

Multidark Lightcones Status

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Lensing and galaxy mocks in the same lightcone

Based on BigMultidark simulation

- DM only with Planck cosmology ($\Omega_m = 0.31$)
- 2.5 Gpc/h periodic box and 3840^3 particles
- 80 snapshots from $z = 10$ to $z = 0$ (69 below $z = 1$)
- Mass range of halos 4.7×10^{11} to 6×10^{15} Msun/h
- Force resolution 10 kpc/h (lowz) and 30 kpc/h (highz)
- Halo catalogs extracted with BDM and FoF

Box remapping code (Carlson&White2010)

Lensing ray tracing with GLAMER (Metcalf et al. 2013, Petkova et al. 2013)

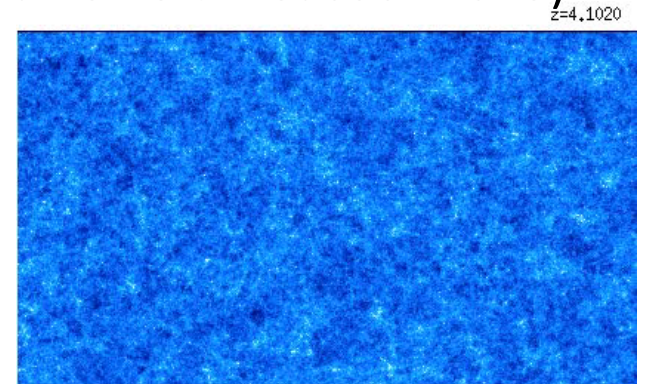
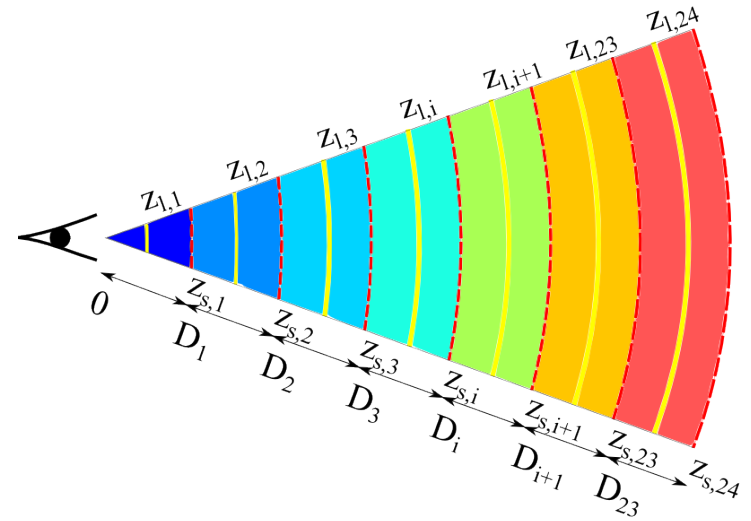
HOD Galaxy clustering with peculiar velocities (de la Torre & Peacock 2013)

⇒ Overlapping Lensing and Clustering mocks:

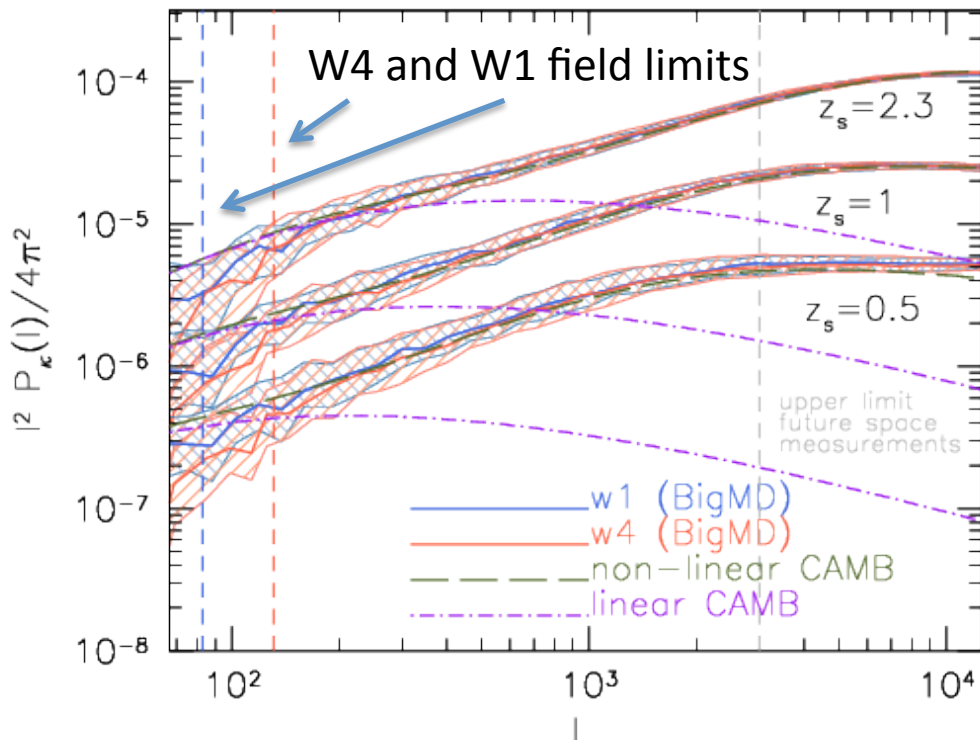
⇒ 54 LC of area $7 \times 4 \text{ deg}^2$ up to $z=2.3$

