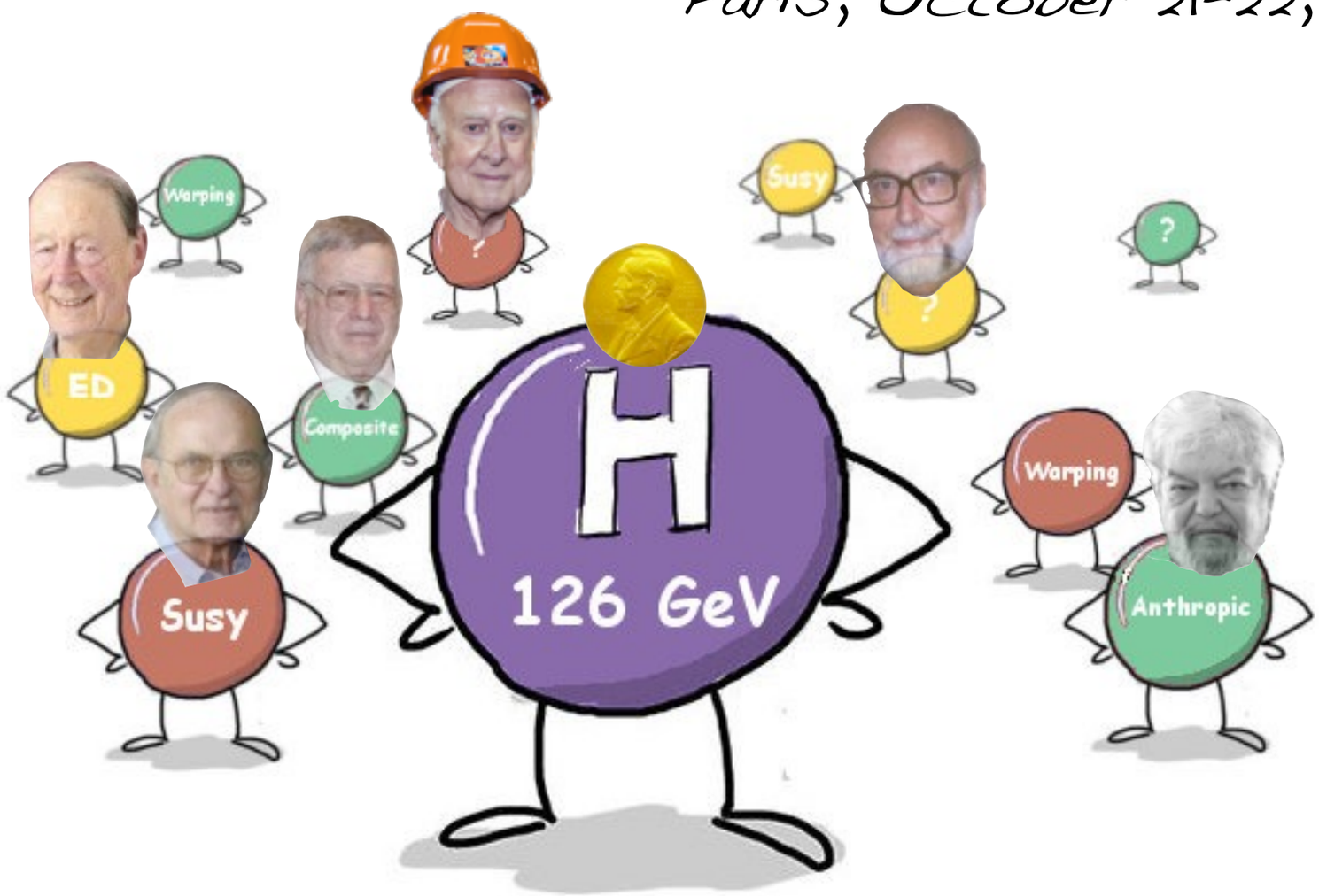
Beyond the Standard Model

Journées prospectives du LLR

Paris, October 21-22, 2015



P. Cámara/C. Grojean adapted from phdcomics.com

Christophe Grojean

DESY (Hamburg)

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The Higgs: Now what? What's Next?

" The experiment worked better than expected and the analysis uncovered a very difficult to find signal"

the words of a string theorist

Great success...

...but the experimentalists haven't found what the BSM theorists told them they will find in addition to the Higgs boson:
no susy, no BH, no extra dimensions, nothing ...

Have the theorists been lying for so many years?

Have the EXP's been too naive to believe the TH's?

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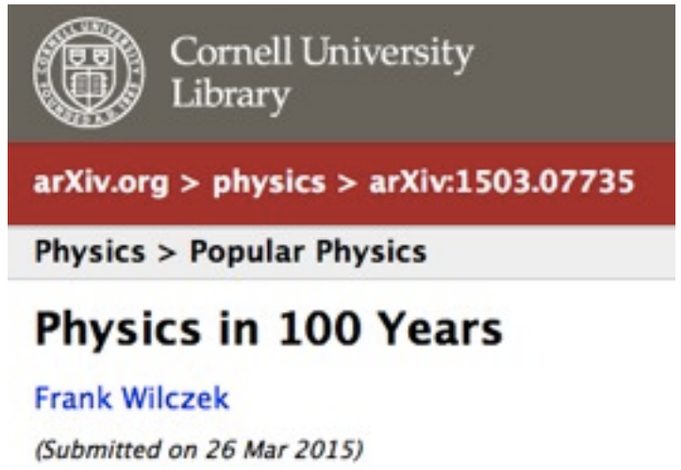
A unique moment in the history of physics

The Higgs discovery is the triumph of XXth century physics
combination of Quantum Mechanism + Special Relativity

For the first time in the history of physics,
we have a *consistent* description of the fundamental constituents of matter and their interactions and this description can be extrapolated to very high energy (up M_{Planck}?)


Cornell University Library

arXiv.org > physics > arXiv:1503.07735

Physics > Popular Physics

Physics in 100 Years

Frank Wilczek

(Submitted on 26 Mar 2015)

The equations of the [SM] have been tested with far greater accuracy, and under far more extreme conditions, than are required for applications in chemistry, biology, engineering, or astrophysics. While there certainly are many things we don't understand, we do understand the Matter we're made from, and that we encounter in normal life - even if we're chemists, engineers, or astrophysicists (sic: DM!)

The SM is not free of inadequacies: (without forgetting flavor and neutrinos)

1) Only a description of EW symmetry breaking, not an explanation
   ➠ What separates the EW scale from the Planck scale?
2) No place for the particle(s) that make up the cosmic DM
   ➠ What are the DM particles?
3) Does not explain the asymmetry matter-antimatter
   ➠ Are the conditions realized to allow for EW baryogenesis?

we do not understand the Matter the Universe is made from

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Where and how does the SM break down?

Which machine(s) will reveal (best) this breakdown?

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we do not understand the Matter the Universe is made from

Which Machine(s)?

Hadrons

- large mass reach ⇛ exploration
- S/B ~ 10^{-10} (w/o trigger)
- S/B ~ 0.1 (w/ trigger)
- requires multiple detectors
  (w/ optimized design)
- ⇛ couplings to quarks and gluons

Leptons

- S/B ~ 1
- polarized beams
  (handle to chose the dominant process)
- limited (direct) mass reach
- identifiable final states
- ⇛ EW couplings

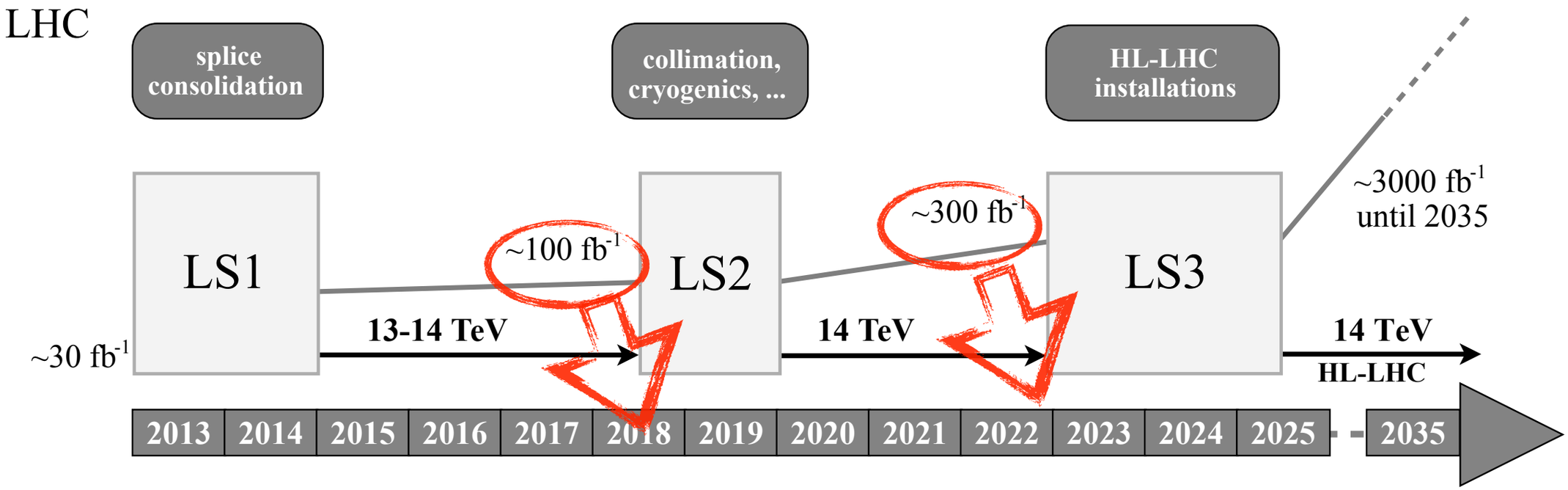
Circular

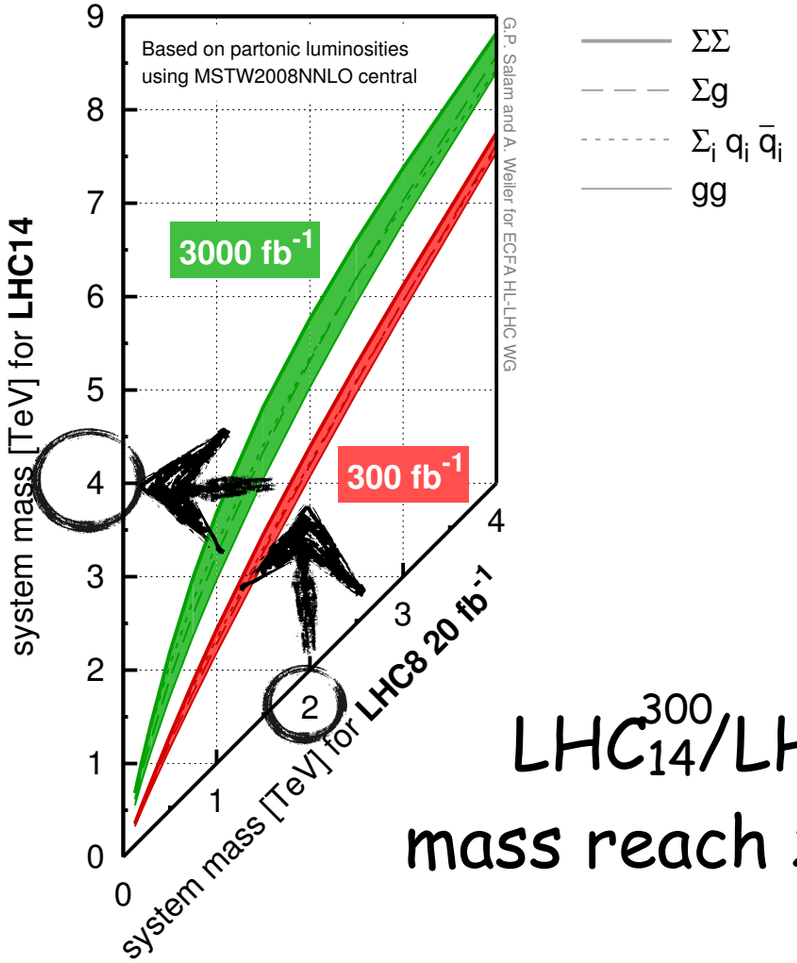
- $\sqrt{s}$ limited by synchroton radiation
- higher luminosity
- several interaction points
- precise E-beam measurement
- but only pdf access to $\sqrt{\hat{s}}$

Linear

- larger $\sqrt{s}$, energy scanning
- easier to upgrade in energy
- greener: less power consumption
- easier polarized beams
- large beamsthralung

The world according to LHC





Salam & Weiler "cern.ch/collider-reach" '14

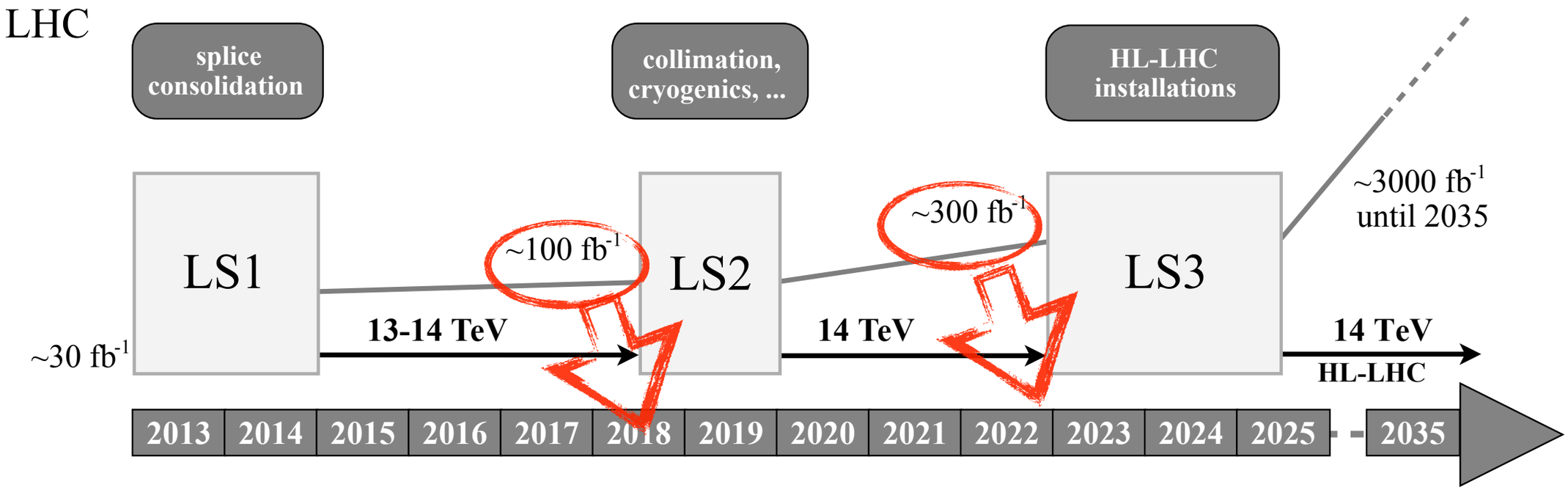
$LHC_{14}^{300}/LHC_8^{20}$
mass reach x O(2)

Direct exploration
of an
uncharted territory

A significative energy step
(maybe the last one before a long time)

What can be discovered @ 100/fb-14TeV
knowing what is excluded @ 20/fb-8TeV?

The world according to LHC



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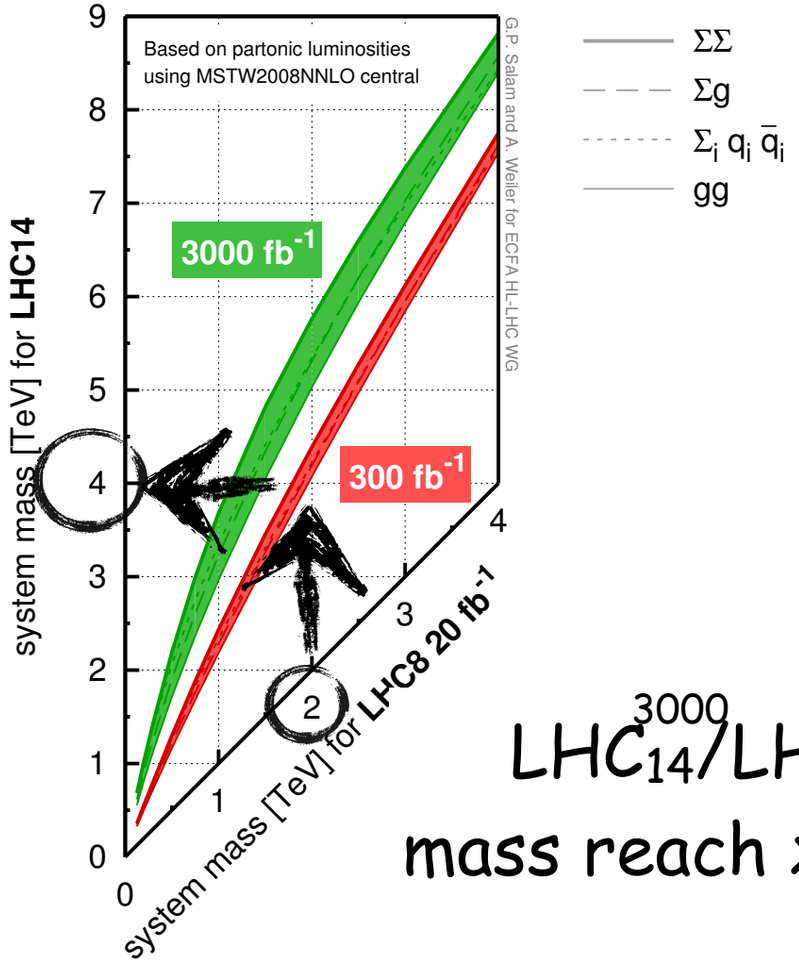


$LHC_{14}^{3000}/LHC_8^{20}$

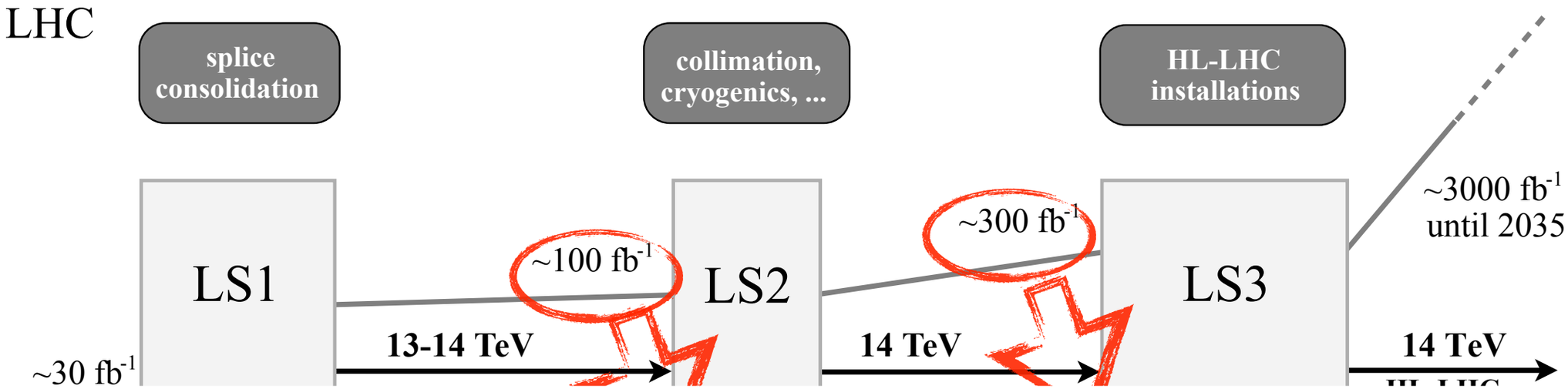
mass reach x O(3)

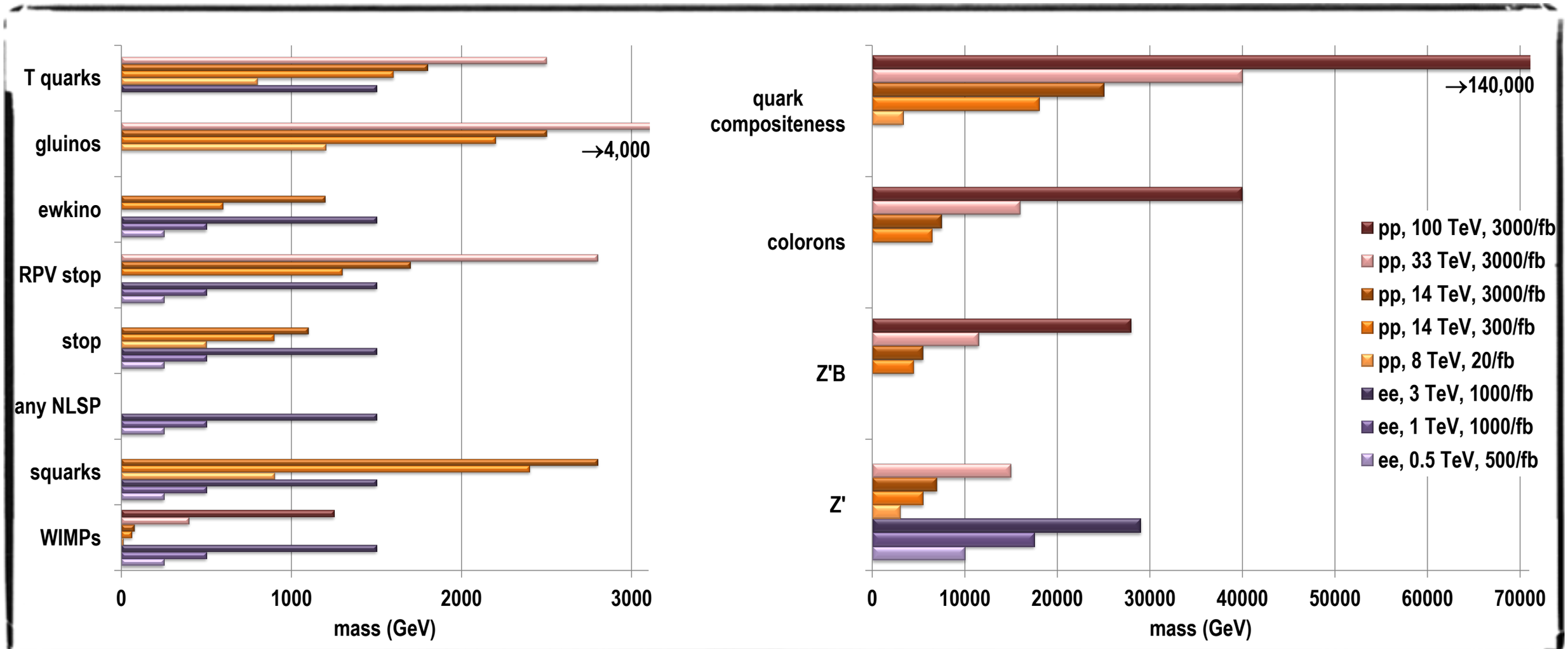
Direct exploration of an uncharted territory

A significative energy step
(maybe the last one before a long time)

What can be discovered @ 3/ab-14TeV
knowing what is excluded @ 100/fb-14TeV?

The world according to LHC





The future collider landscape

1) to which extent the various concepts are competitive, complementary, realistic or redundant, in terms of both physics and technology?

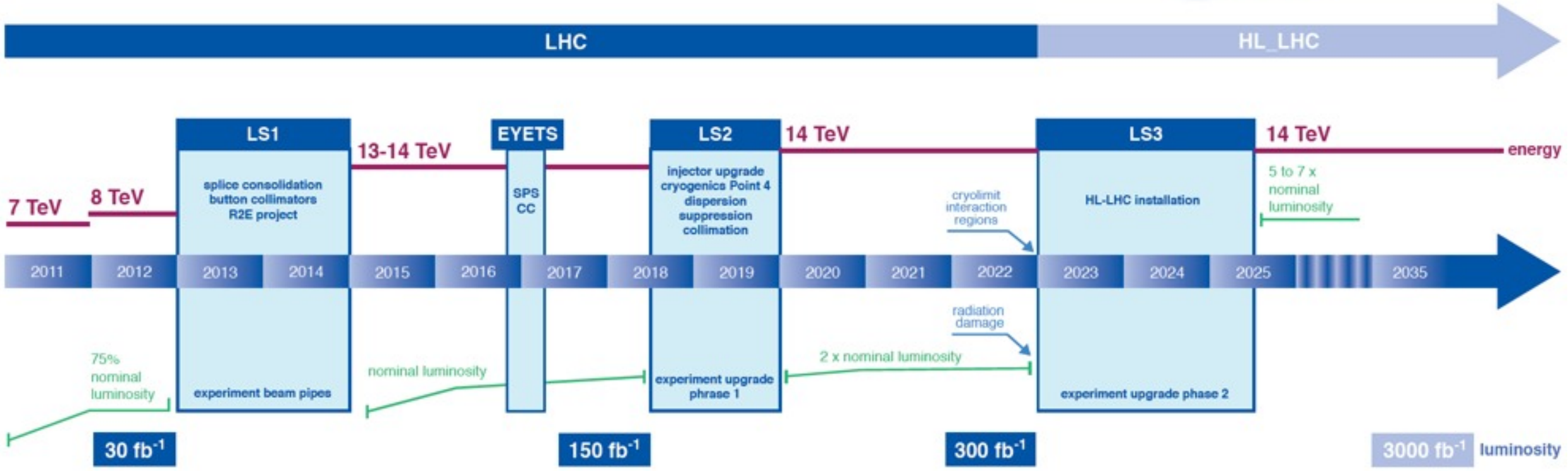
2) should the community continue with its current R&D efforts or consider adopting other programmes?

3) what should be the priorities in view of what we know today and the physics cases?

HL-LHC (2023-2035)

14 TeV - 3/ab

LHC / HL-LHC Plan



	Higgs bosons at √s=14TeV
HL-LHC, 3000fb^{-1}	170M
VBF (all decays)	13M
ttH (all decays)	1.8M
H->Zγ	230k
H->μμ	37k
HH (all)	121k

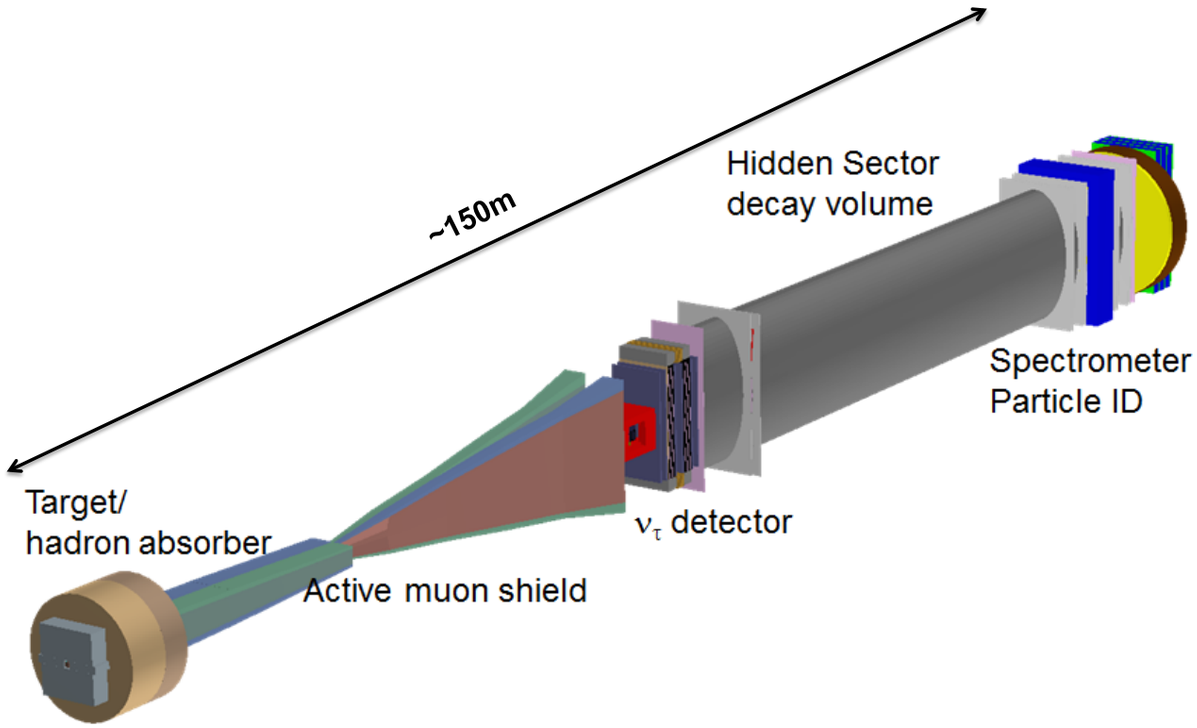
		κ_γ	κ_W	κ_Z	κ_g	κ_b	κ_t	κ_τ	$\kappa_{Z\gamma}$	κ_μ
300fb^{-1}	ATLAS	[9,9]	[9,9]	[8,8]	[11,14]	[22,23]	[20,22]	[13,14]	[24,24]	[21,21]
300fb^{-1}	CMS	[5,7]	[4,6]	[4,6]	[6,8]	[10,13]	[14,15]	[6,8]	[41,41]	[23,23]
3000fb^{-1}	ATLAS	[4,5]	[4,5]	[4,4]	[5,9]	[10,12]	[8,11]	[9,10]	[14,14]	[7,8]
3000fb^{-1}	CMS	[2,5]	[2,5]	[2,4]	[3,5]	[4,7]	[7,10]	[2,5]	[10,12]	[8,8]

$\times 10^{-2}$

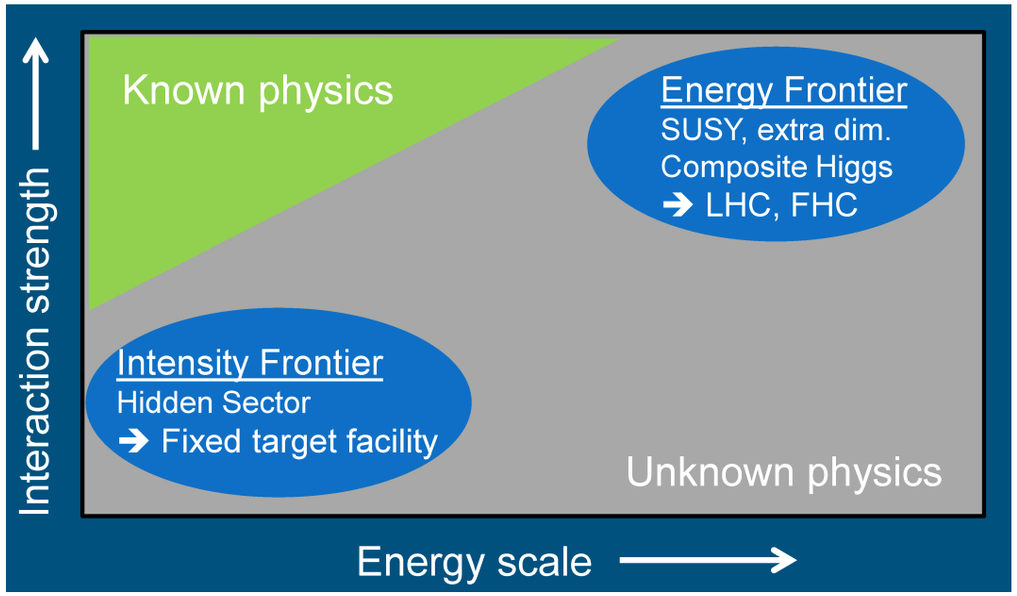
SHiP (TBA: 2018-2030)

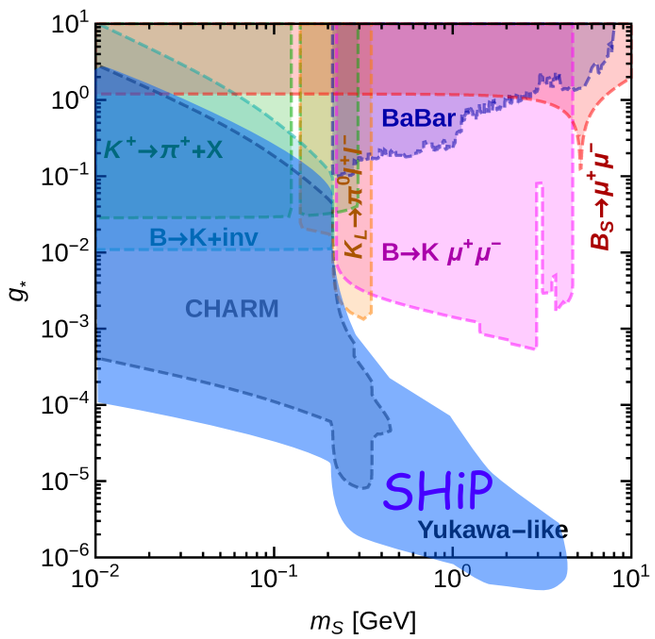
beam dump experiment: 400 GeV SPS protons on fixed target

$\sqrt{s}$~28 GeV 10^{20} protons over 10 years, i.e. $\mathcal{L} = 10^{39}\,\mathrm{cm}^{-2}\,\mathrm{s}^{-1}$



intensity frontier





Higgs portal

$$(\alpha_1 S + \alpha S^2) H^\dagger H + L_{SM} + L_{hidden}$$

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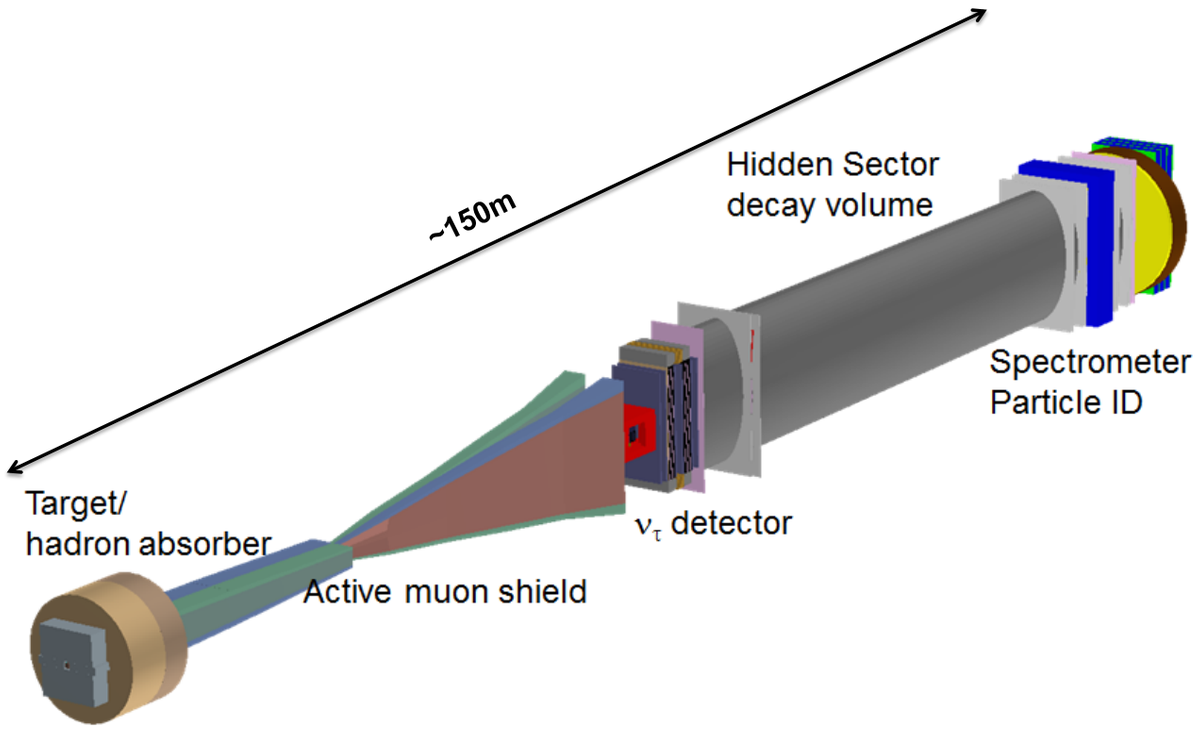

