LFV in B decays

Diego Guadagnoli LAPTh Annecy (France)

LFV in B decays

Diego Guadagnoli LAPTh Annecy (France)



Main line of argument based on Glashow, DG, Lane, PRL 2015

LHCb's $b \rightarrow s$ data

Renewed interest in B-decay LFV is motivated by the following pieces of exp info (LHCb):

LHCb's $b \rightarrow s$ data

Renewed interest in B-decay LFV is motivated by the following pieces of exp info (LHCb):

$$R_K = \frac{BR(B^+ \to K^+ \mu \mu)_{[1,6]}}{BR(B^+ \to K^+ e \, e)_{[1,6]}} = 0.745 \cdot (1 \pm 13 \%)$$
 whereas the SM predicts unity within any foreseeable exp accuracy

LHCb's $b \rightarrow s$ data

Renewed interest in B-decay LFV is motivated by the following pieces of exp info (LHCb):

$$R_K = \frac{BR(B^+ \to K^+ \mu \mu)_{[1,6]}}{BR(B^+ \to K^+ e \, e)_{[1,6]}} = 0.745 \cdot (1 \pm 13\%)$$
 whereas the SM predicts unity within any foreseeable exp accuracy

$$BR(B^+ \to K^+ \mu \mu)_{[1,6]} = (1.19 \pm 0.07) \cdot 10^{-7}$$

VS.

$$BR(B^+ \to K^+ \mu \mu)_{[1,6]}^{SM} = 1.75_{-0.29}^{+0.60} \times 10^{-7}$$

[Bobeth, Hiller, van Dick (2012)]

LHCb's $b \rightarrow s$ data

Renewed interest in B-decay LFV is motivated by the following pieces of exp info (LHCb):

$$R_K = \frac{BR(B^+ \to K^+ \mu \mu)_{[1,6]}}{BR(B^+ \to K^+ e e)_{[1,6]}} = 0.745 \cdot (1 \pm 13\%)$$

whereas the SM predicts unity within any foreseeable exp accuracy

$$BR(B^+ \to K^+ \mu \mu)_{[1,6]} = (1.19 \pm 0.07) \cdot 10^{-7}$$

VS.

$$BR(B^+ \to K^+ \mu \mu)_{[1,6]}^{SM} = 1.75_{-0.29}^{+0.60} \times 10^{-7}$$

[Bobeth, Hiller, van Dick (2012)]

BR
$$(B^+ \rightarrow K^+ e e)_{[1,6]}$$
 agrees with the SM (within large errors)

LHCb's $b \rightarrow s$ data

Renewed interest in B-decay LFV is motivated by the following pieces of exp info (LHCb):

$$R_K = \frac{BR(B^+ \to K^+ \mu \mu)_{[1,6]}}{BR(B^+ \to K^+ e e)_{[1,6]}} = 0.745 \cdot (1 \pm 13\%)$$

whereas the SM predicts unity within any foreseeable exp accuracy

$$BR(B^+ \to K^+ \mu \mu)_{[1,6]} = (1.19 \pm 0.07) \cdot 10^{-7}$$

VS.

$$BR(B^+ \to K^+ \mu \mu)_{[1,6]}^{SM} = 1.75^{+0.60}_{-0.29} \times 10^{-7}$$

[Bobeth, Hiller, van Dick (2012)]

BR
$$(B^+ \rightarrow K^+ e e)_{[1,6]}$$
 agrees with the SM (within large errors)

Note

 muons are among the most reliable objects within LHCb

LHCb's $b \rightarrow s$ data

Renewed interest in B-decay LFV is motivated by the following pieces of exp info (LHCb):

$$R_K = \frac{BR(B^+ \to K^+ \mu \mu)_{[1,6]}}{BR(B^+ \to K^+ e e)_{[1,6]}} = 0.745 \cdot (1 \pm 13\%)$$

whereas the SM predicts unity within any foreseeable exp accuracy

$$BR(B^+ \to K^+ \mu \mu)_{[1,6]} = (1.19 \pm 0.07) \cdot 10^{-7}$$

VS.

$$BR(B^+ \to K^+ \mu \mu)_{[1,6]}^{SM} = 1.75_{-0.29}^{+0.60} \times 10^{-7}$$

[Bobeth, Hiller, van Dick (2012)]

BR
$$(B^+ \rightarrow K^+ e e)_{[1,6]}$$
 agrees with the SM (within large errors)

Note

- muons are among the most reliable objects within LHCb
- the electron channel would be an obvious culprit (brems + low stats).
 But there is no disagreement

LHCb's $b \rightarrow s$ data

Renewed interest in B-decay LFV is motivated by the following pieces of exp info (LHCb):

$$R_K = \frac{BR(B^+ \to K^+ \mu \mu)_{[1,6]}}{BR(B^+ \to K^+ e e)_{[1,6]}} = 0.745 \cdot (1 \pm 13\%)$$

whereas the SM predicts unity within any foreseeable exp accuracy

$$BR(B^+ \to K^+ \mu \mu)_{[1,6]} = (1.19 \pm 0.07) \cdot 10^{-7}$$

VS.

$$BR(B^+ \to K^+ \mu \mu)_{[1,6]}^{SM} = 1.75_{-0.29}^{+0.60} \times 10^{-7}$$

[Bobeth, Hiller, van Dick (2012)]

BR
$$(B^+ \rightarrow K^+ e e)_{[1,6]}$$
 agrees with the SM (within large errors)

Note

- muons are among the most reliable objects within LHCb
- the electron channel would be an obvious culprit (brems + low stats).
 But there is no disagreement

$$\mathbf{0} + \mathbf{2} + \mathbf{6}$$
 \Rightarrow There seems to be BSM LFNU and the effect is in $\mu\mu$, not ee

Actually, after some effective-theory insights, two further pieces of info support the above picture

4 P'_{5} deficit in angular $B \rightarrow K^{*} \mu \mu$ data

it occurs also in the low-q² range

Actually, after some effective-theory insights, two further pieces of info support the above picture

