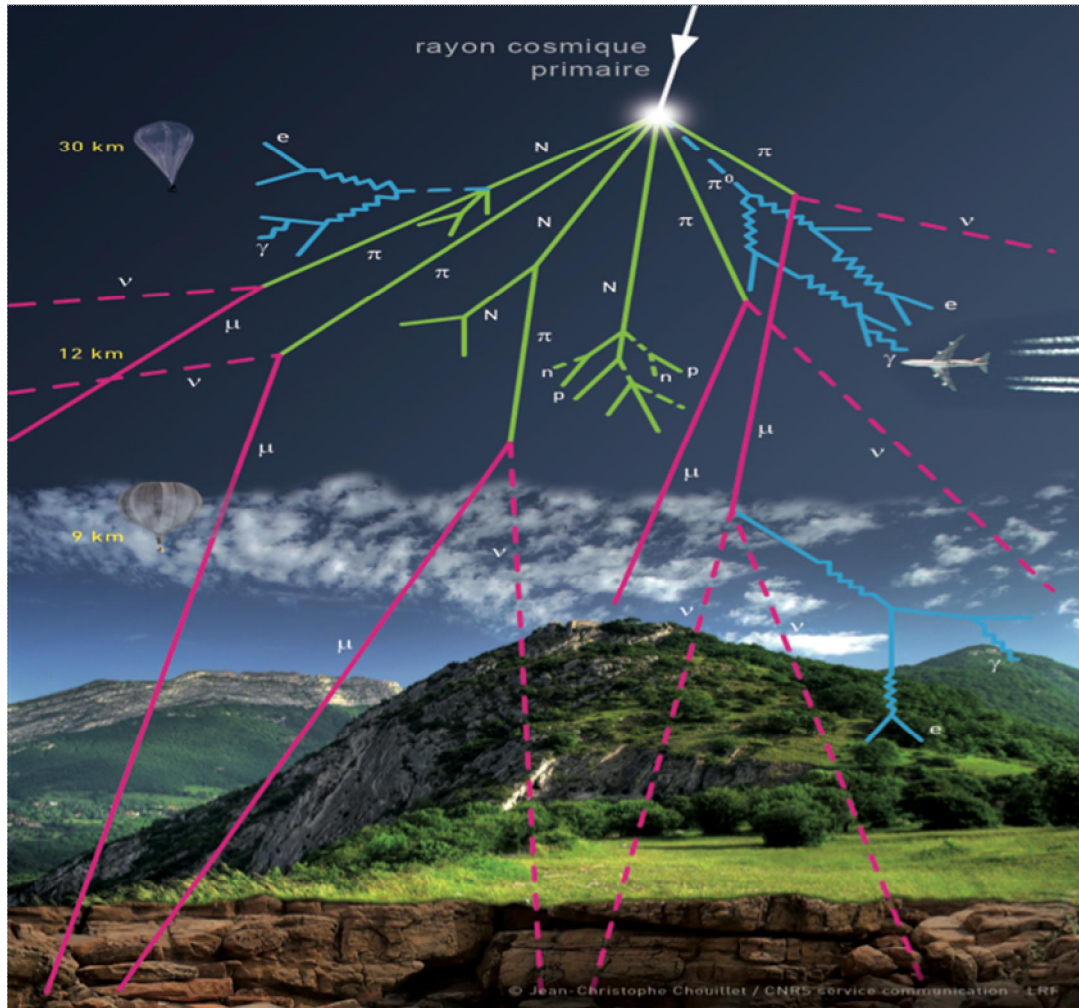


Can we detect a marble tomb with cosmic muons ?

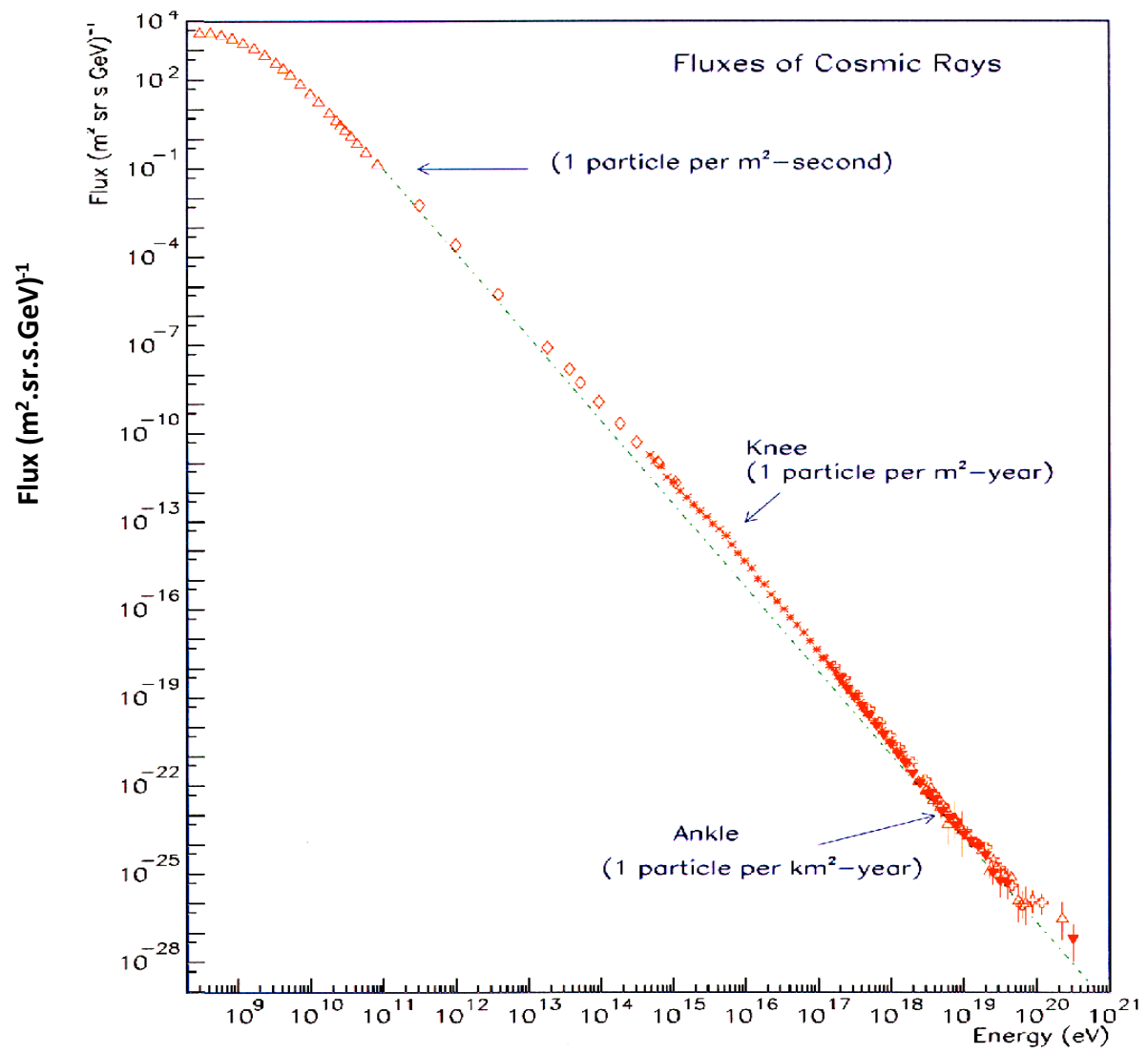
Corinne Goy, Max Chefdeville, Jean Jacquemier, Yannis Karyotakis