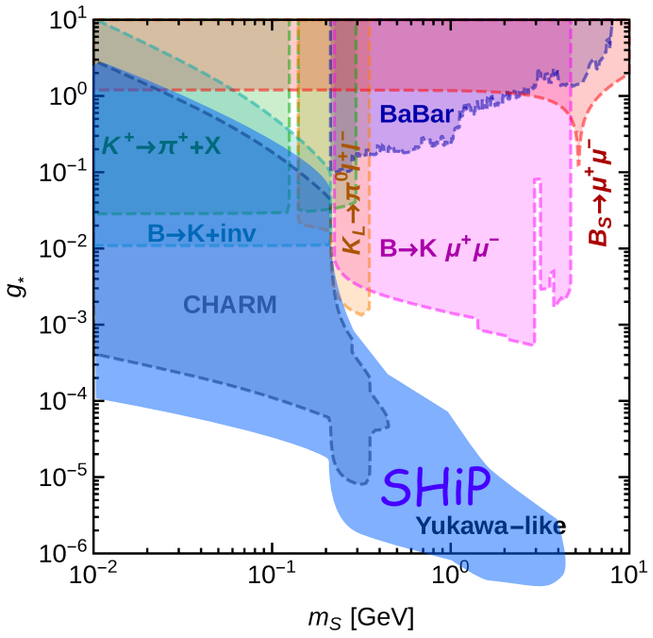

Higgs portal

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Physics case

CERN-SPSC-2015-017 (SPSC-P-350-ADD-1)

A facility to Search for Hidden Particles at the CERN SPS: the SHiP physics case

Sergey Alekhin,[1,2] Wolfgang Altmannshofer,[3] Takehiko Asaka,[4] Brian Batell,[5] Fedor Bezrukov,[6,7] Kyrylo Bondarenko,[8] Alexey Boyarsky*,[8] Nathaniel Craig,[9] Ki-Young Choi,[10] Cristóbal Corral,[11] David Curtin,[12] Sacha Davidson,[13,14] André de Gouvêa,[15] Stefano Dell'Oro,[16] Patrick deNiverville,[17] P. S. Bhupal Dev,[18] Herbi Dreiner,[19] Marco Drewes,[20] Shintaro Eijima,[21] Rouven Essig,[22] Anthony Fradette,[17] Björn Garbrecht,[20] Belen Gavela,[23] Gian F. Giudice,[5] Dmitry Gorbunov,[24,25] Stefania Gori,[3] Christophe Grojean§,[26,27] Mark D. Goodsell,[28,29] Alberto Guffanti,[30] Thomas Hambye,[31] Steen H. Hansen,[32] Juan Carlos Helo,[11] Pilar Hernandez,[33] Alejandro Ibarra,[20] Artem Ivashko,[8,34] Eder Izaguirre,[3] Joerg Jaeckel§,[35] Yu Seon Jeong,[36] Felix Kahlhoefer,[27] Yonatan Kahn,[37] Andrey Katz,[5,38,39] Choong Sun Kim,[36] Sergey Kovalenko,[11] Gordan Krnjaic,[3] Valery E. Lyubovitskij,[40,41,42] Simone Marcocci,[16] Matthew Mccullough,[5] David McKeen,[43] Guenakh Mitselmakher ,[44] Sven-Olaf Moch,[45] Rabindra N. Mohapatra,[46] David E. Morrissey,[47] Maksym Ovchynnikov,[34] Emmanuel Paschos,[48] Apostolos Pilaftsis,[18] Maxim Pospelov§,[3,17] Mary Hall Reno,[49] Andreas Ringwald,[27] Adam Ritz,[17] Leszek Roszkowski,[50] Valery Rubakov,[24] Oleg Ruchayskiy*,[21] Jessie Shelton,[51] Ingo Schienbein,[52] Daniel Schmeier,[19] Kai Schmidt-Hoberg,[27] Pedro Schwaller,[5] Goran Senjanovic,[53,54] Osamu Seto,[55] Mikhail Shaposhnikov*,§,[21] Brian Shuve,[3] Robert Shrock,[56] Lesya Shchutska§,[44] Michael Spannowsky,[57] Andy Spray,[58] Florian Staub,[5] Daniel Stolarski,[5] Matt Strassler,[39] Vladimir Tello,[53] Francesco Tramontano§,[59,60] Anurag Tripathi,[59] Sean Tulin,[61] Francesco Vissani,[16,62] Martin W. Winkler,[63] Kathryn M. Zurek[64,65]

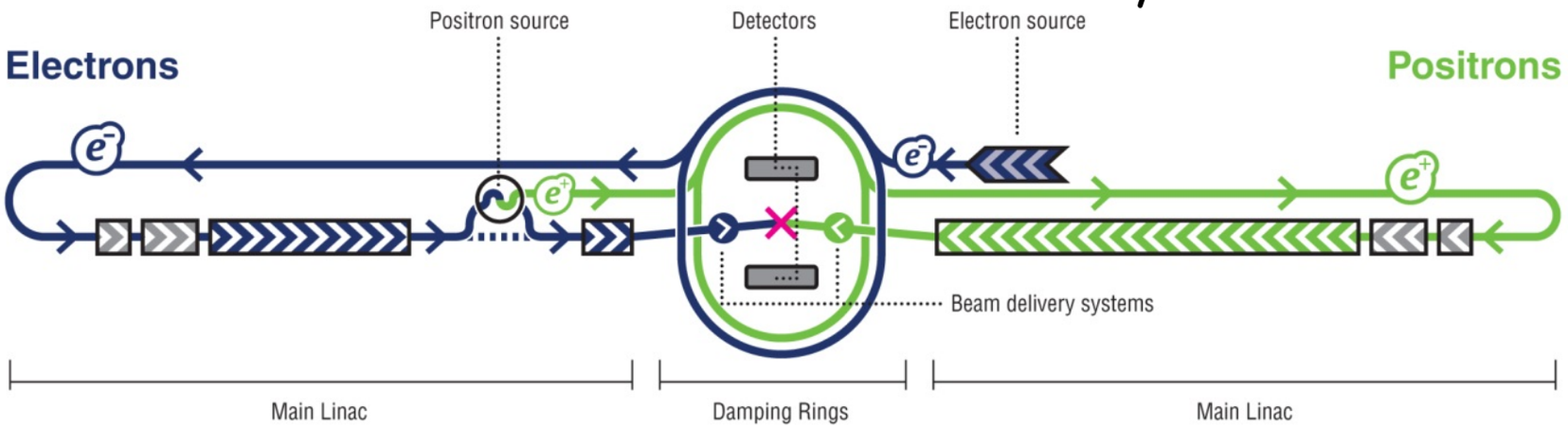
arXiv:1504.04855v1 [hep-ph] 19 Apr 2015

Abstract: This paper describes the physics case for a new fixed target facility at CERN SPS. The SHiP (*Search for Hidden Particles*) experiment is intended to hunt for new physics in the largely unexplored domain of very weakly interacting particles with masses below the Fermi scale, inaccessible to the LHC experiments, and to study tau neutrino physics. The same proton beam setup can be used later to look for decays of tau-leptons with lepton flavour number non-conservation, $\tau \to 3\mu$ and to search for weakly-interacting sub-GeV dark matter candidates. We discuss the evidence for physics beyond the Standard Model and describe interactions between new particles and four different *portals* — scalars, vectors, fermions or axion-like particles. We discuss motivations for different models, manifesting themselves via these interactions, and how they can be probed with the SHiP experiment and present several case studies. The prospects to search for relatively light SUSY and composite particles at SHiP are also discussed. We demonstrate that the SHiP experiment has a unique potential to discover new physics and can directly probe a number of solutions of beyond the Standard Model puzzles, such as neutrino masses, baryon asymmetry of the Universe, dark matter, and inflation.

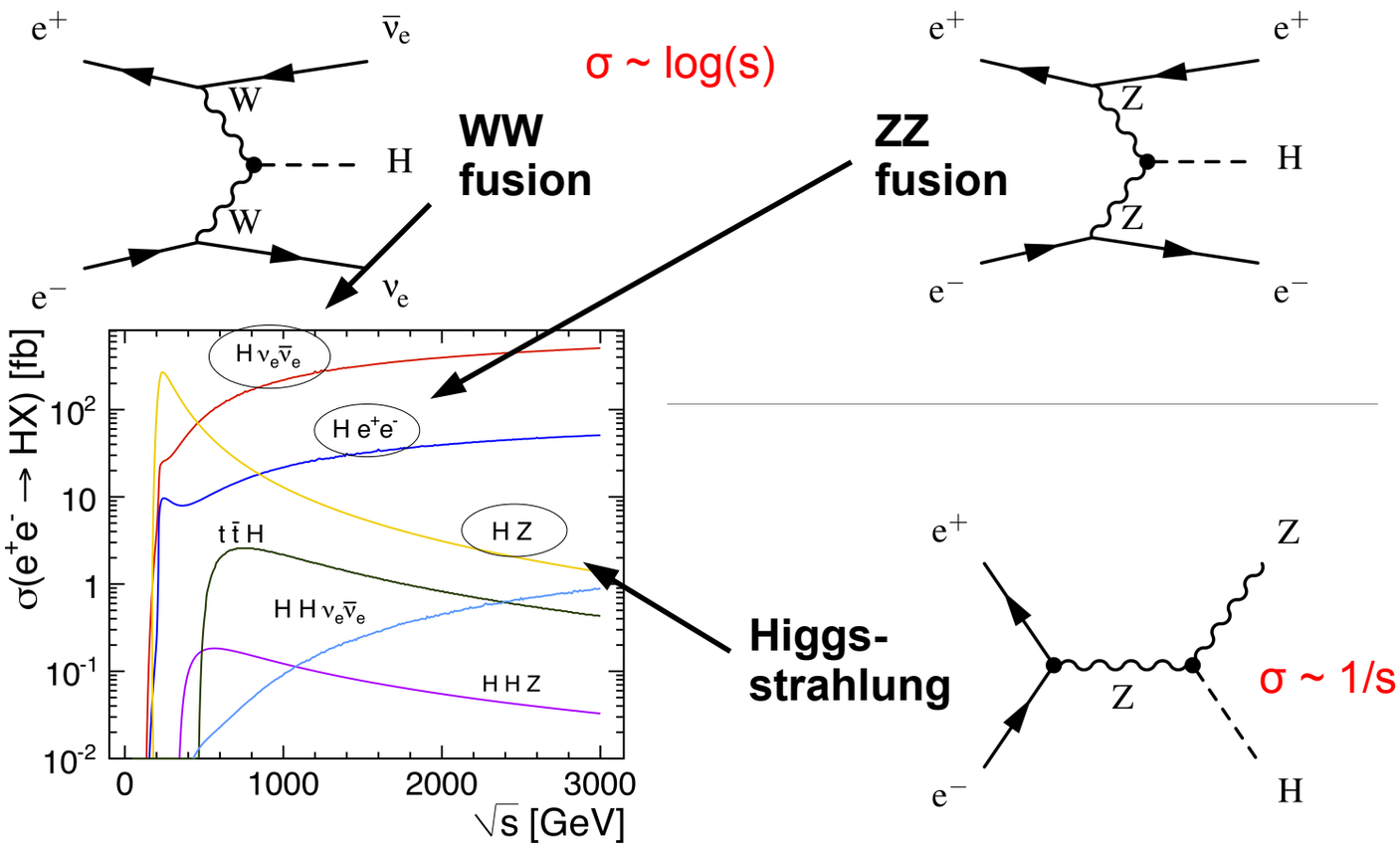
*Editor of the paper
§Convener of the Chapter

ILC (TBA: 2025-2045*)

*ready for construction once approved



(350)/500/1000 GeV - 5/ab



O(10⁶) Higgs produced and reconstructed

ILC (TBA: 2025-2045*)

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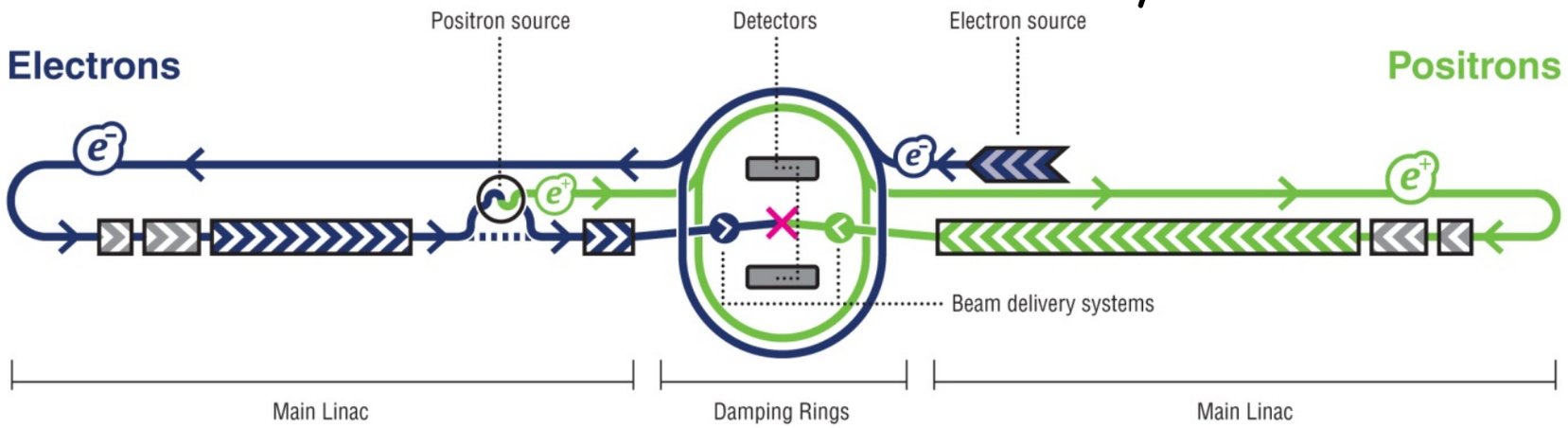

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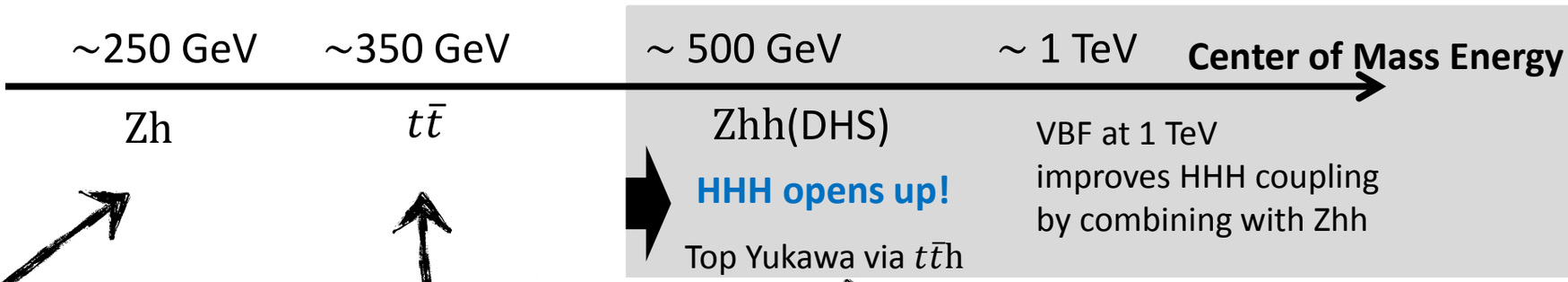


m_h
Γ_h
higgs couplings

m_t
top Yuk.

ZZhh coupling
h^3 coupling
direct top Yuk

h^3 coupling

O(10^6) Higgs produced and reconstructed

ILC (TBA: 2025-2045*)

Precis of the Physics Case for the ILC

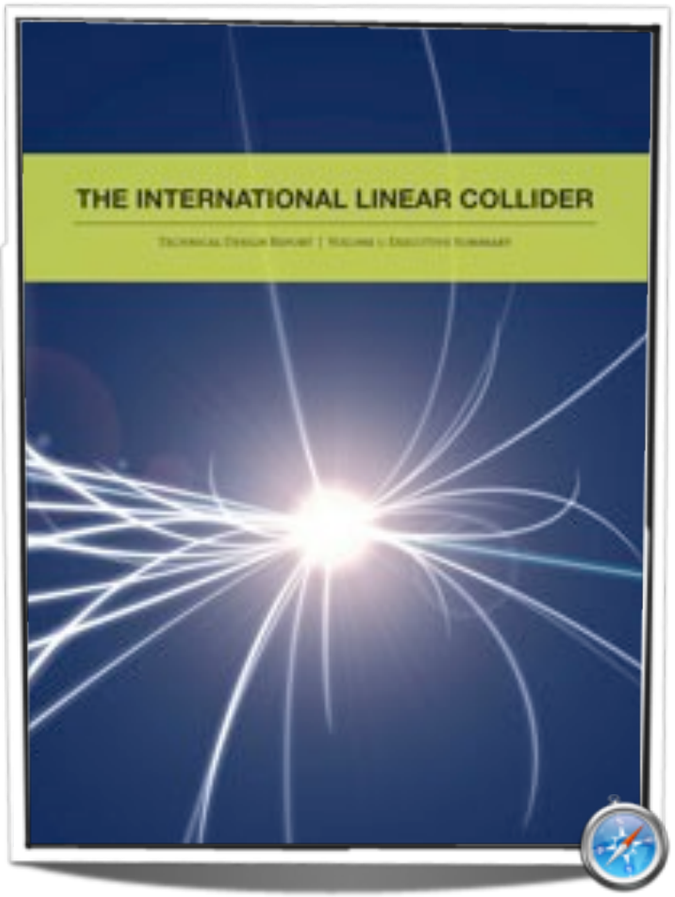
LCC Physics Working Group†

October 2014

Scientific Motivation for the ILC

LCC Physics Working Group†

March 2015



The International Linear Collider

Jim Brau†, Paul Grannis‡, Mike Harrison#, Michael Peskin*, Marc Ross*, Harry Weerts§
for the ILC Collaboration
April 9, 2013

submitted to the Community Summer Study (Snowmass on the Mississippi), July 2013

The Physics Case for an e^+e^- Linear Collider

James E. Brau[a], Rohini M. Godbole[b], Francois R. Le Diberder[c], M.A. Thomson[d],
Harry Weerts[e], Georg Weiglein[f], James D. Wells[g], Hitoshi Yamamoto[h]

A Report Commissioned by the Linear Collider Community†

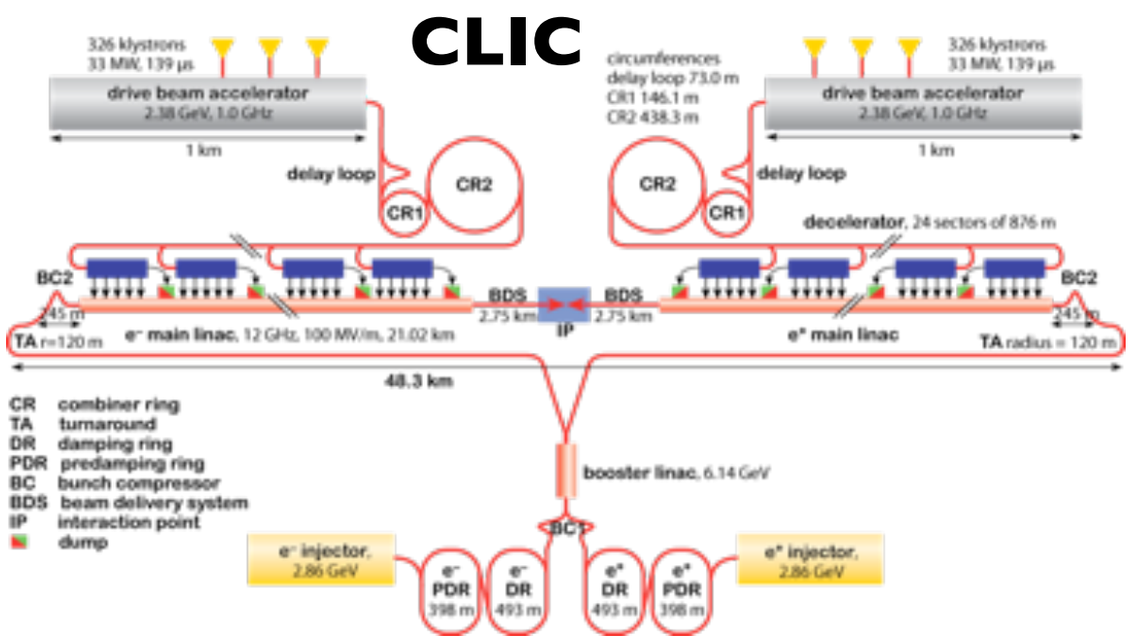
**Physics Case for the ILC Project:
Perspective from Beyond the Standard Model**

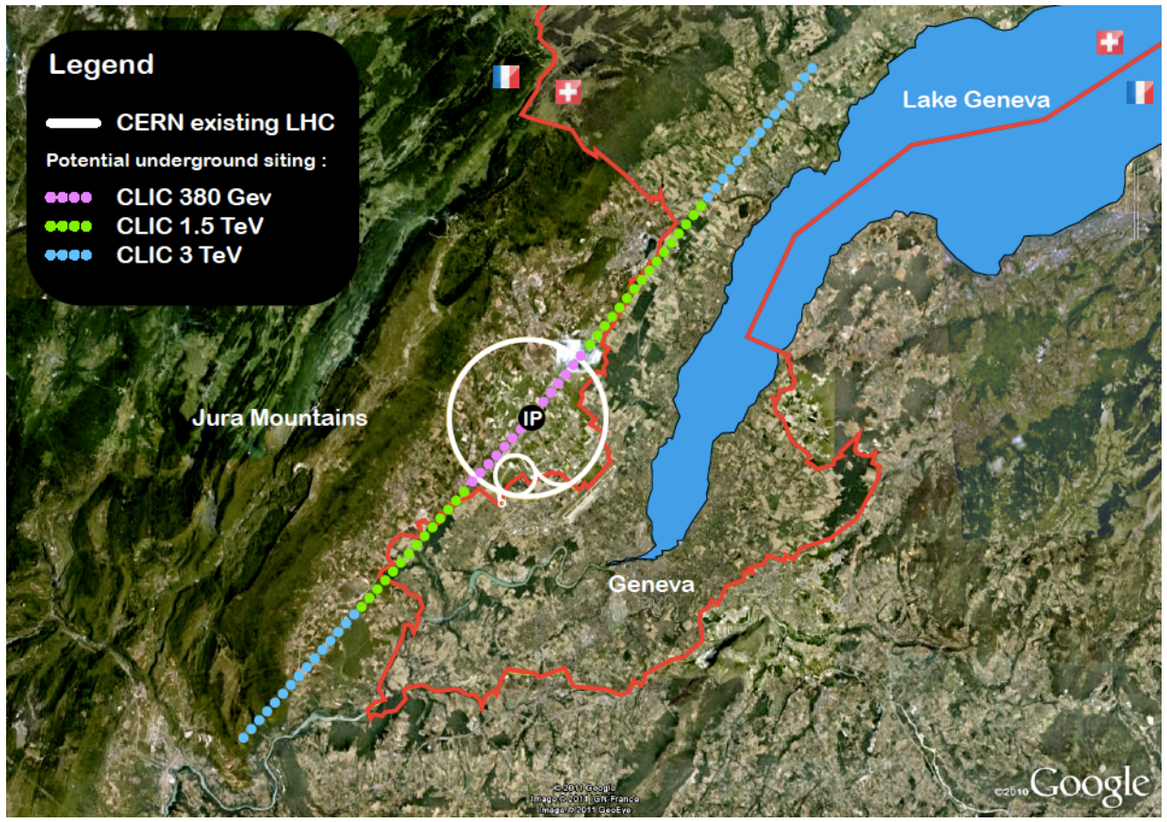
Howard Baer[1], *Mikael Berggren*[2], *Jenny List*[2], *Mihoko M. Nojiri*[3,4],
Maxim Perelstein[5], *Aaron Pierce*[6], *Werner Porod*[7], *Tomohiko Tanabe*[8]

Physics at the e^+e^- Linear Collider

CLIC (TBA: 2025-?)

(350)/1000/3000 GeV - 5/ab



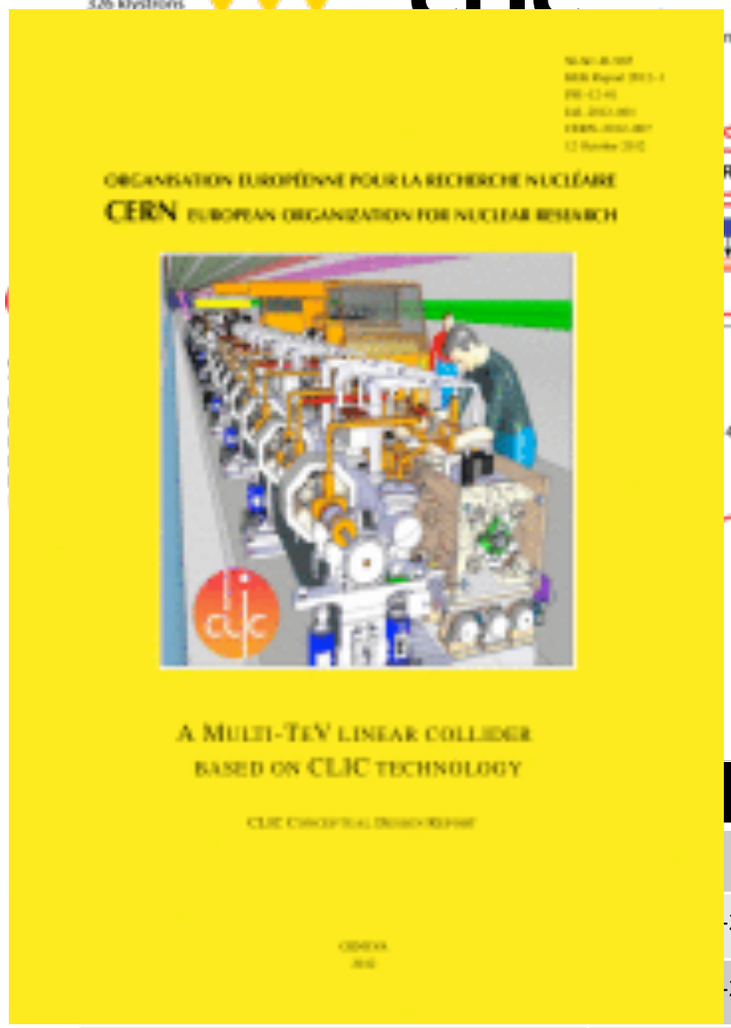


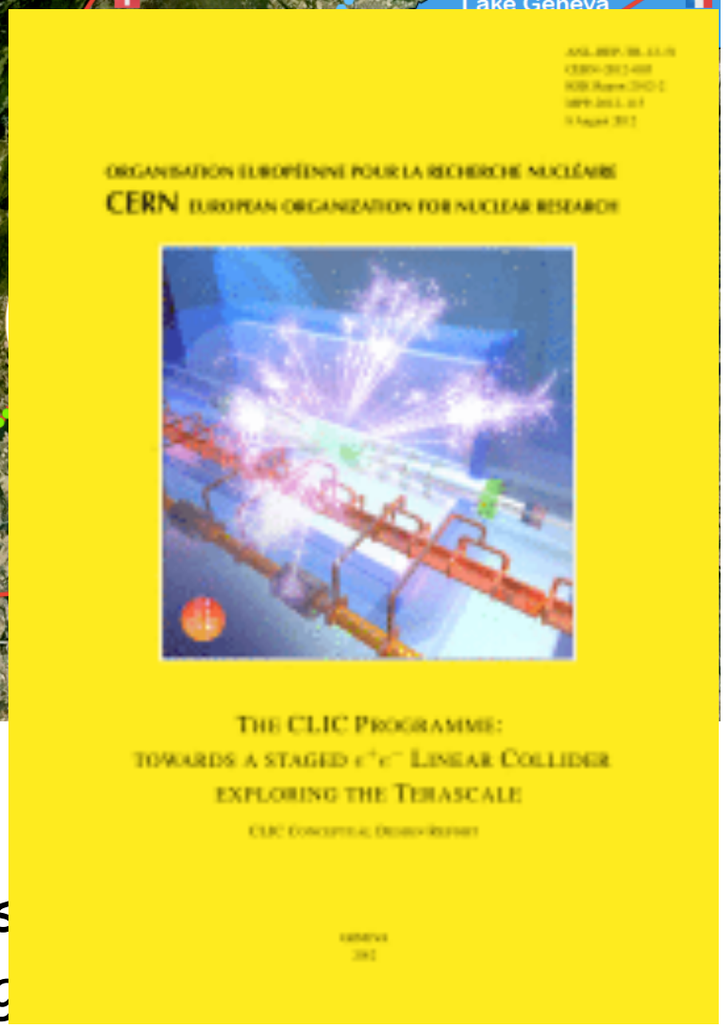
Parameter	Unit	380 GeV	3 TeV
Centre-of-mass energy	TeV	0.38	3
Total luminosity	$10^{34}cm^{-2}s^{-1}$	1.5	5.9
Luminosity above 99% of √s	$10^{34}cm^{-2}s^{-1}$	0.9	2.0
Repetition frequency	Hz	50	50
Number of bunches per train		352	312
Bunch separation	ns	0.5	0.5
Acceleration gradient	MV/m	72	100
Site length	km	11	50

- sub-percent Higgs coupling measurements
- few percents Higgs width
- top mass, top EW couplings
- direct BSM sensitivity in the multi-TeV region (direct and indirectly via precision)

CLIC (TBA: 2025-?)

(350)/1000/3000 GeV - 5/ab





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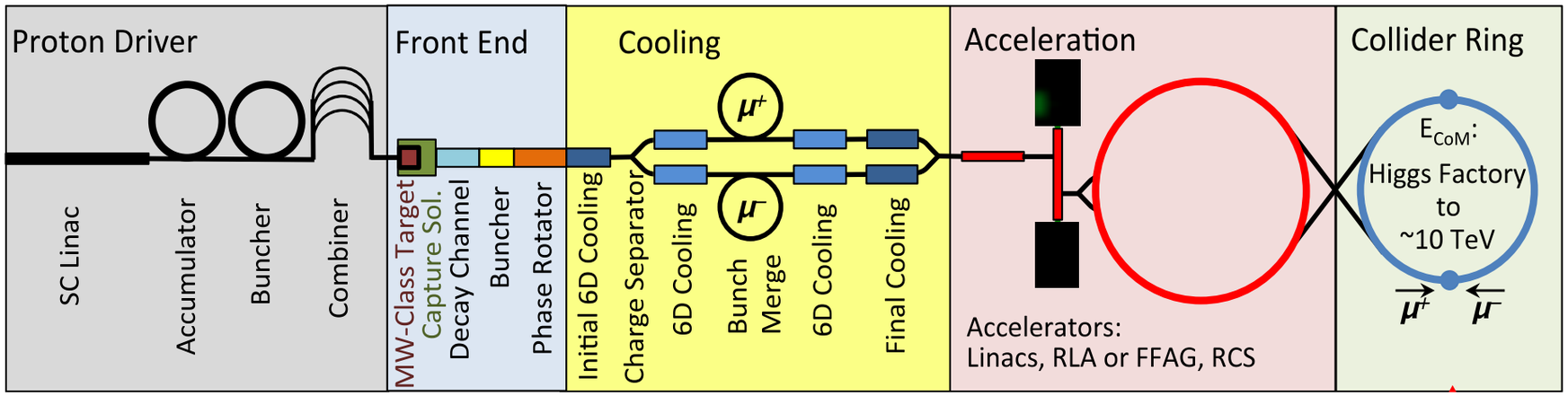
Higgs
Higg

- top mass, top EW couplings
- direct BSM sensitivity in the multi-TeV region (direct and indirectly via precision)

μ collider aka project X (TBD: ?-?)