 P'_{5} deficit in angular $B \rightarrow K^{*} \mu \mu$ data

it occurs also in the low-q² range

$$\frac{BR(B_s \rightarrow \mu \mu)_{\text{exp}}}{BR(B_s \rightarrow \mu \mu)_{\text{SM}}} = 0.77 \pm 0.20$$

Actually, after some effective-theory insights, two further pieces of info support the above picture

4 P'_{5} deficit in angular $B \rightarrow K^{*} \mu \mu$ data

it occurs also in the low-q² range

 $\frac{BR(B_s \rightarrow \mu \mu)_{\text{exp}}}{BR(B_s \rightarrow \mu \mu)_{\text{SM}}} = 0.77 \pm 0.20$

- Each of the above points, taken singly, is at best a 3σ effect
 - ⇒ Early to get excited

Actually, after some effective-theory insights, two further pieces of info support the above picture

4 P'_{5} deficit in angular $B \rightarrow K^* \mu \mu$ data

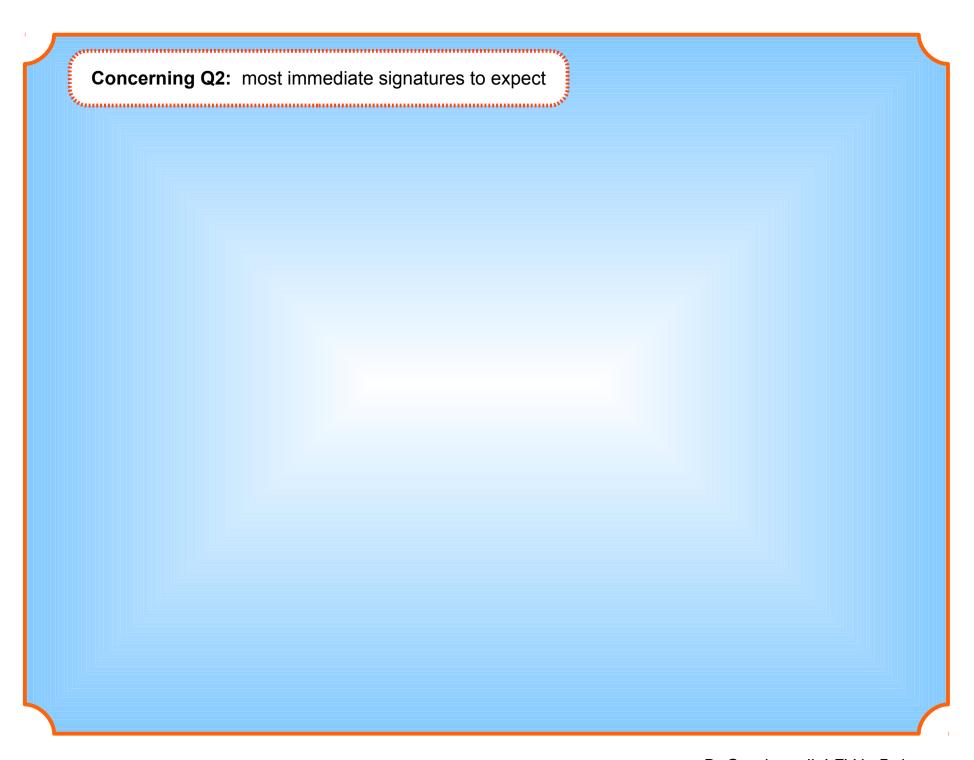
it occurs also in the low-q² range

 $\frac{BR(B_s \rightarrow \mu \mu)_{\text{exp}}}{BR(B_s \rightarrow \mu \mu)_{\text{SM}}} = 0.77 \pm 0.20$

• Each of the above points, taken singly, is at best a 3σ effect

⇒ Early to get excited

- Yet: Q1: Can we (easily) make sense of **1** to **5** ?
 - **Q2:** What are the most immediate signatures to expect?



Basic observation:

• If R_{κ} is signaling LFNU at a non-SM level, we may also expect LFV at a non-SM level.

Basic observation:

• If R_{κ} is signaling LFNU at a non-SM level, we may also expect LFV at a non-SM level.

In fact:

Consider a new, LFNU interaction above the EWSB scale, e.g. with

new vector bosons: $\overline{\ell} \, \mathbf{Z}' \ell$ or leptoquarks: $\overline{\ell} \, \phi \, \mathbf{q}$

Basic observation:

• If R_{κ} is signaling LFNU at a non-SM level, we may also expect LFV at a non-SM level.

In fact:

Consider a new, LFNU interaction above the EWSB scale, e.g. with

new vector bosons: $\overline{\ell} \, \mathbf{Z'} \ell$ or leptoquarks: $\overline{\ell} \, \phi \, \mathbf{q}$

In what basis are quarks and leptons in the above interaction?
 Generically, it's not the mass eigenbasis.
 (This basis doesn't yet even exist. We are above the EWSB scale.)

Basic observation:

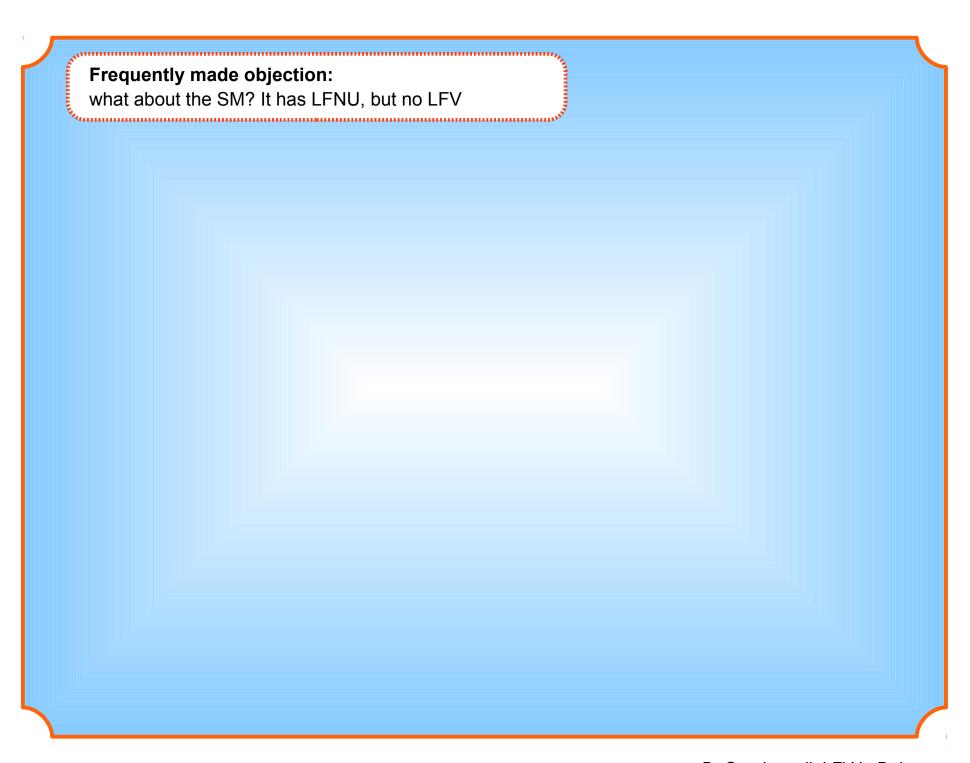
• If R_{κ} is signaling LFNU at a non-SM level, we may also expect LFV at a non-SM level.

In fact:

Consider a new, LFNU interaction above the EWSB scale, e.g. with

new vector bosons: $\overline{\ell} \, \mathbf{Z'} \ell$ or leptoquarks: $\overline{\ell} \, \phi \, \mathbf{q}$

- In what basis are quarks and leptons in the above interaction?
 Generically, it's not the mass eigenbasis.
 (This basis doesn't yet even exist. We are above the EWSB scale.)
- Rotating q and ℓ to the mass eigenbasis generates LFV interactions.



what about the SM? It has LFNU, but no LFV

Take the SM with zero v masses.

 Charged-lepton Yukawa couplings are LFNU, but they are diagonal in the mass eigenbasis (hence no LFV)

what about the SM? It has LFNU, but no LFV

Take the SM with zero v masses.

 Charged-lepton Yukawa couplings are LFNU, but they are diagonal in the mass eigenbasis (hence no LFV)

Or more generally, take the SM plus a minimal mechanism for v masses.

• Physical LFV will appear in W couplings, but it's suppressed by powers of $(m_v/m_w)^2$

what about the SM? It has LFNU, but no LFV

Take the SM with zero v masses.

 Charged-lepton Yukawa couplings are LFNU, but they are diagonal in the mass eigenbasis (hence no LFV)

Or more generally, take the SM plus a minimal mechanism for v masses.

• Physical LFV will appear in W couplings, but it's suppressed by powers of $(m_v/m_w)^2$

Bottom line: in the SM+v there is LFNU, but LFV is nowhere to be seen (in decays)

what about the SM? It has LFNU, but no LFV

Take the SM with zero v masses.

 Charged-lepton Yukawa couplings are LFNU, but they are diagonal in the mass eigenbasis (hence no LFV)

Or more generally, take the SM plus a minimal mechanism for v masses.

• Physical LFV will appear in W couplings, but it's suppressed by powers of $(m_v/m_w)^2$

Bottom line: in the SM+v there is LFNU, but LFV is nowhere to be seen (in decays)

 But nobody ordered that the reason (=tiny m₁) behind the above conclusion be at work also beyond the SM

what about the SM? It has LFNU, but no LFV

Take the SM with zero v masses.

 Charged-lepton Yukawa couplings are LFNU, but they are diagonal in the mass eigenbasis (hence no LFV)

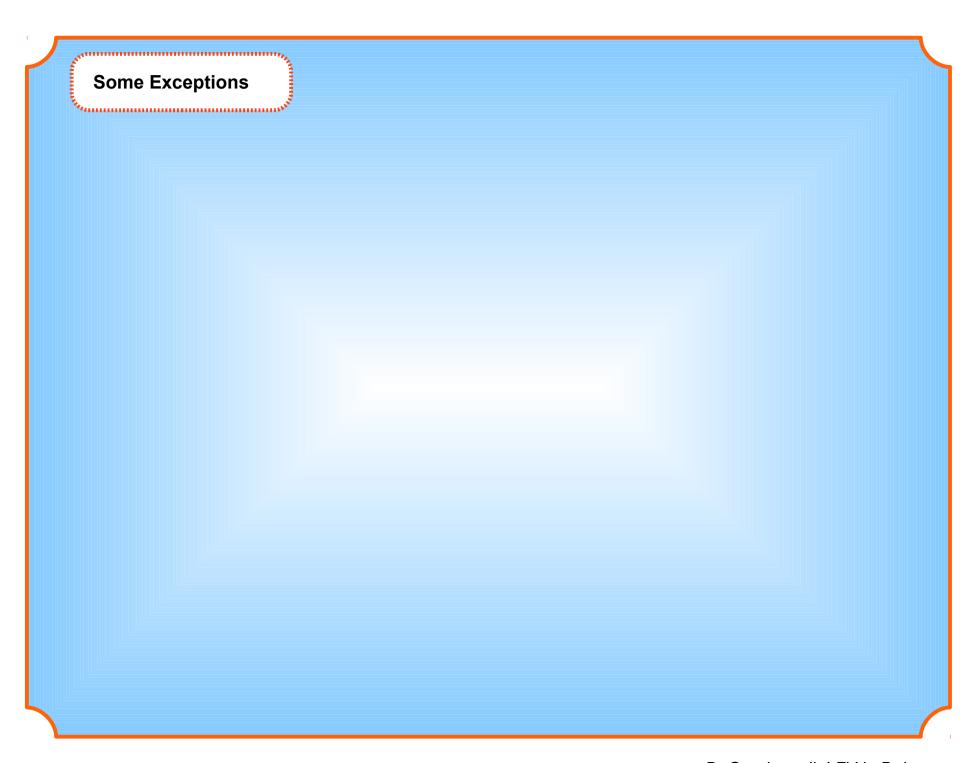
Or more generally, take the SM plus a minimal mechanism for v masses.