21 December 2015

Primary Cosmic rays

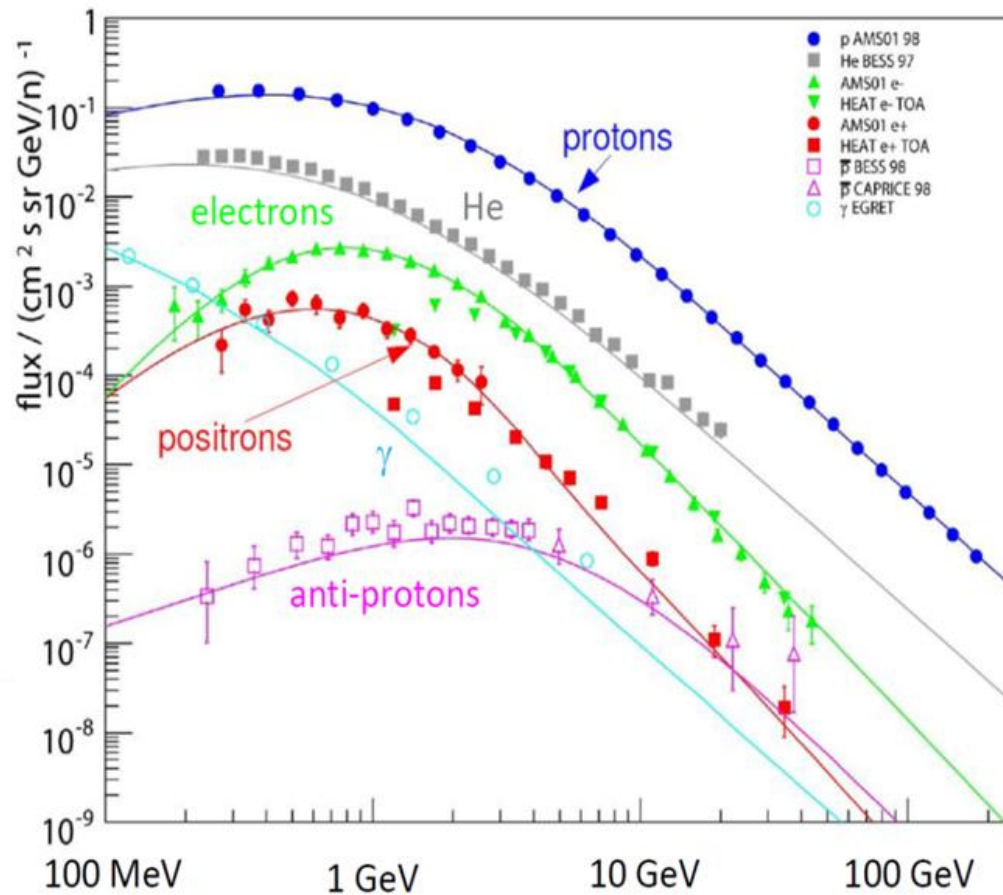


Showering
in
atmosphere
→
Destruction



1 GeV = 10^9 eV

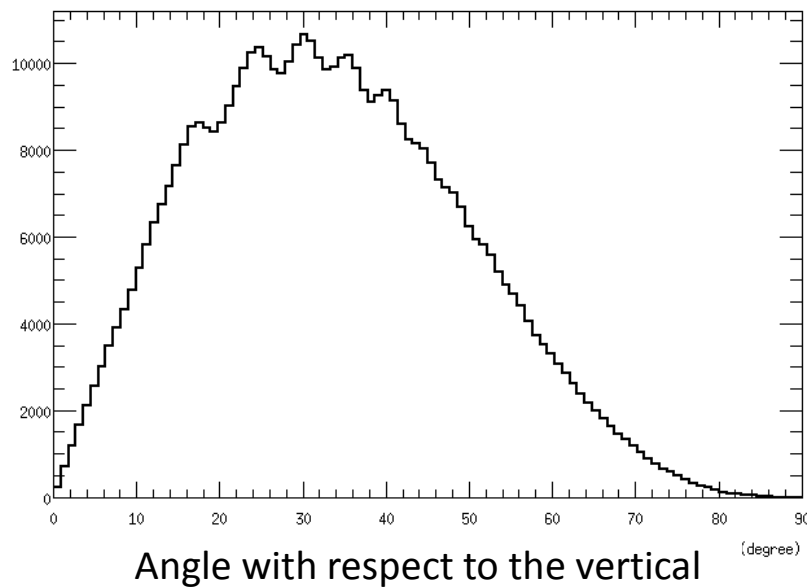
Composition of primary cosmic rays



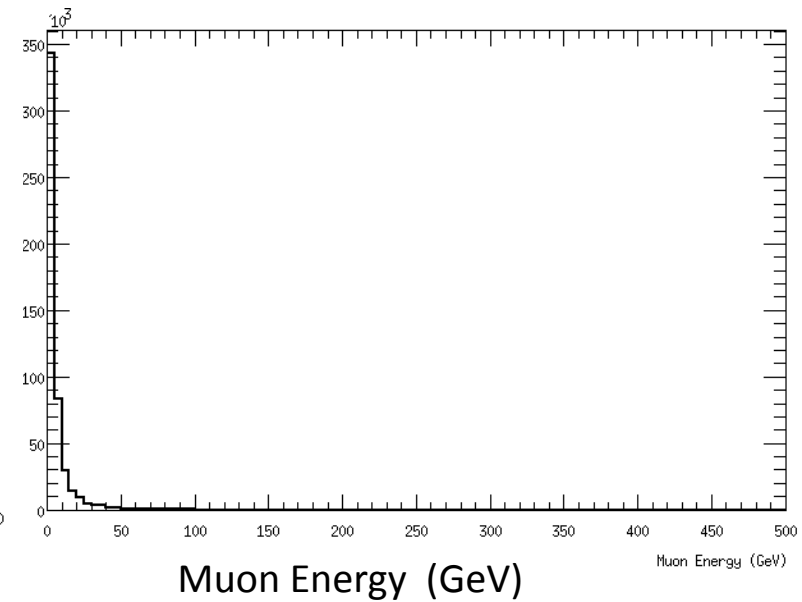
90 % protons
10 % Helium
1 % electrons
0.1 % positrons

Sea – level : remnants of primary CRs – essentially muons (and neutrinos)

