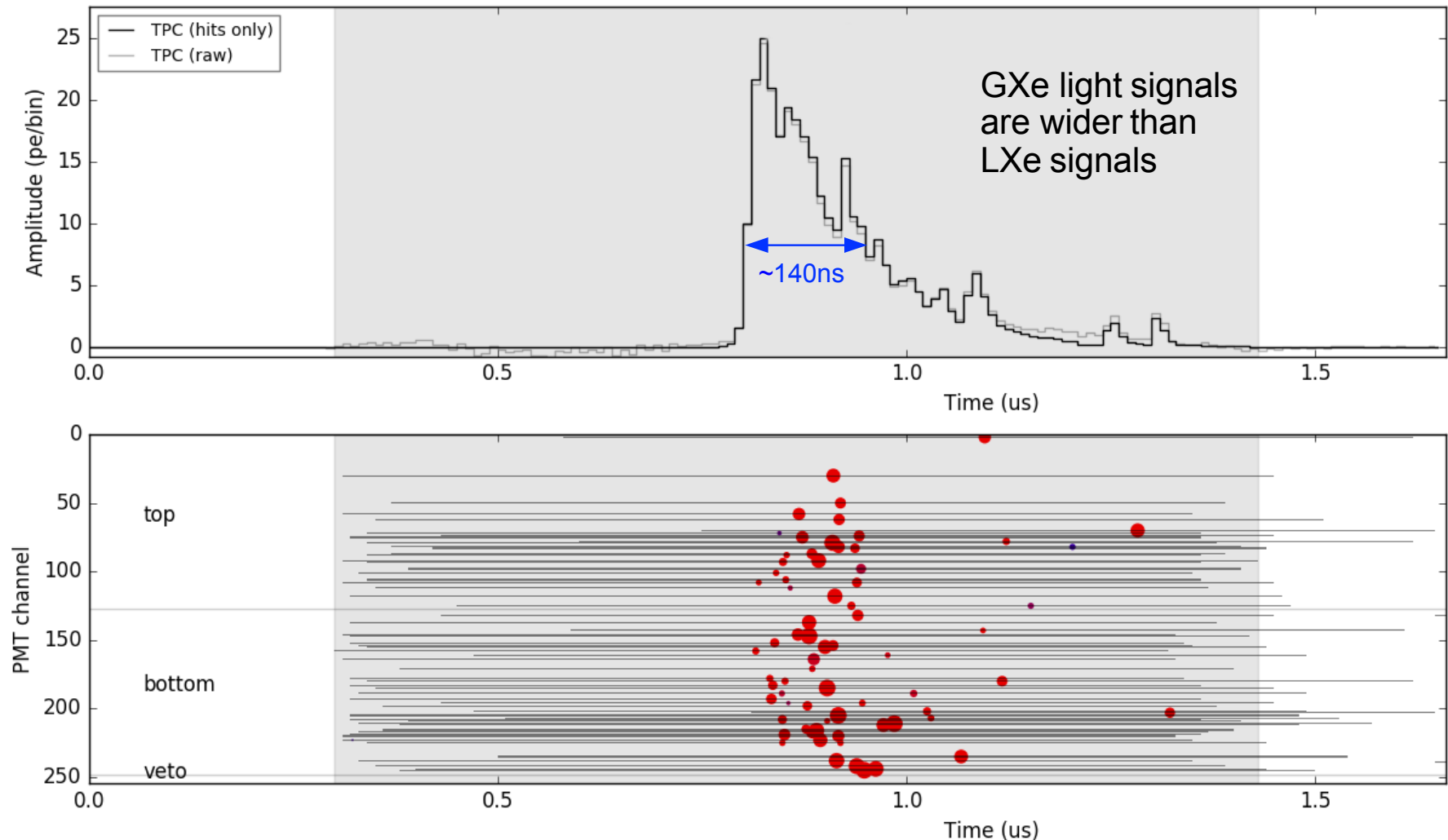


In less than 10 days we expect to reach the sensitivity of the currently running experiments