126/1'000/10'000 GeV - O(1)/ab



M. Palmer, CERN '15

Physics Frontiers

- **Intense and cold muon beams ⇨ unique physics reach**
  - Tests of Lepton Flavor Violation
  - Anomalous Magnetic Moment (g-2)
  - Precision sources of neutrinos
  - Next generation lepton collider

$m_\mu = 105.7\,MeV/c^2$

$\tau_\mu = 2.2\,\mu s$

Colliders

- **Opportunities**
  - s-channel production of scalar objects
  - Strong coupling to particles like the Higgs → $\triangleleft\left(\frac{m_\mu^2}{m_e^2}\right) \cong 4\times10^4$
  - Reduced synchrotron radiation ⇨ multi-pass acceleration feasible
  - Beams can be produced with small energy spread
  - Beamstrahlung effects suppressed at IP
- ***BUT* accelerator complex/detector must be able to handle the impacts of μ decay**

Collider Synergies

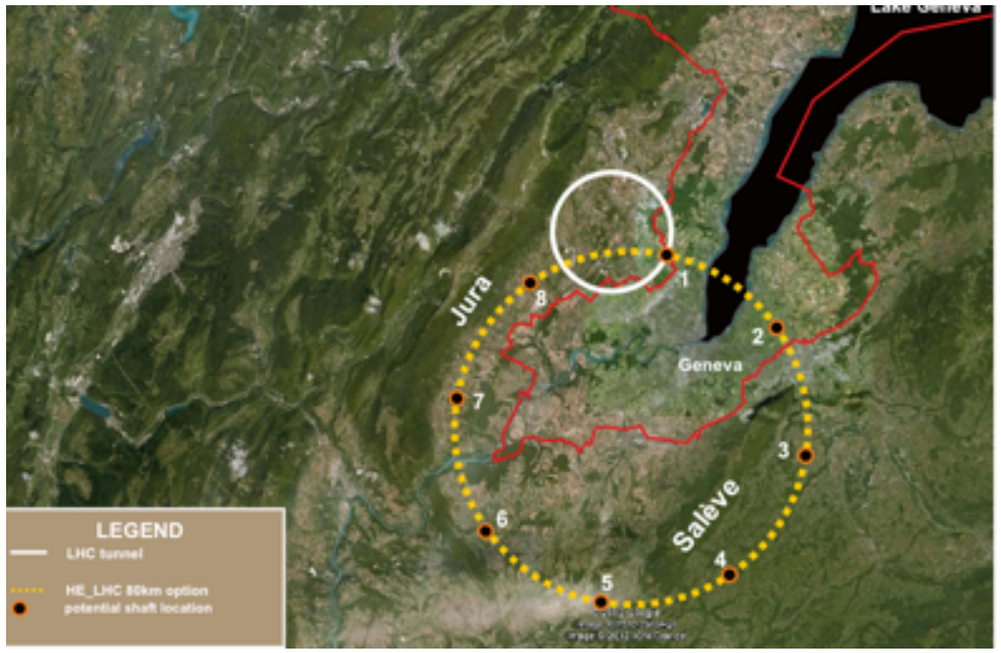
- High intensity beams required for a long-baseline Neutrino Factory are readily provided in conjunction with a Muon Collider Front End
- Such overlaps offer unique staging strategies to guarantee physics output while developing a muon accelerator complex capable of supporting collider operations

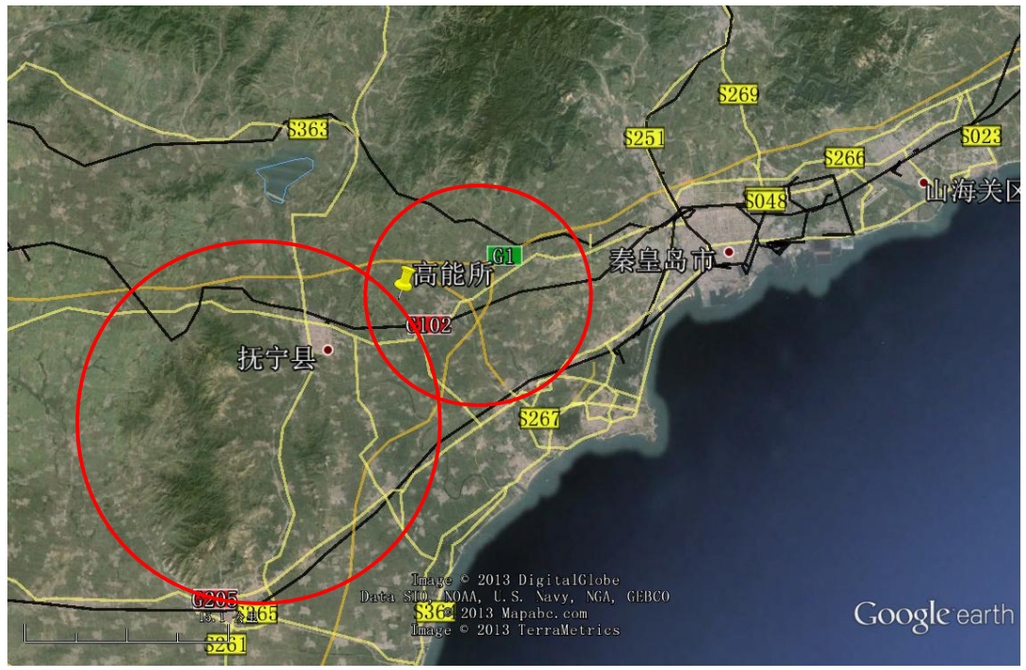
$\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu$

$\mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu$

FCC-ee/CepC (TBA: maybe soon?-?)

240/350/(500) - 10/ab

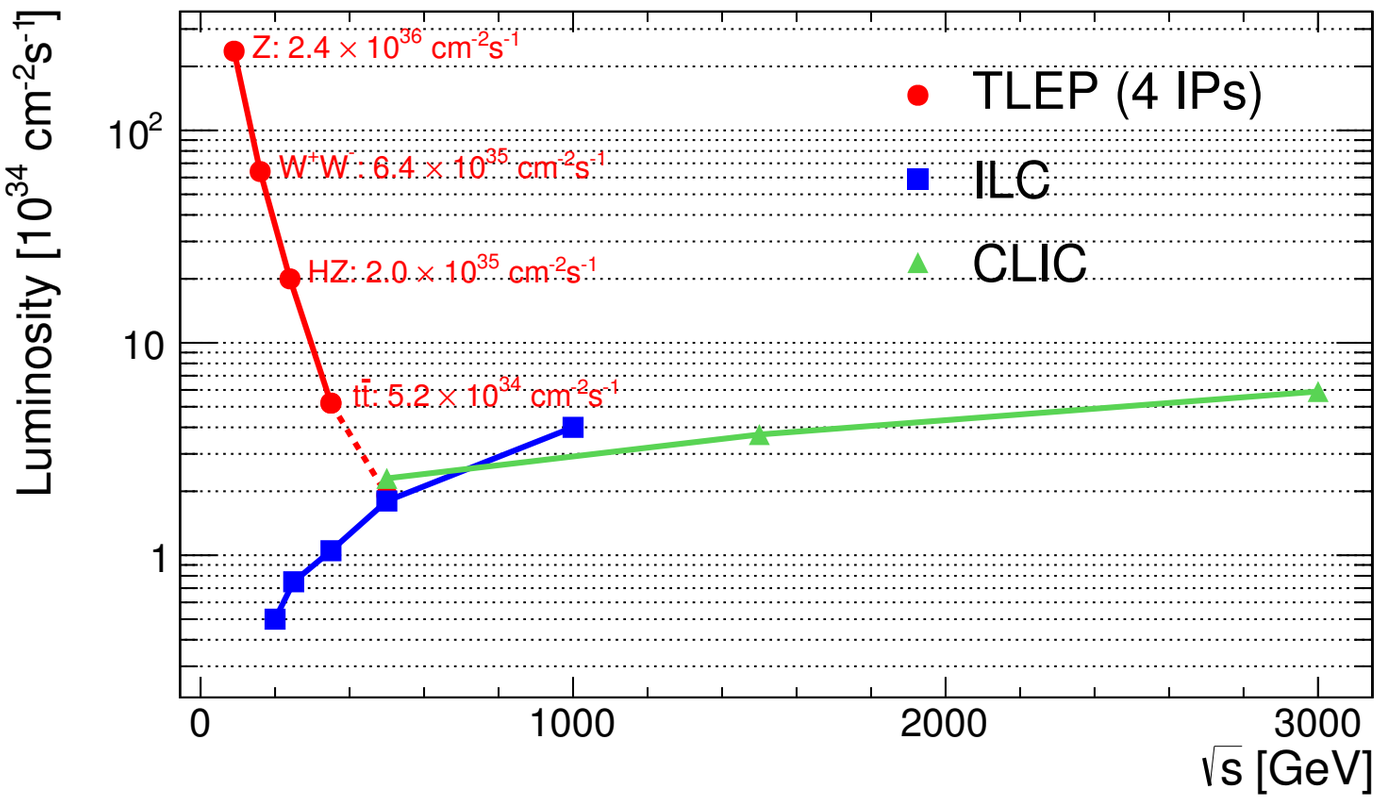




FCC-ee/CepC (TBA: maybe soon?-?)

240/350/(500) - 10/ab

parameter	FCC-ee			CEPC	LEP2
energy/beam [GeV]	**45**	**120**	**175**	**120**	105
bunches/beam	13000-60000	500-1400	51- 98	50	4
beam current [mA]	1450	30	6.6	16.6	3
luminosity/IP x 10^{34} $cm^{-2}s^{-1}$	**21 - 280**	**5 - 11**	**1.5 - 2.6**	**2.0**	0.0012
energy loss/turn [GeV]	0.03	1.67	7.55	3.1	3.34
synchrotron power [MW]	100			103	22
RF voltage [GV]	0.2-2.5	3.6-5.5	11	6.9	3.5



FCC-ee/CepC (TBA: maybe soon?-?)

240/350/(500) - 10/ab

parameter	FCC-ee			CEPC	LEP2
energy/beam [GeV]	**45**	**120**	**175**	**120**	105
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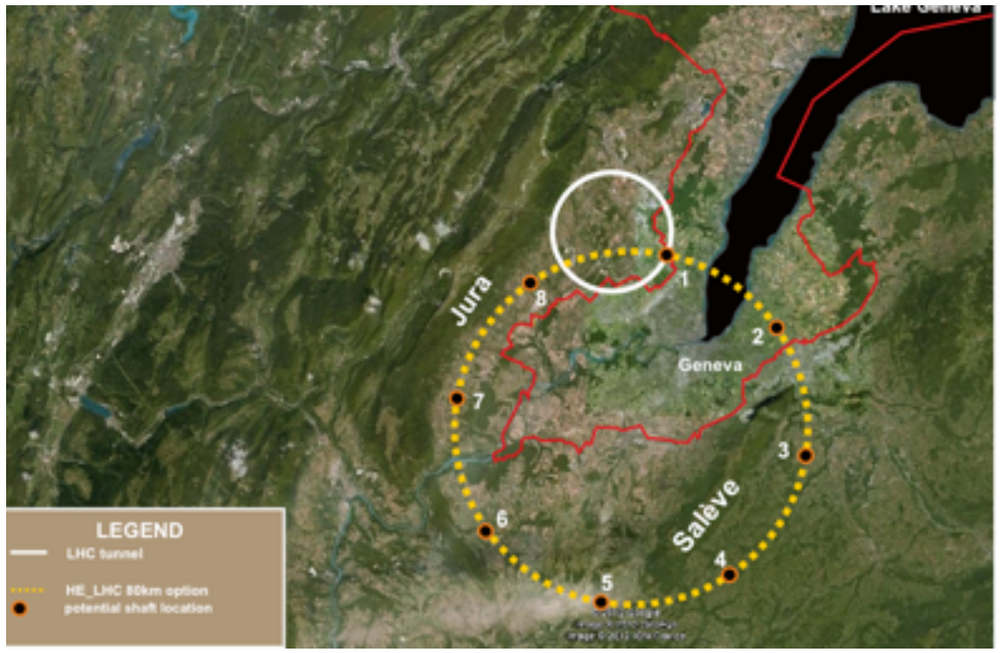
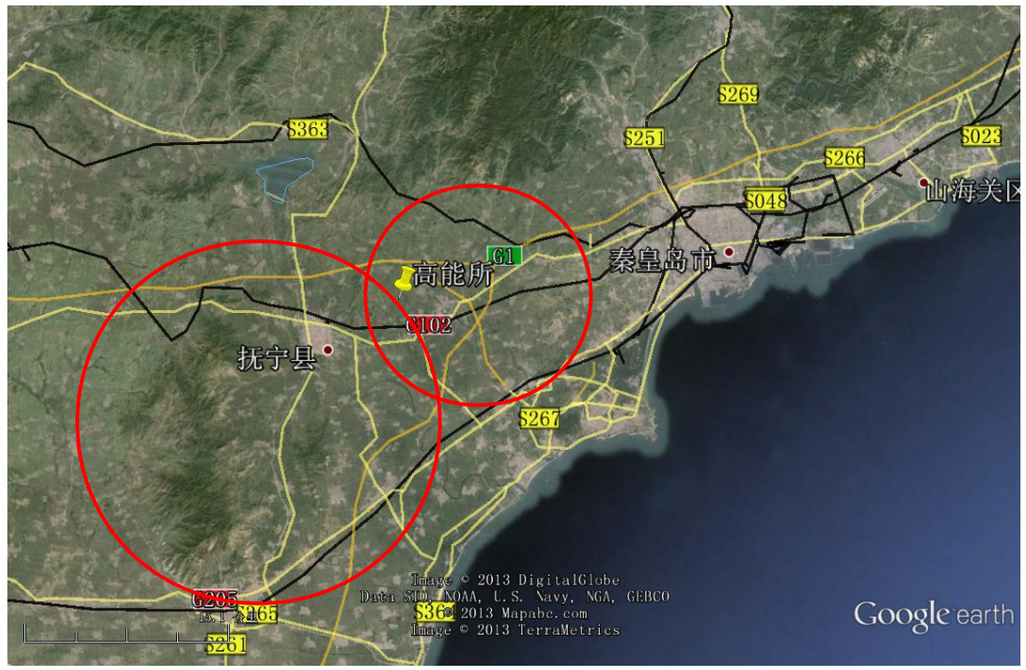
- 10^6 H
- 10^{12} Z (line-shape, mass & width, probe rare (FCNC) decays)
- 10^8 W (mass)
- 3×10^{10} tau/muon pairs
- 2×10^{11} b/c quarks $\Rightarrow$ >20'000 $B_s\rightarrow\tau^+\tau^-$
- TLEP@340/500: 10^6 top pairs (pole mass, probe FCNC decays, top Yukawa)

FCC-ee/CepC (TBA: maybe soon?-?)

240/350/(500) - 10/ab

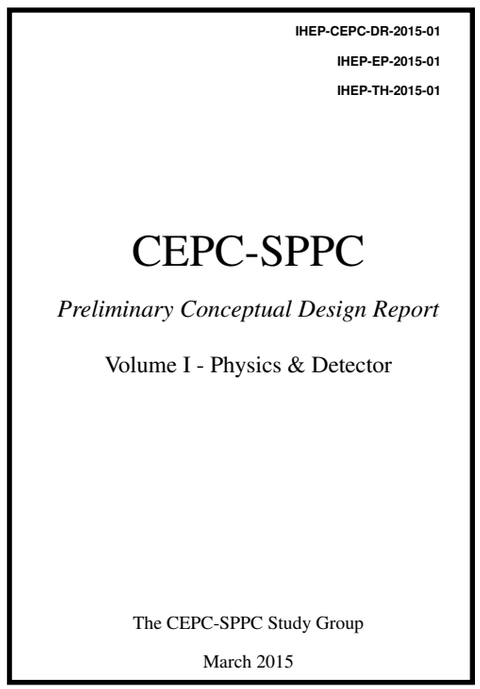




- physics case: arXiv:1308.6176
- CDR and cost review due in 2018

The FCC and CepC are essentially equivalent proposals with different emphasis; FCC – hadrons via e+e-, CepC – e+e- then hadrons

pre-CDR:


IHEP-CEPC-DR-2015-01
IHEP-EP-2015-01
IHEP-TH-2015-01

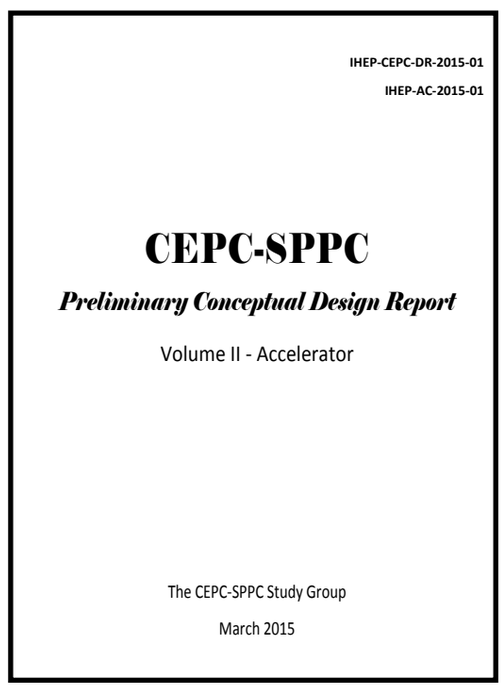
CEPC-SPPC

Preliminary Conceptual Design Report

Volume I - Physics & Detector

The CEPC-SPPC Study Group

March 2015


IHEP-CEPC-DR-2015-01
IHEP-AC-2015-01

CEPC-SPPC

Preliminary Conceptual Design Report

Volume II - Accelerator

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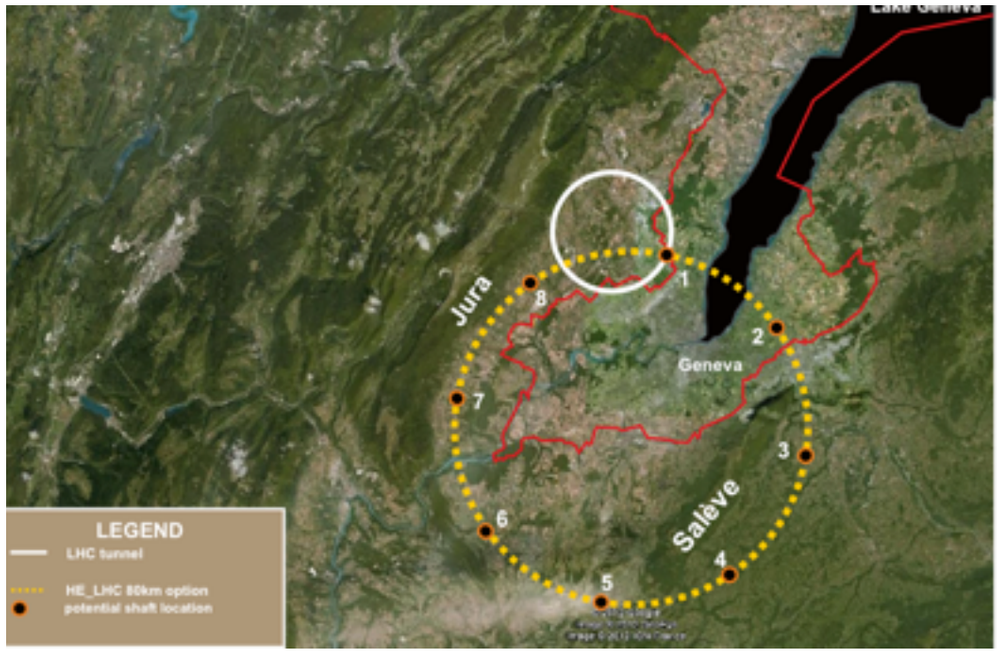
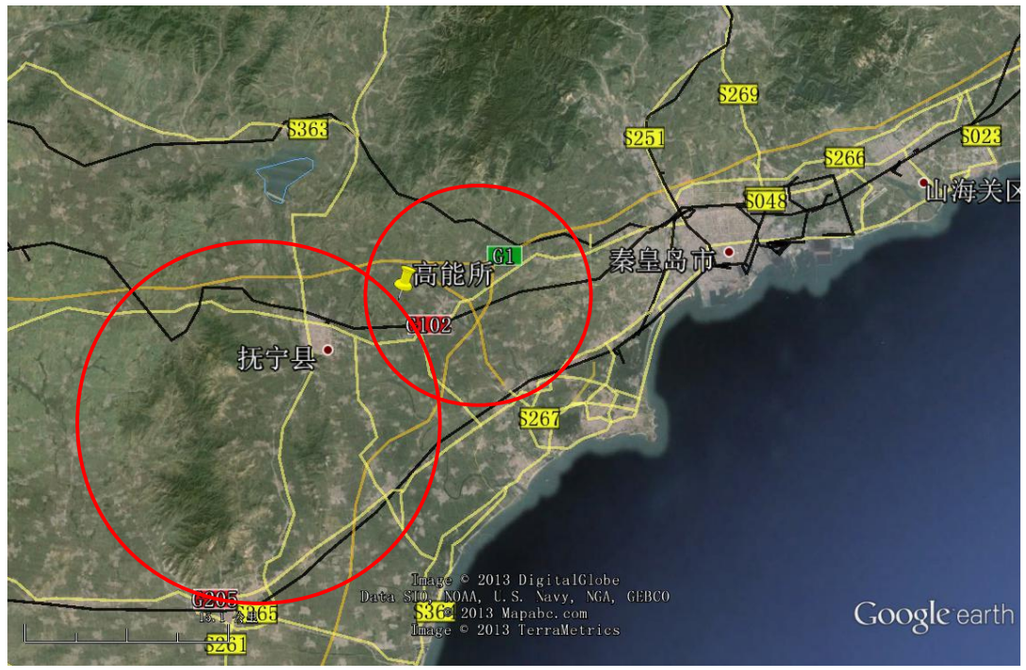
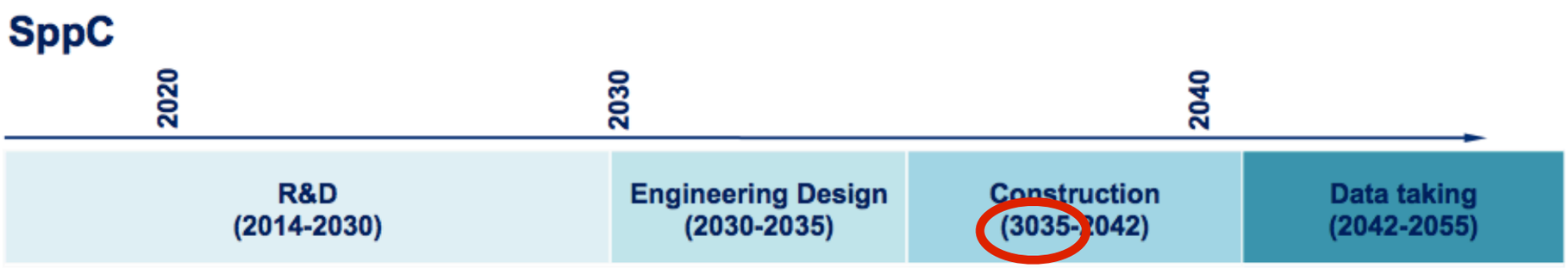
March 2015

FCC-hh/SppC (TBA: 2035-2060)

80/100 TeV - 30/ab

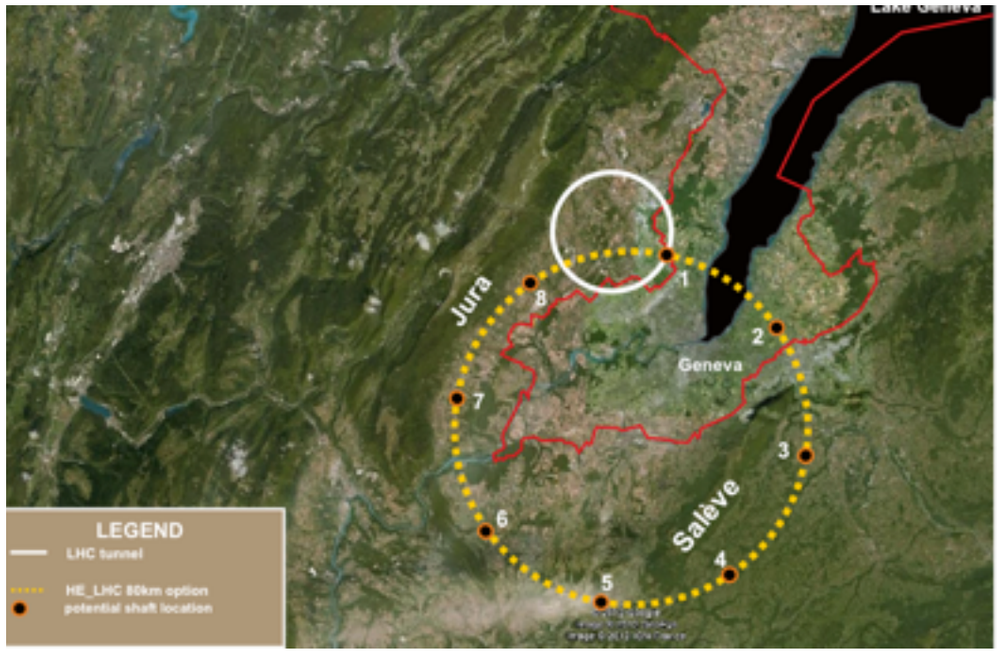


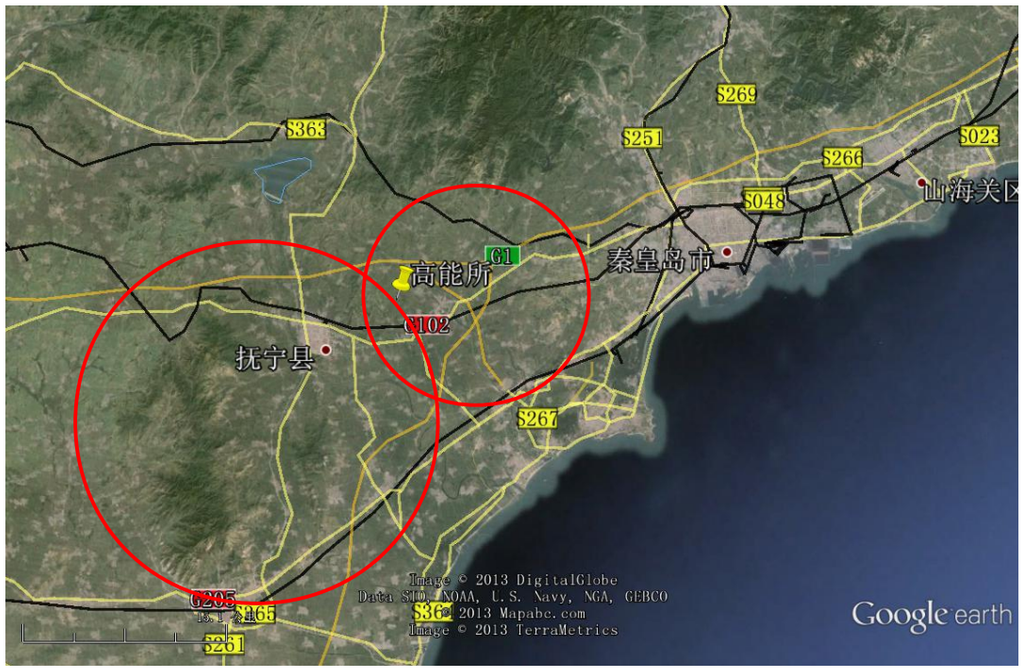




FCC-hh/SppC (TBA: 2035-2060)

80/100 TeV - 30/ab

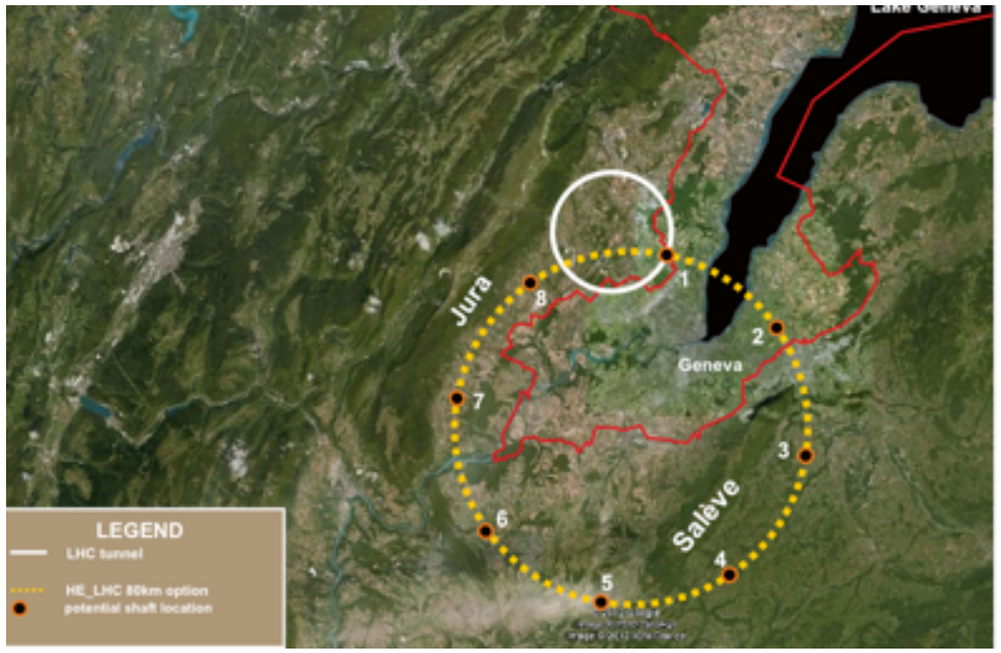


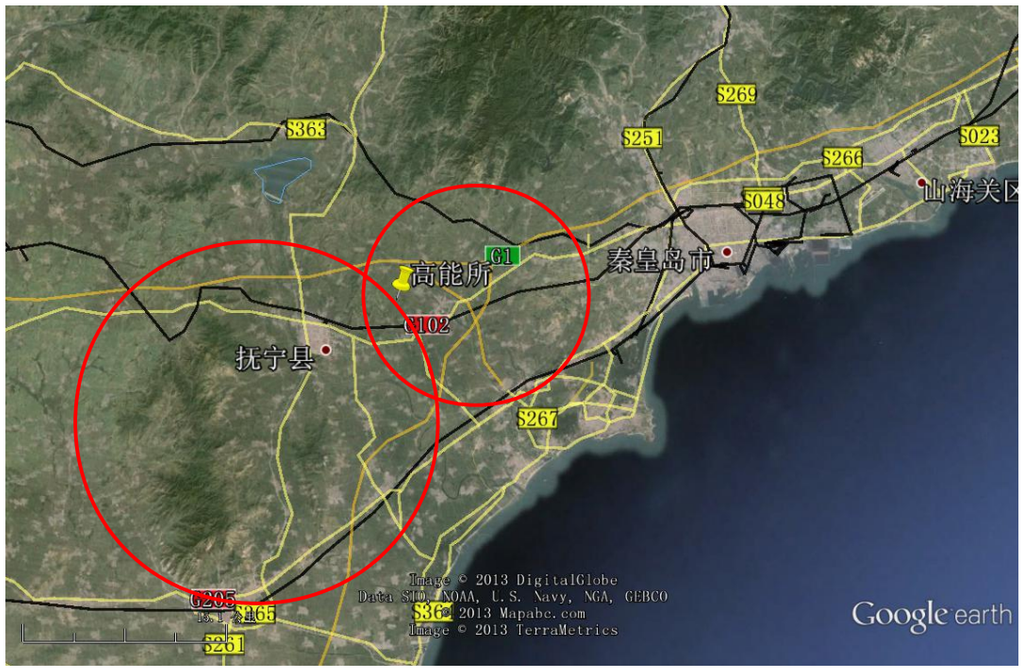


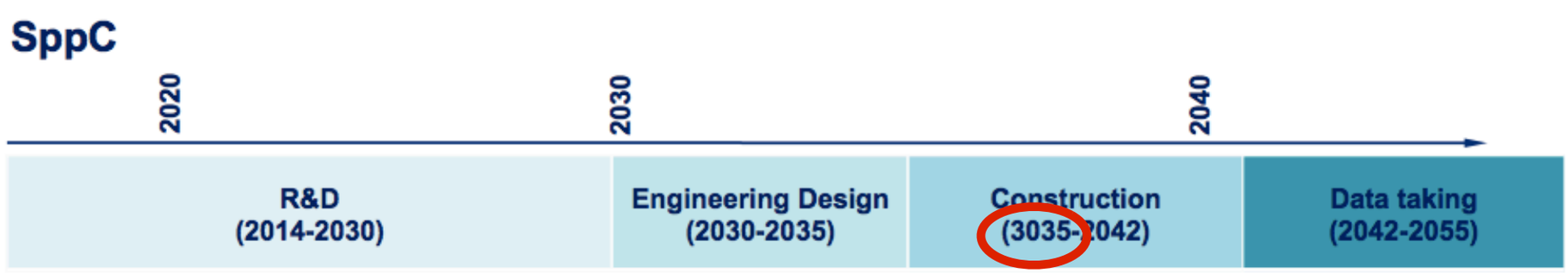
Parameter	FCC-hh		SPPC	LHC	HL LHC
collision energy cms [TeV]	100		71.2	14	
dipole field [T]	16		20	8.3	
# IP	2 main & 2		2	2 main & 2	
bunch intensity [10^{11}]	1	1 (0.2)	2	1.1	2.2
bunch spacing [ns]	25	25 (5)	25	25	25
luminosity/Ip [10^{34} $cm^{-2}s^{-1}$]	5	25	12	1	5
events/bx	170	850 (170)	400	27	135
stored energy/beam [GJ]	8.4		6.6	0.36	0.7
synchr. rad. [W/m/apert.]	30		58	0.2	0.35

FCC-hh/SppC (TBA: 2035-2060)

80/100 TeV - 30/ab







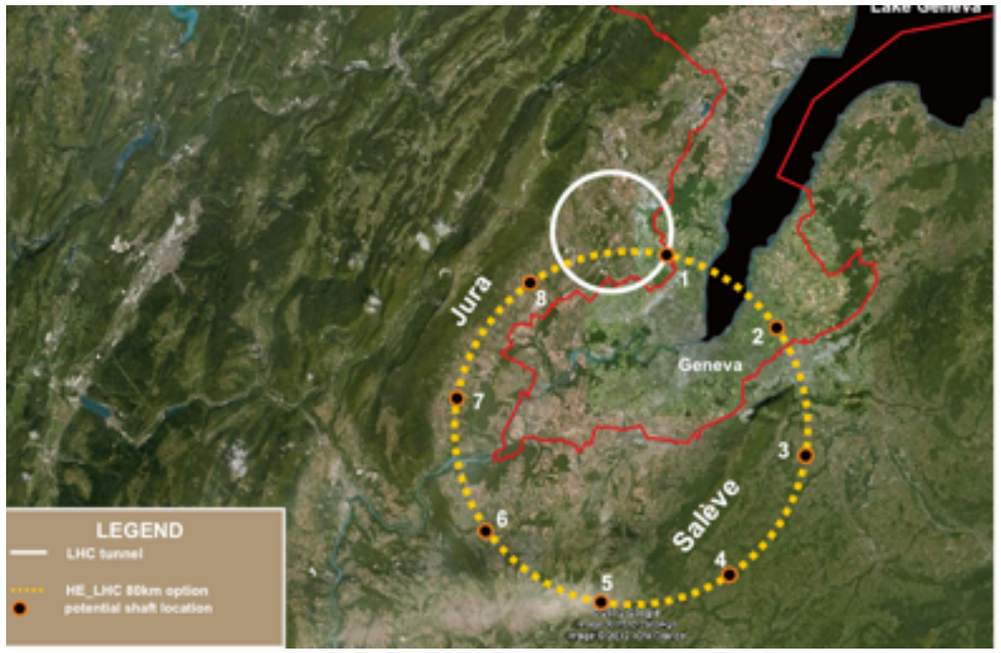
	σ	*N / 10ab⁻¹*
gg→H	740 pb	7.4 G
VBF	82 pb	0.8 G
WH	16 pb	160 M
ZH	11 pb	110 M
ttH	38 pb	380 M
gg→HH	1.4 pb	14 M

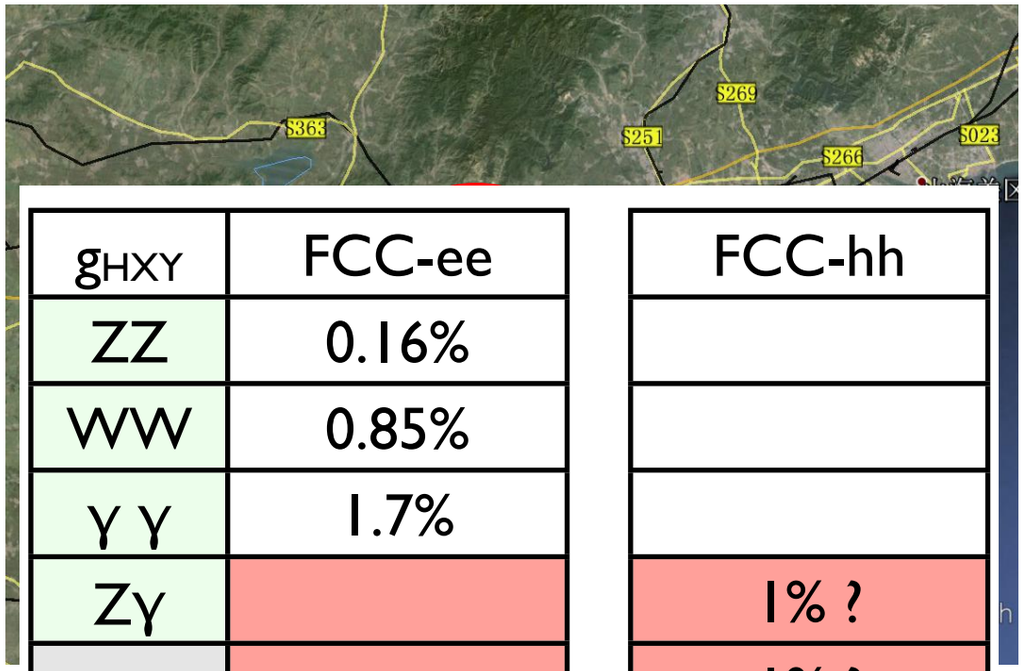
10^{10} Higgs bosons => 10^4 x today

10^{12} top quarks => 5 10^4 x today

FCC-hh/SppC (TBA: 2035-2060)

80/100 TeV - 30/ab





g_{HXY}	FCC-ee	FCC-hh
ZZ	0.16%	
WW	0.85%	
γ γ	1.7%	
Zγ		1% ?
tt		1% ?
bb	0.88%	
τ τ	0.94%	
cc	1.0%	
ss	H→Vγ, in progr.	
μμ	6.4%	2% ?
uu,dd	H→Vγ, in progr.	
ee	e^+e^-→H, in progr.	
HH		5% ?
BR_{exo}	0.48%	< 10^{-6} ?