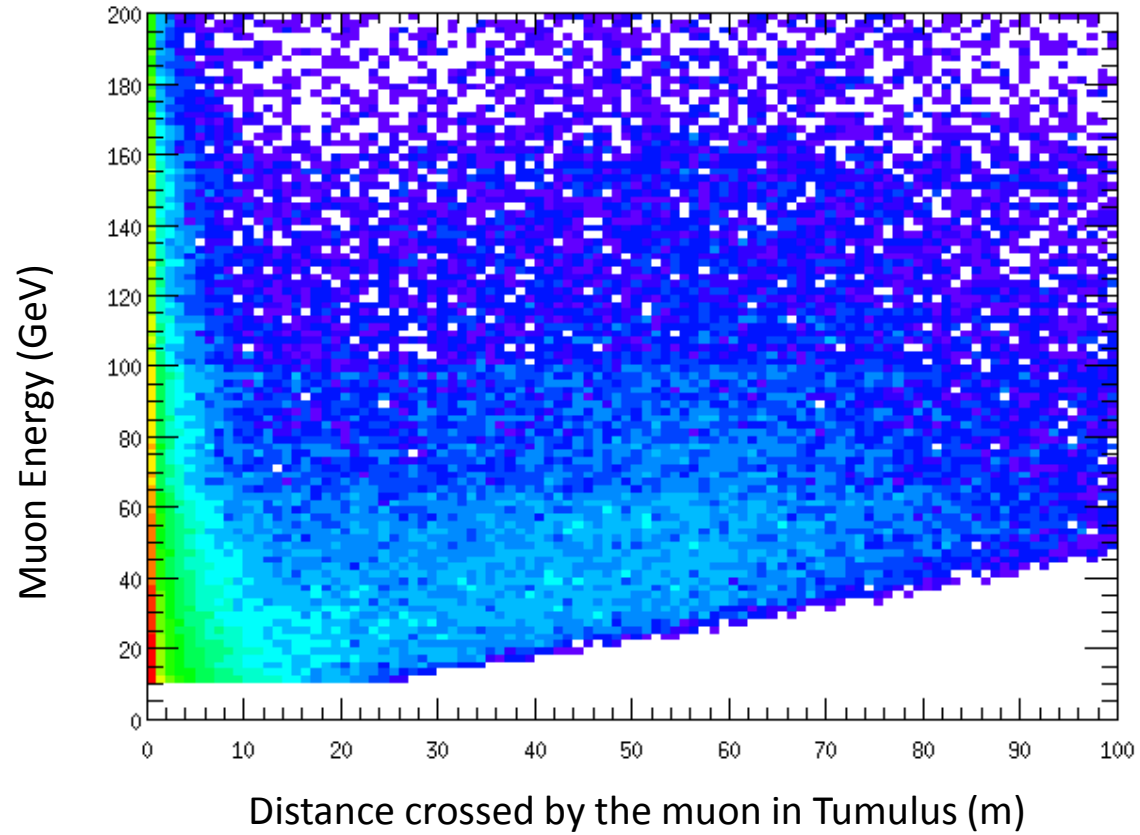
Few Horizontal muon



Low energy



We are interested in rather high energy muons ,
otherwise they are absorbed in the tumulus
before reaching the detector



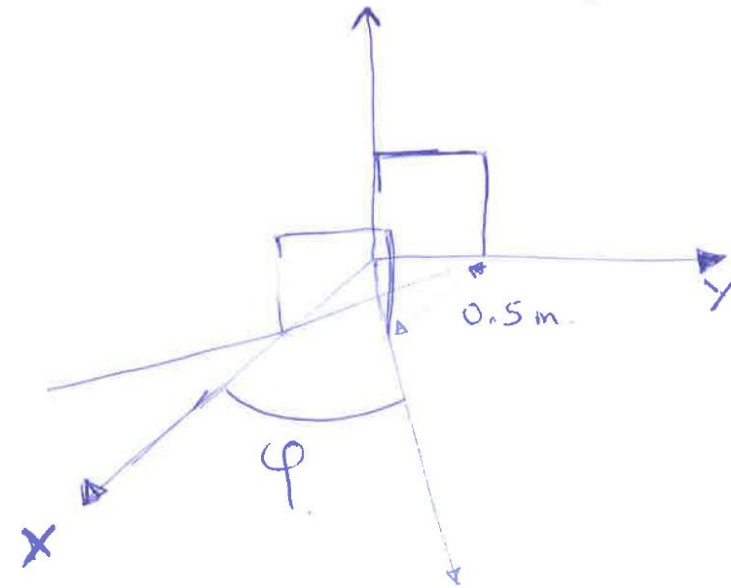
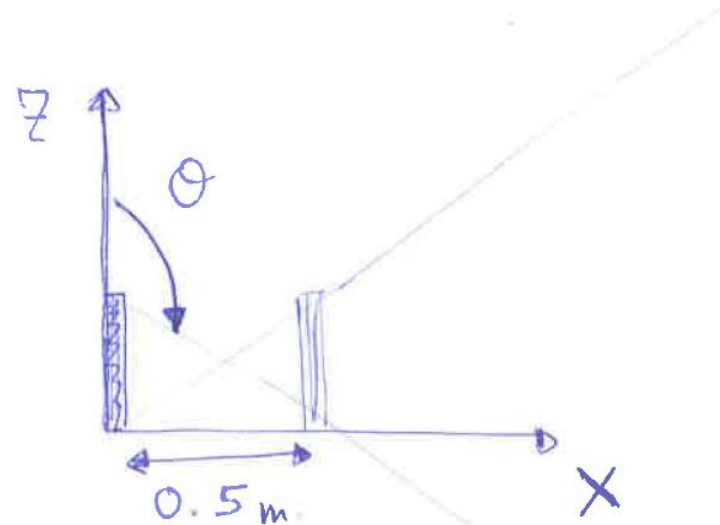
Acceptance Studies

Methods and prelim. Results

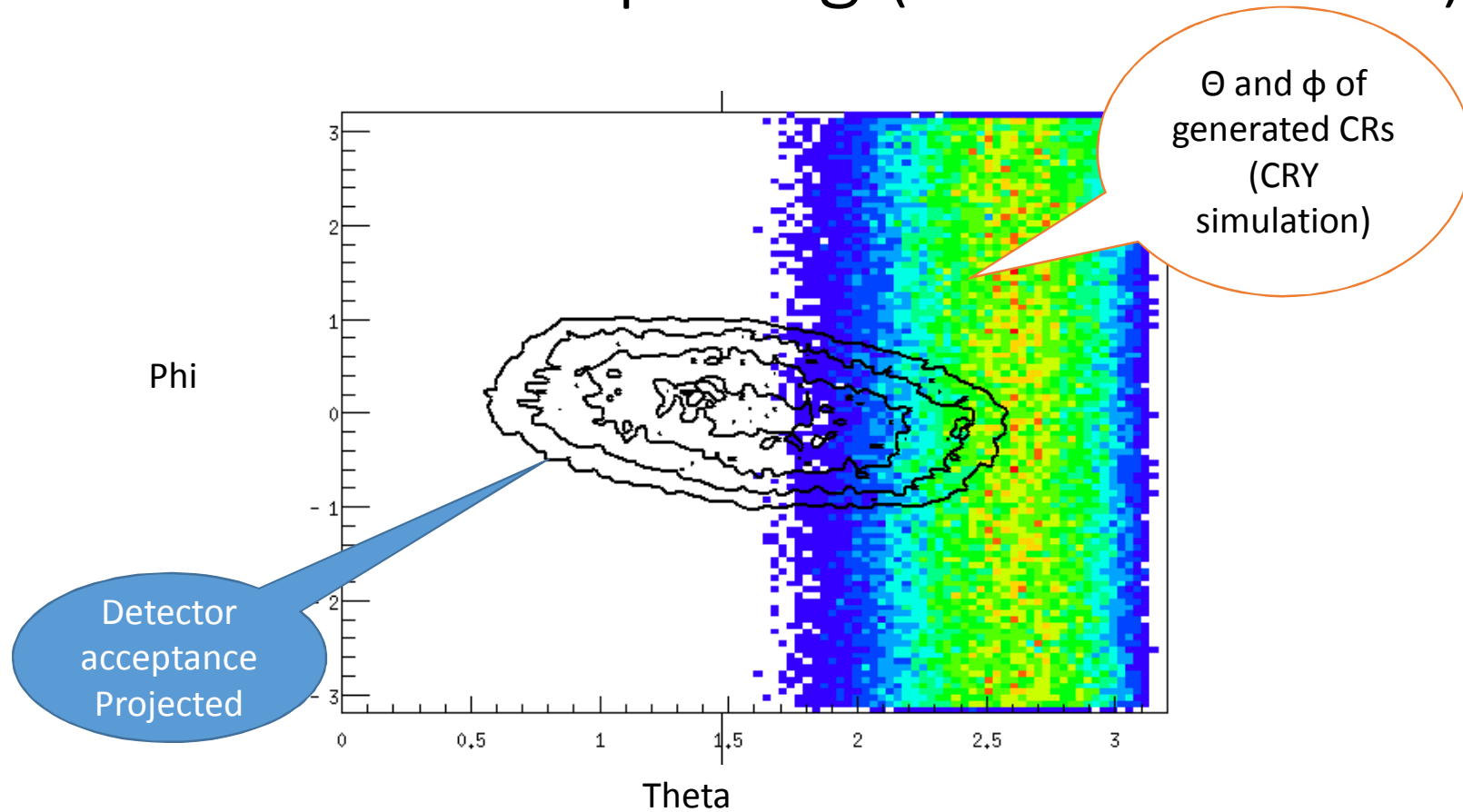
Toy MC

- “ Generate random Y & Z on a 1 square meter. (\sim 1 chamber) at $X = 0$.
- “ Generate ϕ in $[-\pi; +\pi]$
- “ Generate θ in $[0; \pi]$
- “ Extrapolate track to $X = 0.5$ m (resp 2 m)