⇒ 99 LC of area $5.5 \times 1.6 \text{ deg}^2$ up to $z=2.3$

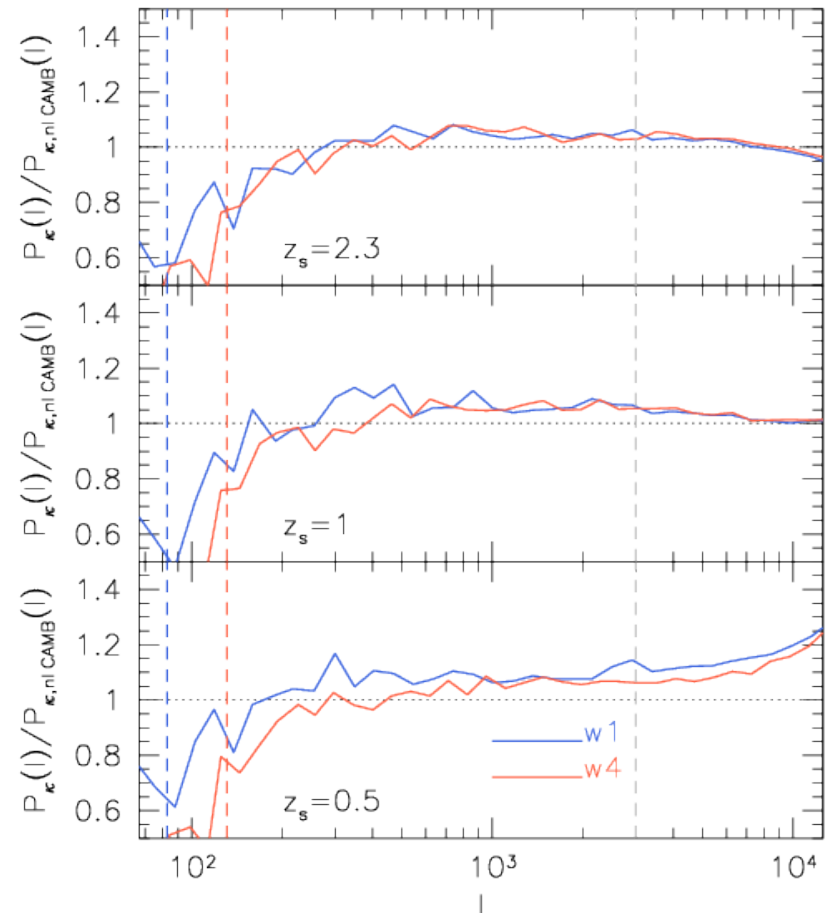
⇒ 4 LC of area $87 \times 2.6 \text{ deg}^2$ up to $z=2$