• Physical LFV will appear in W couplings, but it's suppressed by powers of $(m_v/m_w)^2$

Bottom line: in the SM+v there is LFNU, but LFV is nowhere to be seen (in decays)

 But nobody ordered that the reason (=tiny m_,) behind the above conclusion be at work also beyond the SM

So, BSM LFNU \Rightarrow BSM LFV (i.e. not suppressed by $m_{_{V}}$)



- Take Minimal Flavor Violation (MFV) in the lepton sector
 - By def, in MFV the only sources of flavor violation are the SM ones, i.e. the SM Yukawas

Carrier and the contract of th

- Take Minimal Flavor Violation (MFV) in the lepton sector
 - By def, in MFV the only sources of flavor violation are the SM ones, i.e. the SM Yukawas
 - Tricky to define MFV in the lepton sector:
 we don't know whether LH v are Dirac or Majorana and whether RH v exist at all.
 Must-read ref: Cirigliano-Grinstein-Isidori-Wise, NPB 2005

(a.....

- Take Minimal Flavor Violation (MFV) in the lepton sector
 - By def, in MFV the only sources of flavor violation are the SM ones, i.e. the SM Yukawas
 - Tricky to define MFV in the lepton sector:
 we don't know whether LH v are Dirac or Majorana and whether RH v exist at all.
 Must-read ref: Cirigliano-Grinstein-Isidori-Wise, NPB 2005
- Bottom line: In such scenarios, LFV couplings are related to LH v masses.
 (Neglecting CPV in the LH v mass matrix, the above statement is generic within MLFV.)

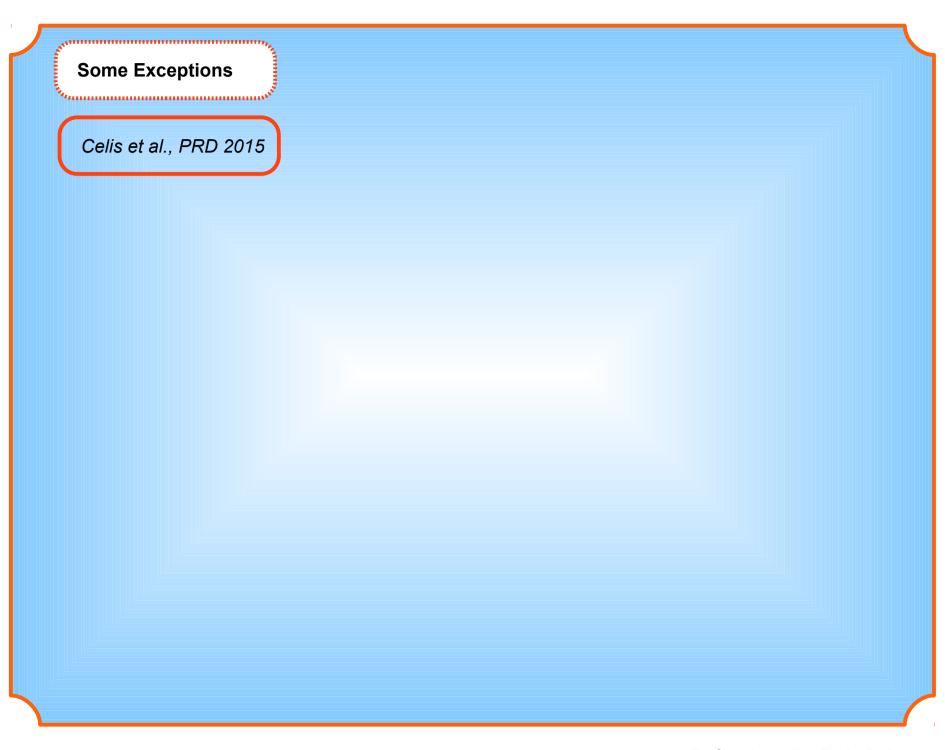
- Take Minimal Flavor Violation (MFV) in the lepton sector
 - By def, in MFV the only sources of flavor violation are the SM ones, i.e. the SM Yukawas
 - Tricky to define MFV in the lepton sector:
 we don't know whether LH v are Dirac or Majorana and whether RH v exist at all.
 Must-read ref: Cirigliano-Grinstein-Isidori-Wise, NPB 2005
- Bottom line: In such scenarios, LFV couplings are related to LH v masses.
 (Neglecting CPV in the LH v mass matrix, the above statement is generic within MLFV.)
 - Low-energy LFV processes are generally small, being suppressed by LH v masses. (This brings back to the previous slide)

4......

Alonso, Grinstein, Martin-Camalich, 1505.05164

- Take Minimal Flavor Violation (MFV) in the lepton sector
 - By def, in MFV the only sources of flavor violation are the SM ones, i.e. the SM Yukawas
 - Tricky to define MFV in the lepton sector:
 we don't know whether LH v are Dirac or Majorana and whether RH v exist at all.
 Must-read ref: Cirigliano-Grinstein-Isidori-Wise, NPB 2005
- Bottom line: In such scenarios, LFV couplings are related to LH v masses.
 (Neglecting CPV in the LH v mass matrix, the above statement is generic within MLFV.)
 - Low-energy LFV processes are generally small, being suppressed by LH v masses. (This brings back to the previous slide)
- "Generally small" means:

Barring MFV models where sizable LFV and small LH ν masses can be engineered to be so by tuning a dimensionful parameter to be small. (Back to fine tuning.)