- “ Retain tracks that crosses the second chamber ie
 - “ $0 < Z < 1$ m
 - “ $0 < Y < 1$ m @ $X = 0.5$ m (resp 2m)

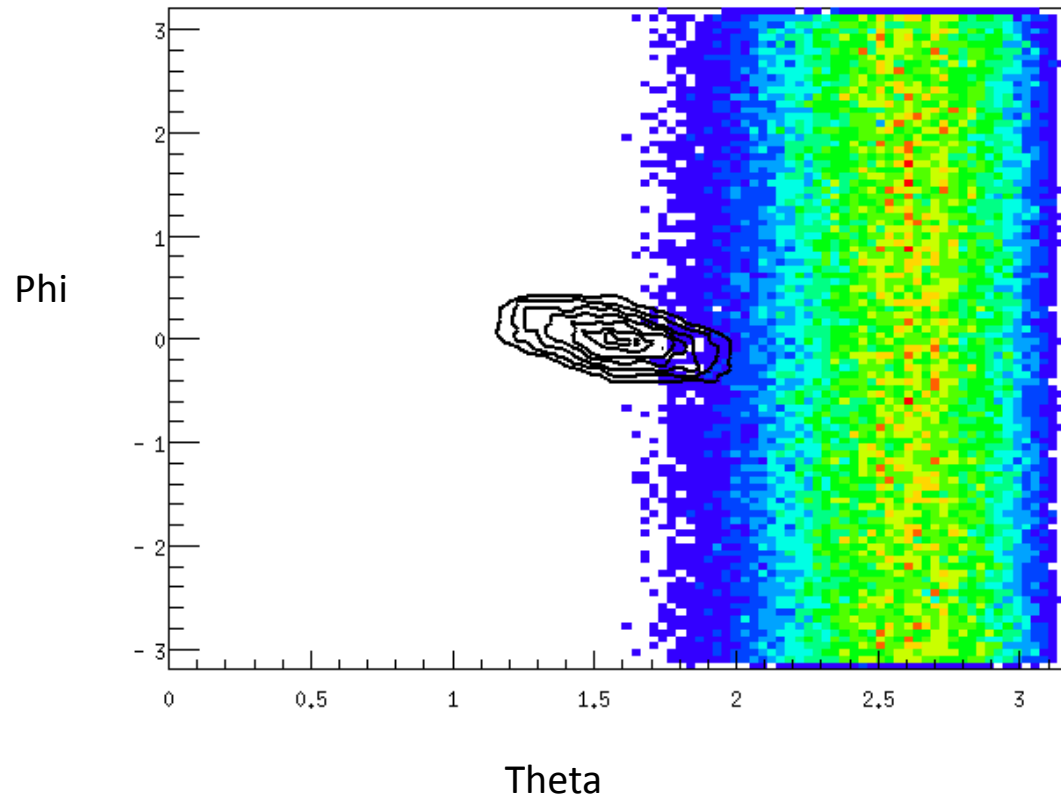


2 det. with 50 cm spacing (for illustration)

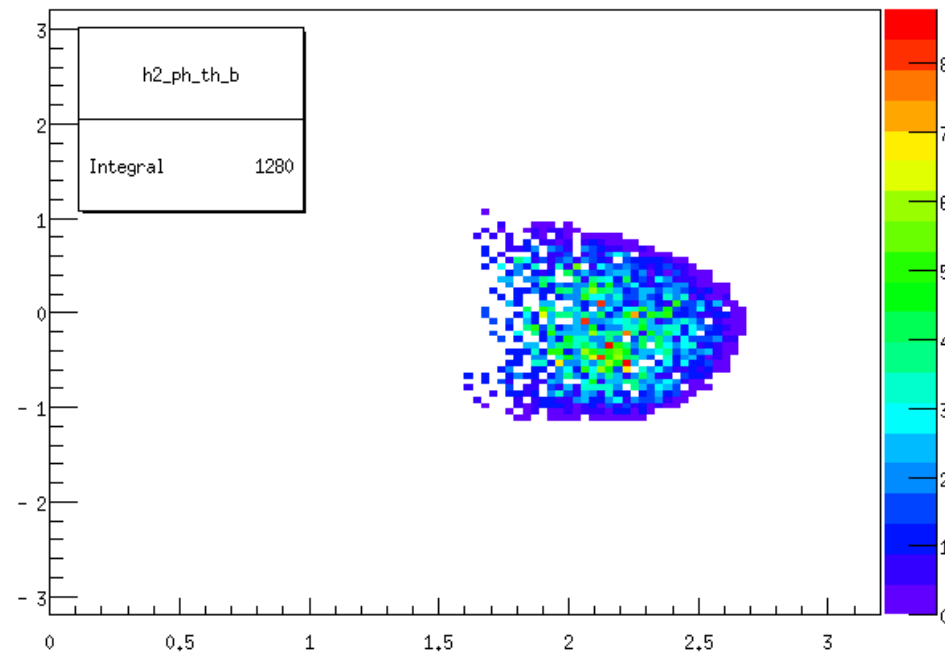


Inclined detector will move the pattern to the right

2 det. with 2m spacing (for illustration)



Number of muons per (θ, ϕ) bin = Number of generated * efficiency of the detector
normalized by the time (gen.SimulatedTime /CRY)



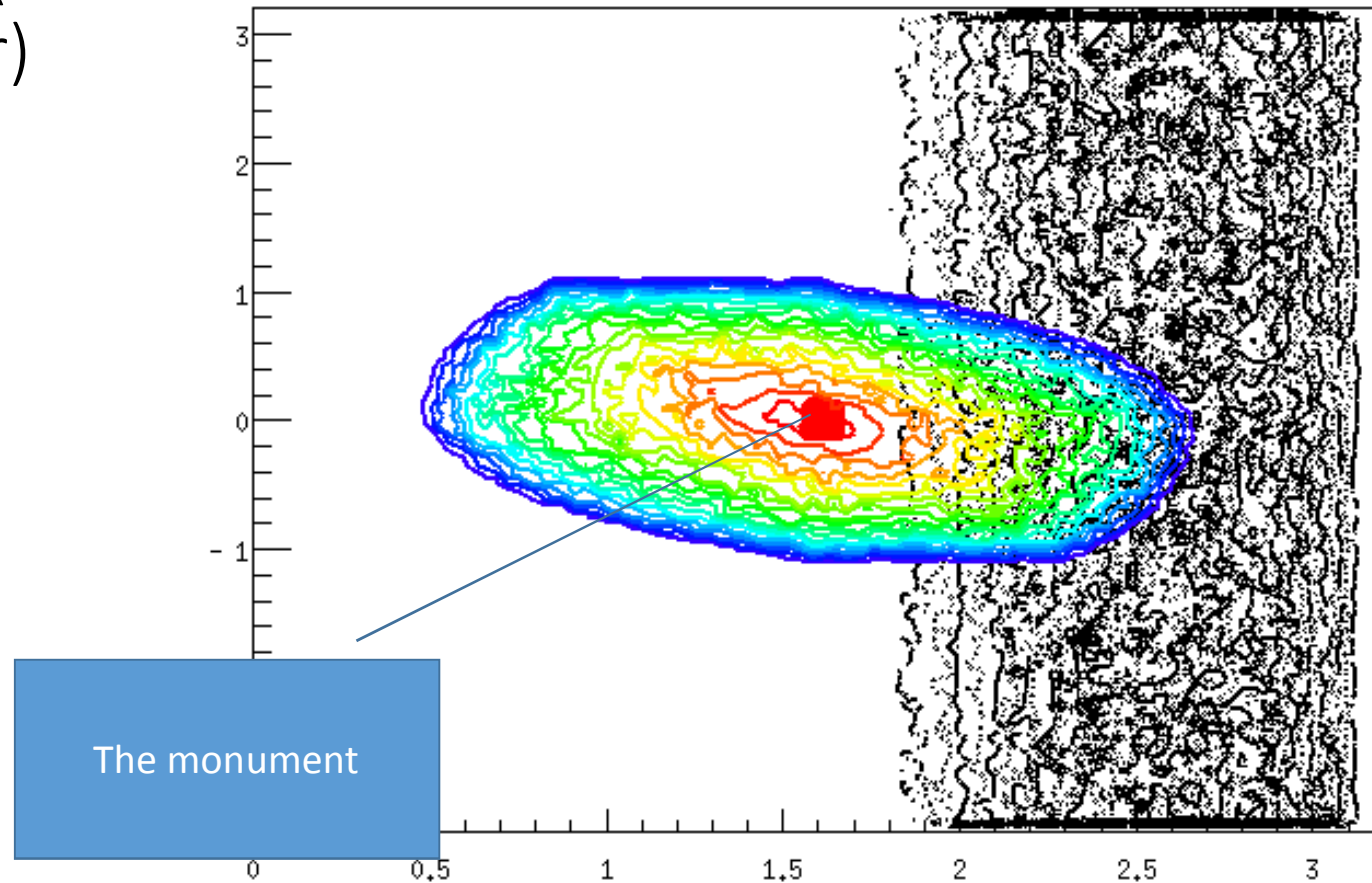
- “ Muons ($E > 20$ GeV) per hour in the “toy” detector
- “ Detector horizontal