SppC

2020 — 2030

R&D (2014-2030)	Engineering Design (2030-2035)

	σ	*N / 10ab⁻¹*
gg→H	740 pb	7.4 G
VBF	82 pb	0.8 G
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10^{10} Hig

10^{12} top

Higgs physics

1) Precision in Higgs coupling measurements

2) Higgs and flavor

3) Exploring extreme phase space regions

4) Multi-Higgs

5) Searching for extended Higgs sector

HEP with a Higgs boson

The successes have been breathtaking

- in O(3) years, the Higgs mass has been measured to 0.2% (vs 0.5% for the 20-year old top)
- some of its couplings, e.g. κ_γ, have been measured with LEP accuracy (10^{-3})

The meaning of the Higgs

Particle physics is not so much about particles but more about fundamental principles

- About 10^{-10}s after the Big Bang, the Universe filled with the Higgs substance because it saved energy by doing so:

"the vacuum is not empty"

(even when $\hbar \to 0$, not a Casimir effect)

- The masses are emergent quantities due to a non-trivial vacuum structure

HEP with a Higgs boson

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- in O(3) years, the Higgs mass has been measured to 0.2% (vs 0.5% for the 20-year old top)
- some of its couplings, e.g. κ_γ, have been measured with LEP accuracy (10^{-3})

Higgs agenda for the LHC-II, HL-LHC, ILC/CLIC, FCC, CepC, SppC, SHiP

multiple independent, synergetic and complementary approaches to achieve **precision** (couplings), **sensitivity** (rare and forbidden decays) and **perspective** (role of Higgs dynamics in broad issues like EWSB and vacuum stability, baryogenesis, inflation, naturalness, etc)

M.L. Mangano, Washington '15

- rare Higgs decays: $h\to\mu\mu$, $h\to\gamma Z$
- Higgs flavor violating couplings: $h\to\mu\tau$ and $t\to hc$
- Higgs CP violating couplings
- exclusive Higgs decays (e.g. $h\to J/\Psi+\gamma$) and measurement of couplings to light quarks
- exotic Higgs decay channels:

  $h\to \not{E}_T$, $h\to 4b$, $h\to 2b2\mu$, $h\to 4\tau, 2\tau 2\mu$, $h\to 4j$, $h\to 2\gamma 2j$, $h\to 4\gamma$, $h\to\gamma/2\gamma+\not{E}_T$,

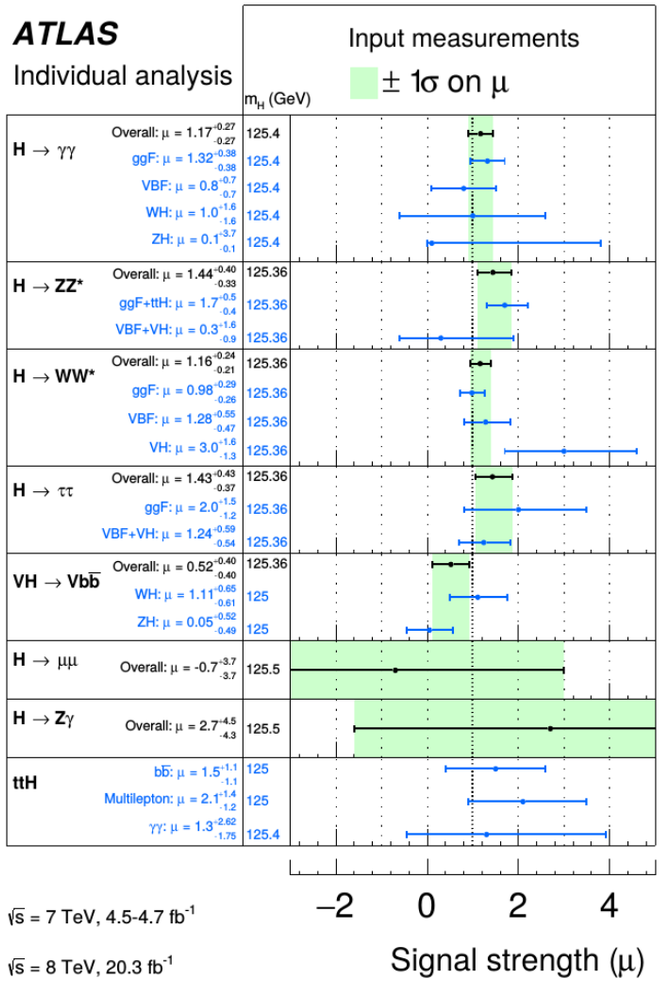
  $h\to$isolated leptons$+\not{E}_T$, $h\to 2l+\not{E}_T$, $h\to$one/two lepton-jet(s)+X, $h\to bb+\not{E}_T$, $h\to\tau\tau+\not{E}_T$ ...

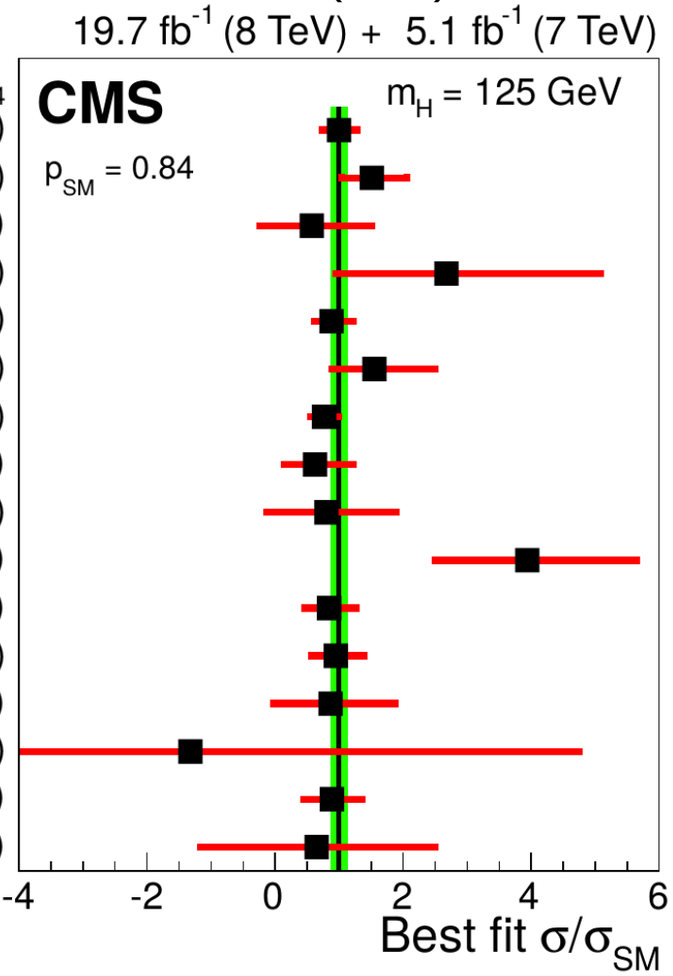
- searches for extended Higgs sectors (H, A, $H^{\pm}$,$H^{\pm\pm}$...)
- Higgs self-coupling(s)
- Higgs width
- Higgs/axion coupling?
- ...

Higgs@LHC run-1 in 1 slide

combination of results: ATLAS and CMS

Results based on integrated luminosities of
~5/fb @ 7 TeV (2011) and ~20/fb @ 8 TeV (2012) per experiment
for the "big five" H→ ZZ, γγ, WW, ττ and bb + some (rare) others





ATLAS Mass: PRL 114, 191803
ATLAS Couplings: arXiv:1507.04548
CMS Mass & Couplings
EPJ 75 (2015) 212

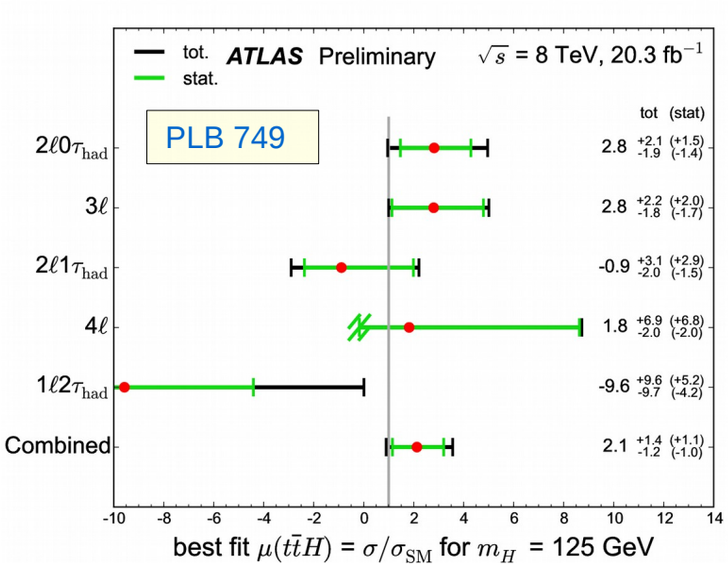
Various production (ggF, VBF, VH, ttH) modes observed
Decays modes to heaviest particles observed
Already meaningful bounds on couplings to lighter particles (H≠Z')

ttH: A channel rich of promises?

Individual Channels: ttH

directly probing the largest Higgs-boson coupling using associated production with $t\bar{t}$

$t\bar{t}H$, multi-leptons (WW, ZZ, ττ)



$\mu = 2.1^{+1.4}_{-1.2},\ \mu < 4.7(2.4)$

$t\bar{t}H$, H → $b\bar{b}$, neural network — EPJC75 349

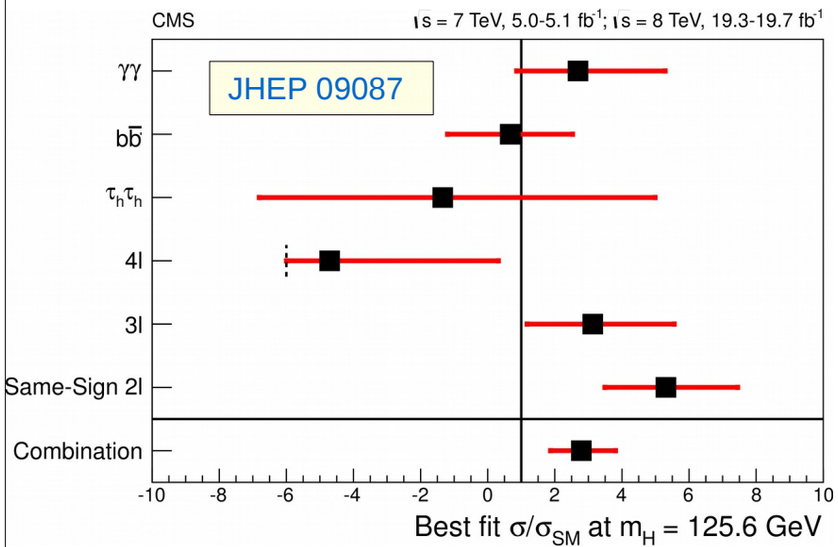
$\mu = 1.5 \pm 1.1,\ \mu < 3.4(2.2)$

$t\bar{t}H$, H → γγ — PLB 740

$\mu < 6.7(4.9)$

ATLAS

$t\bar{t}H$, H → hadrons, photons, leptons



$\mu = 2.9^{+1.1}_{-0.9},\ \mu < 4.5(2.7)$

$t\bar{t}H$, H → $b\bar{b}$, matrix element method — EPJC 75 251

$\mu = 1.2^{+1.6}_{-1.5},\ \mu < 4.2(3.3)$

CMS

~ compatible with SM, but both experiments see an excess at the 2-3σ level, **to be watched !**

ttH: the channel of LHC run-2 and HL-LHC

Higgs & SM: a paradoxical triumph

The Higgs is related to some of the deepest problems of HEP

$$\mathcal{L}_{\rm Higgs} = V_0 - \mu^2 H^\dagger H + \lambda \left(H^\dagger H\right)^2 + \left(y_{ij}\bar{\psi}_{Li}\psi_{Rj}H + h.c.\right)$$

- vacuum energy cosmological constant: $V_0 \approx (2\times 10^{-3}\ \mathrm{eV})^4 \ll M_{\rm PL}^4$
- hierarchy problem: $m_H \approx 100\ \mathrm{GeV} \ll M_{\rm Pl}$
- triviality/stability of EW vacuum
- mass and mixing hierarchy
- flavour & CP

~~ Higgs interactions ~~

many different couplings not set by any gauge symmetry
(are fundamental interactions all linked to gauge symmetry?)
but they obey 3 basic structures

1) proportionality: $g_{hff} \propto m_f \qquad g_{hVV} \propto m_V^2$

➠ test of the Higgs as the unitarizing agent

2) factor of proportionality: $g_{hff}/m_f = \sqrt{2}/v$

➠ test for extended Higgs sectors

➠ test for Higgs compositeness

3) flavor alignment: $g_{hf_if_j} \propto \delta_{ij}$

➠ test for flavor models, origin of fermion masses

Higgs & SM: a paradoxical triumph

The Higgs is related to some of the deepest problems of HEP

$\mathcal{L}_{\rm Higgs}$ = ... $Rj H + h.c.)$

vacuum energy
cosmological constant

$V_0 \approx (2 \times 10^{-3}\ \mathrm{eV})^4 \ll M_{\rm PL}^4$

mixing
hy

flavour & CP

many ...
(are fun...

nmetry

nmetry?)

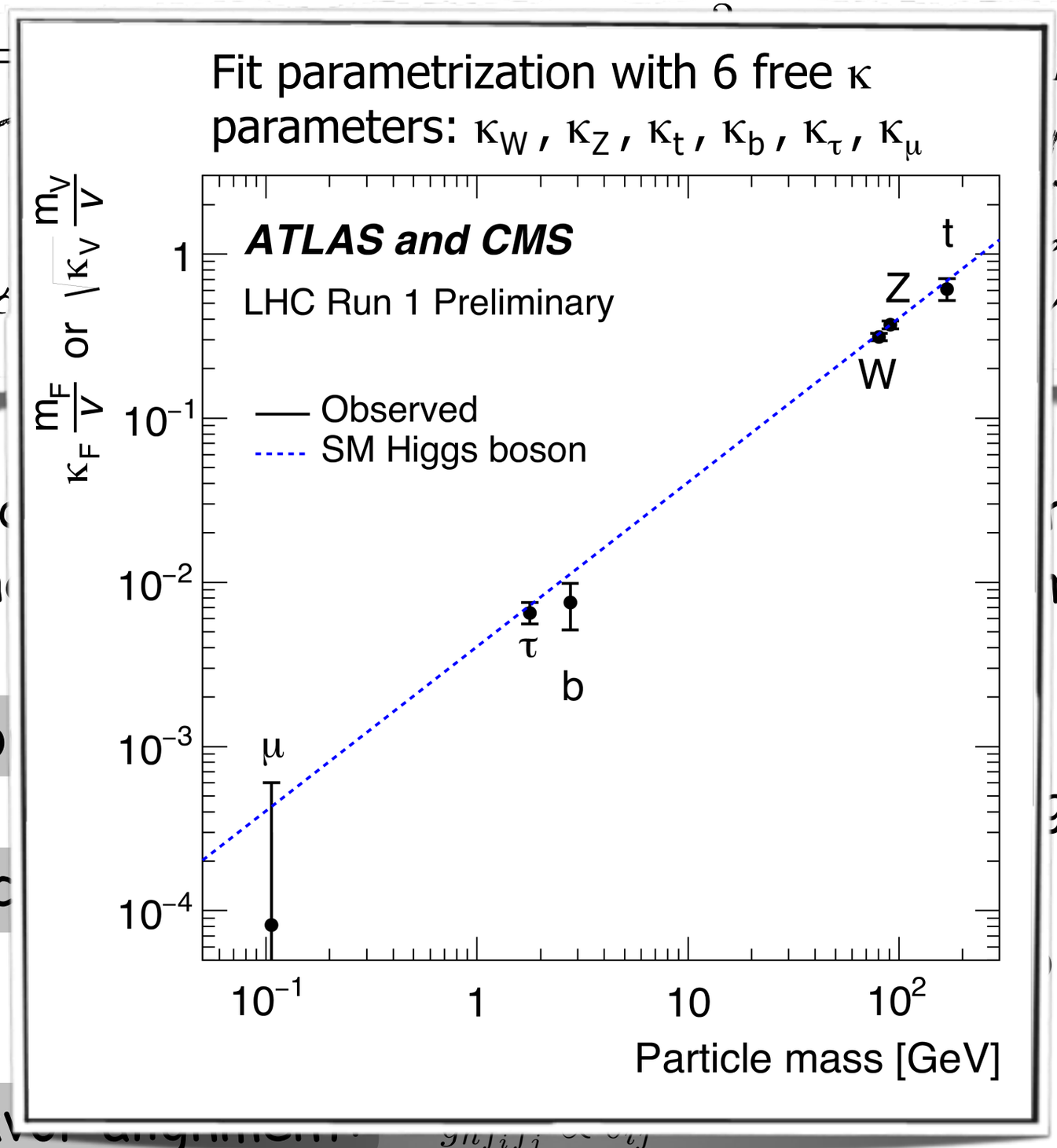
1) pro...

2) fac...

g agent

rs

3) fla...

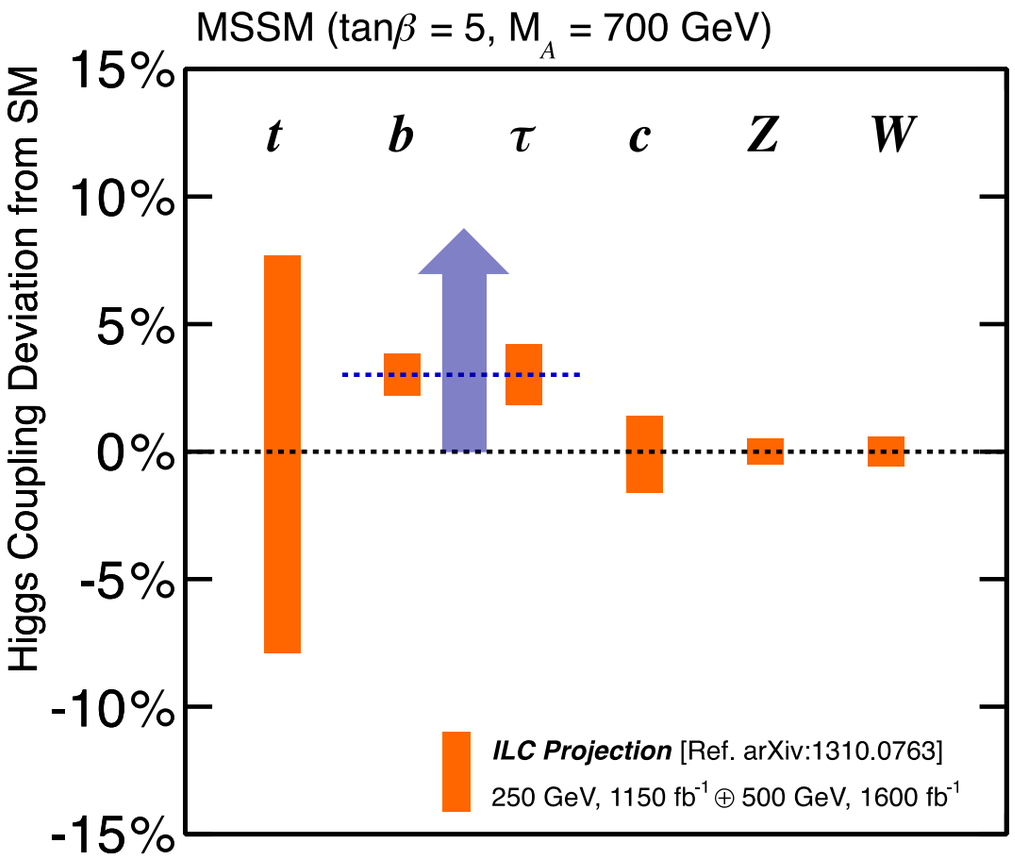


➠ test for flavor models, origin of fermion masses

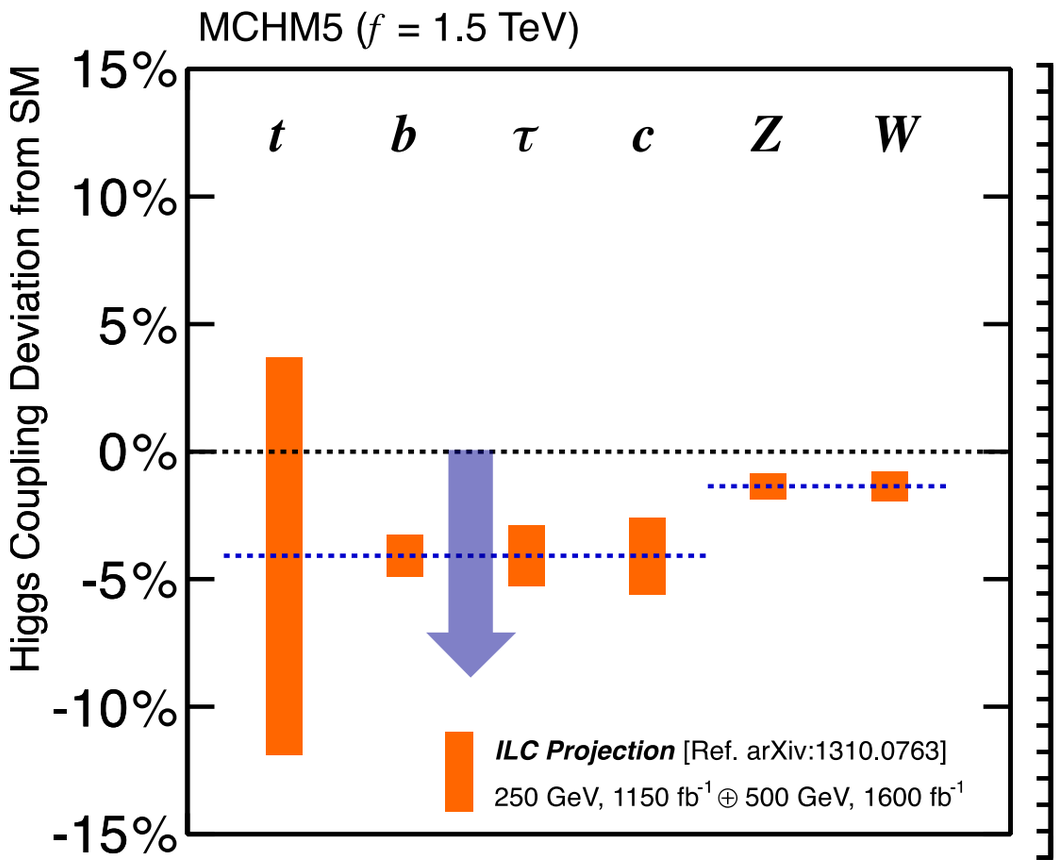
Higgs couplings and model discriminations

The pattern of Higgs coupling deviations is a signature of the underlying dynamics beyond the Standard Model

Supersymmetry
(MSSM)



Composite Higgs
(MCHM5)



ILC Physics WG, '15

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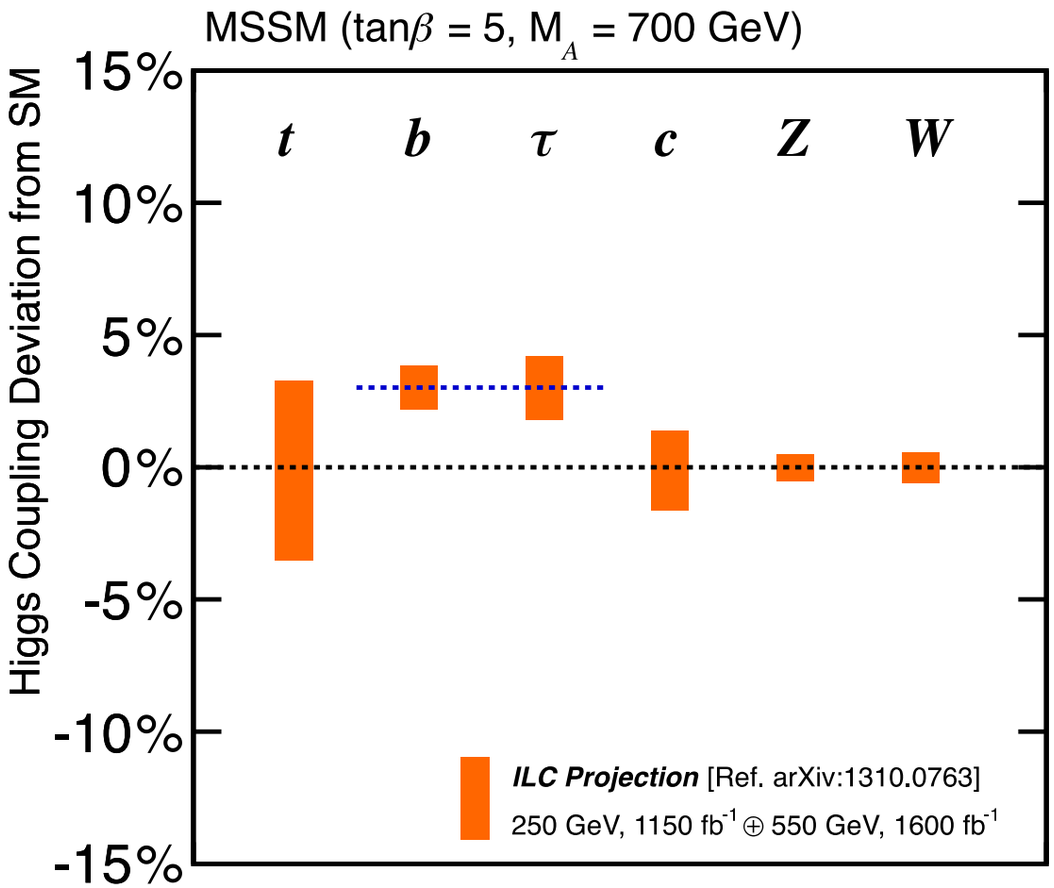

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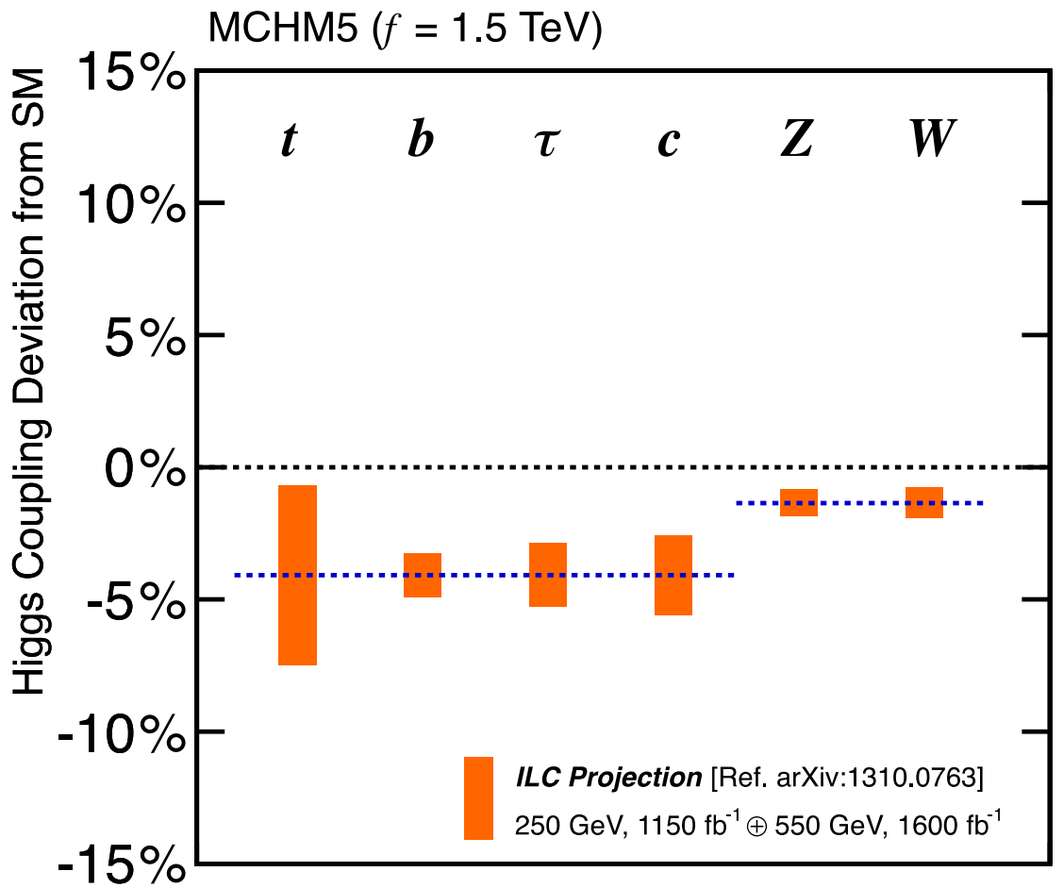


ILC Physics WG, '15

ILC 250+550 LumiUp

Higgs couplings and model discriminations

The pattern of Higgs coupling deviations is a signature of the underlying dynamics beyond the Standard Model

~~ expected largest relative deviations ~~

	hff	hVV	hγγ	hγZ	hGG	h^3
MSSM	√		√	√	√	
NMSSM	√	√	√	√	√	
PGB Composite	√	√		√		√
SUSY Composite	√	√	√	√	√	√
SUSY partly-composite			√	√	√	√
“Bosonic TC”						√
Higgs as a dilaton			√	√	√	√

]

A. Pomarol, Naturalness ’15

Higgs couplings and model discriminations

The pattern of Higgs coupling deviations is a signature of the underlying dynamics beyond the Standard Model

~~ expected largest relative deviations ~~

	hff	hVV	hγγ	hγZ	hGG	h^3
MSSM	√		√	√	√	
NMSSM	√	√			√	
PGB Composite						√
				√	√	√
SUS			√	√	√	√
						√
Higgs as a dilaton			√	√	√	√

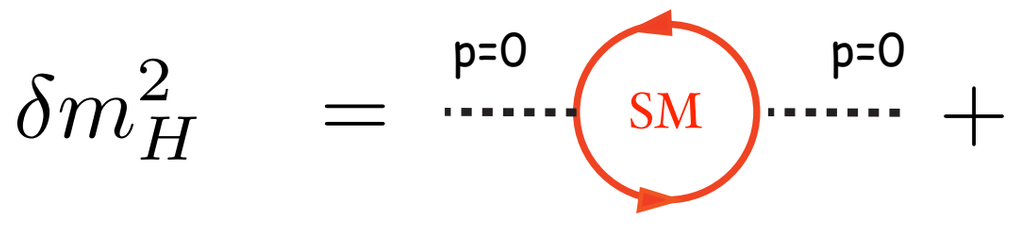
There exist some physics scenarios where high energy thresholds are important and complementary to physics near Higgs threshold (HH & ttH productions)

]