Section of the sectio

Celis et al., PRD 2015

- Take a Branco-Grimus-Lavoura (BGL) global symmetry.
 - BGL models are a proposal to solve the monstrous flavor problem of general 2HDM (tree-level FCNCs)

Celis et al., PRD 2015

- Take a Branco-Grimus-Lavoura (BGL) global symmetry.
 - BGL models are a proposal to solve the monstrous flavor problem of general 2HDM (tree-level FCNCs)
 - They engineer an Abelian global symmetry that relates all Higgs-quark flavor-changing couplings to CKM entries

Celis et al., PRD 2015

- Take a Branco-Grimus-Lavoura (BGL) global symmetry.
 - BGL models are a proposal to solve the monstrous flavor problem of general 2HDM (tree-level FCNCs)
 - They engineer an Abelian global symmetry that relates all Higgs-quark flavor-changing couplings to CKM entries
- Gauge this symmetry, and require anomaly cancellation.

Celis et al., PRD 2015

(a.....)

- Take a Branco-Grimus-Lavoura (BGL) global symmetry.
 - BGL models are a proposal to solve the monstrous flavor problem of general 2HDM (tree-level FCNCs)
 - They engineer an Abelian global symmetry that relates all Higgs-quark flavor-changing couplings to CKM entries
- Gauge this symmetry, and require anomaly cancellation.
- This requirement yields diagonal charged-lepton Yukawa couplings.



BSM LFNU but no BSM LFV

Celis et al., PRD 2015

(a.....

- Take a Branco-Grimus-Lavoura (BGL) global symmetry.
 - BGL models are a proposal to solve the monstrous flavor problem of general 2HDM (tree-level FCNCs)
 - They engineer an Abelian global symmetry that relates all Higgs-quark flavor-changing couplings to CKM entries
- Gauge this symmetry, and require anomaly cancellation.
- This requirement yields diagonal charged-lepton Yukawa couplings.



Plausible mechanism?

Let's now turn to Q1: Can we (easily) make sense of data **1** to **5** ? It is highly non-trivial that a simple consistent BSM picture exists to describe the above data • to •

Can we (easily) make sense of data **1** to **5** ?

It is highly non-trivial that a simple consistent BSM picture exists to describe the above data • to •

Consider the following Hamiltonian

$$H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4 G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4 \pi} \left[\bar{b}_L \gamma^{\lambda} s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_{\lambda} \gamma_5 \mu \right) \right]$$

Can we (easily) make sense of data **1** to **5** ?

¢2.....

It is highly non-trivial that a simple consistent BSM picture exists to describe the above data 0 to 6

Consider the following Hamiltonian

purely vector lepton current

purely axial lepton current

$$H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4 G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4 \pi} \left[\bar{b}_L \gamma^{\lambda} s_L \cdot \left(C_9^{(\mu)} (\bar{\mu} \gamma_{\lambda} \mu) + C_{10}^{(\mu)} (\bar{\mu} \gamma_{\lambda} \gamma_5 \mu) \right) \right]$$

Can we (easily) make sense of data **1** to **5** ?

·

It is highly non-trivial that a simple consistent BSM picture exists to describe the above data 0 to 9

Consider the following Hamiltonian

 $H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4\pi} \left[\bar{b}_L \gamma^{\lambda} s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu \right) + C_{10}^{(\mu)} \bar{\mu} \gamma_{\lambda} \gamma_5 \mu \right]$

• Note:
$$C_9^{
m SM}(m_b) pprox +4.2$$
 $C_{10}^{
m SM}(m_b) pprox -4.4$

[Bobeth, Misiak, Urban, 99] [Khodjamirian et al., 10]

Can we (easily) make sense of data **0** to **5** ?

·

It is highly non-trivial that a simple consistent BSM picture exists to describe the above data 10 to 15

Consider the following Hamiltonian

$$H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \, \mu \, \mu) \, = \, -\frac{4 \, G_F}{\sqrt{2}} \, V_{tb}^* V_{ts} \, \frac{\alpha_{\text{em}}}{4 \, \pi} \left[\bar{b}_L \, \gamma^\lambda s_L \cdot \left(C_9^{(\mu)} \, \bar{\mu} \, \gamma_\lambda \mu \right) + \, C_{10}^{(\mu)} \, \bar{\mu} \, \gamma_\lambda \gamma_5 \mu \right]$$

Note:
$$C_9^{\rm SM}(m_b) \approx +4.2$$
 i.e. in the SM also the lepton current has nearly V – A structure

$$C_9^{\rm SM}(m_b) \approx -C_{10}^{\rm SM}(m_b)$$

[Bobeth, Misiak, Urban, 99] [Khodjamirian et al., 10]

Can we (easily) make sense of data • to • ?

.....

It is highly non-trivial that a simple consistent BSM picture exists to describe the above data **0** to **6**

Consider the following Hamiltonian

purely vector lepton current

purely axial lepton current

$$H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4\pi} \left[\bar{b}_L \gamma^{\lambda} s_L \cdot \left(C_9^{(\mu)} (\bar{\mu} \gamma_{\lambda} \mu) + C_{10}^{(\mu)} (\bar{\mu} \gamma_{\lambda} \gamma_5 \mu) \right) \right]$$

$$C_{10}^{\rm SM}(m_b) \approx -4.4$$

Note: $C_9^{\rm SM}(m_b) \approx +4.2$ i.e. in the SM also the lepton current has nearly V – A structure

[Bobeth, Misiak, Urban, 99] [Khodjamirian et al., 10]

We assume the above V – A structure to hold also beyond the SM, namely

$$C_9^{(t)} \approx -C_{10}^{(t)}$$

$$C_9^{(t)} \approx -C_{10}^{(t)}$$
 with $C_{9,10}^{(t)} = C_{9,10}^{\text{SM}} + C_{9,10}^{(t),\text{NP}}$

cf. also Hiller, Schmaltz; Ghosh, Nardecchia, Renner; Hurth, Mahmoudi, Neshatpour

Such an hypothesis provides a successful fit to the discussed data. See Altmannshofer-Straub, EPJC 2015.