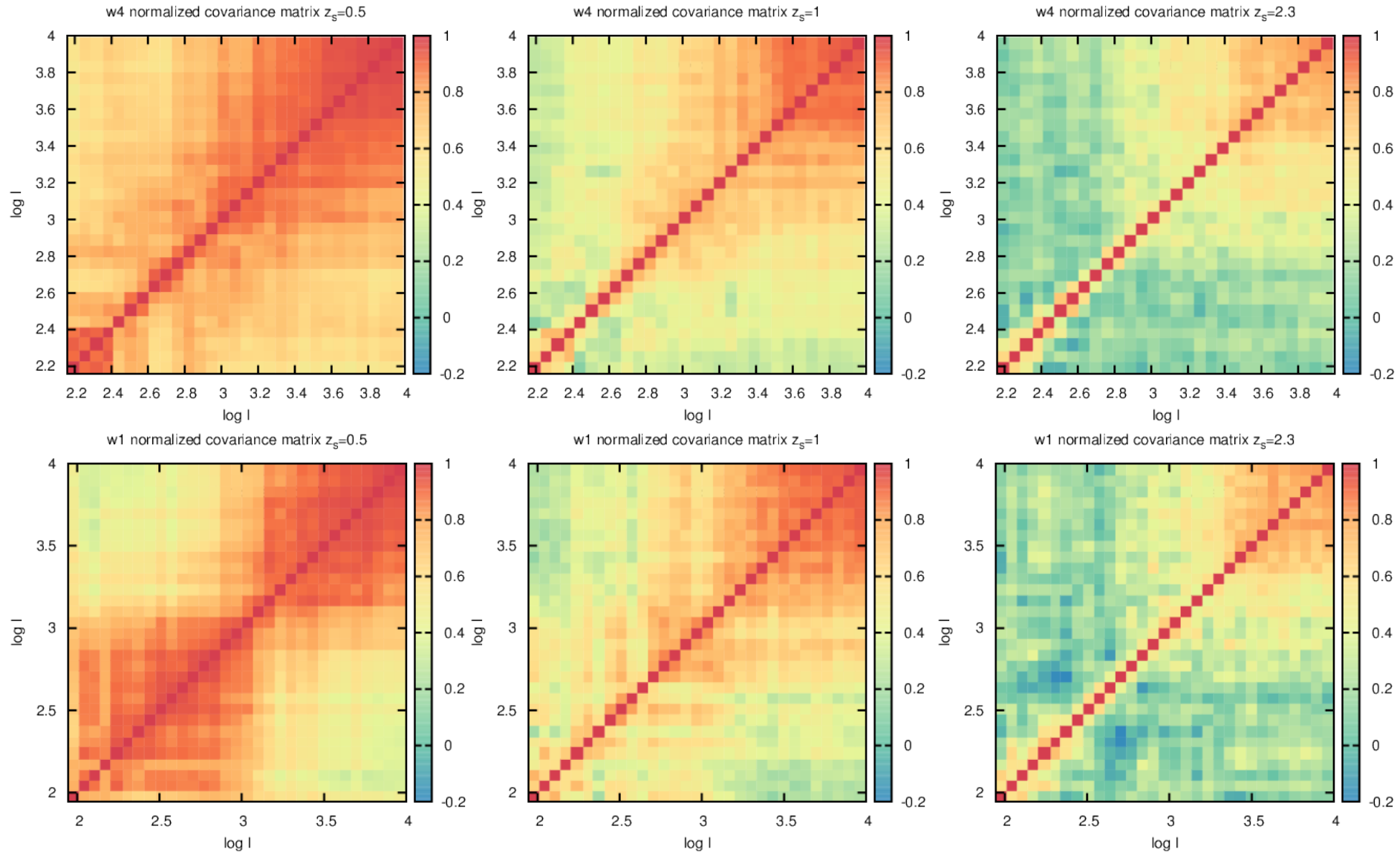
Lensing power spectrum



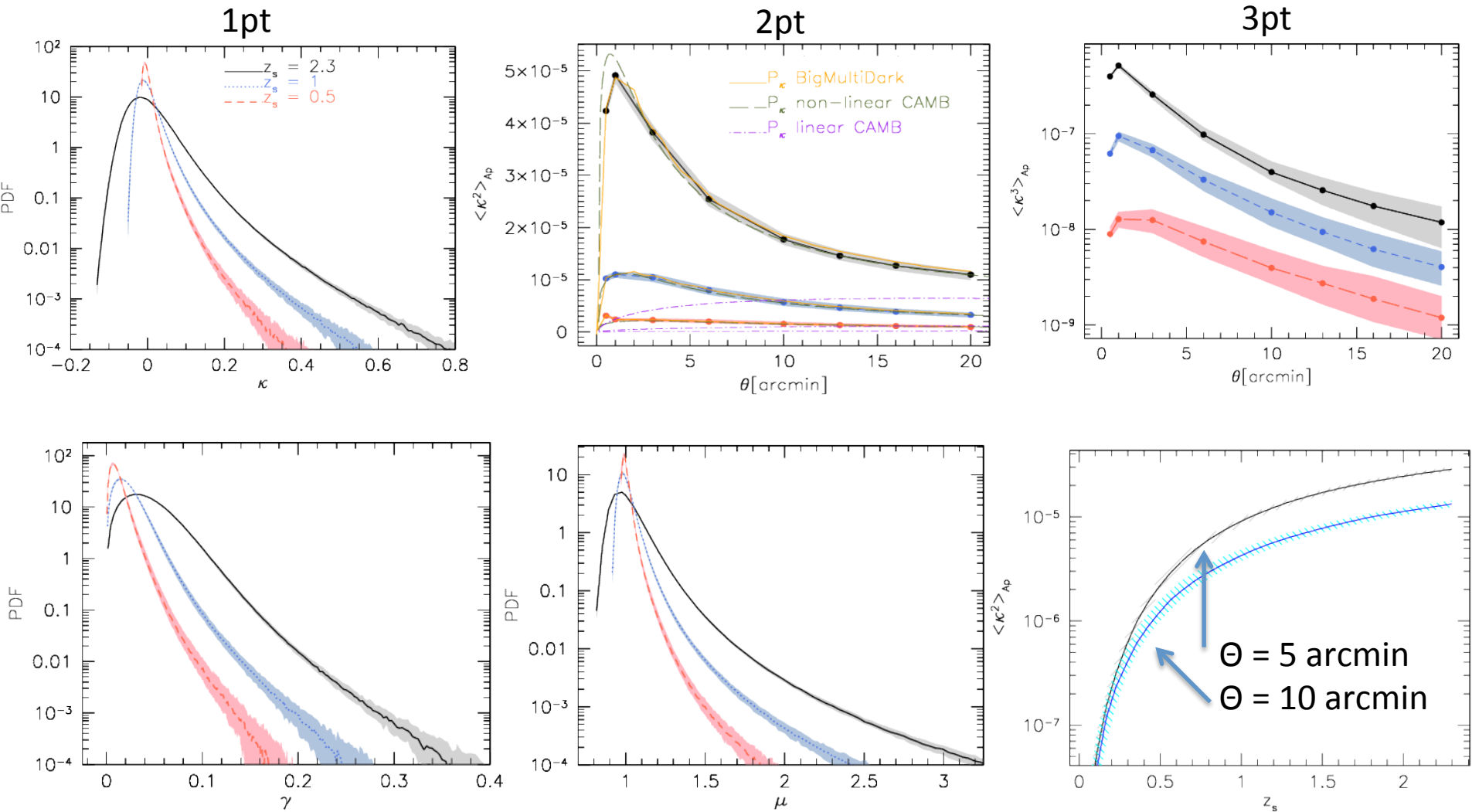
→ Good accuracy (<10% bias) up to $l \sim 3000$



Covariance matrices on $P_{\kappa}(l)$