A. Pomarol, Naturalness '15

Higgs couplings as a test of naturalness

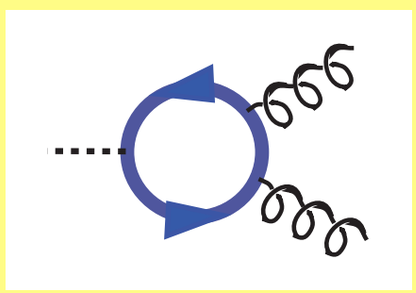
$$\delta m_H^2 = -(125\ \text{GeV})^2 \left(\frac{\Lambda}{600\ \text{GeV}}\right)^2 + \frac{g_*^2}{16\pi^2}\Lambda^2 \sim m_H^2$$



charged particles

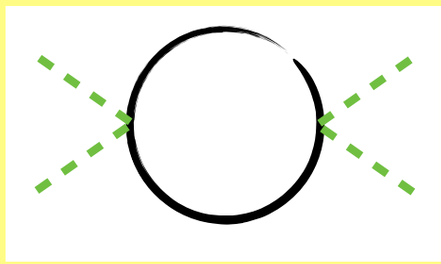
generically

neutral particles



$$\frac{g_s^2 g_*^2}{16\pi^2}\frac{1}{m_*^2}|H|^2 G_{\mu\nu}^2 \qquad \frac{e^2 g_*^2}{16\pi^2}\frac{1}{m_*^2}|H|^2 F_{\mu\nu}^2$$

$$\frac{\Delta BR(h\to\gamma\gamma, Z\gamma, gg)}{\text{SM}} \sim \frac{g_*^2 v^2}{m_*^2}$$



$$\frac{g_*^2}{16\pi^2}\frac{1}{m_*^2}\left(\partial_\mu |H|^2\right)^2$$

$$BR(h\to ii) = BR_{\text{SM}} \qquad \Gamma = \left(1 - \frac{g_*^2 v^2}{16\pi^2 m_*^2}\right)\Gamma_{\text{SM}}$$

$$\delta\sigma_{Zh} = -\frac{g_\star^2}{8\pi^2}\frac{v^2}{m_\star^2}$$

nice to be able to measure Zh & Γ

Colorful naturalness probed @ LHC

Neutral naturalness (invisible?) @ LHC

Higgs couplings measurement projections

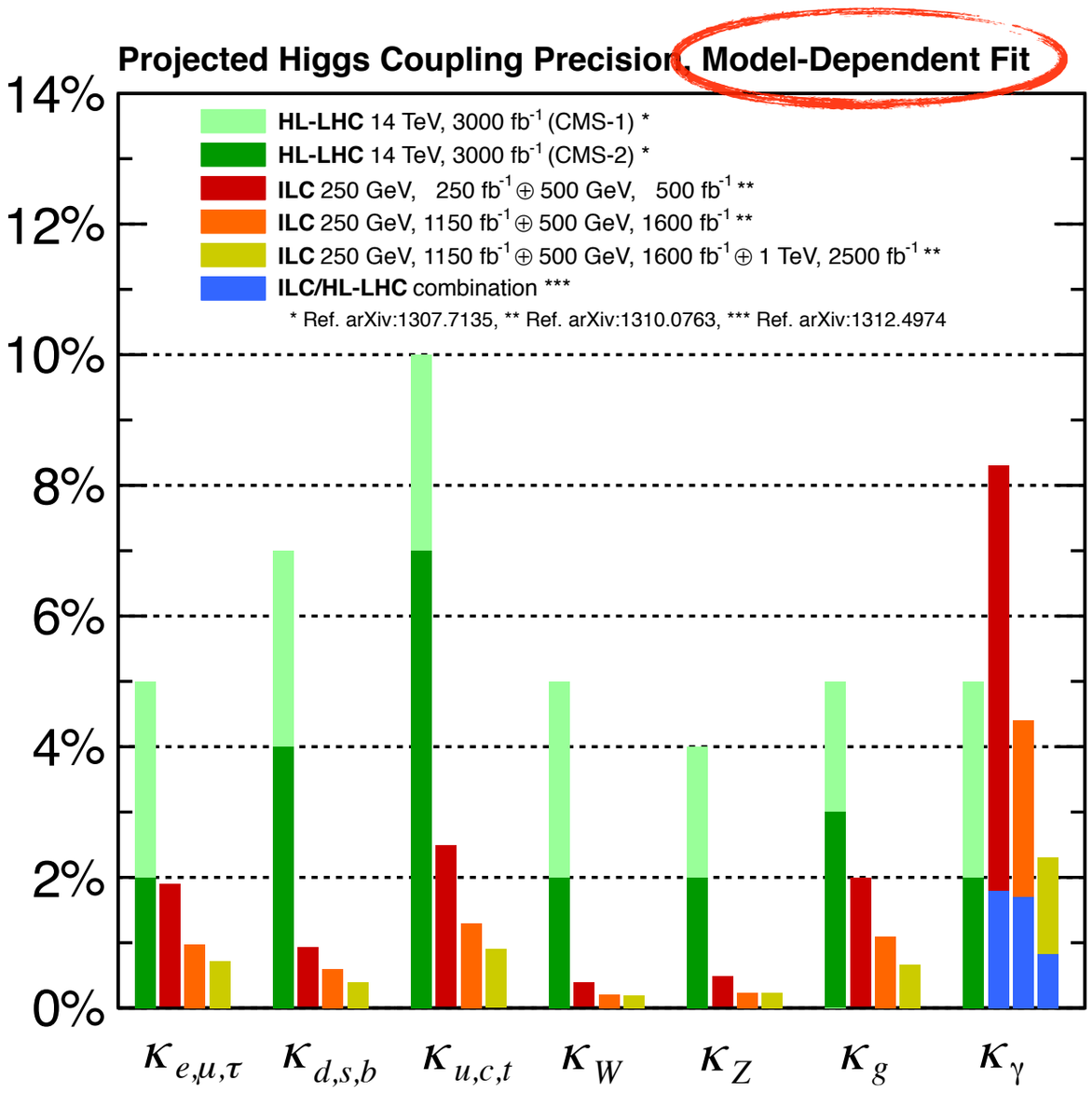
Table 1-20. *Expected precisions on the Higgs couplings and total width from a constrained 7-parameter fit assuming no non-SM production or decay modes. The fit assumes generation universality ($\kappa_u \equiv \kappa_t = \kappa_c$, $\kappa_d \equiv \kappa_b = \kappa_s$, and $\kappa_\ell \equiv \kappa_\tau = \kappa_\mu$). The ranges shown for LHC and HL-LHC represent the conservative and optimistic scenarios for systematic and theory uncertainties. ILC numbers assume (e^-, e^+) polarizations of $(-0.8, 0.3)$ at 250 and 500 GeV and $(-0.8, 0.2)$ at 1000 GeV, plus a 0.5% theory uncertainty. CLIC numbers assume polarizations of $(-0.8, 0)$ for energies above 1 TeV. TLEP numbers assume unpolarized beams.*

Facility	LHC	HL-LHC	ILC500	ILC500-up	ILC1000	ILC1000-up	CLIC	TLEP (4 IPs)
$\sqrt{s}$ (GeV)	14,000	14,000	250/500	250/500	250/500/1000	250/500/1000	350/1400/3000	240/350
$\int \mathcal{L}dt$ (fb^{-1})	300/expt	3000/expt	250+500	1150+1600	250+500+1000	1150+1600+2500	500+1500+2000	10,000+2600
κ_γ	$5-7\%$	$2-5\%$	8.3%	4.4%	3.8%	2.3%	−/5.5/<5.5%	1.45%
κ_g	$6-8\%$	$3-5\%$	2.0%	1.1%	1.1%	0.67%	3.6/0.79/0.56%	0.79%
κ_W	$4-6\%$	$2-5\%$	0.39%	0.21%	0.21%	0.2%	1.5/0.15/0.11%	0.10%
κ_Z	$4-6\%$	$2-4\%$	0.49%	0.24%	0.50%	0.3%	0.49/0.33/0.24%	0.05%
κ_ℓ	$6-8\%$	$2-5\%$	1.9%	0.98%	1.3%	0.72%	3.5/1.4/<1.3%	0.51%
$\kappa_d = \kappa_b$	$10-13\%$	$4-7\%$	0.93%	0.60%	0.51%	0.4%	1.7/0.32/0.19%	0.39%
$\kappa_u = \kappa_t$	$14-15\%$	$7-10\%$	2.5%	1.3%	1.3%	0.9%	3.1/1.0/0.7%	0.69%

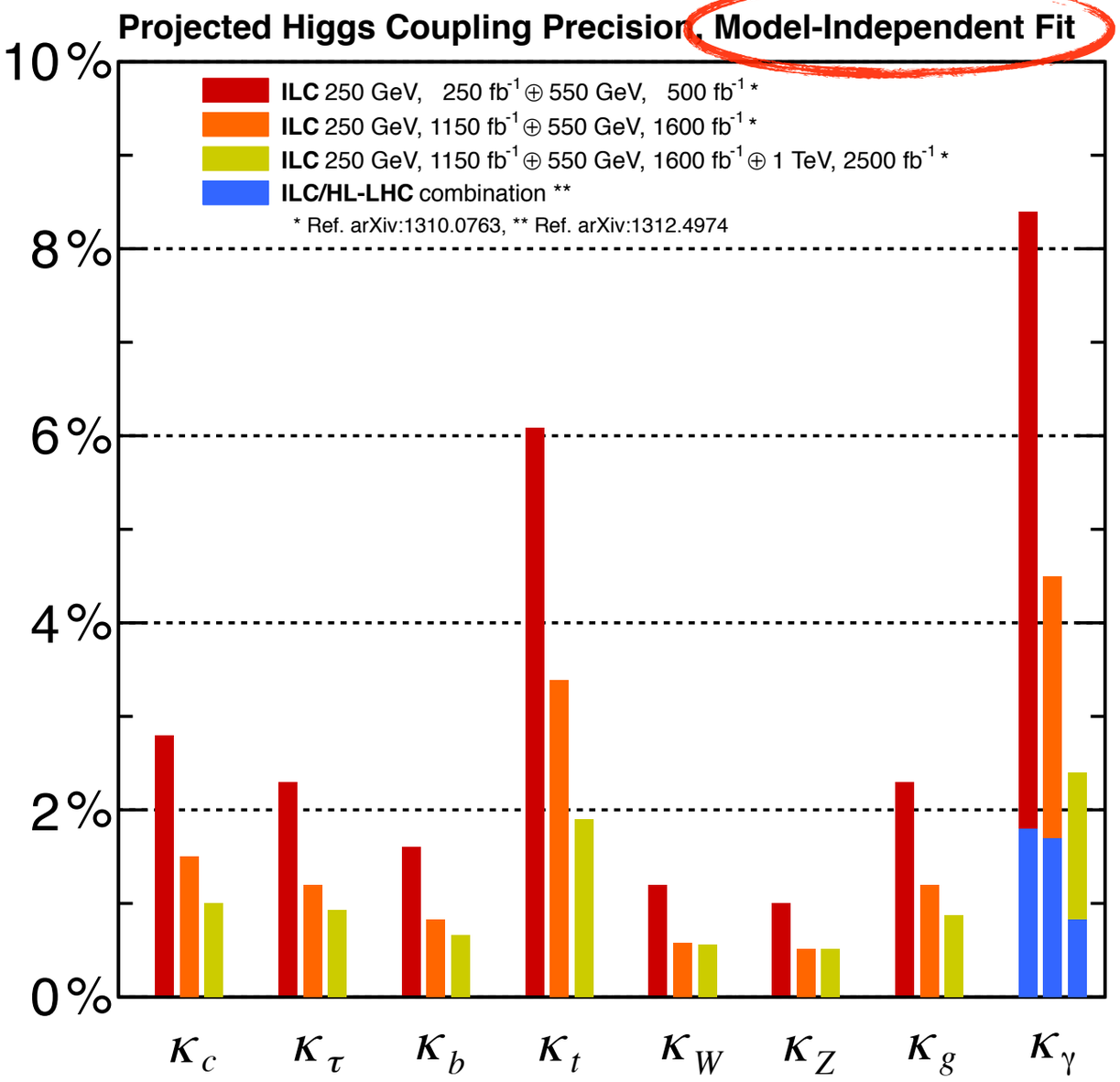
Rich experimental program of (sub)percent precision

Higgs couplings measurement projections

assumption about Higgs width



no assumption about Higgs width



Rich experimental program of (sub)percent precision
Nice synergy/complementarity LHC-ILC (hγγ)

use BR ratios from hh with absolute precise BR from ee

to export ee precision to Higgs decays that are limited by statistics in ee

Higgs couplings measurement projections

g_{HXY}	FCC-ee	FCC-hh
ZZ	0.16%	
WW	0.85%	
γ γ	1.7%	<1% ?
Zγ		1% ?
tt		1% ?
bb	0.88%	
τ τ	0.94%	
cc	1.0%	
ss	H→Vγ, in progr.	
μμ	6.4%	2% ?
uu,dd	H→Vγ, in progr.	
ee	e^+e^-→H, in progr.	
HH		5% ?
BR_{exo}	0.48%	< 10^{-6} ?

# of Higgses in 3 ab^{-1}

100 TeV > 2 billion

33 TeV > 500 million

14 TeV > 150 million

In comparison, O(million) Higgs at Higgs factories

@ 100 TeV

	σ	N / 10ab^{-1}
gg→H	740 pb	7.4 G
VBF	82 pb	0.8 G
WH	16 pb	160 M
ZH	11 pb	110 M
ttH	38 pb	380 M
gg→HH	1.4 pb	14 M

M. Mangano, HXSWG '15

the Higgses produced are not the Higgses reconstructed but fantastic playground for rare/exotic decays

Flavor alignment

In SM, the Yukawa interactions are the only source of the fermion masses

$$y_{ij}\bar{f}_{L_i}Hf_{R_j} = \frac{y_{ij}v}{\sqrt{2}}\bar{f}_{L_i}f_{R_j} + \frac{y_{ij}}{\sqrt{2}}h\bar{f}_{L_i}f_{R_j}$$

mass ↑ ↑ higgs-fermion interactions

both matrices are simultaneously diagonalizable

↓ ↓ ↓

no tree-level Flavor Changing Current induced by the Higgs

Not true anymore if the SM fermions mix with vector-like partners(*) or for non-SM Yukawa

$$y_{ij}\left(1+c_{ij}\frac{|H|^2}{f^2}\right)\bar{f}_{L_i}Hf_{R_j} = \frac{y_{ij}v}{\sqrt{2}}\left(1+c_{ij}\frac{v^2}{2f^2}\right)\bar{f}_{L_i}f_{R_j} + \left(1+3c_{ij}\frac{v^2}{2f^2}\right)\frac{y_{ij}}{\sqrt{2}}h\bar{f}_{L_i}f_{R_j}$$

Look for SM forbidden Flavor Violating decays $h \to \mu\tau$ and $t \to hc$

- weak indirect constrained by flavor data (e.g. $\mu \to e\gamma$): BR<10%
- ATLAS and CMS have the sensitivity to set bounds O(1%)
- ILC/CLIC/FCC-ee can certainly do much better

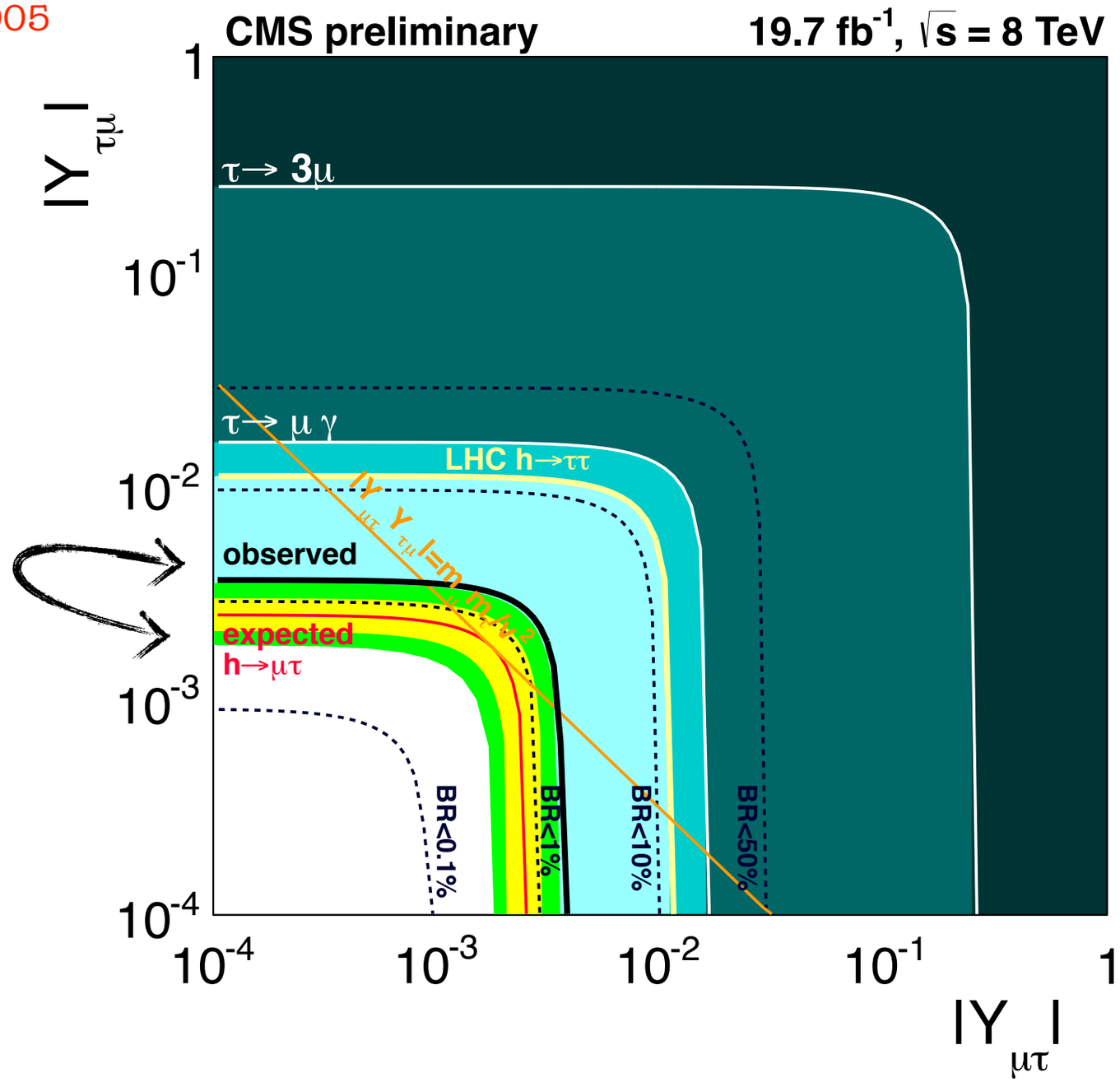
Blankenburg, Ellis, Isidori '12

Harnik et al '12
Davidson, Verdier '12
CMS-PAS-HIG-2014-005

(*) e.g. Buras, Grojean, Pokorski, Ziegler '11

Flavor alignment

CMS-PAS-HIG-2014-005



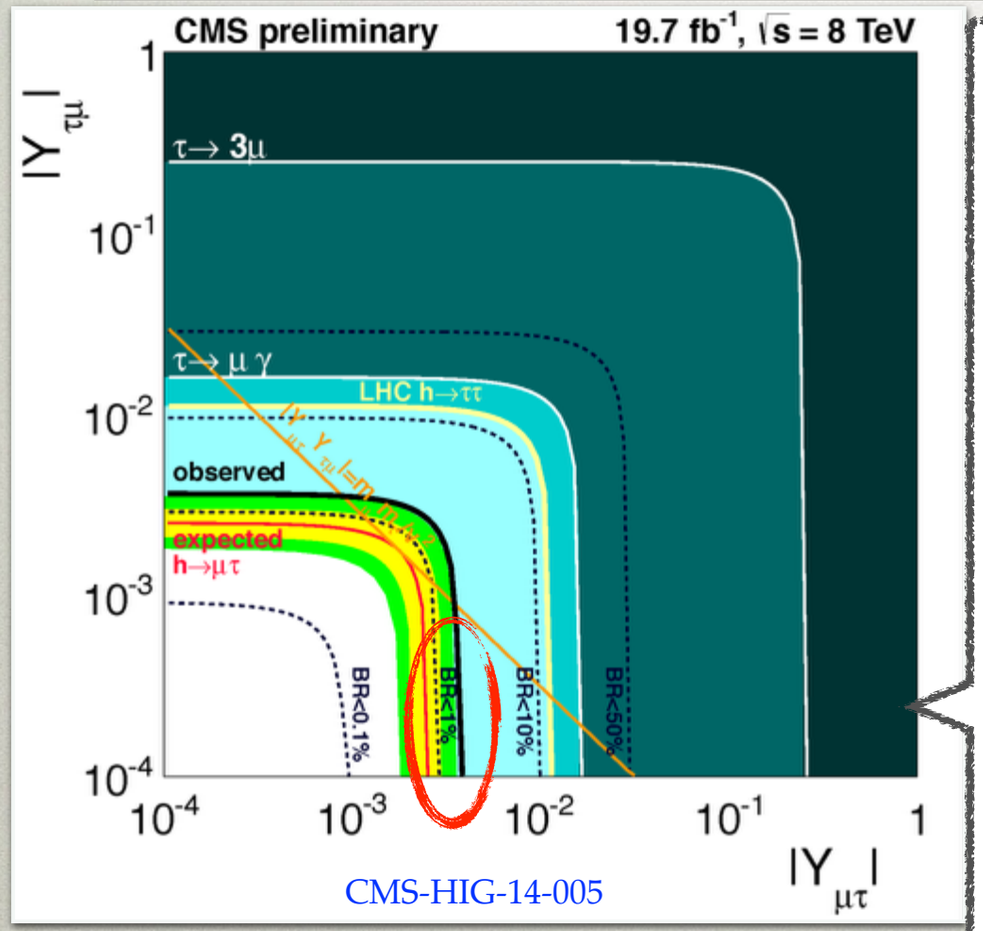
by the way:
2.3σ excess!

Off-diagonal Higgs couplings can reveal the origin of flavor

Importance of efficient flavor tagging!

Flavor alignment @ FCC

$h \to \tau\mu$



- right now: 2j channel statistics limited, 0j+1j not
- how about with ~10^9 h?
  LHC8 $\Rightarrow$ 100 TeV 3 ab^{-1}
- assume same scaling for signal and bckg
  - $Br \sim 10^{-2} \Rightarrow Br \sim 10^{-4}$
  - $\Lambda \sim 0.2\ TeV \Rightarrow \Lambda \sim 2 TeV$
- if bckg free
  - $Br \sim 10^{-2} \Rightarrow Br \sim 10^{-6}$
  - $\Lambda \sim 0.2\ TeV \Rightarrow \Lambda \sim 20 TeV$
    $(Y_{\mu\tau}\, Y_{\tau\mu} = m_\mu m_\tau / \Lambda^2)$

J. Zupan, FCC@Washington '15

sensitivity in $h \to \mu e$ channel not competitive with flavor measurements

Flavor changing Higgs couplings @ LHC

		ATLAS	CMS
$h\to\mu\tau$		✔	✔
$t\to hq$ w/	$h\to\gamma\gamma$	✔	✔
	h→multilepton		✔
	h→bb	✔	

Assuming a simple universal scaling: $Y_{ij} \sim \sqrt{(m_i m_j/v^2)}$,

BR($h\to\mu\tau$)=(0.89±0.40)% implies BR(t→hc)~0.25%

while direct constraint is currently ~0.5%, but can improve by combining various channels

Flavor changing Higgs couplings @ LHC

In SM, the Yukawa interactions are the only source of the fermion masses

Not the case in e.g. generic 2HDM

Omura, Senaha and Tobe '15

Botella, Branco, Nebot, Rebelo '15

$$y_{hij} = \frac{m_f^i}{v} s_{\beta\alpha}\delta_{ij} + \frac{\rho_f^{ij}}{\sqrt{2}} c_{\beta\alpha},$$

$$y_{Hij} = \frac{m_f^i}{v} c_{\beta\alpha}\delta_{ij} - \frac{\rho_f^{ij}}{\sqrt{2}} s_{\beta\alpha},$$

$$y_{Aij} = \begin{cases} -\frac{i\rho_f^{ij}}{\sqrt{2}}, & (f = u), \\ \frac{i\rho_f^{ij}}{\sqrt{2}}, & (f = d,\ e), \end{cases}$$

even small mixing ($c_{\beta\alpha}$~0.1)
and moderate FCYC (ρ~0.3)
can explain
BR(h→$\mu\tau$)=(0.89±0.40)%

consequences:

- $(g\text{-}2)_\mu$
- $\tau \to \mu\gamma$

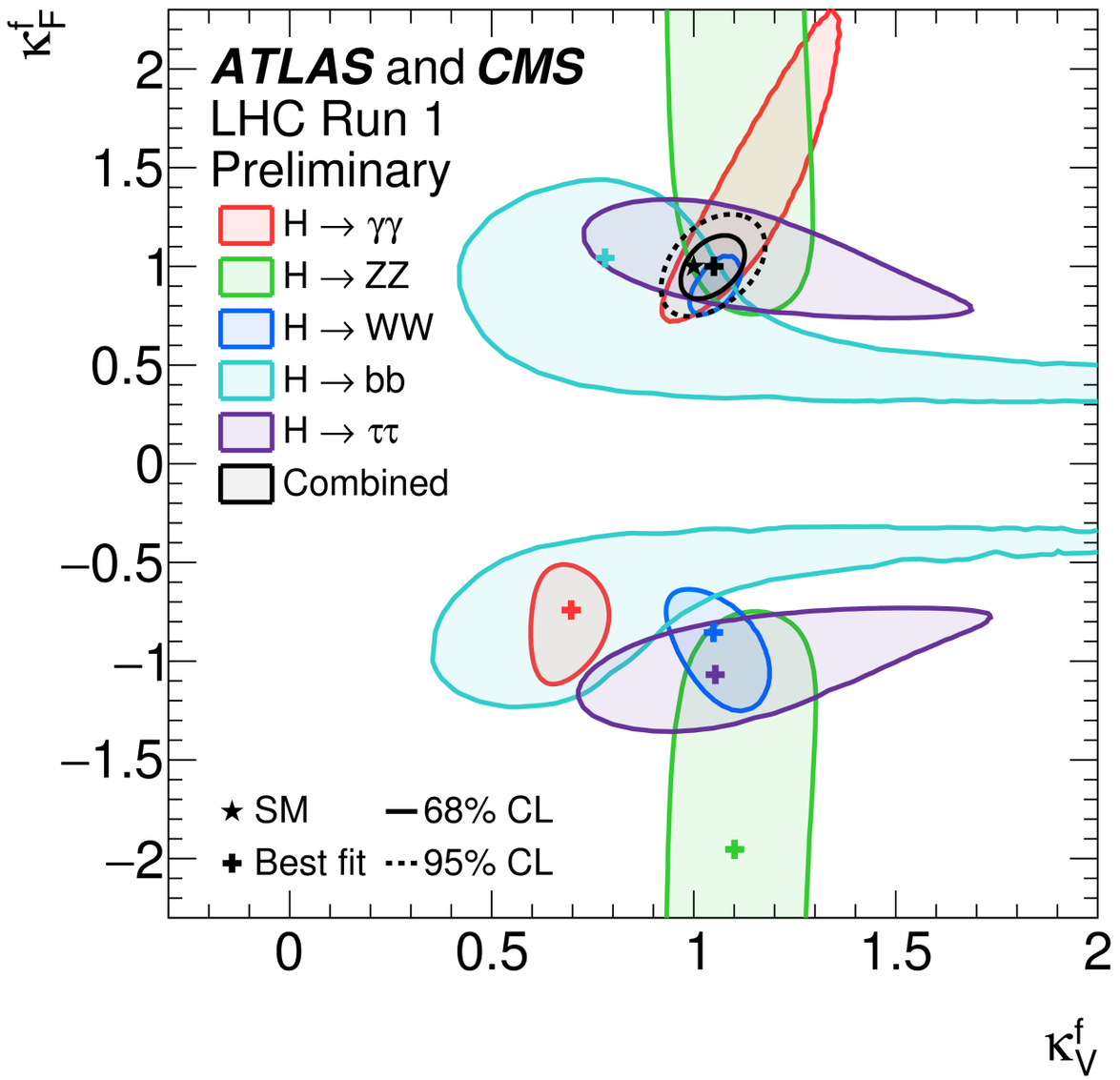
NB: Flavor Changing decays of H and A can easily dominate
would explain why even light H/A hasn't been found at LHC directly

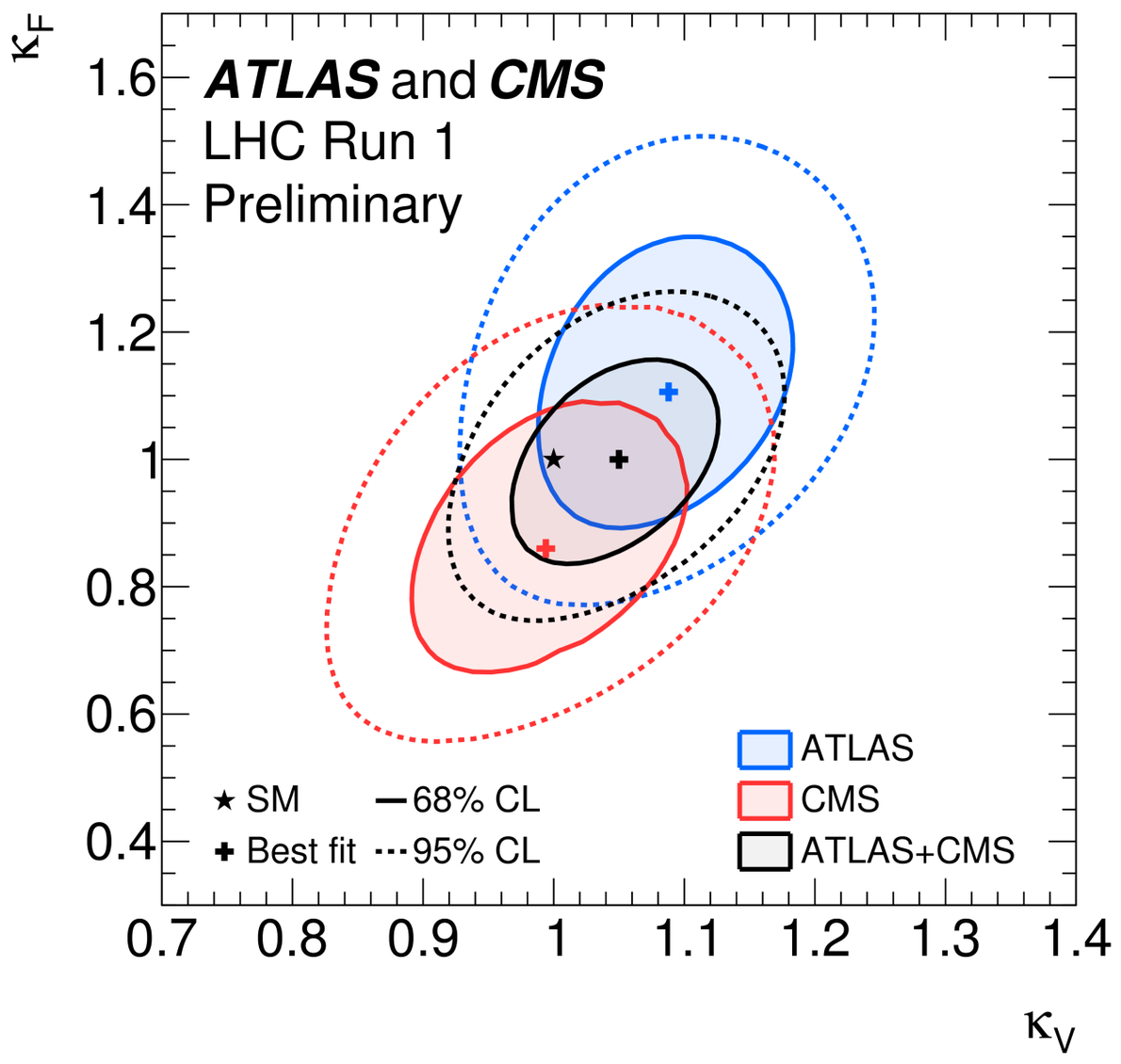
Why going beyond inclusive Higgs processes?

So far the LHC has mostly produced Higgses on-shell
in processes with a characteristic scale $\mu \approx m_H$

⇩ ⇩

access to Higgs couplings @ m_H





Why going beyond inclusive Higgs processes?

So far the LHC has mostly produced Higgses on-shell
in processes with a characteristic scale $\mu \approx m_H$

⇩ ⇩

access to Higgs couplings @ m_H

Producing a Higgs with boosted additional particle(s)
probe the Higgs couplings @ large energy
(important to check that the Higgs boson ensures perturbative unitarity)

Probing new corrections to the SM Lagrangian?

on-shell Z @ LEP1	off-shell Z @ LEP2
constraints on S and T oblique corrections	constraints on W and Y oblique corrections (same order as S and T but cannot be probed @ LEP1)

high-p_T Higgs (either in VH, H+jet) are easy to trigger on
even at high pileup, contrary to inclusive gg → h

Boosted Higgs

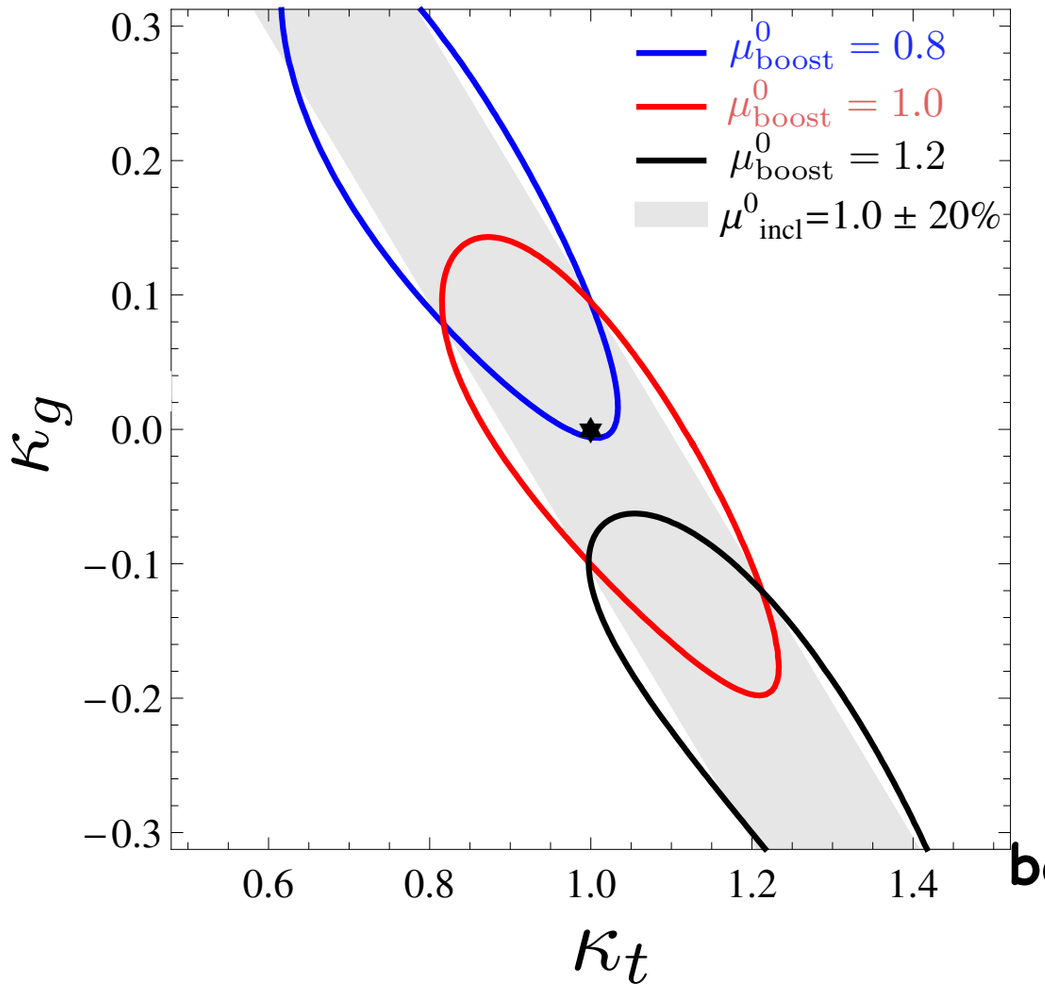
see also Azatov, Paul '13

high p_T tail discriminates short and long distance physics contribution to $gg \to h$

$$\sqrt{s} = 14 \text{ TeV}, \int dt\, \mathcal{L} = 3\text{ab}^{-1}, p_T > 650 \text{ GeV}$$

(partonic analysis in the boosted "ditau-jets" channel)

see Schlaffer et al '14 for a more complete analysis including WW channel



Grojean, Salvioni, Schlaffer, Weiler '13

high p_T ≈ Higgs off-shell
we "see" the details of the particles running inside the loops

Baur, Glover '90
Langenegger, Spira, Starodumov, Trueb '06

10-20% precision on κ_t

⇓ ⇓

competitive/complementary to htt channel for the measure the top-Higgs coupling

boosted h+jet easy to trigger on even at high pile-up (contrary to inclusive $gg \to h$)

Are the NLO_m QCD corrections (not known) going to destroy all the sensitivity?
Frontier priority: N^3LO_∞ for inclusive xs or NLO_{mt} for pT spectrum?