- In short, our model requirements are: $C_9^{(t)} pprox C_{10}^{(t)}$ (V A structure)
 - $|C_{9,{
 m NP}}^{(\mu)}| \gg |C_{9,{
 m NP}}^{(e)}|$ (LFNU)

- In short, our model requirements are: $-C_9^{(t)} pprox -C_{10}^{(t)}$ (V A structure)
 - $|C_{9,{
 m NP}}^{(\mu)}| \gg |C_{9,{
 m NP}}^{(e)}|$ (LFNU)
- This structure can be generated from a purely 3rd-generation interaction of the kind

$$H_{\mathrm{NP}} = G \, \bar{b}'_L \gamma^{\lambda} b'_L \, \bar{\tau}'_L \gamma_{\lambda} \tau'_L$$
with $G = 1/\Lambda_{\mathrm{NP}}^2 \ll G_F$

expected e.g. in topcolor models [see C.T. Hill, PLB 1995]

4.....

- In short, our model requirements are: $-C_9^{(t)} pprox -C_{10}^{(t)}$ (V A structure)
 - $|C_{9,{
 m NP}}^{(\mu)}| \gg |C_{9,{
 m NP}}^{(e)}|$ (LFNU)
- This structure can be generated from a purely 3rd-generation interaction of the kind

$$H_{\mathrm{NP}} = G \, \bar{b}'_L \gamma^{\lambda} b'_L \, \bar{\tau}'_L \gamma_{\lambda} \tau'_L$$
 with $G = 1/\Lambda_{\mathrm{NP}}^2 \ll G_F$

expected e.g. in topcolor models [see C.T. Hill, PLB 1995]

- Note: primed fields
 - Fields are in the gauge basis (= primed)

- In short, our model requirements are: $-C_9^{(t)} \approx -C_{10}^{(t)}$ (V A structure)
 - $|C_{9,NP}^{(\mu)}| \gg |C_{9,NP}^{(e)}|$ (LFNU)
- This structure can be generated from a purely 3rd-generation interaction of the kind

$$H_{\mathrm{NP}} = G \, \bar{b} \, '_L \gamma^{\lambda} b \, '_L \, \bar{\tau} \, '_L \gamma_{\lambda} \tau \, '_L$$
 with $G = 1/\Lambda_{\mathrm{NP}}^2 \ll G_F$

expected e.g. in topcolor models [see C.T. Hill, PLB 1995]

- Note: primed fields
 - Fields are in the gauge basis (= primed)
 - They need to be rotated to the mass eigenbasis



$$b'_L \equiv (d'_L)_3 = (U_L^d)_{3i} (d_L)_i$$

$$\tau'_L \equiv (\ell'_L)_3 = (U_L^t)_{3i} (\ell_L)_i$$

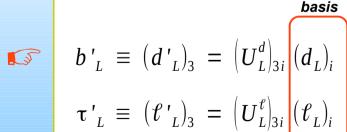
mass

- In short, our model requirements are: $-C_9^{(t)} pprox -C_{10}^{(t)}$ (V A structure)
 - $|C_{9,{
 m NP}}^{(\mu)}| \gg |C_{9,{
 m NP}}^{(e)}|$ (LFNU)
- This structure can be generated from a purely 3rd-generation interaction of the kind

$$H_{\mathrm{NP}} = G \, \bar{b}'_L \gamma^{\lambda} b'_L \, \bar{\tau}'_L \gamma_{\lambda} \tau'_L$$
 with $G = 1/\Lambda_{\mathrm{NP}}^2 \ll G_F$

expected e.g. in topcolor models [see C.T. Hill, PLB 1995]

- Note: primed fields
 - Fields are in the gauge basis (= primed)
 - They need to be rotated to the mass eigenbasis
 - This rotation induces <u>LFNU and LFV</u> effects



mass

Recalling our full Hamiltonian

$$H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \mu \mu) = -\frac{4 G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4 \pi} \left[\bar{b}_L \gamma^{\lambda} s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_{\lambda} \gamma_5 \mu \right) \right]$$

the shift to the C_9 Wilson coeff. in the $\mu\mu$ -channel becomes

$$k_{\text{SM}} C_9^{(\mu)} = k_{\text{SM}} C_{9,\text{SM}} + \frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} |(U_L^t)_{32}|^2$$

Recalling our full Hamiltonian

k_{SM} (SM norm. factor)

$$H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \mu \mu) = \left[-\frac{4 G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4 \pi} \right] \left[\bar{b}_L \gamma^{\lambda} s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_{\lambda} \gamma_5 \mu \right) \right]$$

the shift to the C_9 Wilson coeff. in the $\mu\mu$ -channel becomes

$$k_{\text{SM}} C_9^{(\mu)} = k_{\text{SM}} C_{9,\text{SM}} + \frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} |(U_L^t)_{32}|^2$$

Recalling our full Hamiltonian

·

k_{SM} (SM norm. factor)

$$H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \mu \mu) = \left[-\frac{4 G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4 \pi} \right] \left[\bar{b}_L \gamma^{\lambda} s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_{\lambda} \gamma_5 \mu \right) \right]$$

the shift to the C_9 Wilson coeff. in the $\mu\mu$ -channel becomes

$$k_{\text{SM}} C_9^{(\mu)} = k_{\text{SM}} C_{9,\text{SM}} + \frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} |(U_L^t)_{32}|^2$$

$$= \beta_{\text{SM}}$$

Recalling our full Hamiltonian

k_{SM} (SM norm. factor)

$$H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \mu \mu) = \left[-\frac{4 G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4 \pi} \right] \left[\bar{b}_L \gamma^{\lambda} s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_{\lambda} \gamma_5 \mu \right) \right]$$

the shift to the C_9 Wilson coeff. in the $\mu\mu$ -channel becomes

$$k_{\text{SM}} C_9^{(\mu)} = k_{\text{SM}} C_{9,\text{SM}} + \underbrace{\frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} |(U_L^t)_{32}|^2}_{+}$$

$$= \beta_{\text{SM}} + \beta_{\text{NP}}$$

The NP contribution has opposite sign than the SM one if

$$G\left(U_L^d\right)_{32} < 0$$

Recalling our full Hamiltonian

Variani and a 1 and 1

k_{SM} (SM norm. factor)

$$H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \mu \mu) = \left[-\frac{4 G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4 \pi} \right] \left[\bar{b}_L \gamma^{\lambda} s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_{\lambda} \gamma_5 \mu \right) \right]$$

the shift to the C_9 Wilson coeff. in the $\mu\mu$ -channel becomes

$$k_{\text{SM}} C_9^{(\mu)} = k_{\text{SM}} C_{9,\text{SM}} + \frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} |(U_L^t)_{32}|^2$$

$$= \beta_{\text{SM}} + \beta_{\text{NP}}$$

The NP contribution has opposite sign than the SM one if

$$G\left(U_L^d\right)_{32} < 0$$

On the other hand, in the ee-channel

$$k_{\text{SM}} C_9^{(e)} = k_{\text{SM}} C_{9,\text{SM}} + \frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} |(U_L^t)_{31}|^2$$