XENON1T: First GXe event

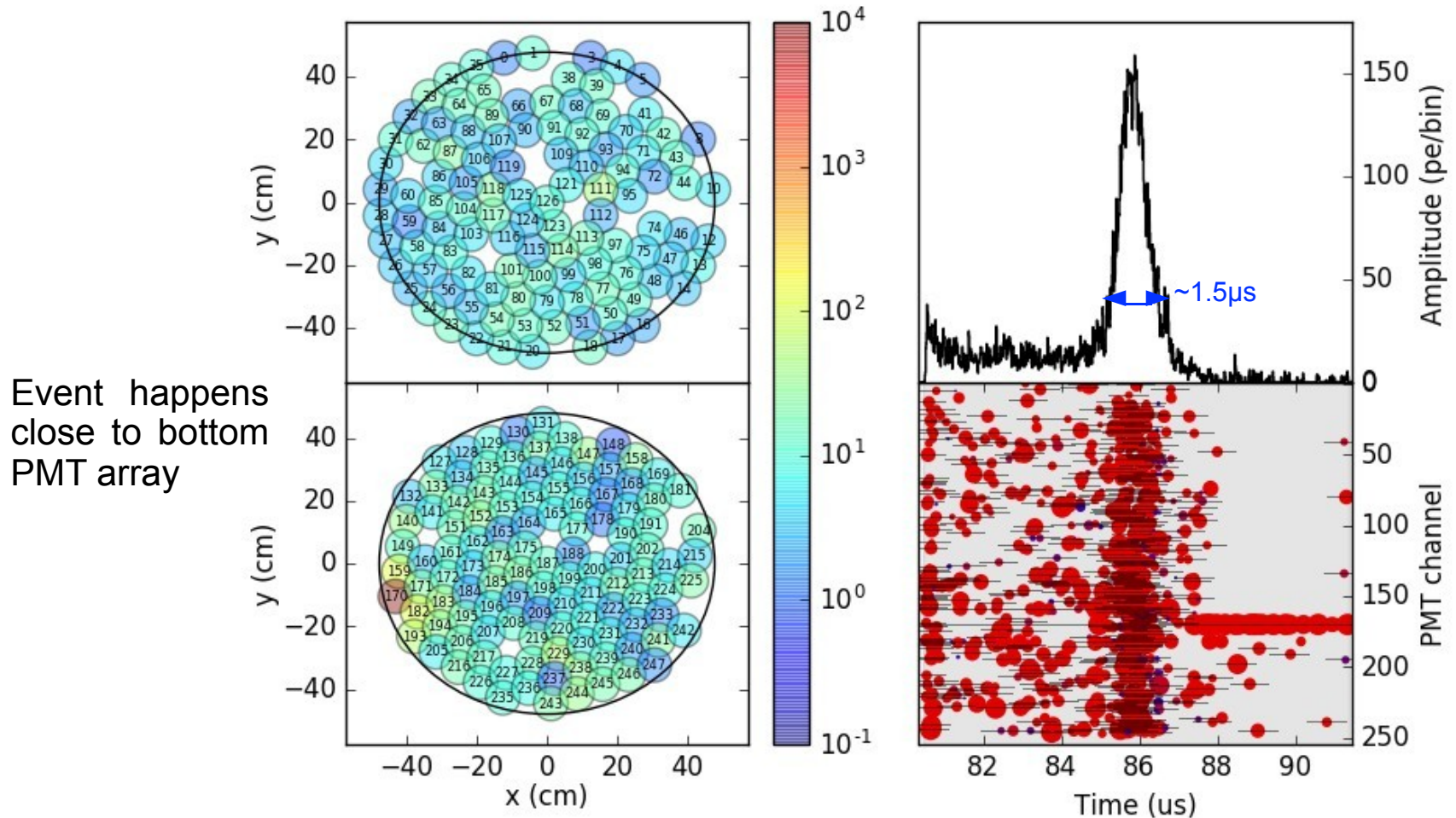
March 2016: TPC filled with GXe

→ start commissioning of **PMTs**, **DAQ** and **analysis** with particle interactions