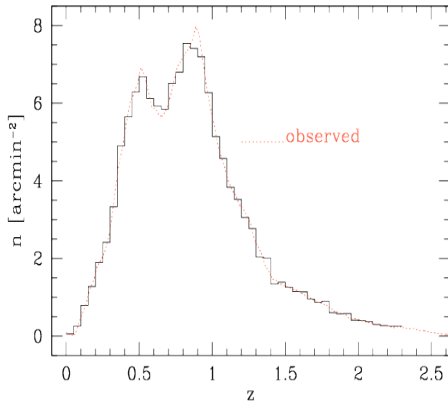


Lensing statistics checks



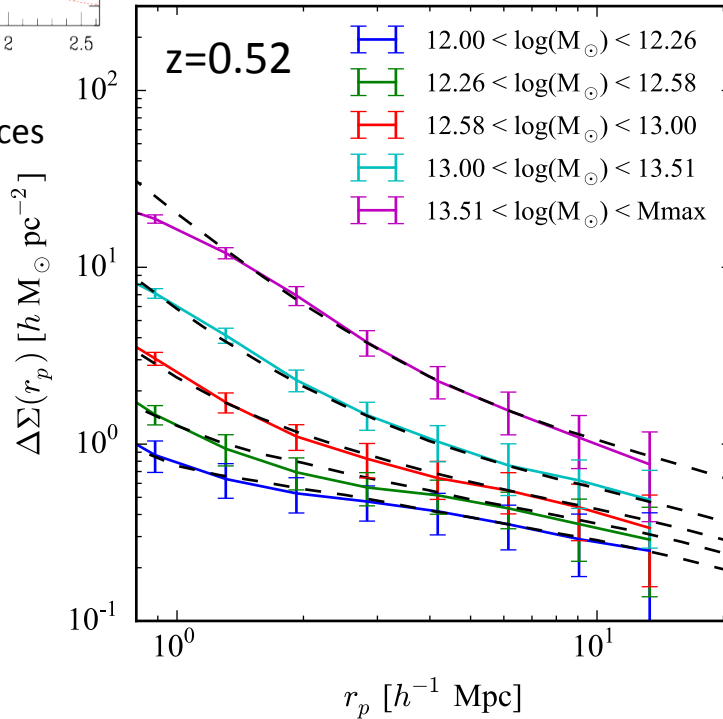
➔ We verify that our lensing measurements are in agreement with theory