Recalling our full Hamiltonian

Variani and a 1 and 1

 k_{SM} (SM norm. factor)

$$H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \mu \mu) = \left[-\frac{4 G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4 \pi} \right] \left[\bar{b}_L \gamma^{\lambda} s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_{\lambda} \gamma_5 \mu \right) \right]$$

the shift to the C_α Wilson coeff. in the μμ-channel becomes

$$k_{\text{SM}} C_9^{(\mu)} = k_{\text{SM}} C_{9,\text{SM}} + \frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} |(U_L^t)_{32}|^2$$

$$= \beta_{\text{SM}} + \beta_{\text{NP}}$$

The NP contribution has opposite sign than the SM one if

$$G(U_L^d)_{32} < 0$$

On the other hand, in the ee-channel

$$k_{\text{SM}} C_9^{(e)} = k_{\text{SM}} C_{9,\text{SM}} + \frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} [(U_L^t)_{31}]^2$$

The NP contrib. in the eechannel is negligible, provided

$$\left|\left(\boldsymbol{U}_{L}^{t}\right)_{31}\right|^{2} \ll \left|\left(\boldsymbol{U}_{L}^{t}\right)_{32}\right|^{2}$$

Recalling our full Hamiltonian

Variani and a 1 and 1

k_{SM} (SM norm. factor)

$$H_{\text{SM+NP}}(\bar{b} \rightarrow \bar{s} \mu \mu) = \left[-\frac{4 G_F}{\sqrt{2}} V_{tb}^* V_{ts} \frac{\alpha_{\text{em}}}{4 \pi} \right] \left[\bar{b}_L \gamma^{\lambda} s_L \cdot \left(C_9^{(\mu)} \bar{\mu} \gamma_{\lambda} \mu + C_{10}^{(\mu)} \bar{\mu} \gamma_{\lambda} \gamma_5 \mu \right) \right]$$

the shift to the C_α Wilson coeff. in the μμ-channel becomes

$$k_{\text{SM}} C_9^{(\mu)} = k_{\text{SM}} C_{9, \text{SM}} + \frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} |(U_L^t)_{32}|^2$$

$$= \beta_{\text{SM}} + \beta_{\text{NP}}$$

The NP contribution has opposite sign than the SM one if

$$G\left(U_L^d\right)_{32} < 0$$

On the other hand, in the ee-channel

$$k_{\text{SM}} C_9^{(e)} = k_{\text{SM}} C_{9,\text{SM}} + \frac{G}{2} (U_L^d)_{33}^* (U_L^d)_{32} [(U_L^t)_{31}]^2$$

 $\simeq \beta_{\text{SM}}$

The NP contrib. in the eechannel is negligible, provided

$$\left|\left(\boldsymbol{U}_{L}^{t}\right)_{31}\right|^{2} \ll \left|\left(\boldsymbol{U}_{L}^{t}\right)_{32}\right|^{2}$$

• The above shifts to the C_{9,10} Wilson coeffs. imply

$$R_K \approx \frac{|C_9^{(\mu)}|^2 + |C_{10}^{(\mu)}|^2}{|C_9^{(e)}|^2 + |C_{10}^{(e)}|^2} = \frac{2 \cdot (\beta_{SM} + \beta_{NP})^2}{2 \cdot \beta_{SM}^2}$$

The above shifts to the C_{9,10} Wilson coeffs. imply

$$R_{K} \approx \frac{|C_{9}^{(\mu)}|^{2} + |C_{10}^{(\mu)}|^{2}}{|C_{9}^{(e)}|^{2} + |C_{10}^{(e)}|^{2}} = \frac{2 \cdot (\beta_{SM} + \beta_{NP})^{2}}{2 \cdot \beta_{SM}^{2}}$$

factors of 2:

equal contributions from $|C_9|^2$ and $|C_{10}|^2$

4..........

• The above shifts to the C_{9.10} Wilson coeffs. imply

$$R_{K} \approx \frac{|C_{9}^{(\mu)}|^{2} + |C_{10}^{(\mu)}|^{2}}{|C_{9}^{(e)}|^{2} + |C_{10}^{(e)}|^{2}} = \frac{2 \cdot (\beta_{SM} + \beta_{NP})^{2}}{2 \cdot \beta_{SM}^{2}}$$

factors of 2: equal contributions from $|C_g|^2$ and $|C_{10}|^2$

Approximations

- phase-space factor is about the same
 in the μμ- and in the ee-channel
- dominance of the $|C_{9,10}|^2$ contributions in the concerned q^2 region

• The above shifts to the C_{9.10} Wilson coeffs. imply

$$R_{K} \approx \frac{|C_{9}^{(\mu)}|^{2} + |C_{10}^{(\mu)}|^{2}}{|C_{9}^{(e)}|^{2} + |C_{10}^{(e)}|^{2}} = \frac{2 \cdot (\beta_{SM} + \beta_{NP})^{2}}{2 \cdot \beta_{SM}^{2}}$$

factors of 2: equal contributions from $|C_g|^2$ and $|C_{10}|^2$

Approximations

- phase-space factor is about the same in the μμ- and in the ee-channel
- dominance of the $|C_{9,10}|^2$ contributions in the concerned q^2 region