Exp Muons per hour in the det acceptance (w/o tumulus)

E _{min} (GeV)	L _{box} = 300 m Horizontal	L _{box} = 10m Horizontal	L _{box} = 300 m Det @ 20deg
8	3830	4920	9107
10	2980	3910	6920
15	1830	2470	4095
20	1280	1740	2785
25	900	1220	1907
50	300	430	600
100	70	110	135

Solid Angle of the monument (seen from the detector)

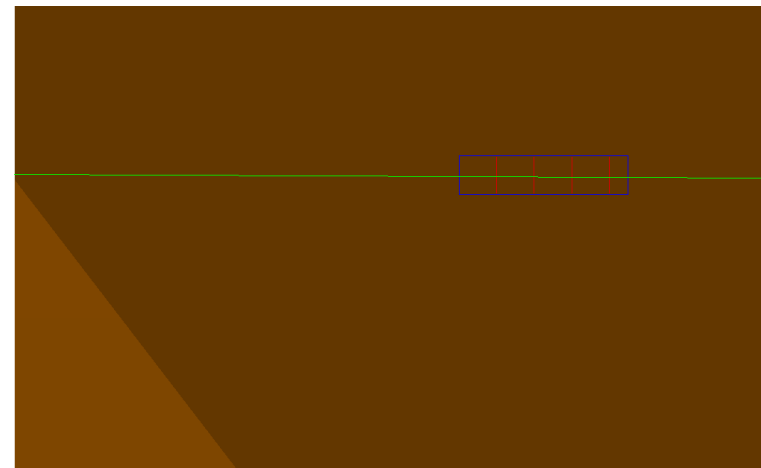
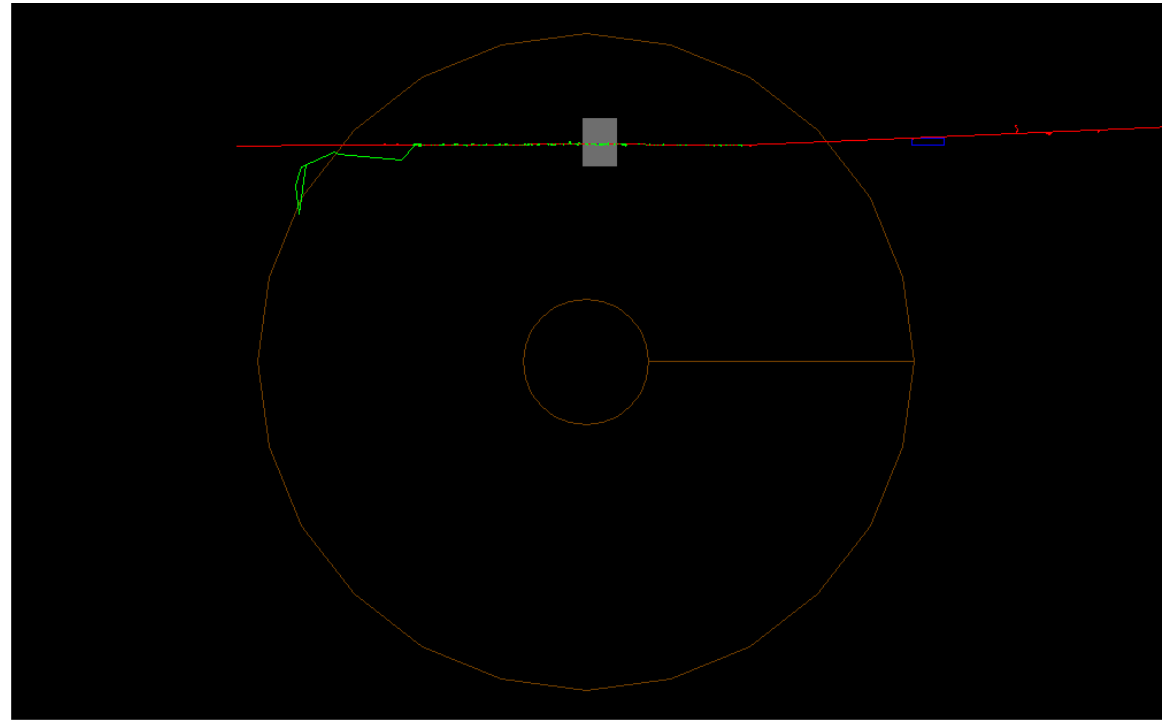
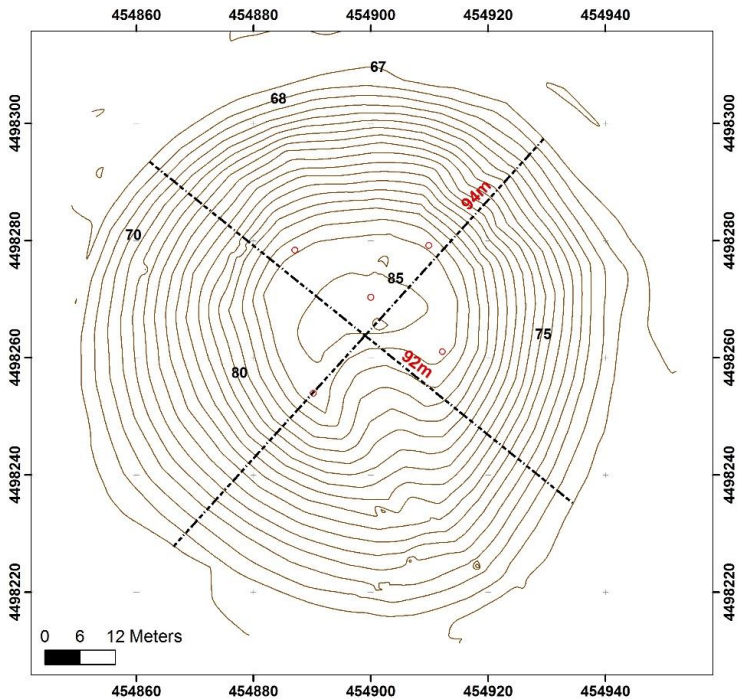
“ Det @ $y = 31.5$ m
 $x = 49.0$ m
 $z = -0.5$ m