Boosted Higgs

see also Azatov, Paul '13

high p_T tail discriminates short and long distance physics contribution to $gg \to h$

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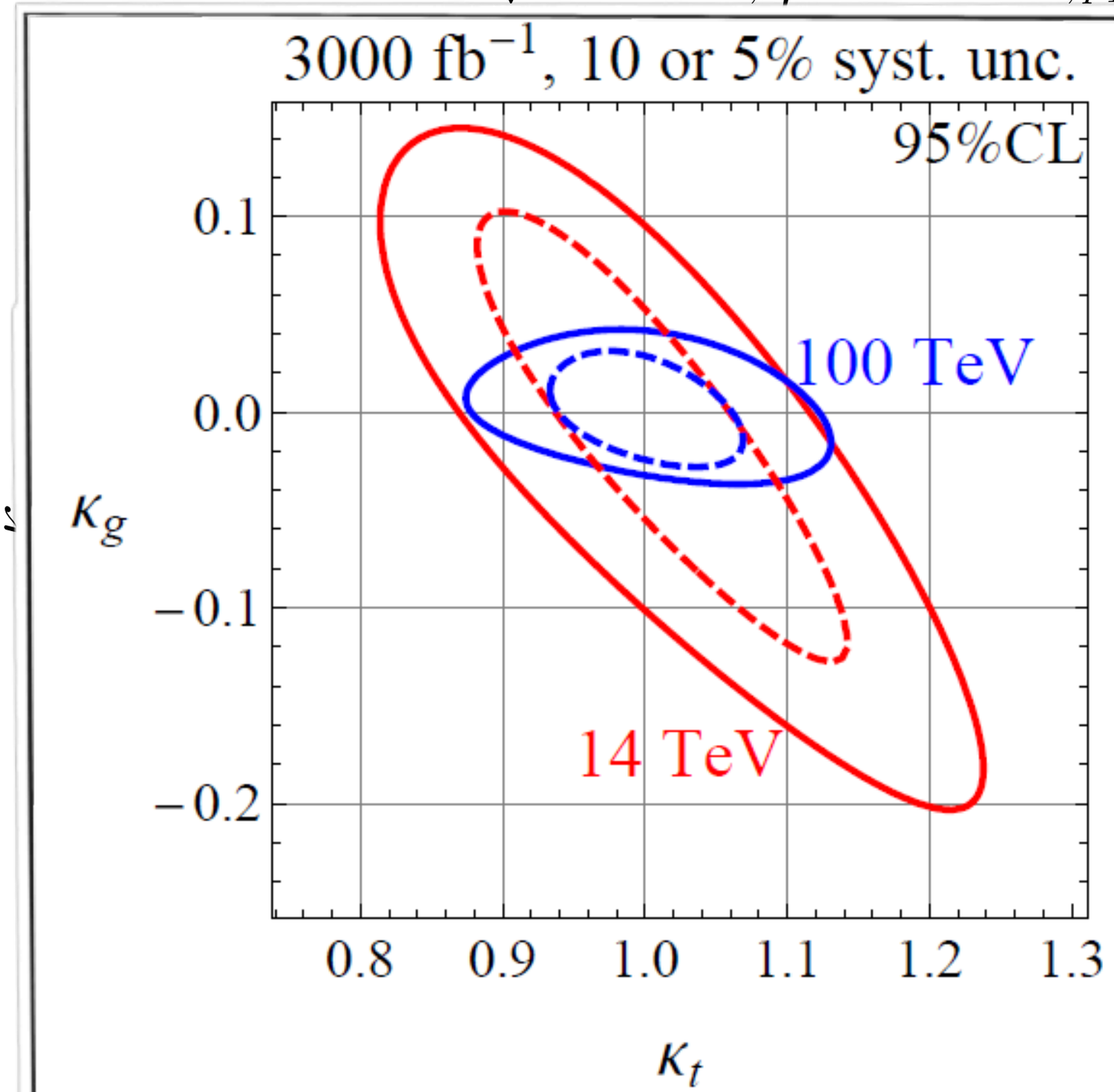


Grojean, Salvioni, Schlaffer, Weiler '13

A perfect case for a very energetic machine

tth increases by 10 from 14 to 100TeV

h+$j_{pT>600GeV}$ increases by 210

$$\mathcal{R}_{14} = \frac{\sigma(p_T > 650\,\text{GeV})}{\sigma(p_T > 150\,\text{GeV})}$$

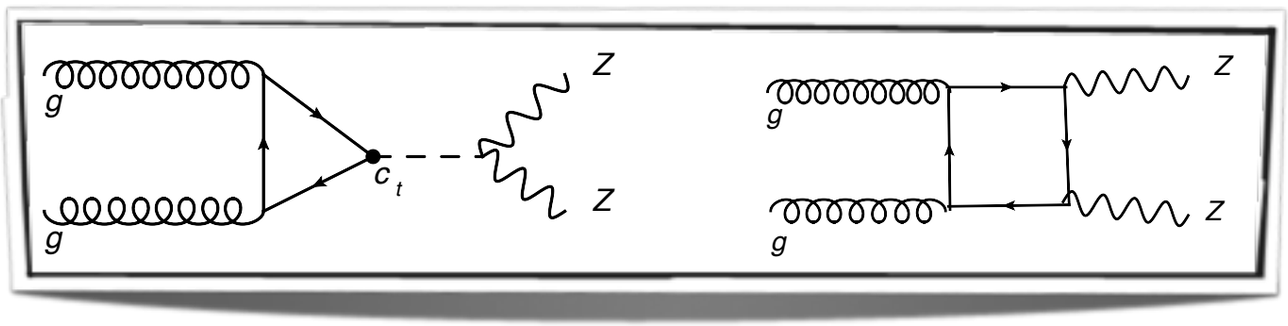
$$\mathcal{R}_{100} = \frac{\sigma(p_T > 2000\,\text{GeV})}{\sigma(p_T > 500\,\text{GeV})}$$

p

Off-shell Higgs: $gg \to h^* \to ZZ \to 4l$

off-shell effects enhanced by the particular couplings of H to V_L

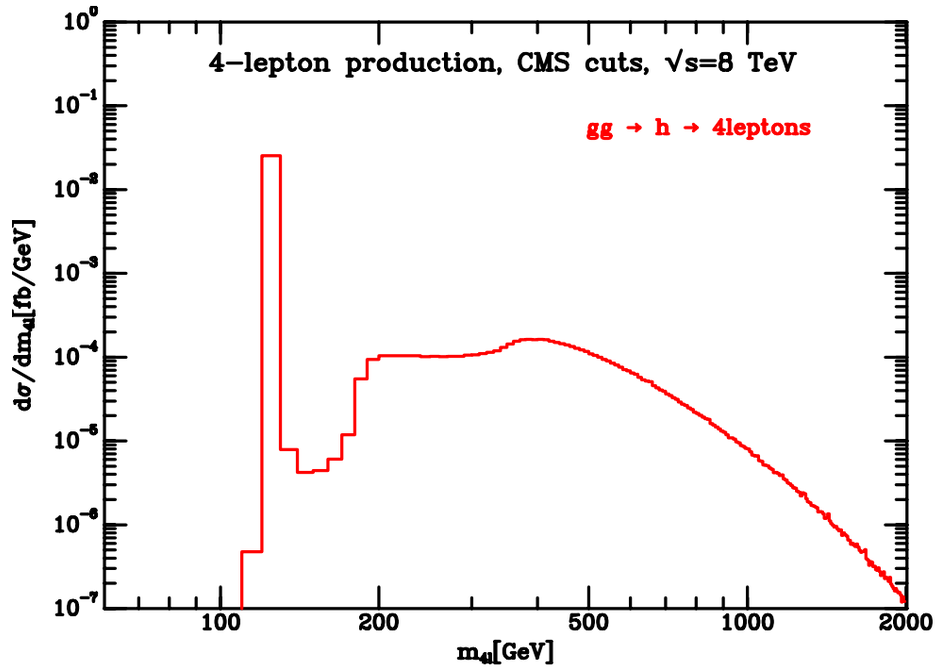
Glover, van der Bij '89



$$\mathcal{M}^{++00}_{\rm Higgs} \sim \log^2 \frac{\hat{s}}{m_t^2} \qquad \mathcal{M}^{++00}_{\rm box} \sim -\log^2 \frac{\hat{s}}{m_t^2}$$

SM: cancelation forced by unitarity

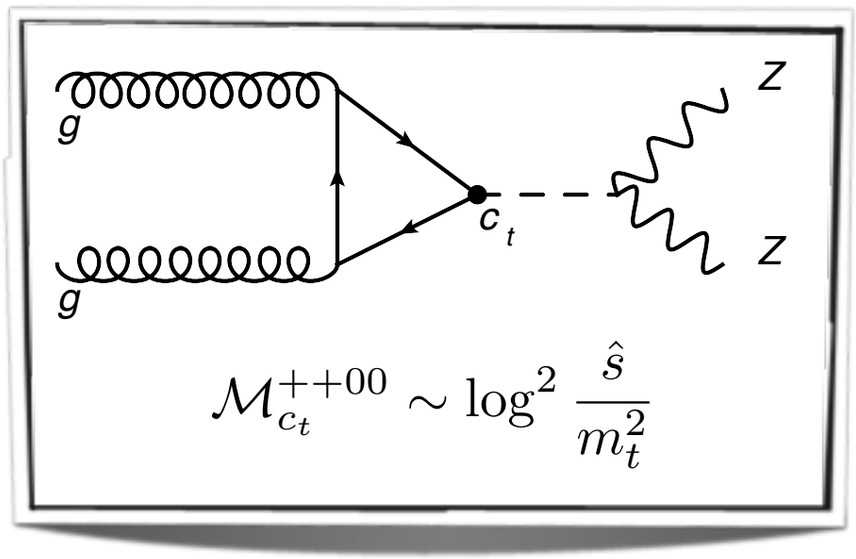
BSM: deviations of Higgs couplings at large $\hat{s}$ will be amplified

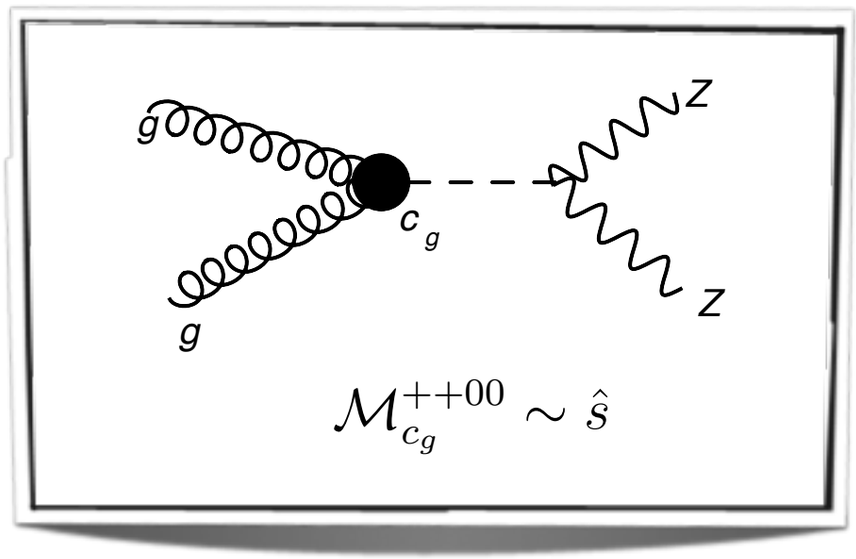


Azatov, Grojean, Paul, Salvioni '14

interpretations in terms of bounds of the Higgs width are limited/model-dependent

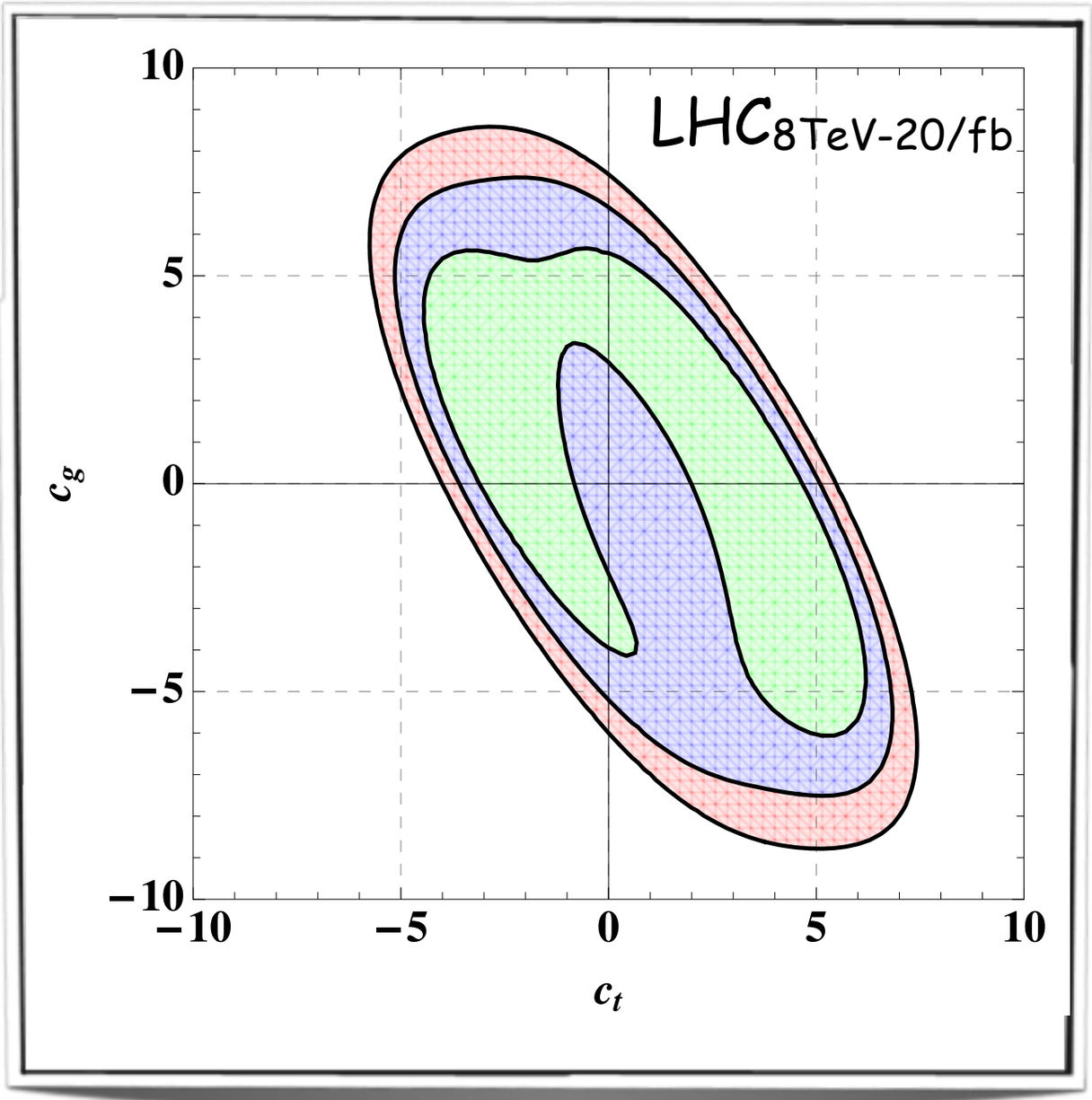
but data can be better used to measure the structure of the couplings at high $\sqrt{s}$

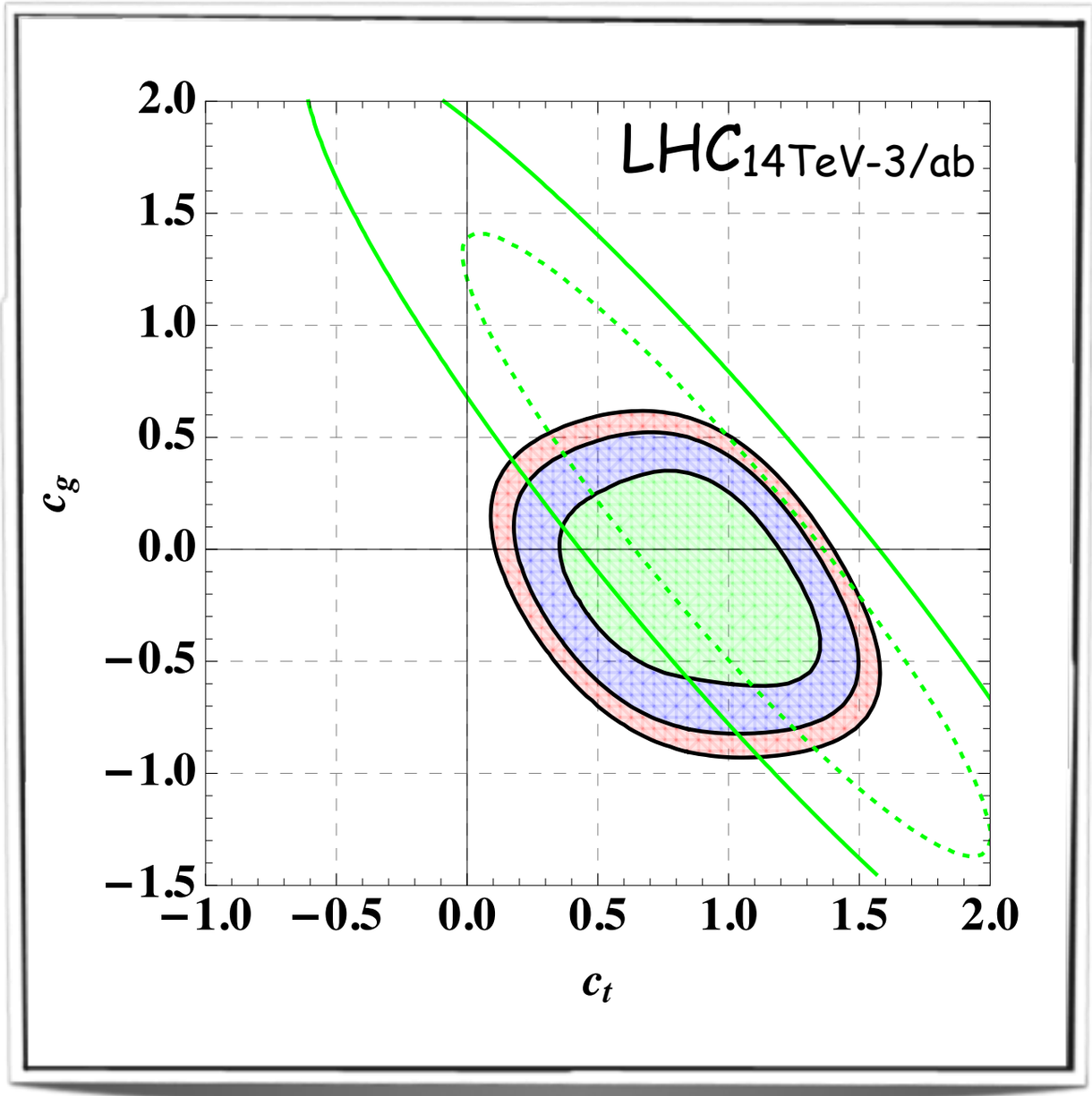




Off-shell Higgs: $gg \to h^* \to ZZ \to 4l$

off-shell effects enhanced by the particular couplings of H to V_L





Azatov, Grojean, Paul, Salvioni '14

Higgs self-couplings

The Higgs self-couplings plays important roles

1) controls the stability of the EW vacuum

2) dictates the dynamics of EW phase transition and potentially conditions the generation of a matter-antimatter asymmetry via EW baryogenesis

Does it need to be measured with high accuracy?

difficult to design new physics scenarios that dominantly affect the Higgs self-couplings and leave the other Higgs coupling deviations undetectable

Higgs self-coupling prospects

	HL LHC 3/ab	ILC/CLIC	FCC 100TeV
Precision on λ_{HHH}	$b\bar{b}\gamma\gamma$: poor, only $\sim O(1)$ determination Other channels: needs more detailed studies	ILC • DHS alone at 500 GeV and 1TeV gives only $\sim O(1)$ determination • ~28% via VBF at 1TeV, 1/ab CLIC at 3TeV, 2/ab • ~12% via VBF	$b\bar{b}\gamma\gamma$: golden channel. 5-10% determination might be possible with 30/ab. ~3x less sensitivity with 3/ab
Comments	Combining various channels might be important	The role of VBF is important High CM energy and high luminosity are crucial	Improvements on heavy flavor tagging, fakes, mass resolution etc are crucial to achieve our goal

M. Son, Washington '15

Higgs self-couplings and Naturalness

In the SM, $|H|^2$ is the only relevant operator
and it is the source of the hierarchy/naturalness/fine-tuning problem
It presence has never been tested!

Reconstructing the Higgs potential before EW symmetry breaking
from measurements around the vacuum is difficult in general
but we can easily test gross features, like the presence of the relevant operator

SM

$$V = -\mu^2|H|^2 + \lambda|H|^4 \quad \Rightarrow \quad V(h) = \frac{1}{2}m_h^2 h^2 + \frac{1}{6}\frac{3m_h^2}{v}h^3 + \dots$$

EWSB W/O H^2

$$V = -\lambda|H|^4 + \frac{1}{\Lambda^2}|H|^6 \quad \Rightarrow \quad V(h) = \frac{1}{2}m_h^2 h^2 + \frac{1}{6}\frac{7m_h^2}{v}h^3 + \dots$$

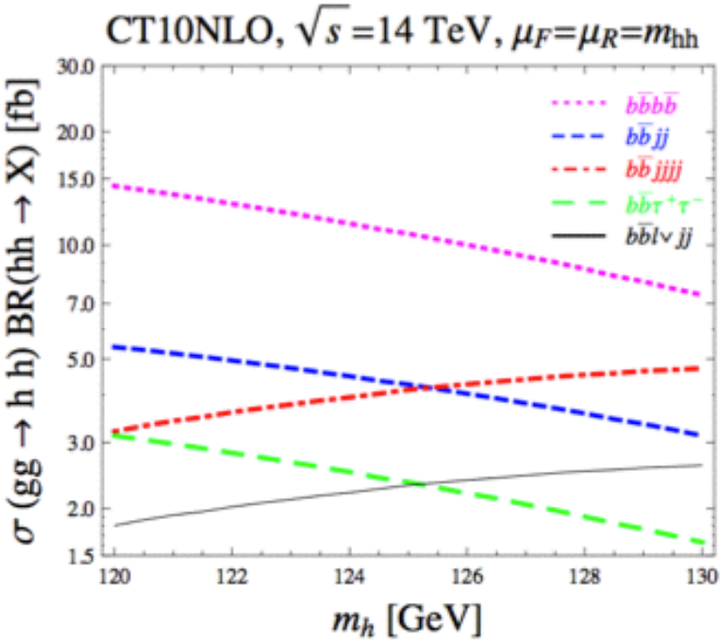
200% correction
to SM prediction
+
allows 1st phase transition

HH@LHC

Measuring this small cross section in an inclusive search is very challenging at the HL-LHC: compromise between branching ratio and cleanliness of the signal

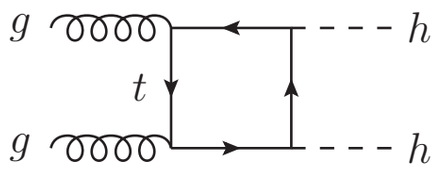
M. Spannowsky, Mainz '15

Channel	BR (%)	Events/3 ab
$bbWW$	24.7	30000
$bb\tau\tau$	7.3	9000
$WWWW$	4.3	5200
$bb\gamma\gamma$	0.27	330
$bbZZ(\to e^+e^-\mu^+\mu^-)$	0.015	19
$\gamma\gamma\gamma\gamma$	0.00052	1

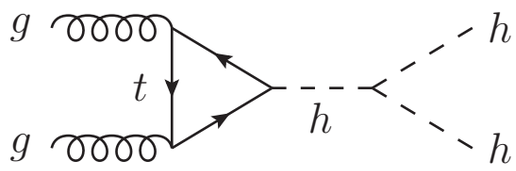


Decay	Issues	Expectation 3000 ifb	References
$b\bar{b}\gamma\gamma$	• Signal small • BKG large & difficult to asses • Simple reconst.	$S/B \simeq 1/3$ $S/\sqrt{B} \simeq 2.5$	[Baur, Plehn, Rainwater] [Yao 1308.6302] [Baglio et al. JHEP 1304]
$b\bar{b}\tau^+\tau^-$	• tau rec tough • largest bkg tt • Boost+MT2 might help	**differ a lot** $S/B \simeq 1/5$ $S/\sqrt{B} \simeq 5$	[Dolan, Englert, MS] [Barr, Dolan, Englert, MS] [Baglio et al. JHEP 1304]
$b\bar{b}W^+W^-$	• looks like tt • Need semilep. W to rec. two H • Boost + BDT proposed	**differ a lot** best case: $S/B \simeq 1.5$ $S/\sqrt{B} \simeq 8.2$	[Dolan, Englert, MS] [Baglio et al. JHEP 1304] [Papaefstathiou, Yang, Zurita 1209.1489]
$b\bar{b}b\bar{b}$	• Trigger issue (high pT kill signal) • 4b background large difficult with MC • Subjets might help	$S/B \simeq 0.02$ $S/\sqrt{B} \leq 2.0$	[Dolan, Englert, MS] [Ferreira de Lima, Papaefstathiou, MS] [Wardrope et al, 1410.2794]
others	• Many taus/W not clear if 2 Higgs • Zs, photons no rate		

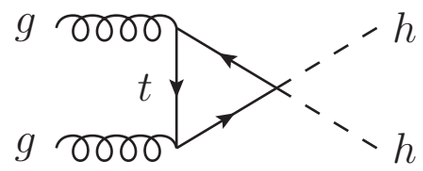
HH production as a probe of HE couplings



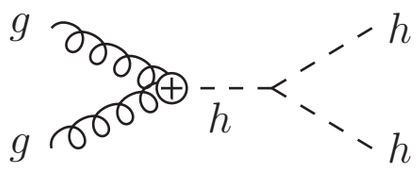
$\sim c_t^2 \times const.$



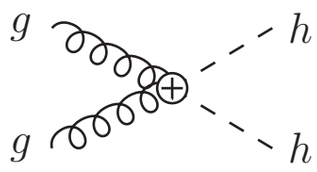
$\sim c_t c_3 \times \frac{m_h^2}{\hat{s}} \log^2\left(\frac{m_t^2}{\hat{s}}\right)$



$\sim c_{2t} \times \log^2\left(\frac{m_t^2}{\hat{s}}\right)$

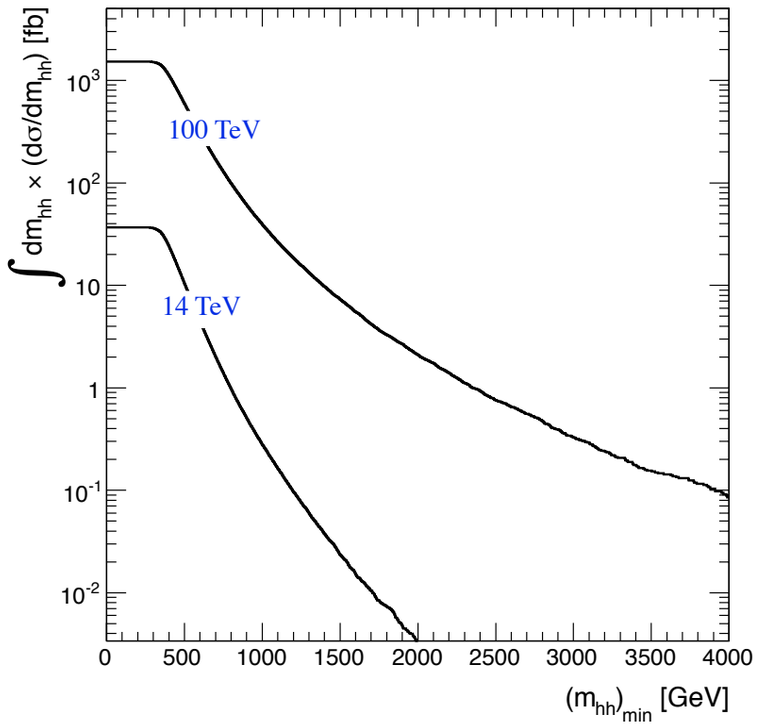


$\sim c_g c_3 \frac{\alpha_s}{4\pi} \times const.$

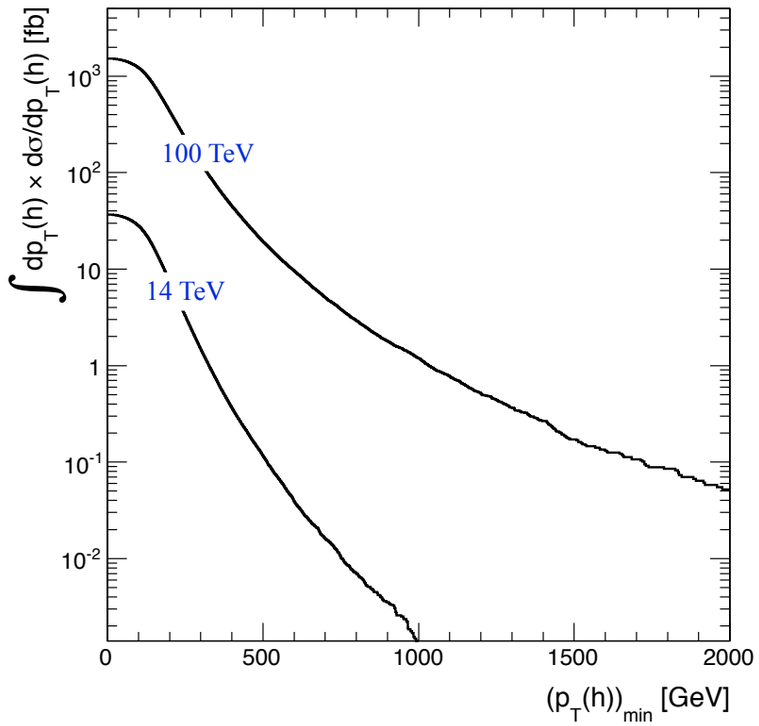


$\sim c_{2g} \frac{\alpha_s}{4\pi} \frac{\hat{s}}{v^2}$

Signal (SM)

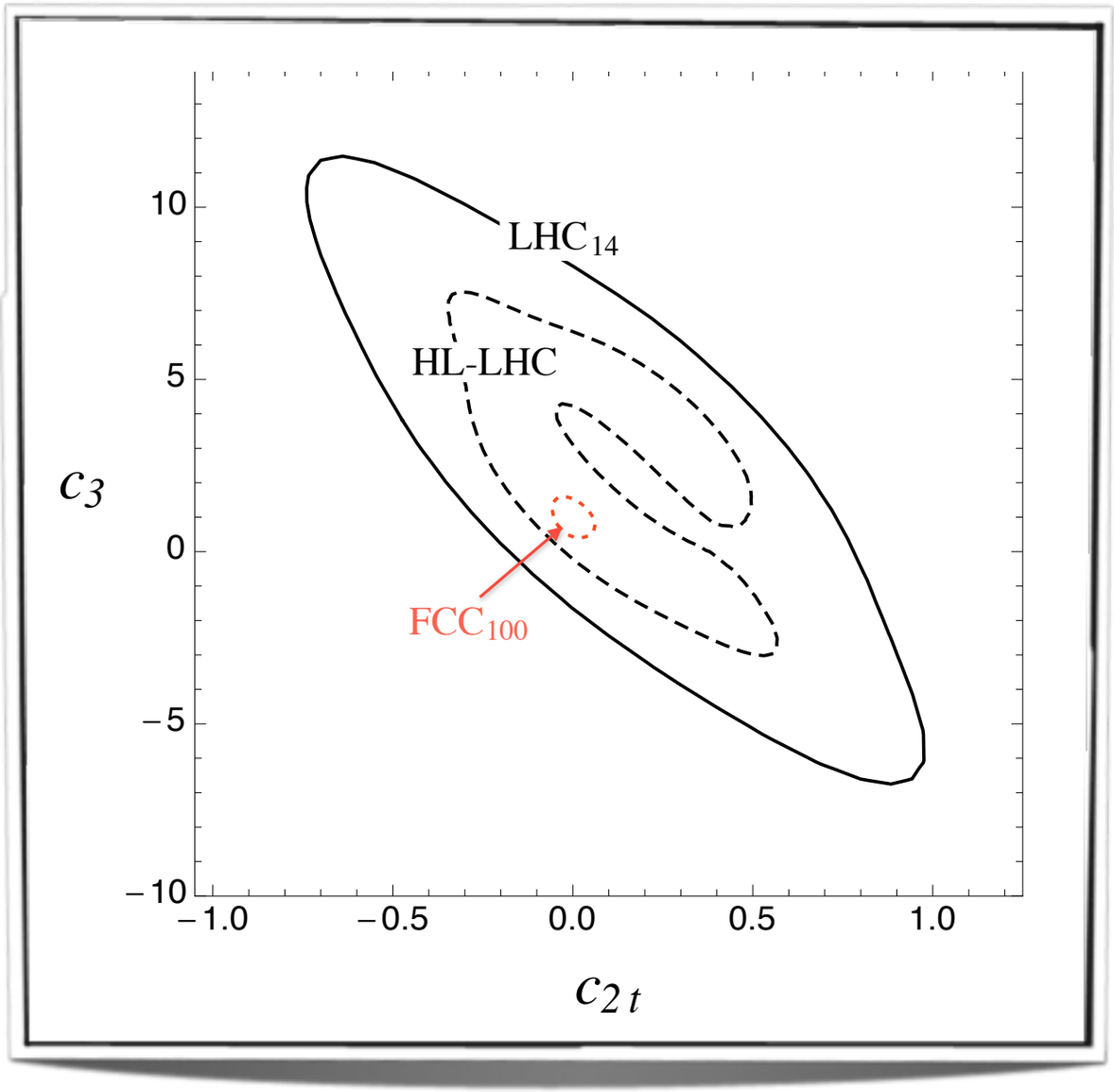


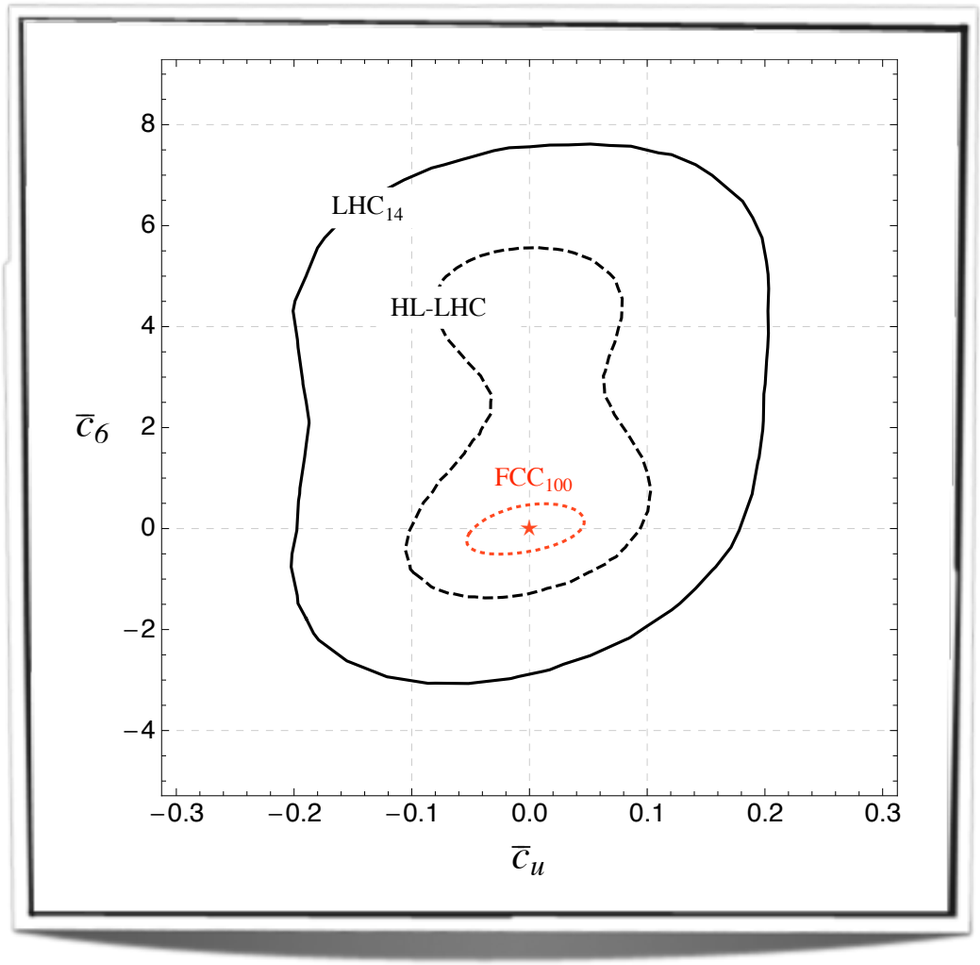
Signal (SM)



Azatov, Contino, Panico, Son '15

HH production as a probe of HE couplings





Azatov, Contino, Panico, Son '15
see also Goertz, Papaefstathiou, Yang, Zurita '14

Remarks:

- unique access to c_3 but sensitivity is limited (within the validity of EFT?).
- statistically limited, with more luminosity

  ⇛ access to distribution

  ⇛ discriminating power c_3 vs. c_{2t} vs c_g

Searching for more Higgses

Giudice, HL-LHC '15

Why ~~should~~ shouldn't new Higgses exist?