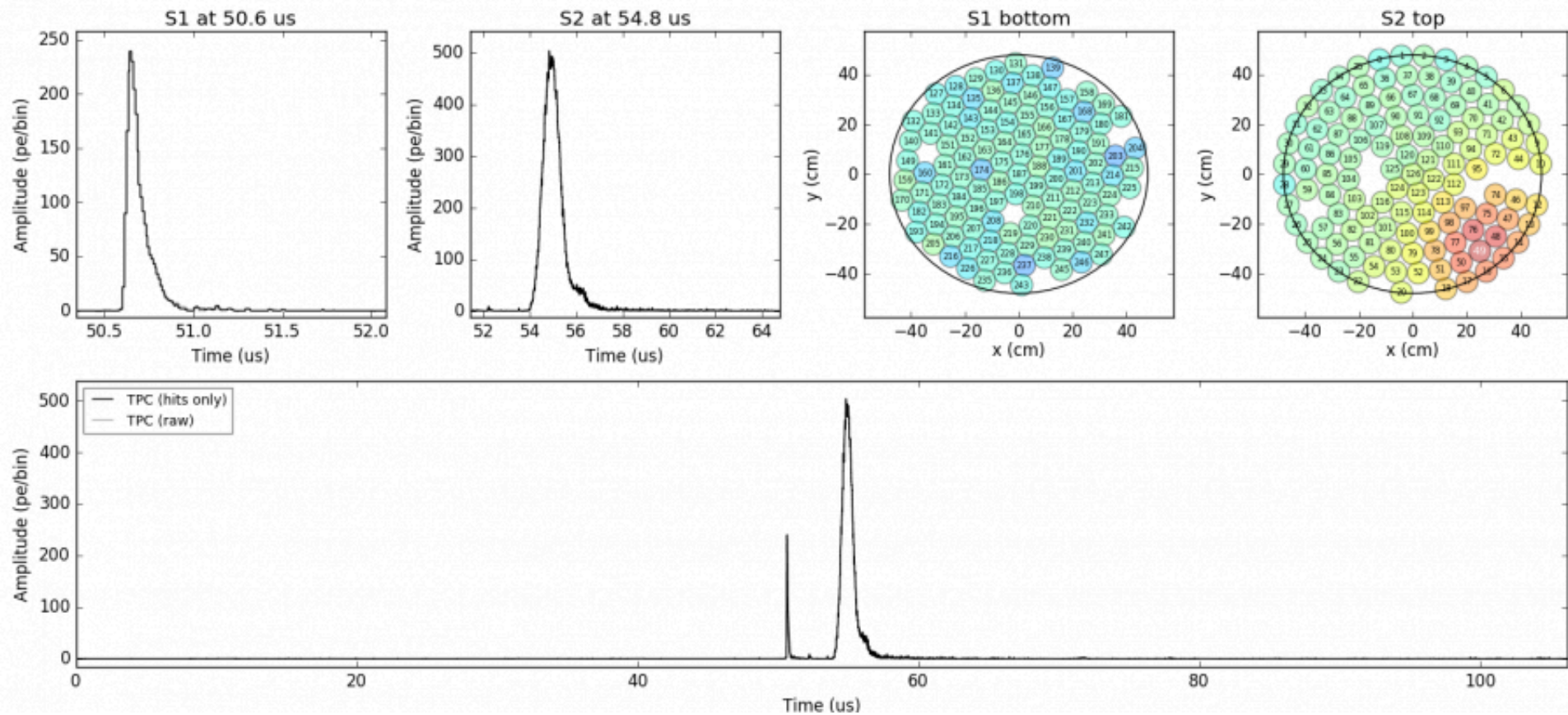
XENON1T: First Charge Signal

the main TPC electrodes were not yet biased
→ we still see first S2 peaks, thanks to the fields
between the PMTs and the screening electrodes (~ 2.1 kV/cm)



XENON1T: First S1 + S2 LXe Signal!

Event 1



XENON1T is coming...

The hunt is about to start !