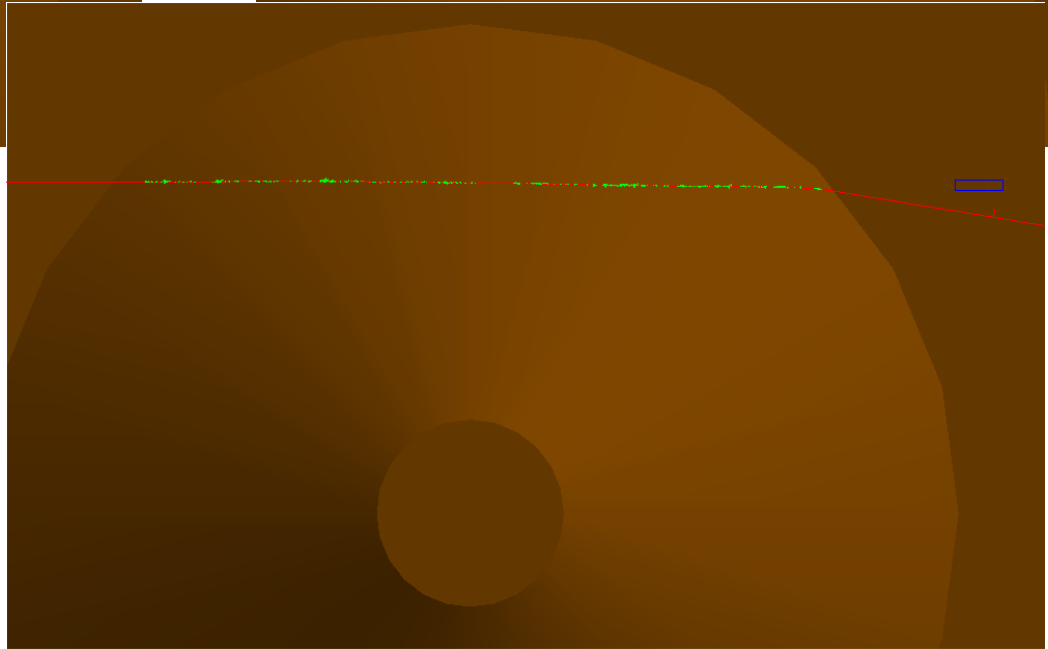
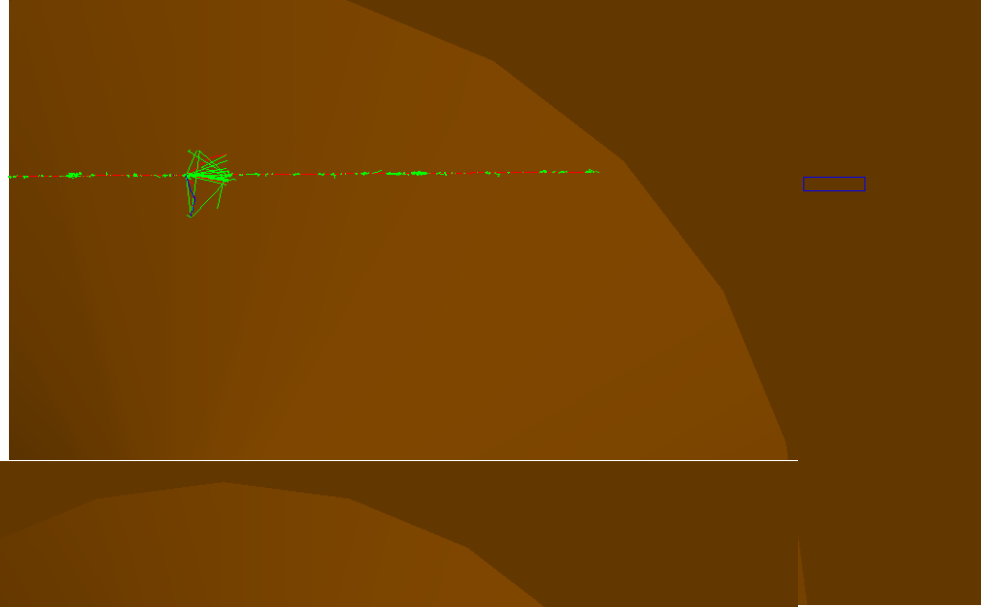
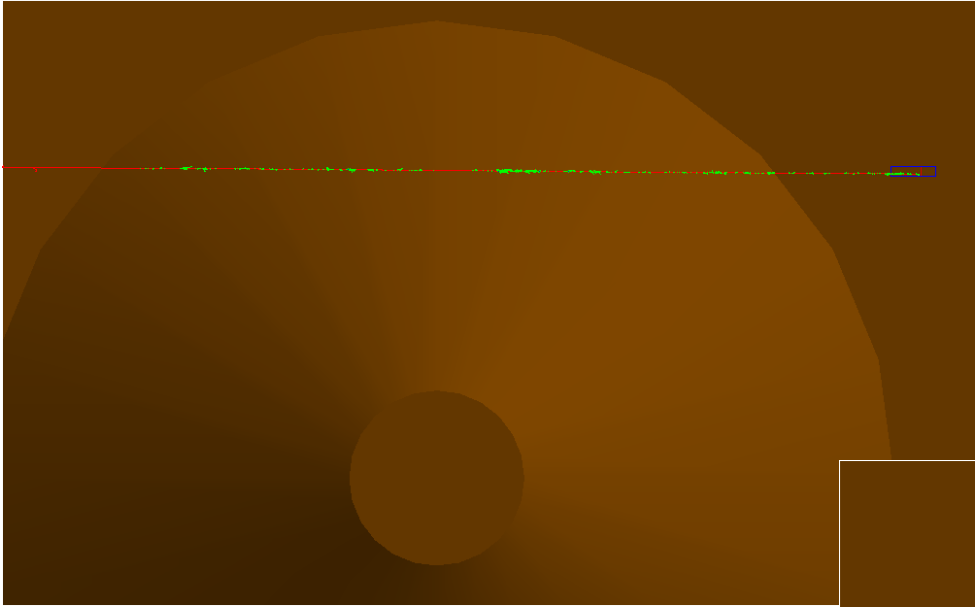
Full simulation

- “ Define geometry : Tumulus, monument, detectors
- “ Generate muons using an energy parametrization from CRY, and limited θ and ϕ range
- “ Track each muon through matter, tumulus soil, marble air, detector material etc and record hits on the detectors if any.
- “ Reconstruct the muon track if detected inside the detector layers

Simulation layout



Tumulus : Cone $R_{\text{bottom}} = 47\text{m}$ $R_{\text{top}} = 9\text{m}$ Dirt with $d = 2.3\text{gr/cm}^3$
Tomb: Size $5 \times 5 \times 7\text{ m}$, walls 1 m thick (CaCO_3 or Pb !) placed inside tumulus at $R = 31.5\text{m}$ and 2.5m below the ground. Air inside tomb
Detector : 4 $1 \times 1\text{ m}^2$ layers placed at $R = 49\text{m}$ in front of the tomb



Two tombs were simulated :

