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Based on works realised in collaboration with K. Benakli, M. Goodsell and P. Slavich (1312.5220, 1508.02534 and 1511.02044)

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Introduction

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Hierarchy problem

- Need to keep all SUSY particles at EW scale to control the Higgs mass (if not, little hierarchy problem)
- ► Gauge coupling unification
- ▶ Dark Matter candidate (WIMP) ∫

Split SUSY idea \longrightarrow keep only the SUSY fermions (higgsinos and gauginos) at EW scale, but have all scalar superpartners heavy at SUSY scale M_S

 \blacktriangleright Can experimental measurements say something about M_S ?

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- \blacktriangleright M_S is still constrained by

The Higgs mass measurement (Section 1)

DM and cosmology (Section 2)

- \blacktriangleright In usual Split SUSY models, "Mini-Split" $\longrightarrow M_S \lesssim 10^6$ GeV
- ▶ In Fake Split SUSY models (FSSM) , "Mega-Split":

 $\blacksquare Higgs mass \longrightarrow No constraint$

Assuming standard Cosmology $\longrightarrow M_S \lesssim 10^{10} \text{ GeV}$

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How high could SUSY go ? └─ Recovering the Higgs mass

Recovering the Higgs mass

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How high could SUSY go ? Recovering the Higgs mass Split SUSY and Mini-Split

Split SUSY and Higgs mass



- ► A "prediction" for the SM-like Higgs mass $M_H^2 = 2\lambda v^2$ since:
 - experiment gives us
 SM-Higgs VEV v
 - boundary conditions at M_S gives us
 - $\lambda = \frac{1}{4} \left[g'^2 + g^2 \right] \cos^2 2\beta$ at tree level.
- Fine-tuning a light Higgs gives $\tan \beta = \sqrt{\frac{m_{H_d}^2 + \mu^2}{m_{H_u}^2 + \mu^2}} \approx \sqrt{\frac{m_{H_d}^2}{m_{H_u}^2}}$

• If
$$m_{H_d} = m_{H_u}$$
 at the GUT scale, then $\tan \beta \gtrsim 2-3$

Recovering the Higgs mass

Split SUSY and Mini-Split

Higgs mass prediction - Split SUSY



Higgs-mass function of M_S for Split SUSY for $\tan \beta = 1, 2$ or 20.

How high could SUSY go ? Recovering the Higgs mass FSSM and Mega-Split

Fake Split SUSY Models (FSSM)



- Replace higgsinos (and possibly gauginos) by F-higgsinos (and F-gauginos) with same quantum numbers but suppressed Yukawa couplings.
 - "Real" higgsinos (and gauginos) are heavy.
- Approximate global symmetry to protect the splitting.
 - A U(1)_F for both F-gauginos and F-higgsinos (FSSM-I)
 - An R symmetry for F-higgsinos only (FSSM-II)

Recovering the Higgs mass

FSSM and Mega-Split



Higgs-mass function of M_S for the FSSM. $\tan \beta = 1, 2$ or 20.

How high could SUSY go ? Recovering the Higgs mass FSSM and Mega-Split

Origin of the new behaviour

 Suppressed Yukawas for F-particles



• "Real" higgsinos are heavy: $m_{H_{d,u}} \sim \mu \sim \mathcal{O}(M_S)$: • $\tan \beta = \sqrt{\frac{m_{H_d}^2 + \mu^2}{m_H^2 + \mu^2}} \sim 1$



 $\tan\beta$ function of M_S and of the trilinear at $M_S.$

Running of λ in Split SUSY and FSSM. $M_S = 2 \times 10^{12}$ GeV.

FSSM and Mega-Split

Summary - Higgs mass



Higgs-mass function of M_S for the FSSM (tan $\beta = 1$)

and Split SUSY ($\tan \beta = 2$).

FSSM models: simple modification of the effective theory below M_S

- FSSM-I, SM, F-higgsinos and F-gauginos
- FSSM-II, SM, F-higgsino and gauginos
- In both cases, Higgs mass
 "prediction" drastically
 different than Split SUSY.

SUSY scale and cosmology

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How high could SUSY go ?
-SUSY scale and cosmology
└─(F-)gluinos life-time

- Long-lived gluinos are basically ruled out by observations (heavy isotopes searches, CMB, diffuse gamma ray background and BBN).
- ▶ In FSSM-II, gluinos are long-lived as Split SUSY (since decay through squarks are suppressed).

 In FSSM-I, F-gluinos are even more long-lived (couplings suppressed by the approximate symmetry which keeps them light).

$$\begin{array}{l} \bullet \quad \tau_{\tilde{g}'} \ \sim \ 4 \ {\rm sec} \times \left(\frac{M_S}{10^7 {\rm GeV}} \right)^6 \times \left(\frac{1 \ {\rm TeV}}{m_{fg}} \right)^7 \\ \hline & \quad {\rm Imply} \ M_S \lesssim 10^8 \ {\rm GeV} \end{array}$$



- Since FSSM have the same gauge interactions than Split SUSY, we expect rather similar DM candidates. Using micrOMEGAs, we find the correct relic density for:
 - F-Higgsino LSP with mass 1.1 TeV
 - (F-)Wino LSP with mass 2.4 TeV
- Interesting issue
 - F-higgsinos LSP are inelastic Dark Matter

- Splitting between F-higgsinos suppressed by the approximate symmetry which keeps them light.
 - Inelastic scattering over nucleons allowed → direct detection possible!
 - We took a conservative bound of 300 keV from LUX analysis
- F-higgsinos mass splitting is