Halo lensing

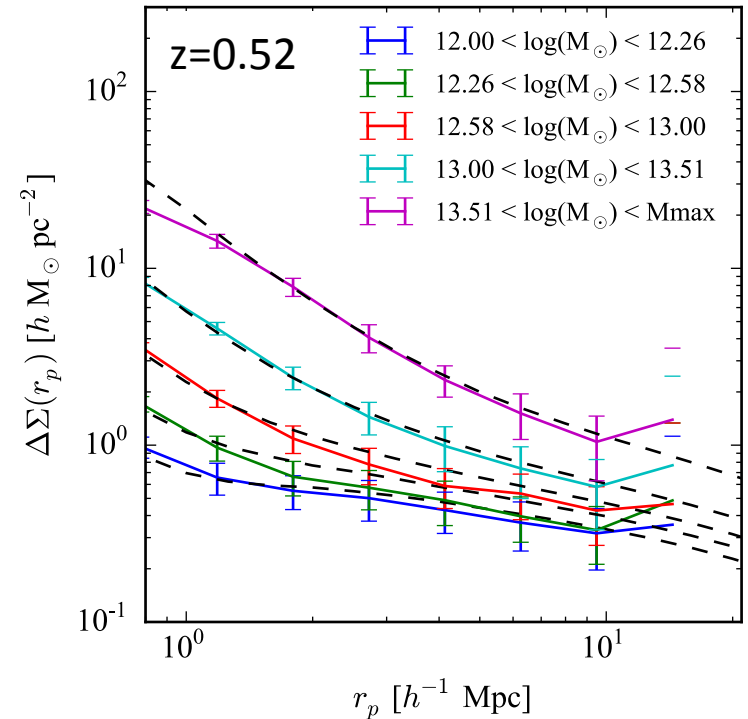


Simulated $n(z)$ from CFHTLenS for the sources

Lensing measurements

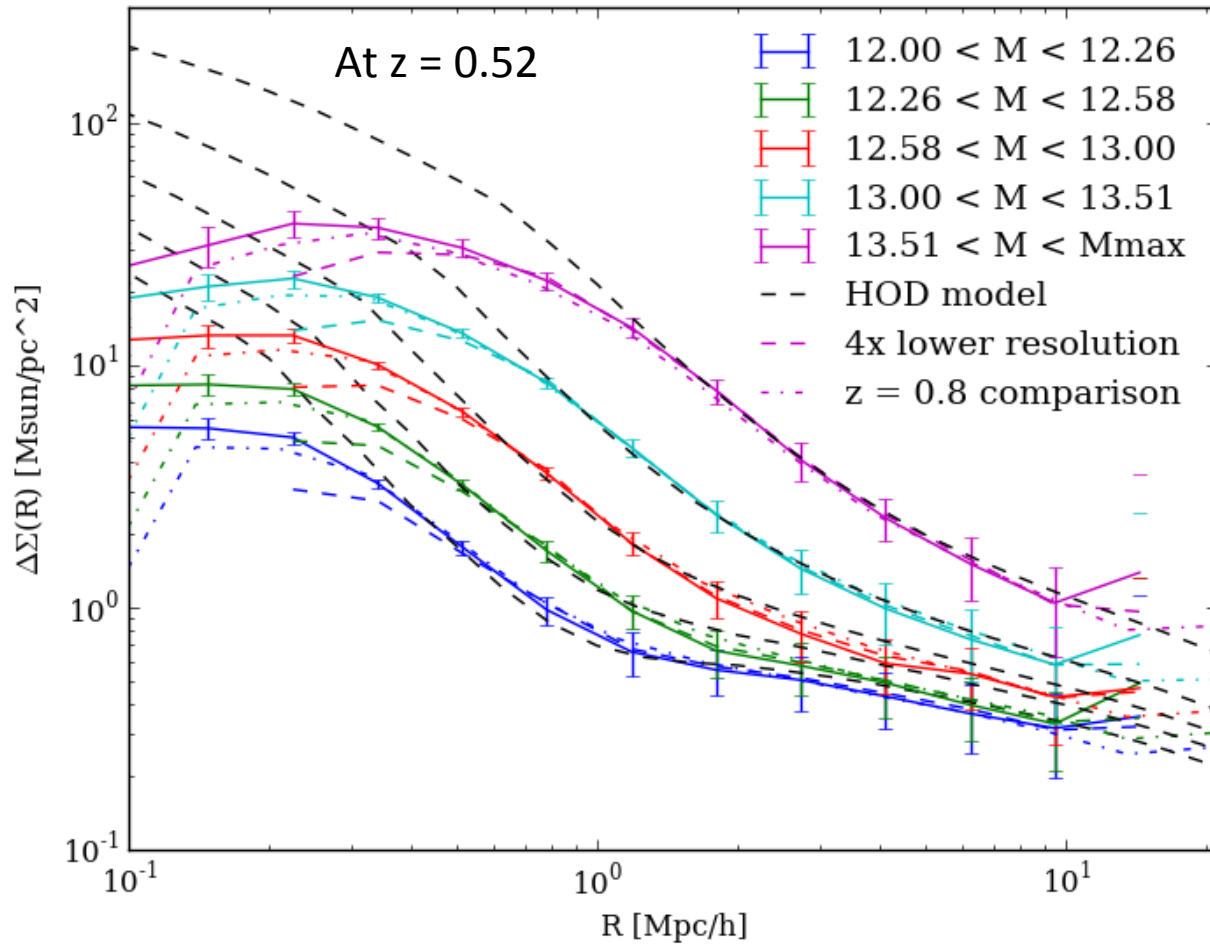


Equivalent density projection



- ➔ Well reproduced with standard HOD model (van den Bosch 2013)