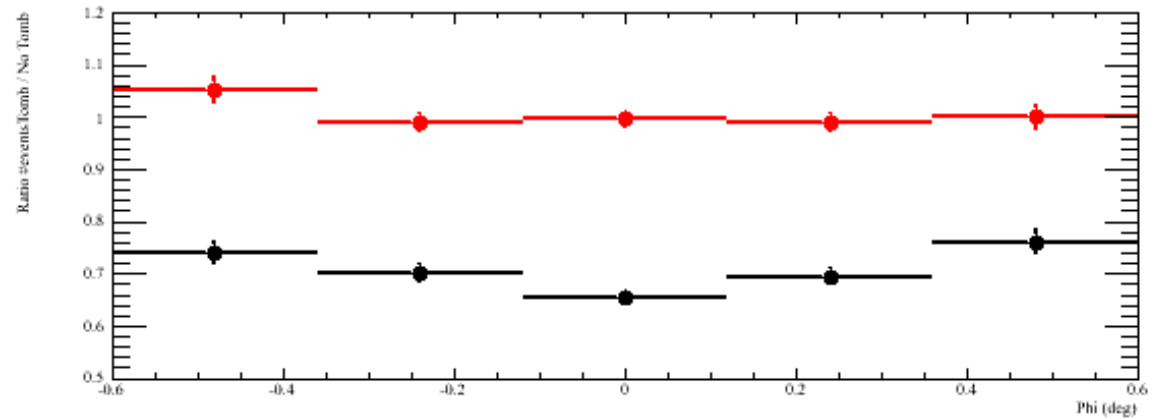
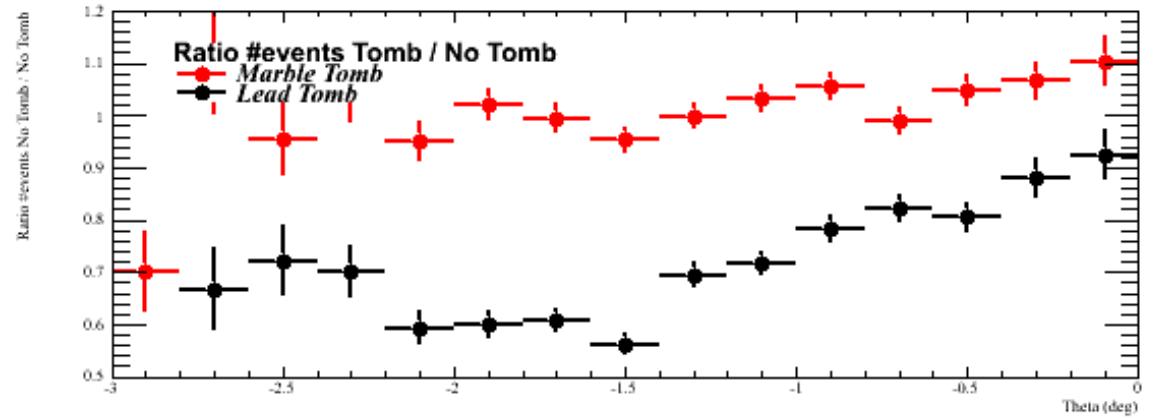
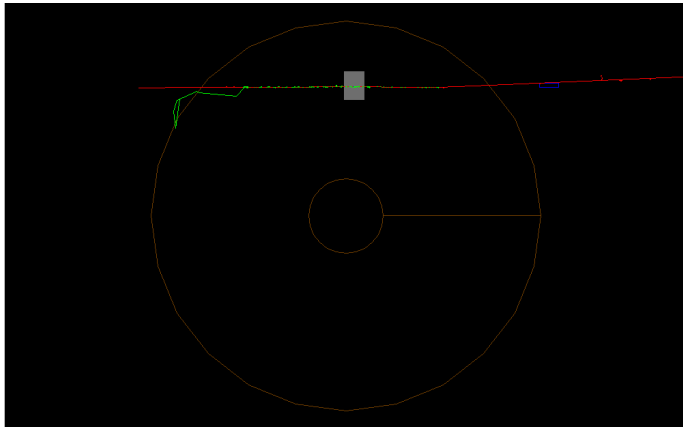
“ A marble tomb

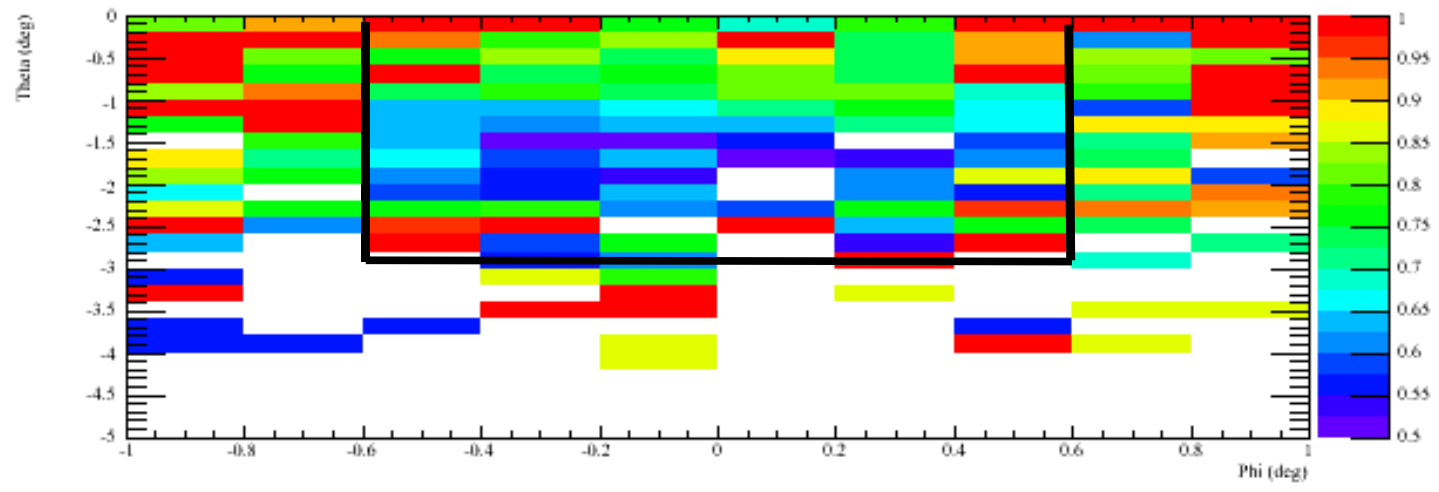
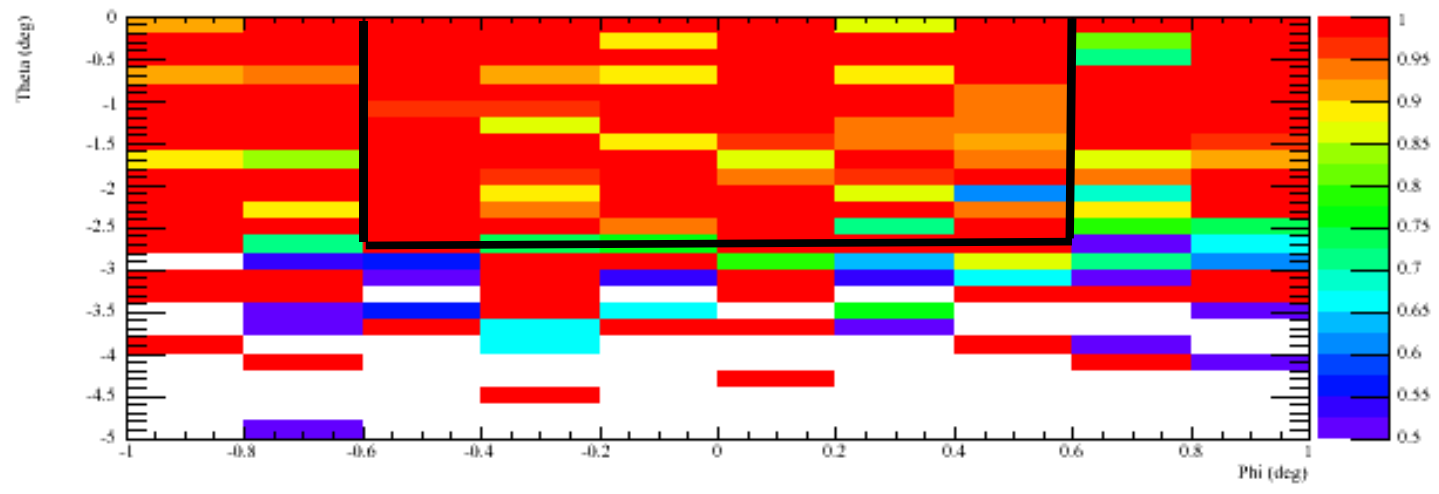
“ A lead tomb !

