## Star-Galaxy classification in SDSS DR12

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## Aim

#### To automatically separate stars from galaxies by using the objects colours in absence of spectroscopical data.

## Nearby galaxies

Iuminosity spread on CCD: stars ~ 1 arcsec galaxies ~ 10 arcsec



# Far galaxies

**luminosity spread on CCD:** 

stars ~ 1 arcsec

galaxies ~ 1 arcsec

### SDSS DR12 spectroscopic sample ~ 2,300,000 stars+galaxies



#### Colour indices as "features" for classification



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### Supervised Classification



Parameters of the separating curve are derived by the logistic regression method.

# Logistic regression

- We take into account 10 colours (c1=u-g, c2=u-r, ...) plus one magnitude (u) and their quadratic function (ci.cj) to have 77 features.
- The separation region is constrained by a 11 dimension hyper parabola defined by 78 parameters.
- From ~600,000 stars and ~1,500,000 galaxies, we randomly put 20000 stars and 20000 galaxies into the training sample.

### Results from the fit



### Results from the classification

- Classification efficiency for the whole sample: 92% galaxies: 93% stars: 90%
- Mean redshift of the galaxies classified wrongly: 0.14 correctly: 0.25
- Mean colour (extinction corrected) of the stars classified wrongly: u-z = 4 (redder stars) correctly: u-z = 2

### Conclusions & Perspectives

- in SDSS DR12, more than 90% of galaxies and stars can be correctly separated using their colours and implementing Logistic Regression.
- Nearby galaxies can be mis-classified as stars.
- Redder stars can be mis-classified as galaxies.
- Quasars of the SDSS will be included
- Classifying the simulated objects according to the LSST observation ability (higher redshifts and fainter objects).
- Comparison with other classifiers

### **Conclusions & Perspectives**

with LSST, MW-like galaxies can be resolved by morphology but not for faint galaxies (dwarfs)