EXAMPLE FSSM-I,
$$\delta \sim 200 \text{ keV} \cdot \left(\frac{400 \text{ TeV}}{M_S}\right)^2 \left(\frac{m_{fg}}{4 \text{ TeV}}\right)$$

EXAMPLE FSSM-II, $\delta \sim 200 \text{ keV} \cdot \left(\frac{10^7 \text{ GeV}}{M_S}\right) \left(\frac{\mu}{1 \text{ TeV}}\right) \left(\frac{4 \text{ TeV}}{m_{fg}}\right)$

- ▶ Therefore if F-higgsinos Dark Matter, M_S is bounded:
 - FSSM-I, $M_S \lesssim 10^6~{
 m GeV}$
 - $\blacksquare~{\rm FSSM-II},~M_S \lesssim 10^8~{\rm GeV}$

How high could SUSY go ? └─ Conclusions

Conclusions

- ► In FSSM, the measured Higgs mass is not a constraint anymore, stark contrast with usual Split SUSY models.
- Cosmology constraints are very relevant for this class of models, especially gluino life-time.
 - We still can have a "Mega-Split" spectrum with M_S up to 10^{10} GeV
- If new light fermions discovered at LHC, we need to study their Yukawas!
- Collider phenomenology ?

How high could SUSY go ? └─Backup slides

Backup slides

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Well-tempered Bino/Higgsinos



Relic density as a function of F-Higgsino pole mass and Bino pole mass.

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Gauge coupling unification



From 1508.02534



Tachyonic Higgs soft masses



RGE for Higgs sector soft masses for $M_S = 10^5 {\rm \ GeV}$ in type-I and type-II FSSM.

Higgs mass prediction - High scale SUSY



Higgs-mass function of M_S for High-scale SUSY, Split SUSY and FSSM. $M_{\tilde{g}'}=\mu=2$ TeV (when relevant) and $\tan\beta=1$ or 40.

Stability in FSSM



Contour plot of the Higgs mass on the $M_S - \tan \beta$ plane, $M_{\tilde{g}'} = \mu = 2$ TeV. Yellowshaded region indicates where λ becomes negative during its running.

Toward a microscopic description:

- ► Above M_S, add to MSSM chiral multiplets in the adjoint representation of each gauge group (Dirac gauginos) → fermions are called *F-gauginos*¹.
- ▶ Also two Higgs-like $SU(2)_W$ doublets \rightarrow fermions are called *F*-higgsinos
- ▶ Suppose no R-symmetry (\neq from Split-SUSY style) protects gauginos and higgsinos masses \rightarrow get masses at M_S

¹Also arXiv:1312.2011 by E. Dudas, M. Goodsell, L.=Heurgier, 🖓 Tziveloglou 🔗 ແ

Realisation of FSSM-I: Getting light sfermions

- ► Use an approximate U(1) symmetry with only the "fake" particles charged under it (Froggatt-Nielsen style).
- The gaugino $(\lambda)/F$ -gaugino (χ) mass terms are of the form.

$$-\Delta \mathcal{L}_{\text{gauginos}} = M_S \left[\frac{1}{2} \lambda \lambda + \mathcal{O}(\varepsilon) \lambda \chi + \mathcal{O}(\varepsilon^2) \chi \chi + \text{h.c.} \right],$$

leading to light F-gaugino-like eigenstate and heavy gauginos-like eigenstate.

- ► SUSY-breaking mass terms of the usual MSSM scalars, fake adjoint scalars and F-Higgs scalars are not protected → heavy.
- ▶ We need to fine-tune the weak scale.

► The SM Higgs boson H is a linear combination of the original Higgs H_u, H_d and F-Higgs H'_u, H'_d doublets.

 $H \approx \cos\beta i\sigma^2 H_d^* + \sin\beta H_u + \mathcal{O}(\varepsilon)i\sigma^2 H_u'^* + \mathcal{O}(\varepsilon)H_d'$

- We have F-higgsinos and F-gauginos at low energy instead of Higgsinos and gauginos:
- $\varepsilon \sim \sqrt{\frac{\text{TeV}}{M_S}}$ determined by requiring TeV-scale F-particles.
- Same particle content as Split SUSY, but suppressed Higgsino and gauginos couplings. Realisation of weakly-coupled FSSM.

- ▶ We want the RG evolution of λ between M_Z and the GUT scale.
- ▶ Make use of perturbative situation: algorithmic procedure
 - 1. Start with SM values and crude approximations at M_Z , evolve it up to GUT scale, fix unified input there.
 - 2. Run it down to M_S , apply FSSM boundary conditions
 - 3. Iterate this procedure above M_S until CV (very fast)
 - 4. Run it down to M_Z , apply again SM boundary conditions
 - 5. Iterate the procedure below M_S until CV (very fast).
 - 6. Iterate the whole thing until CV (again very fast)
 - 7. Calculate Higgs mass at M_{top}

▶ Concerning our precision, the most salient points are:

- Two-loop QCD contributions while converting top pole mass to its MS counterpart.
- Full Split-SUSY two-loop RGEs.
- Dominant two-loop Higgs self-energy to obtain the pole mass, we also include leading-log contribution from three-loop diagrams with (F-)gluinos.
- We did not included threshold corrections at M_S but have estimated them as a GeV effect for $M_S \gtrsim 10^8$ GeV.