- ➔ We verify that our lensing measurements are unbiased estimators of the underlying matter density

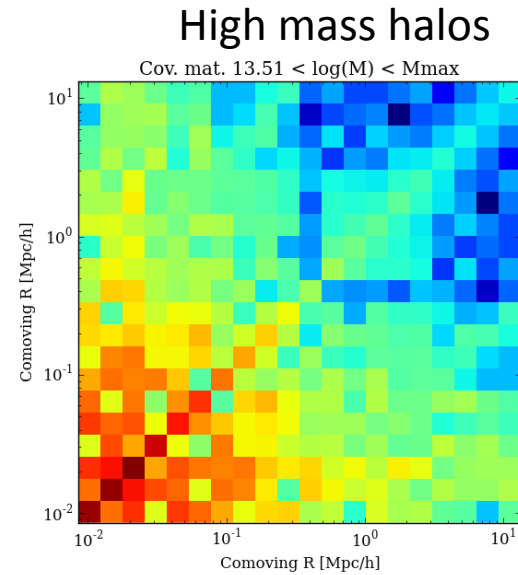
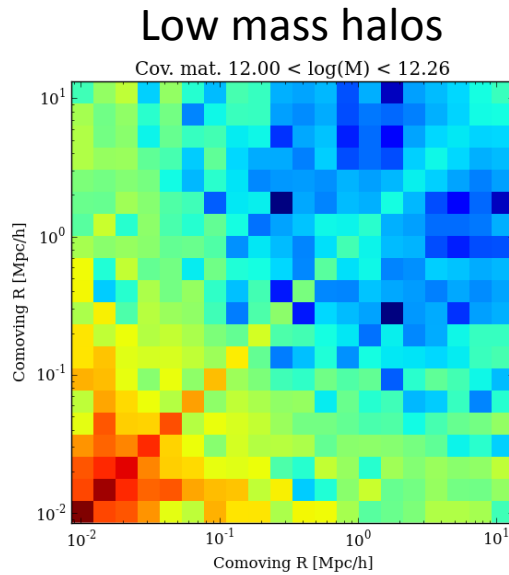
Impact of the map resolution