- Dynamical explanations of hierarchy $\Rightarrow$ enlarged Higgs sector
- Discovery of new Higgses $\Rightarrow$ mortal blow for anthropic explanation
- Relaxation mechanisms probably prefer single Higgs

GRAHAM
KAPLAN
RAJENDRAN

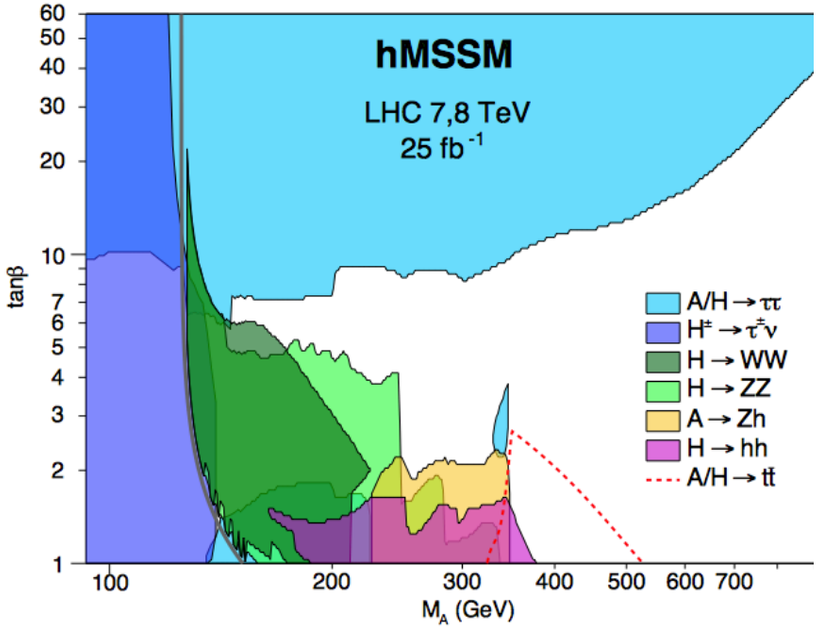
Discovery or exclusion of new Higgses at the EW scale

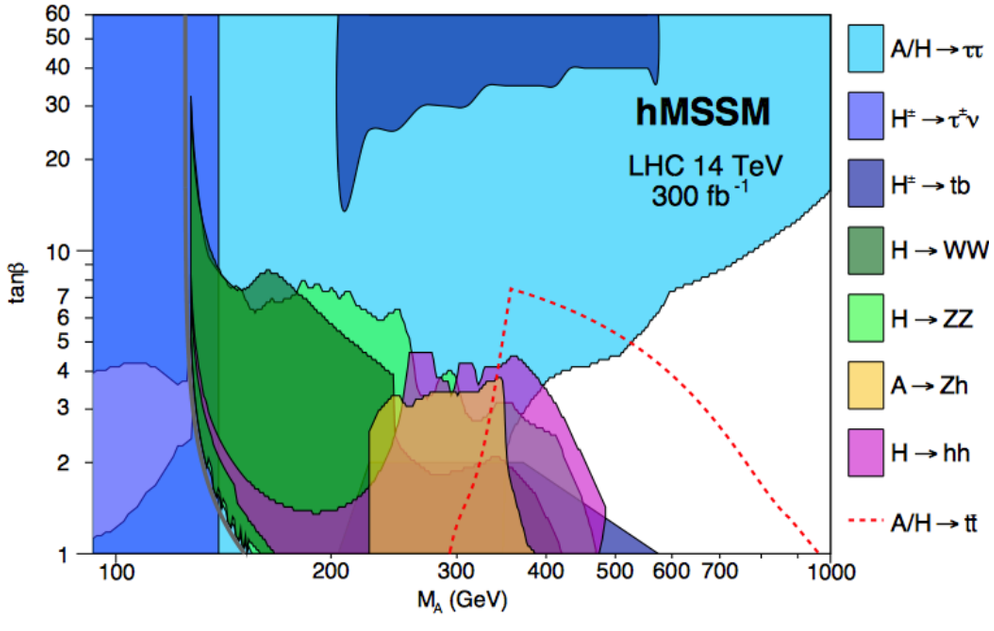
$\Rightarrow$

essential element to understand how nature dealt with naturalness

Searching for more Higgses

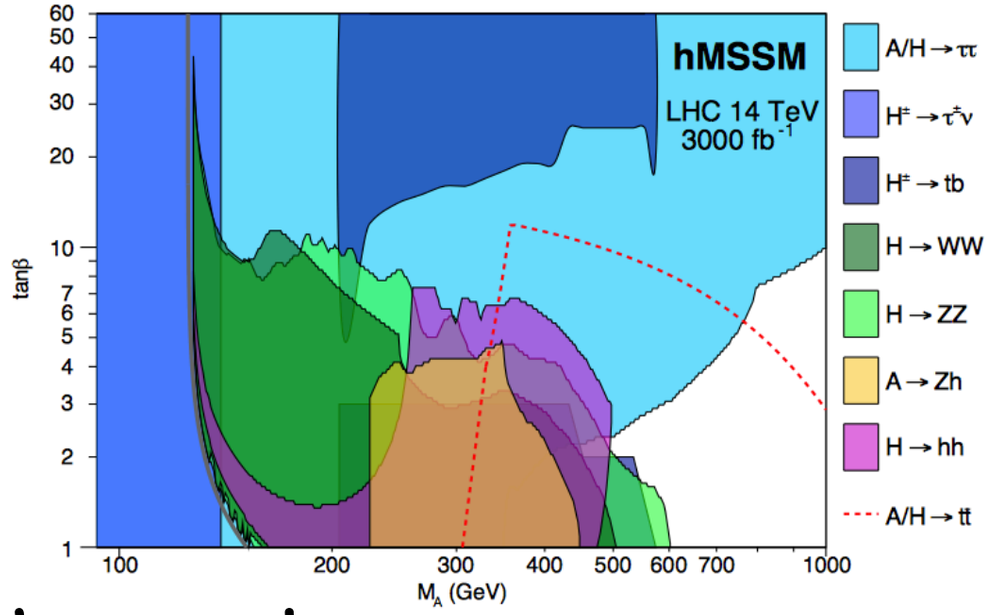
Djouadi et al '15





Direct searches are complementary to bounds from h(125) coupling measurements

(decoupling ⇒ alignment
but in general
alignment ~~⇏~~ decoupling)



Theory homework:

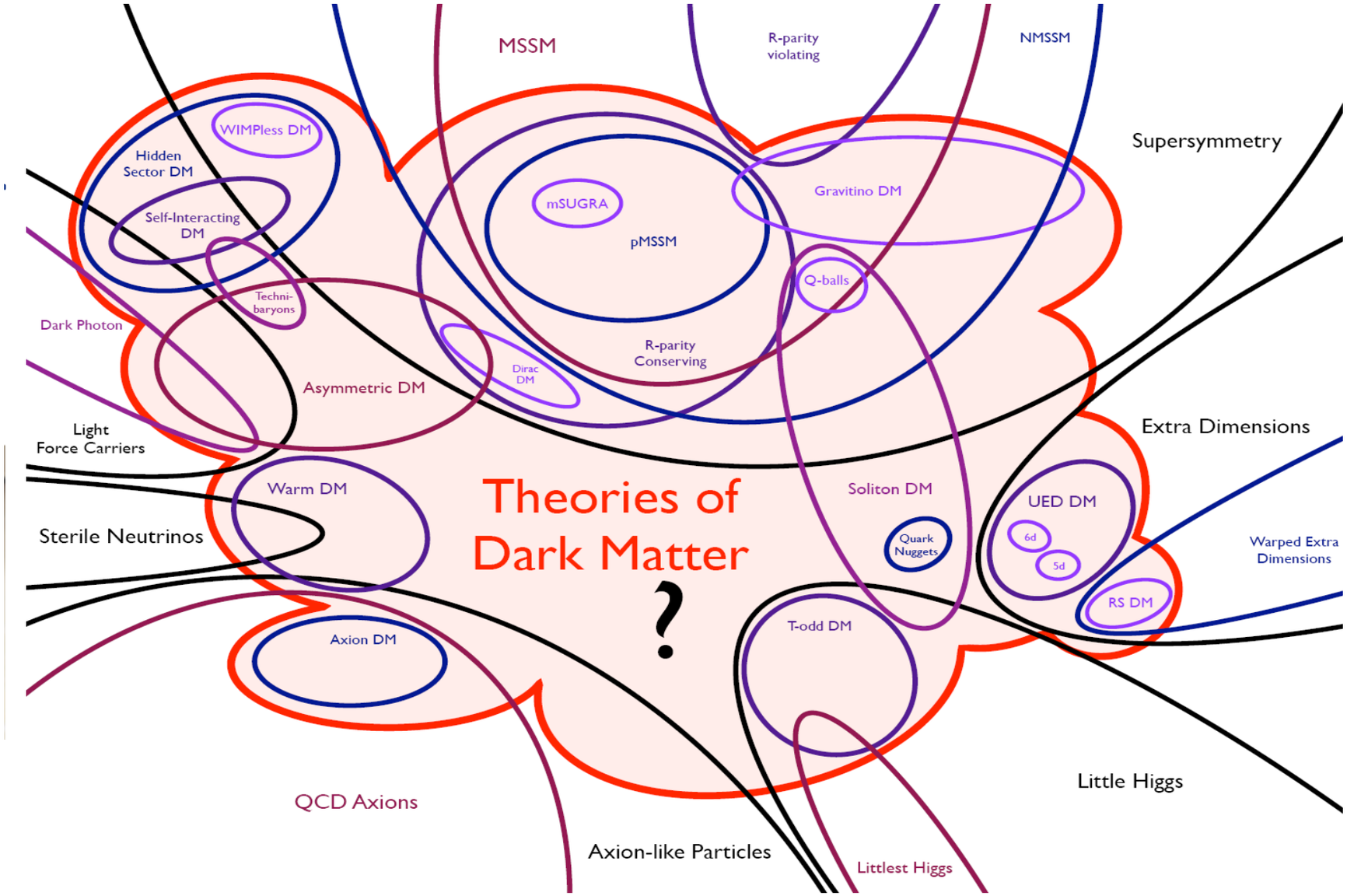
⊙ what is the expected mass scale of an extended Higgs sector?

⊙ which guiding principle (naturalness, signatures, anomalies...)
in the exploration of an extended Higgs sector (2HDM, singlet, triplet...)?

2

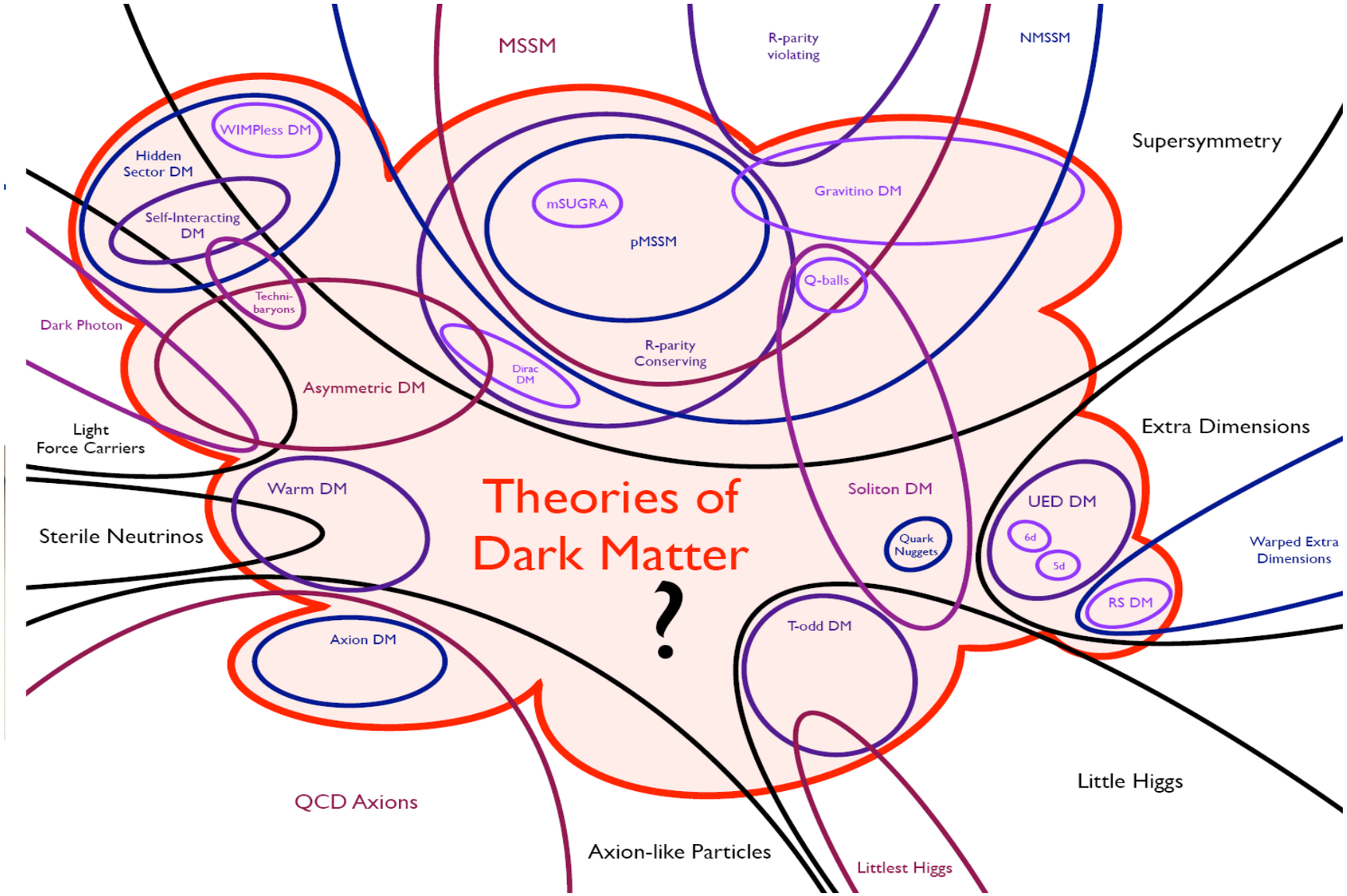
Dark Matter

The energy scale(s) of new physics



T. Tait, DM@LHC '14

The energy scale(s) of new physics



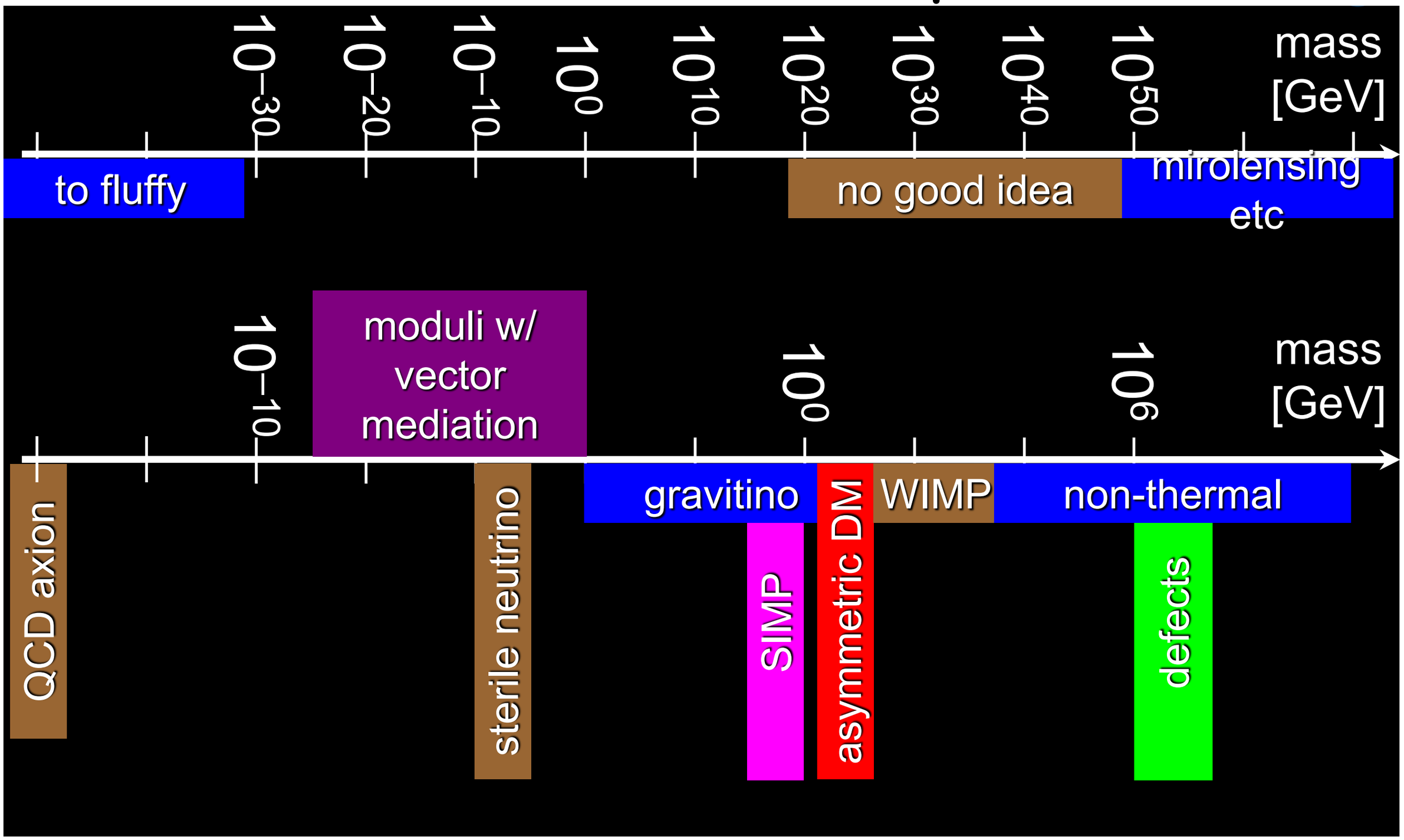
T. Tait, DM@LHC '14

The prediction about the mass scale of DM comes with large error bars:

$$10^{-22}\,\mathrm{eV} < m_{DM} < 10^{20}\,\mathrm{GeV}$$

(ALPs) (Wimpzillas, Q-balls)

The DM landscape



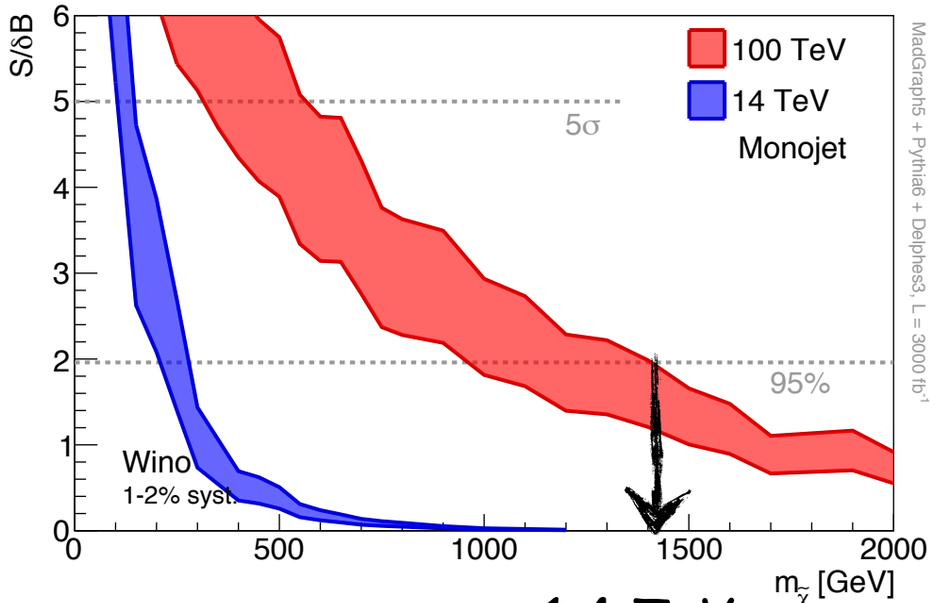
H. Murayama @ CERN '15

The WIMP hypothesis is particularly appealing (hierarchy, GUT, String...) and can be probed thoroughly @ colliders and in astroparticle experiments

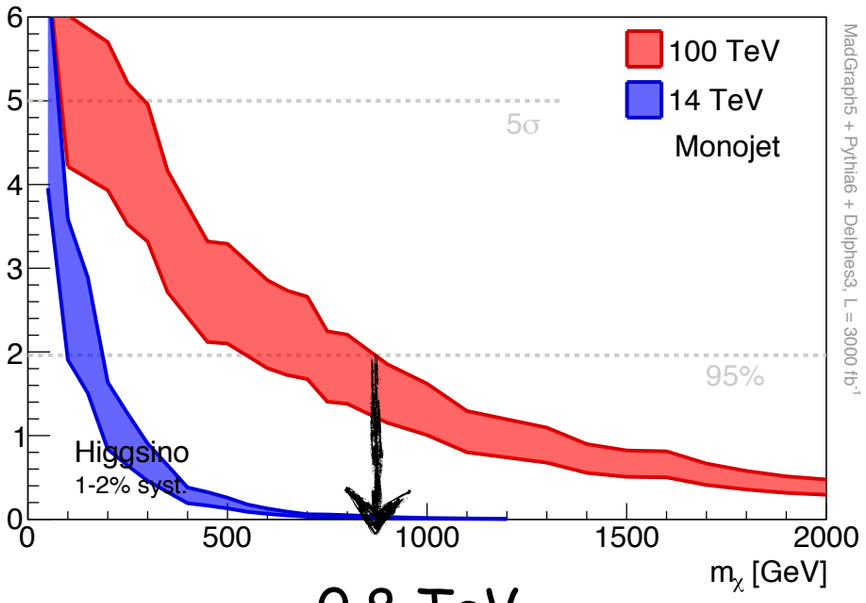
Exploring TeV-scale DM

- monojet searches
- soft lepton searches (compressed spectra)
- disappearing track searches (long chargino lifetime)

Low, Wang '14



1.4 TeV



0.8 TeV

$$M_{\rm WIMP} \leq 1.8 \text{ TeV} \left(\frac{g^2}{0.3}\right)$$ to avoid overclosing the Universe

- LHC only coverage very limited. Rate, systematics...
- 100 TeV pp collider can probe the "bulk" of WIMP parameter space.

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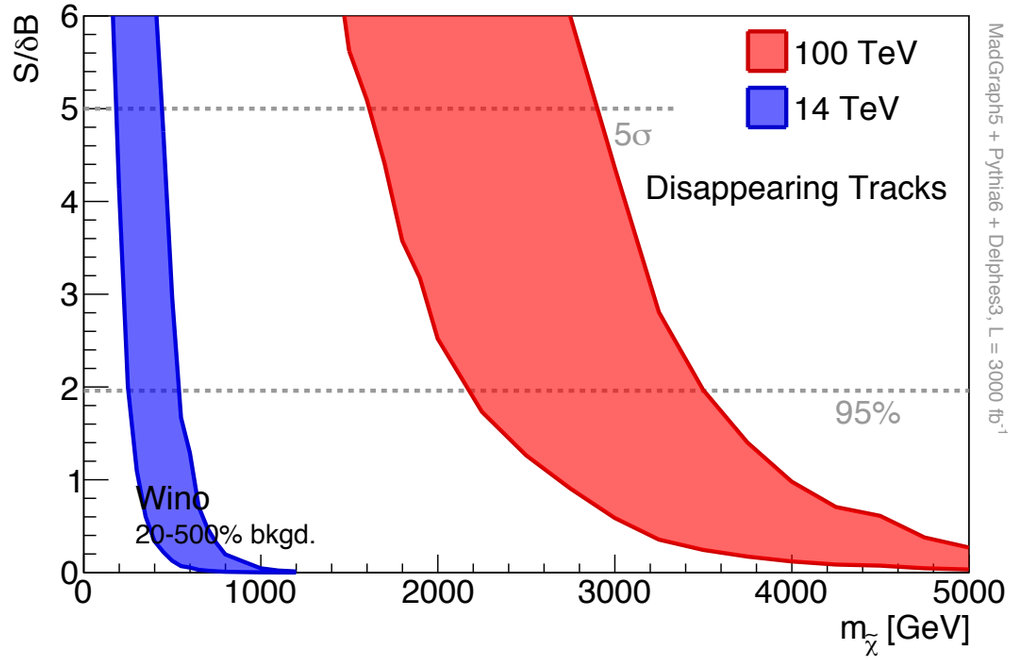

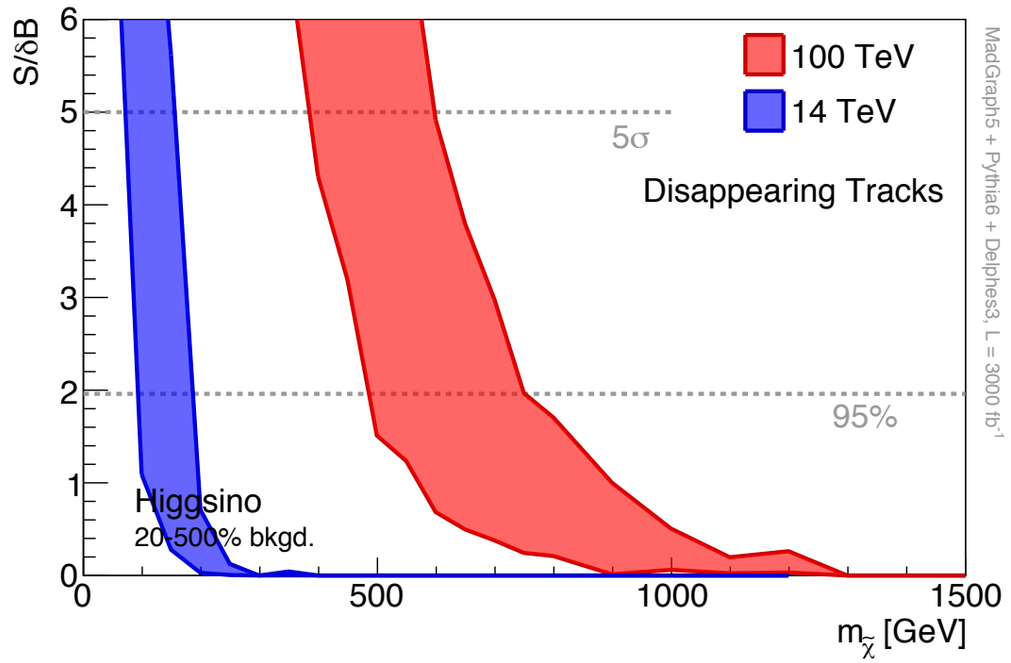

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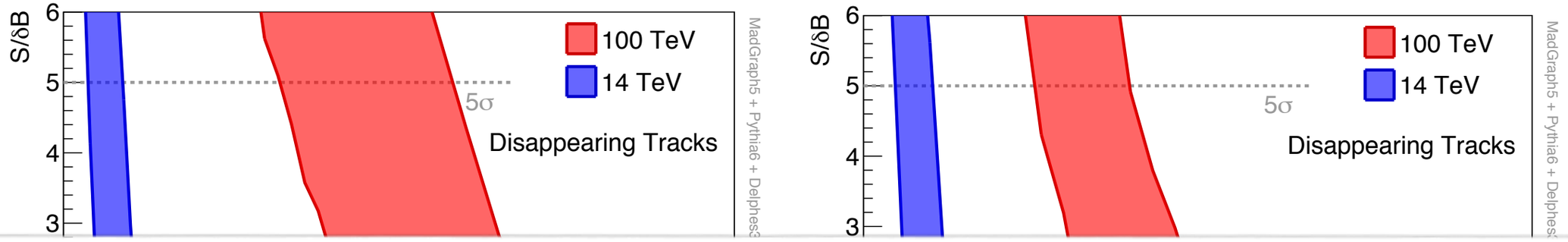


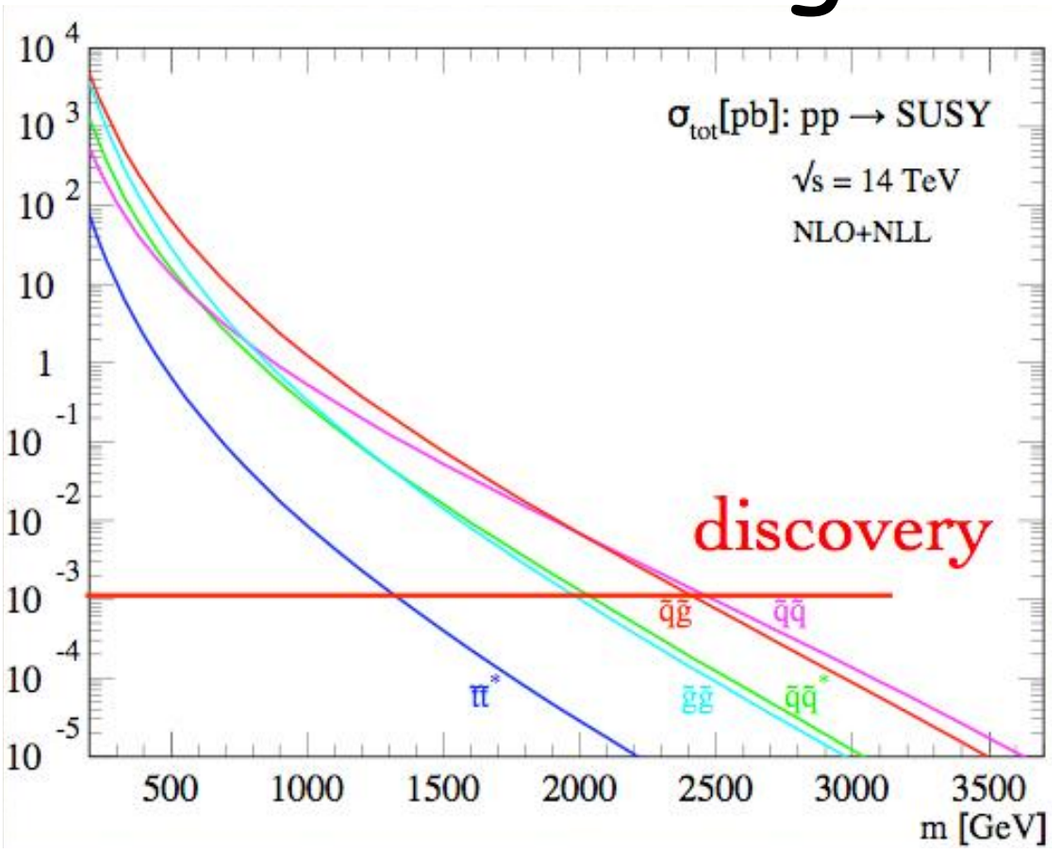
Once a DM particle is discovered
it remains to be seen if its relic abundance
can match the cosmological data
➠ needs to precisely "weight" DM

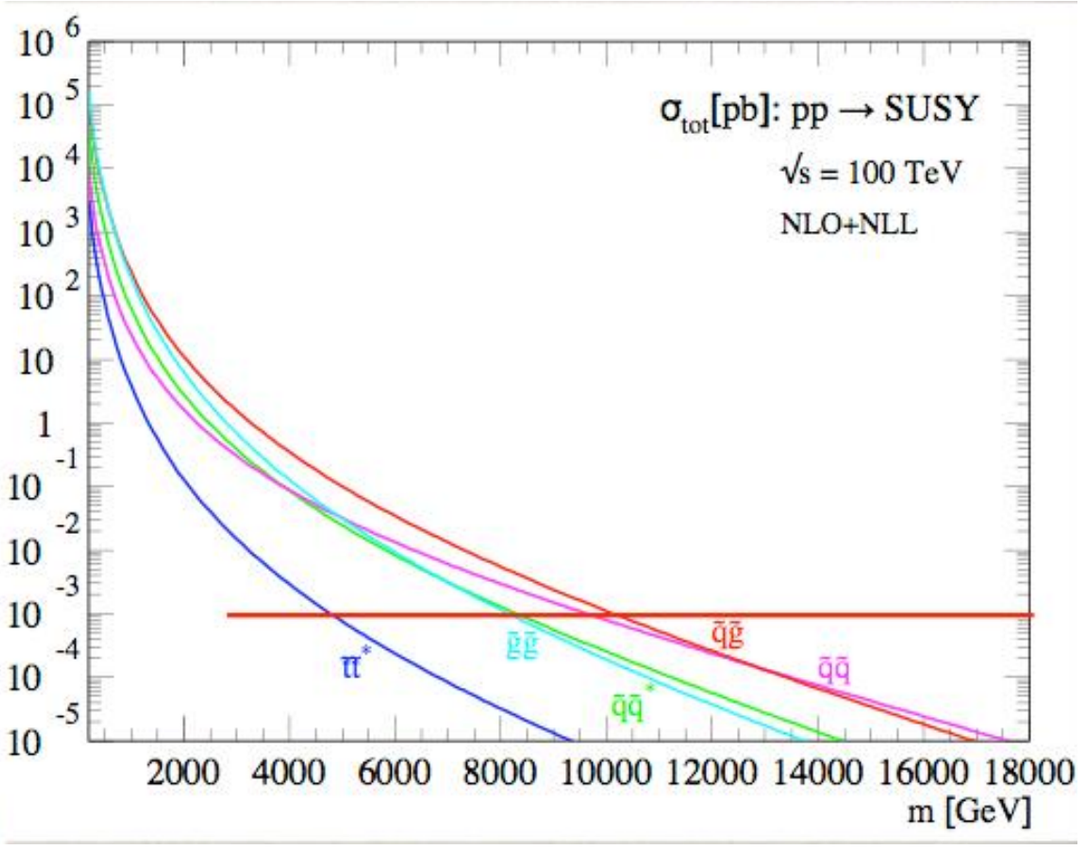
– 100 TeV pp collider can probe the "bulk" of WIMP parameter space.