“ 1.2 M events generated to cross the monument if it is there

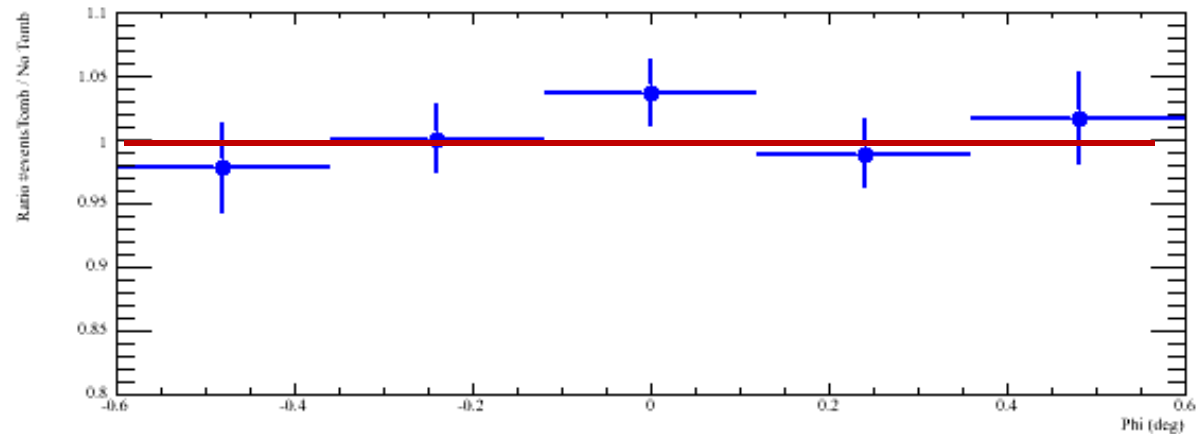
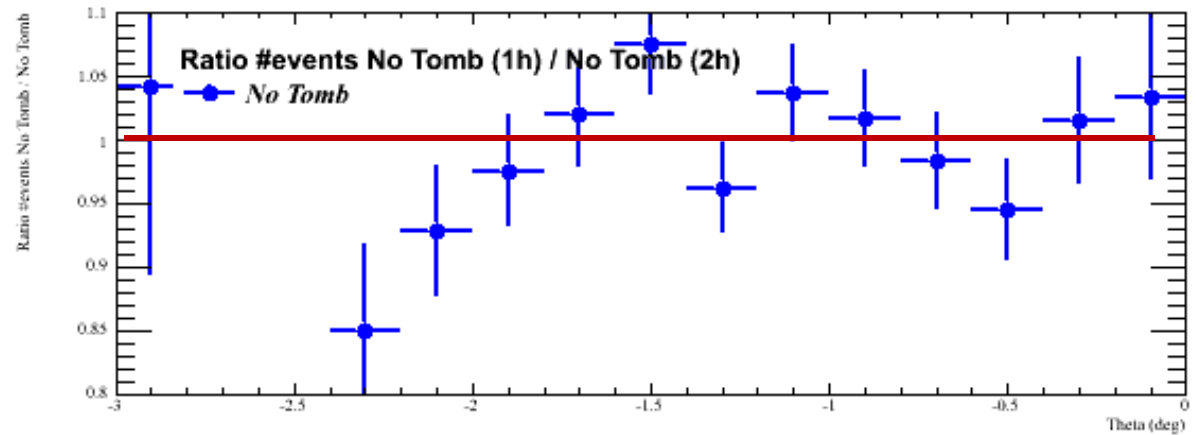
“ $87^\circ < \theta < 90^\circ$ and $-0.6^\circ < \phi < 0.6^\circ$

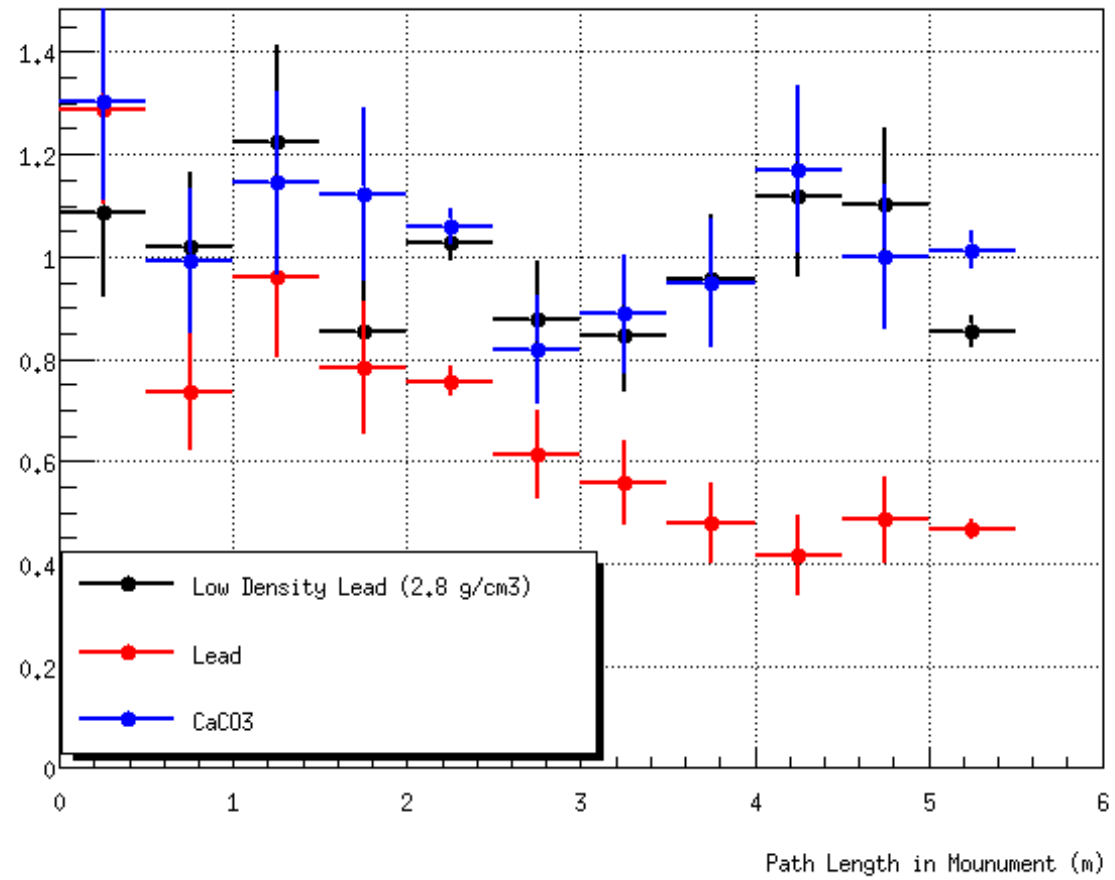
“ 3 runs : 2 with tombs and 1 without any monument





We have 1.2 M events without any monument.
Using 1st half (1h) of them to compare with second half (2h)





Conclusion

- “ Detecting a monument inside a tumulus will not be an easy task
 - “ Low statistics
 - “ Standard dirt and marble not very different
- “ Improve simulation
 - “ Use a different model (Fluka) for cross check
 - “ Input real soil composition. Need some chemical analysis.
 - “ Compare with data
- “ My proposal :
 - “ We have already one or two set of detectors. It is worthwhile to deploy them after March and make an exploratory run until summer.
 - “ Depending on findings, improve simulation and decide on a longer campaign