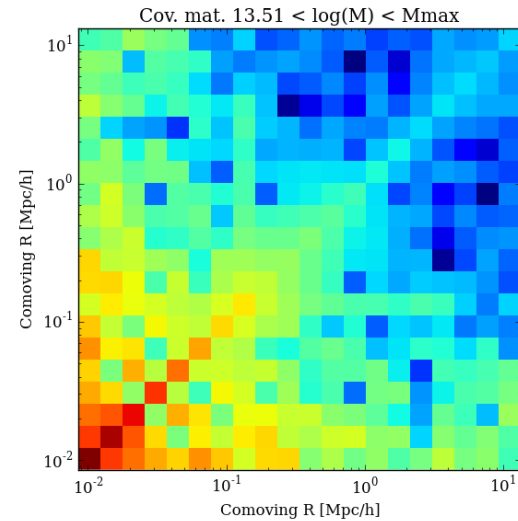
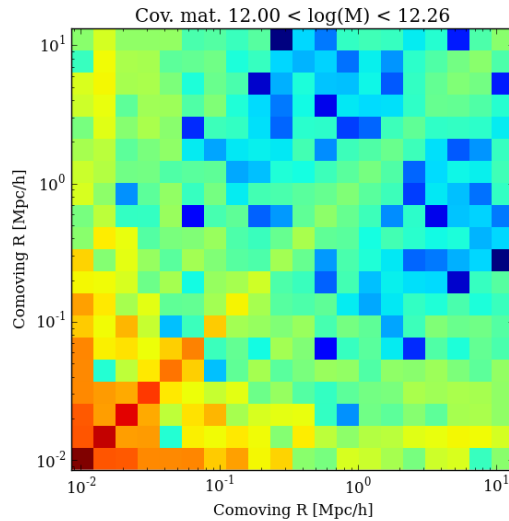
- ➔ Map resolution effect starting below $R \sim 0.3 \text{ Mpc}/h$
- ➔ Halo mass dependent mismatch with theory