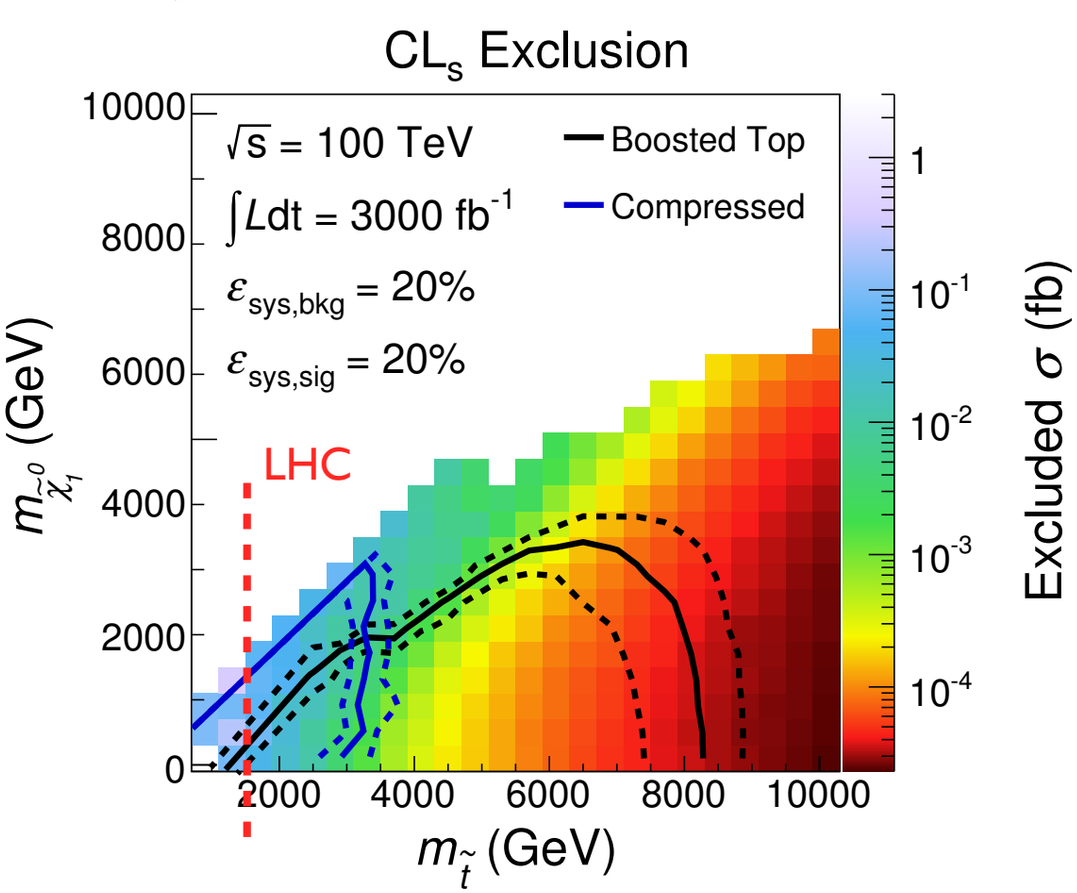
3

Searching for New Physics directly

Pushing the boundaries







Mangano, Plehn '14

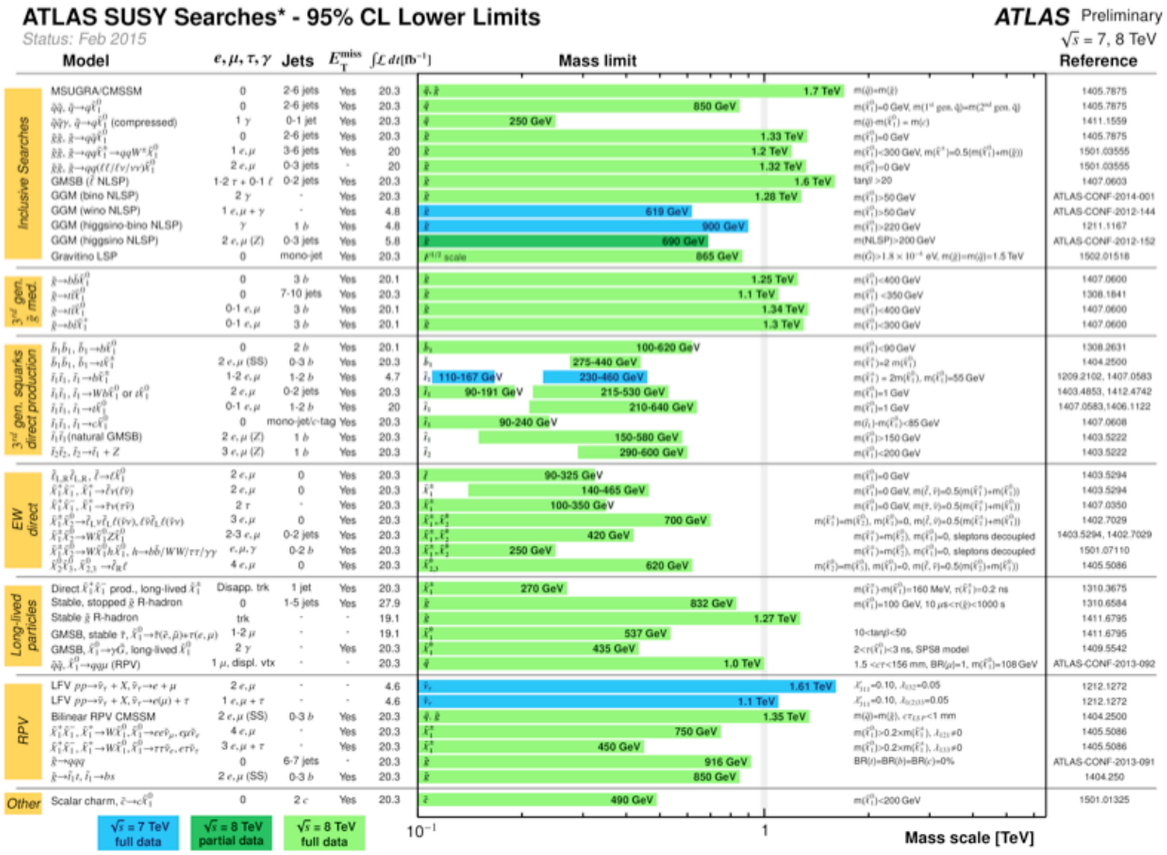
HL-LHC will explore O(2)TeV natural sparticles

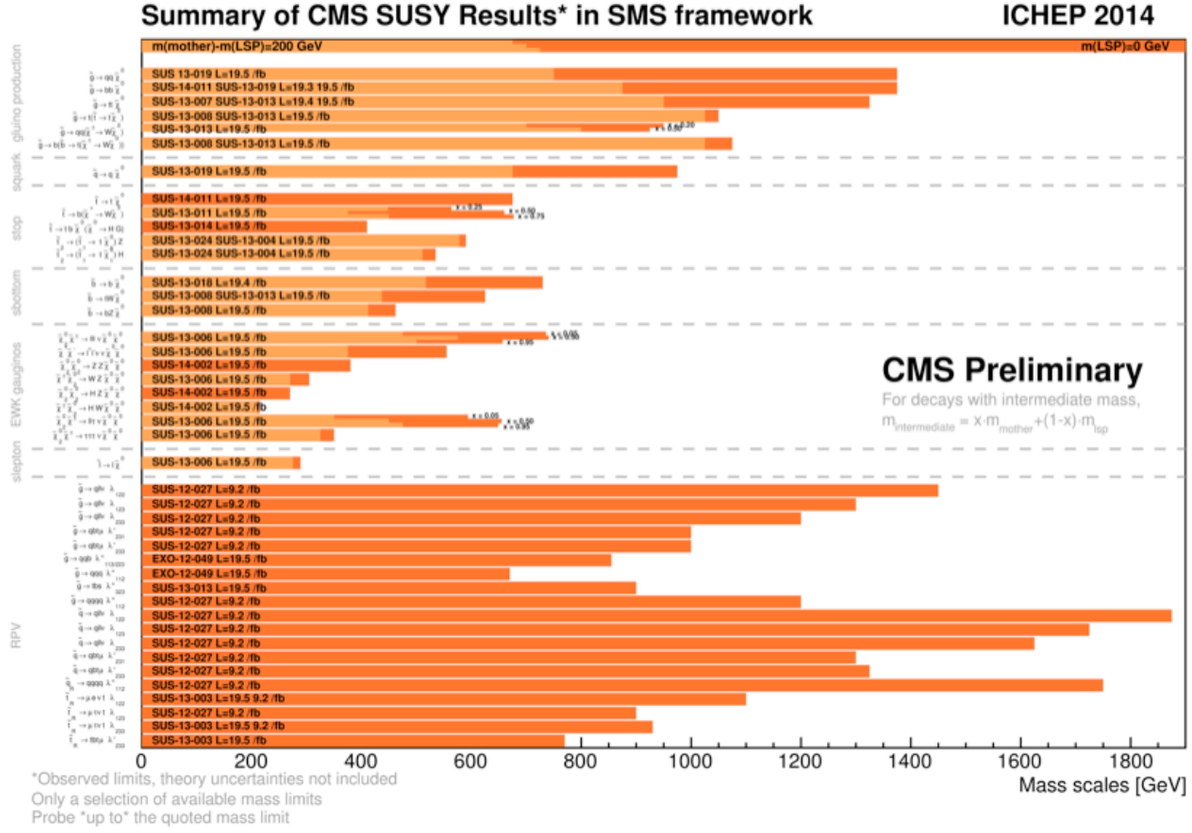
~~

FCC-hh can explore the O(10)TeV domain

Cohen, d'Agnolo, Hance, Lou, Wacker '14

Saving SUSY





No sign of super-partners @ LHC as of today
but most searches involve MET, stable charged particles

There are ways to hide low-scale superpartners

Saving SUSY

" SUSY anywhere is better than SUSY nowhere" J. Ellis' mantra

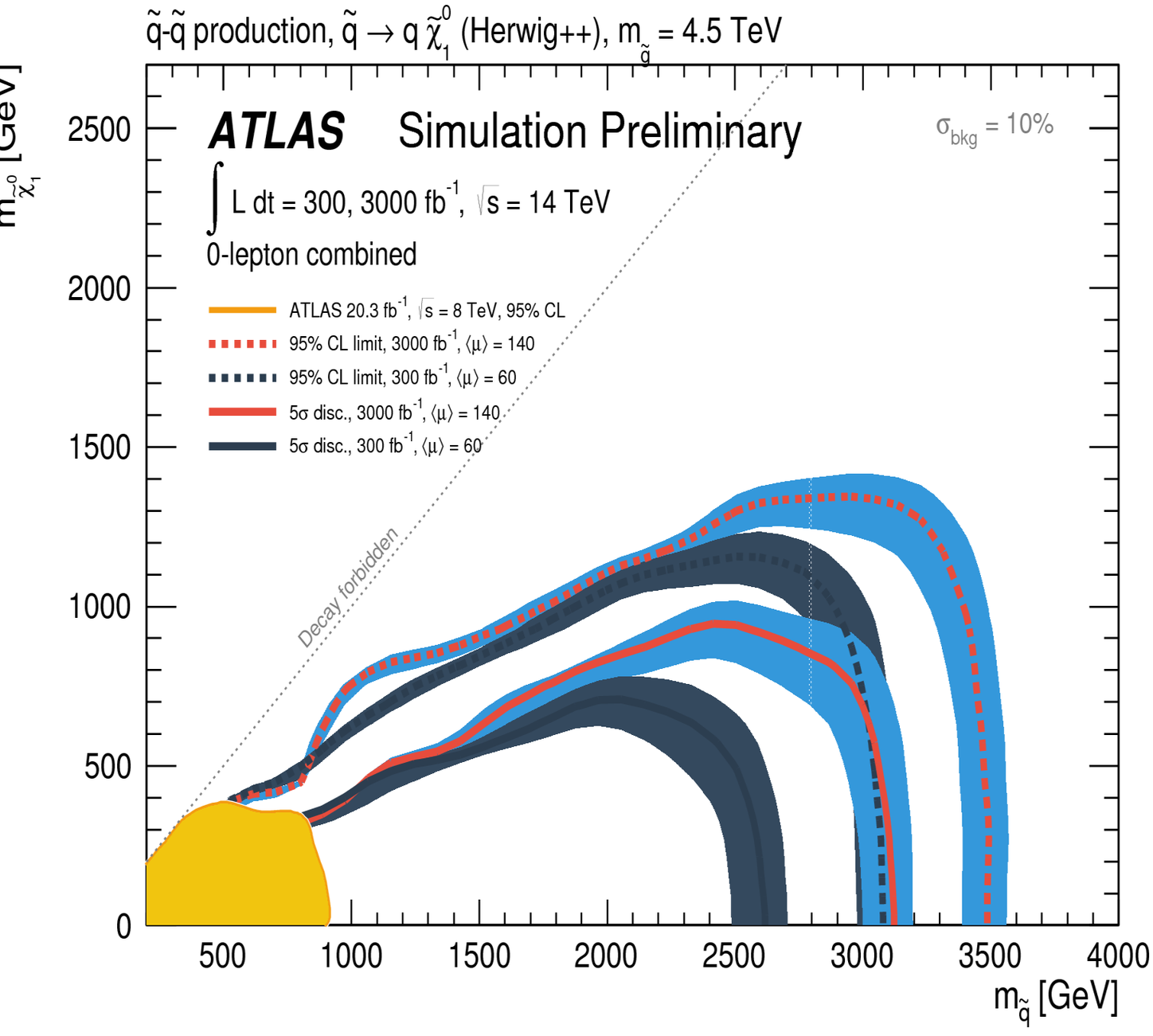
SUSY is Natural but not plain vanilla

- ~~CMSSM~~
- pMSSM
- NMSSM
- Hide SUSY, e.g. smaller phase space
  - reduce production (eg. split families) Mahbubani et al
  - reduce MET (e.g. ~~R-parity~~, compressed spectrum) Csaki et al
  - dilute MET (decay to invisible particles with more invisible particles)
  - soften MET (stealth susy, stop -top degeneracy) Fan et al
  - long-lived particles and displaced vertices

Good coverage of hidden natural susy

- mono-top searches (DM, flavored naturalness - mixing among different squark flavors-, stop-higgsino mixings)
- mono-jet searches with ISR recoil (compressed spectra)
- precise tt inclusive measurement+ spin correlations

Fully exclude SUSY @ weak scale?



HL-LHC can exclude squark up to 3.5 TeV

but there are holes

in particular for compressed spectra (that are particularly relevant for DM)

and/or for EWkinos.

ILC can complement and close these holes

Is there a loophole for TeV-scale new physics?

Naturalness & TeV scale new physics

Following the arguments of Wilson, 't Hooft (and others):
only small numbers associated to the breaking of a symmetry survive quantum corrections
(others are not necessarily theoretically inconsistent
but they require some conspiracy at different scales)

Field	Symmetry as $m \to 0$	Implication
Spin-1/2 $m\Psi\bar{\Psi}$	$\Psi \to e^{i\theta}\Psi$ $\bar{\Psi} \to e^{-i\theta}\bar{\Psi}$ (chiral symmetry)	$\delta m \propto m$ **Natural!**
Spin-1 $m^2 A_\mu A^\mu$	$A_\mu \to A_\mu + \partial_\mu \alpha$ (gauge invariance)	$\delta m \propto m$ **Natural!**

courtesy to N. Craig @ Blois '15

The Higgs mass in the SM doesn't break any (quantum*) symmetry

* it does break classical scale invariance, as the running of the gauge couplings does too!

Naturalness principle @ work

Following the arguments of Wilson, 't Hooft (and others):
only small numbers associated to the breaking of a symmetry survive quantum corrections
(others are not necessarily theoretically inconsistent
but they require some conspiracy at different scales)

Beautiful examples of naturalness to understand the need of "new" physics

see for instance Giudice '13 (and refs. therein) for a recent account

- the need of the positron to screen the electron self-energy: $\Lambda < m_e/\alpha_{\rm em}$
- the rho meson to cutoff the EM contribution to the charged pion mass: $\Lambda^2 < \delta m_\pi^2/\alpha_{\rm em}$
- the kaon mass difference regulated by the charm quark: $\Lambda^2 < \dfrac{\delta m_K}{m_K}\dfrac{6\pi^2}{G_F^2 f_K^2 \sin^2\theta_C}$
- the light Higgs boson to screen the EW corrections to gauge bosons self-energies
- ...
- new physics at the weak scale to cancel the UV sensitivity of the Higgs mass?

The Darwinian solution to the Hierarchy

Other origin of small/large numbers according to Weyl and Dirac:
hierarchies are induced/created by the time evolution/the age of the Universe

Graham, Kaplan, Rajendran '15

- Higgs mass-squared promoted to a field
- The field evolves in time in the early universe
- The mass-squared relaxes to a small negative value
- The electroweak symmetry breaking stops the time-dependence

Self-organized criticality

when the Higgs mass becomes negative, it
back-reacts and generates a potential barrier
that stops the evolution of the scanning field

see also Espinosa, Grojean, Panico, Pomarol, Pujolas, Servant '15

Higgs-axion cosmological relaxation

Graham, Kaplan, Rajendran '15

ϕ slowly rolling field (inflation provides friction) that scans the Higgs mass

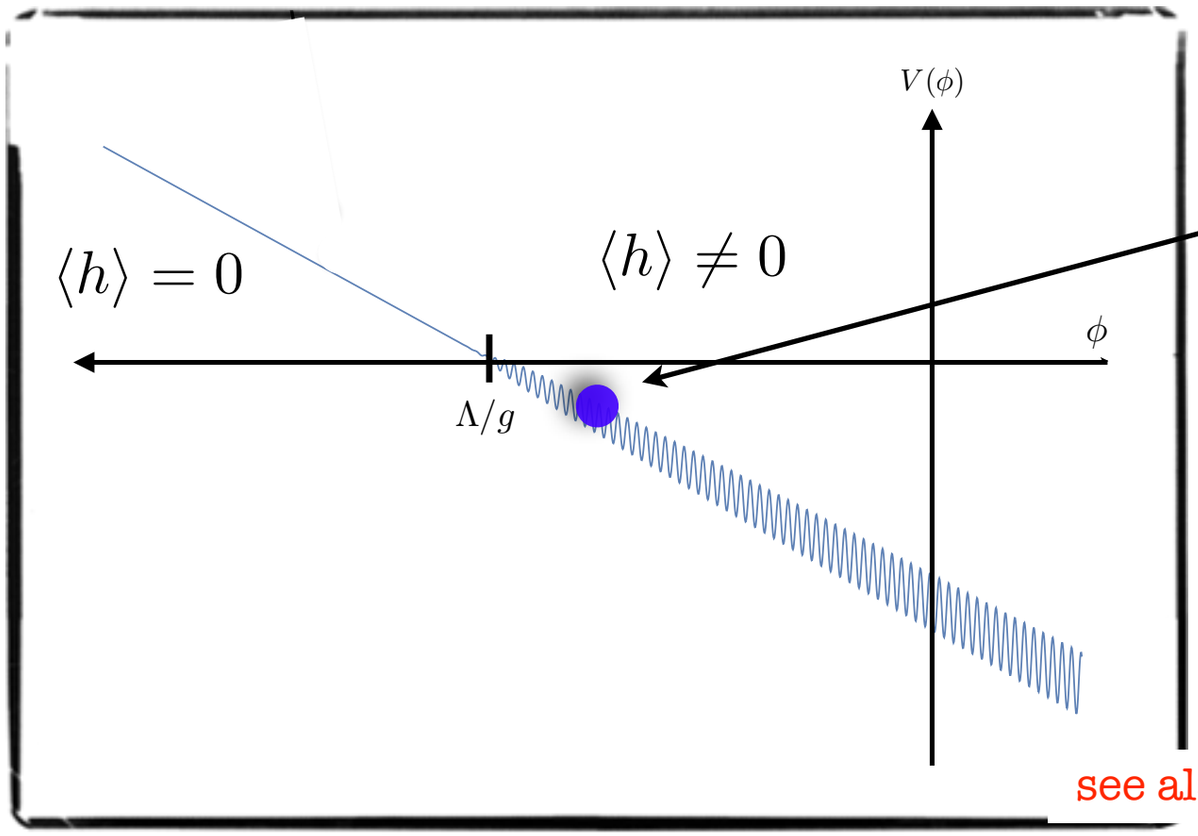
$$\Lambda^2\left(-1+f\left(\frac{g\phi}{\Lambda}\right)\right)|H|^2+\Lambda^4 V\left(\frac{g\phi}{\Lambda}\right)+\frac{1}{32\pi^2}\frac{\phi}{f}\tilde{G}^{\mu\nu}G_{\mu\nu}$$

Higgs mass depends on ϕ

potential needed to force ϕ to roll-down in time (during inflation)

axion-like coupling that will seed the potential barrier stopping the rolling when the Higgs develops its vev

$$\Lambda_{\rm QCD}^3\, h\, \cos\frac{\phi}{f}$$



see also Espinosa, Grojean, Panico, Pomarol, Pujolas, Servant '15

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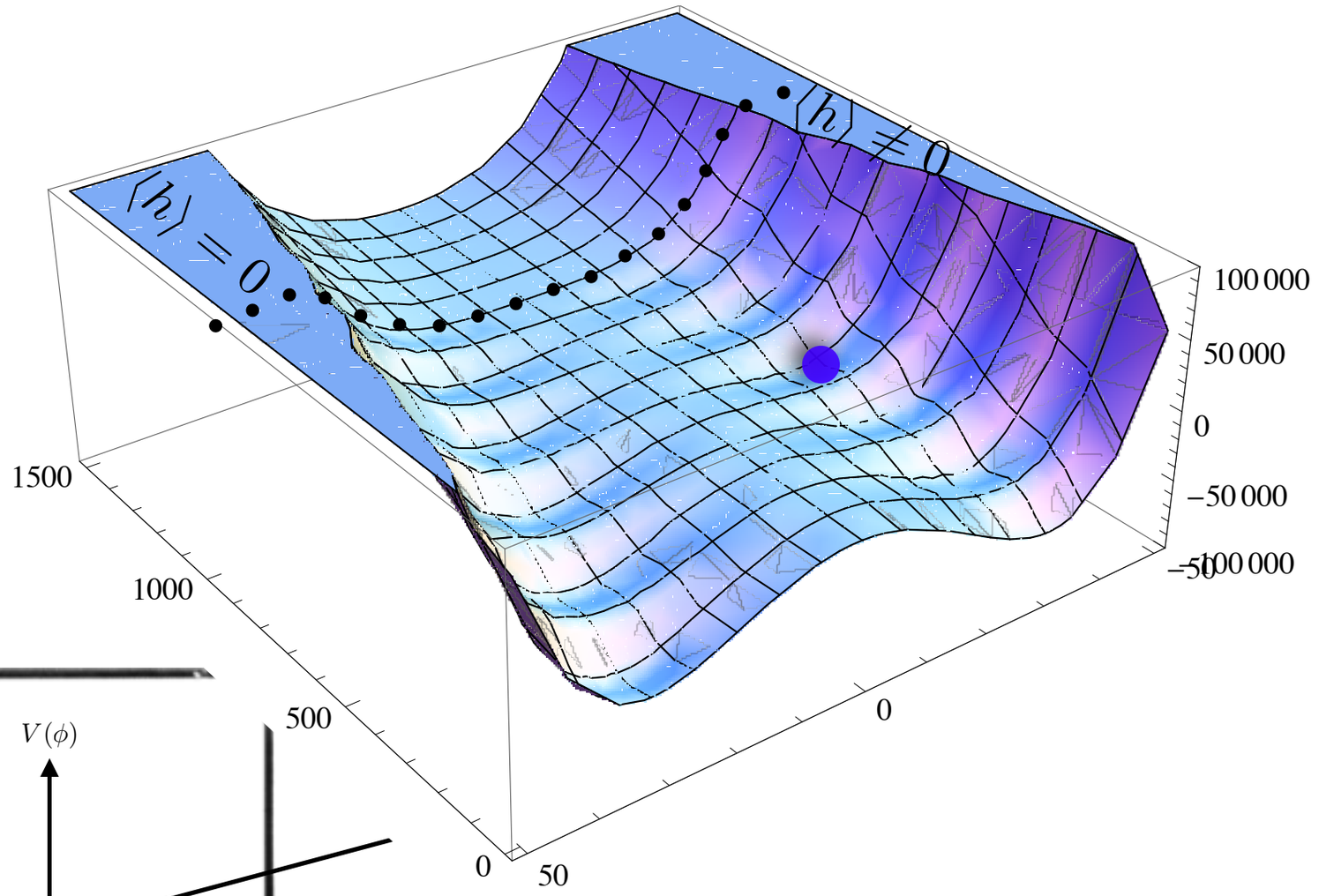
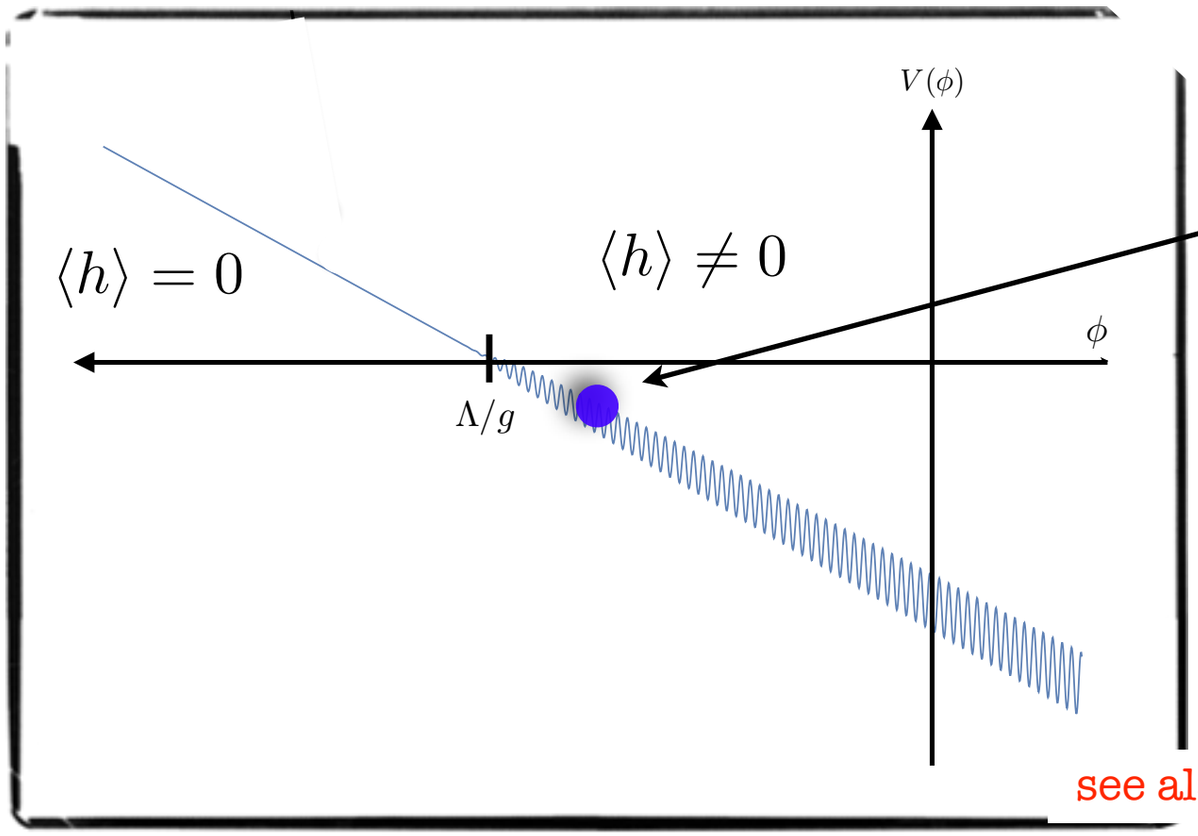
Graham, Kaplan, Rajendran '15

$$\phi$$

$$\Lambda^2\left(-1+f\left(\frac{g\phi}{\Lambda}\right)\right)$$

Higgs mass depends on ϕ





If ϕ continues rolling, the Higgs vev increases, the potential barrier increases and ultimately prevents ϕ from rolling down further

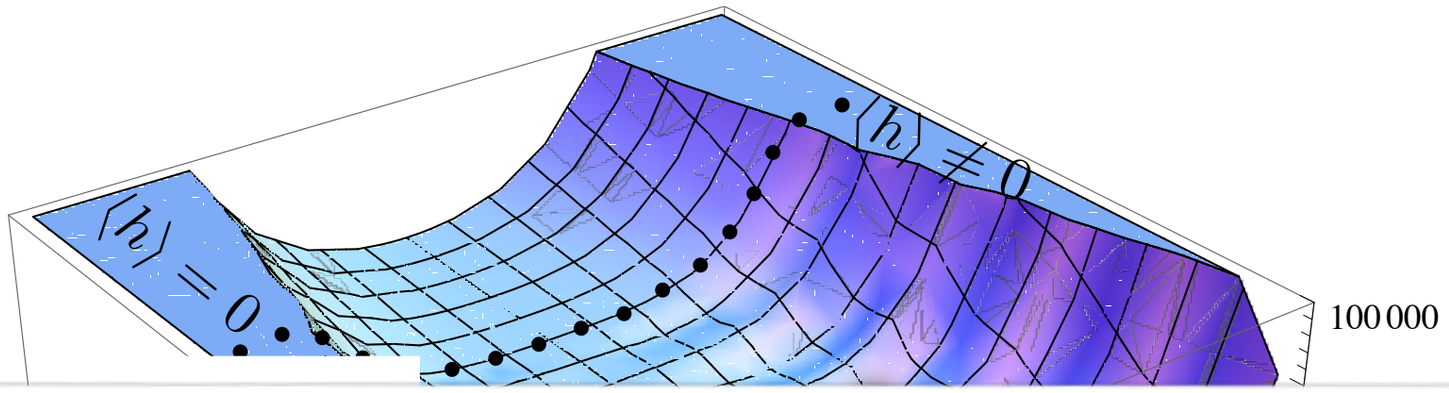
see also Espinosa, Grojean, Panico, Pomarol, Pujolas, Servant '15

Higgs-axion cosmological relaxation

Graham, Kaplan, Rajendran '15

$$\phi$$

$$\Lambda^2 \left(-1 + f\left(\frac{g\phi}{\Lambda}\right)\right)$$



Hierarchy problem solved
by light weakly coupled new physics
and not by TeV scale physics

~interesting cosmology signatures~

- BBN constraints
- decaying DM signs in γ-rays background
- ALPs
- superradiance

~interesting signatures @ SHiP~

- production of light scalars

by B and K decays

see also Espinosa, Grojean, Panico, Pomarol, Pujolas, Servant '15

Phenomenological signatures

Nothing to be discovered at the LHC/ILC/CLIC/CepC/SppC/FCC!

only BSM physics below Λ

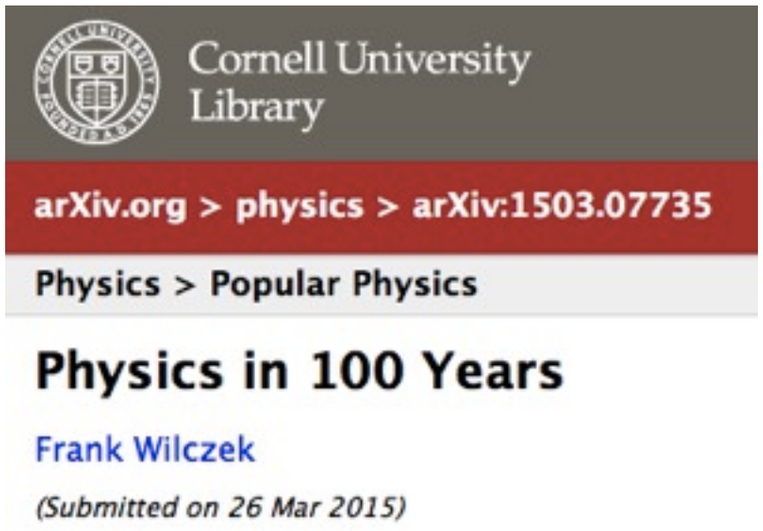
two (very) light and very weakly coupled axion-like scalar fields

$$m_\phi \sim (10^{-20} - 10^{2})\,\mathrm{GeV}$$

$$m_\sigma \sim (10^{-45} - 10^{-2})\,\mathrm{GeV}$$

interesting signatures in cosmology

Conclusions



- What are the weak points in our current understanding and practices?
- What are the growth areas in technique and capability?
- Where are the sweet spots where those two meet?

F. Gianoti EPS '15

More than ever: importance of the synergy and complementarity of the experimental programme

Main questions and main approaches to address them

	High-E colliders	Dedicated high-precision experiments	Neutrino experiments	Dedicated searches	Cosmic surveys
H, EWSB	x	x		x	
Neutrinos	x (ν_R)		x	x	x
Dark Matter	x			x	x
Flavour, CP, matter/antimatter	x	x	x	x	x
New particles, forces, symmetries	x	x		x	
Universe acceleration					x

Combination of these complementary approaches is crucial to explore the largest range of E scales (directly and indirectly) and couplings, and properly interpret signs of new physics → hopefully build a coherent picture of the underlying theory.