Halo lensing covariance matrices

$z=0.8$



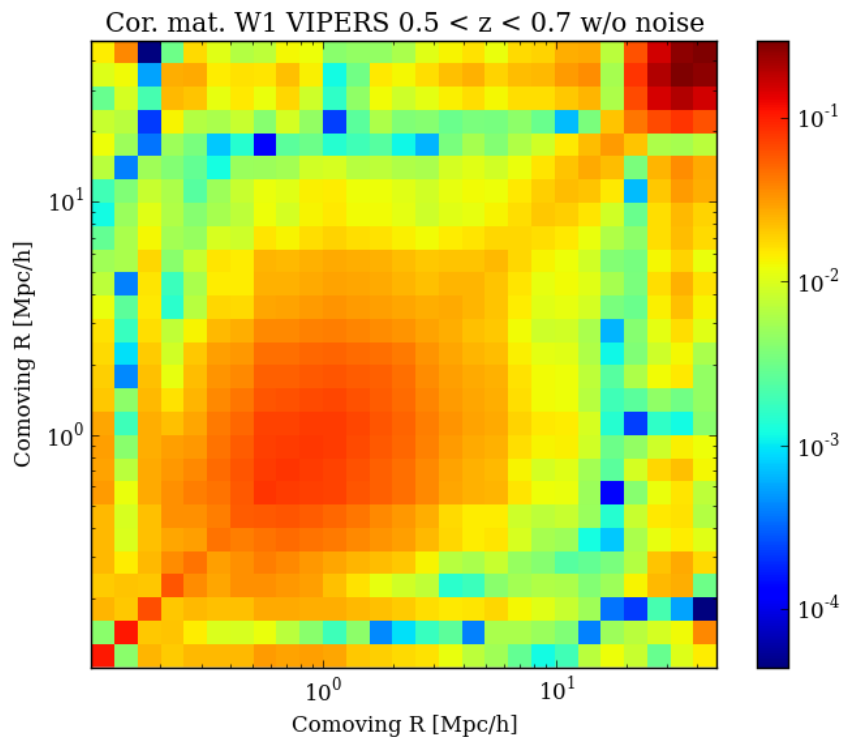
$z=0.52$



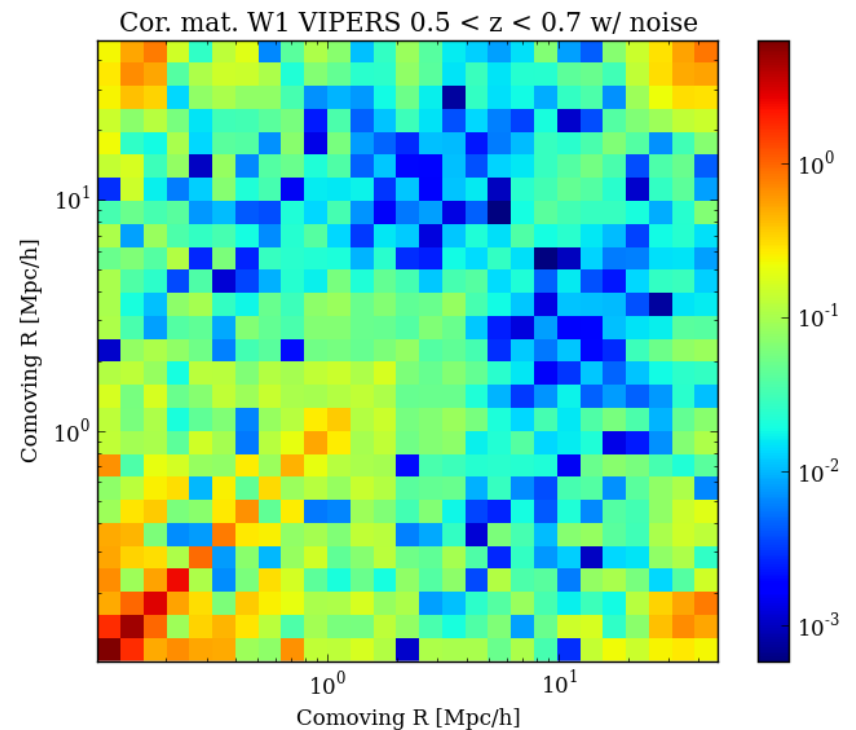
VIPERS covariance matrices

HOD model taken from de la Torre et al. 2013

Without WL shape noise



With WL shape noise



→ We need more mocks to decrease the noise

Future projects

- With



VIPERS – W1 fields

[Realization 1](#) (shear and convergence maps)

All realizations (54) (shear and convergence maps):

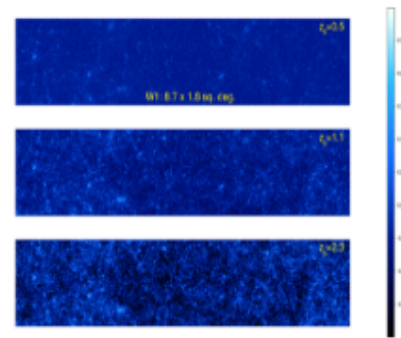
[1-9](#) [10-19](#) [20-29](#) [30-39](#) [40-49](#) [50-54](#)

[Lensing Catalogue for the realization 1](#) (8 random samples of the field of view)

[Lensing Catalogues for all realizations \(54 x 8\)](#) – each field of view and source redshift distribution are randomly sampled 8 times

[Power Spectra](#) all source redshifts and realizations

[Halo Catalogs](#)



VIPERS – W4 fields

[Realization 1](#) (shear and convergence maps)

All realizations (99) (shear and convergence maps):

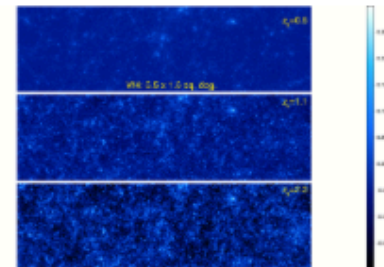
[1-9](#) [10-19](#) [20-29](#) [30-39](#) [40-49](#) [50-59](#) [60-69](#) [70-79](#) [80-89](#) [90-99](#)

[Lensing Catalogue for the realization 1](#) (8 random samples of the field of view)

[Lensing Catalogues for all realizations \(99 x 8\)](#) – each field of view and source redshift distribution are randomly sampled 8 times

[Power Spectra](#) all source redshifts and realizations

[Halo Catalogs](#)



Stripe82 – coming soon

- Lightcones uploaded on the Bologna lens Factory website for public access <https://bolognalensfactory.wordpress.com/home-2/muldarklens/>
- Acknowledgement to Giocoli et al. 2015

Future project

- CESAM is developing a web-service hosted by LAM to automatically produce lensing and clustering catalogs from a given redshift distribution
- We are working on producing catalogs for any field size without an N-body simulation (much faster and less resource intensive)

Conclusion

- We produced lensing lightcones from the BigMultidark cosmological boxes
- We extensively tested the lensing catalogs and found our measurements to be reliable down to scale $R \sim 300$ kpc/h
- We computed covariance matrices for halos, and we are also preparing for galaxy mocks (VIPERS, CMASS, etc)
- Covariance matrices are still noisy and we need to increase the number of lightcones