Note as well

$$0.77 \pm 0.20 = \frac{BR(B_s \rightarrow \mu \mu)_{\text{exp}}}{BR(B_s \rightarrow \mu \mu)_{\text{SM}}} = \frac{BR(B_s \rightarrow \mu \mu)_{\text{SM+NP}}}{BR(B_s \rightarrow \mu \mu)_{\text{SM}}} = \frac{(\beta_{\text{SM}} + \beta_{\text{NP}})^2}{\beta_{\text{SM}}^2}$$

The above shifts to the C_{9.10} Wilson coeffs. imply

$$R_{K} \approx \frac{|C_{9}^{(\mu)}|^{2} + |C_{10}^{(\mu)}|^{2}}{|C_{9}^{(e)}|^{2} + |C_{10}^{(e)}|^{2}} = \frac{2 \cdot (\beta_{SM} + \beta_{NP})^{2}}{2 \cdot \beta_{SM}^{2}}$$

factors of 2: equal contributions from $|C_0|^2$ and $|C_{10}|^2$

Approximations

- phase-space factor is about the same in the μμ- and in the ee-channel
- dominance of the $|C_{9,10}|^2$ contributions in the concerned q^2 region

Note as well

$$0.77 \pm 0.20 = \frac{BR(B_s \rightarrow \mu \mu)_{\text{exp}}}{BR(B_s \rightarrow \mu \mu)_{\text{SM}}} = \frac{BR(B_s \rightarrow \mu \mu)_{\text{SM+NP}}}{BR(B_s \rightarrow \mu \mu)_{\text{SM}}} = \frac{(\beta_{\text{SM}} + \beta_{\text{NP}})^2}{\beta_{\text{SM}}^2}$$

implying (within our model) the correlations

$$\frac{BR(B_s \to \mu \mu)_{\text{exp}}}{BR(B_s \to \mu \mu)_{\text{SM}}} \simeq R_K \simeq \frac{BR(B^+ \to K^+ \mu \mu)_{\text{exp}}}{BR(B^+ \to K^+ \mu \mu)_{\text{SM}}}$$

• The above shifts to the C_{9.10} Wilson coeffs. imply

$$R_{K} \approx \frac{|C_{9}^{(\mu)}|^{2} + |C_{10}^{(\mu)}|^{2}}{|C_{9}^{(e)}|^{2} + |C_{10}^{(e)}|^{2}} = \frac{2 \cdot (\beta_{SM} + \beta_{NP})^{2}}{2 \cdot \beta_{SM}^{2}}$$

factors of 2: equal contributions from $|C_g|^2$ and $|C_{10}|^2$

Approximations

- phase-space factor is about the same in the μμ- and in the ee-channel
- dominance of the $|C_{9,10}|^2$ contributions in the concerned q^2 region

Note as well

$$0.77 \pm 0.20 = \frac{BR(B_s \rightarrow \mu \mu)_{\text{exp}}}{BR(B_s \rightarrow \mu \mu)_{\text{SM}}} = \frac{BR(B_s \rightarrow \mu \mu)_{\text{SM+NP}}}{BR(B_s \rightarrow \mu \mu)_{\text{SM}}} = \frac{(\beta_{\text{SM}} + \beta_{\text{NP}})^2}{\beta_{\text{SM}}^2}$$

implying (within our model) the correlations

$$\frac{BR(B_s \to \mu \mu)_{\text{exp}}}{BR(B_s \to \mu \mu)_{\text{SM}}} \simeq R_K \simeq \frac{BR(B^+ \to K^+ \mu \mu)_{\text{exp}}}{BR(B^+ \to K^+ \mu \mu)_{\text{SM}}}$$

Another good reason to pursue accuracy in the B_s → µµ measurement

Marian Commission of the Commi

$$\frac{BR(B^{+} \to K^{+} \mu e)}{BR(B^{+} \to K^{+} \mu \mu)} = \frac{\beta_{NP}^{2}}{(\beta_{SM} + \beta_{NP})^{2}} \cdot \frac{|(U_{L}^{t})_{31}|^{2}}{|(U_{L}^{t})_{32}|^{2}} \cdot 2$$

·

$$\frac{BR(B^{+} \rightarrow K^{+} \mu e)}{BR(B^{+} \rightarrow K^{+} \mu \mu)} = \begin{bmatrix} \beta_{NP}^{2} \\ (\beta_{SM} + \beta_{NP})^{2} \\ = 0.159^{2} \\ \text{according to } R_{K} \end{bmatrix} \frac{|(U_{L}^{\ell})_{31}|^{2}}{|(U_{L}^{\ell})_{32}|^{2}} \cdot 2$$

·

$$\frac{BR(B^{+} \to K^{+} \mu e)}{BR(B^{+} \to K^{+} \mu \mu)} = \begin{bmatrix} \beta_{NP}^{2} & \frac{|(U_{L}^{t})_{31}|^{2}}{|(\beta_{SM} + \beta_{NP})^{2}|} & \frac{|(U_{L}^{t})_{31}|^{2}}{|(U_{L}^{t})_{32}|^{2}} & \frac{2}{\mu^{+}e^{-} \& \mu^{-} e^{+} modes} \end{bmatrix}$$

$$= 0.159^{2} \text{ according to } R_{K}$$

4

$$\frac{BR(B^{+} \to K^{+} \mu e)}{BR(B^{+} \to K^{+} \mu \mu)} = \begin{bmatrix} \beta_{NP}^{2} & \frac{|(U_{L}^{t})_{31}|^{2}}{|(\beta_{SM} + \beta_{NP})^{2}|} & \frac{|(U_{L}^{t})_{31}|^{2}}{|(U_{L}^{t})_{32}|^{2}} & \frac{2}{\mu^{+}e^{-} \& \mu^{-} e^{+} modes} \end{bmatrix}$$

$$= 0.159^{2} \text{ according to } R_{K}$$

BR(
$$B^+ \to K^+ \mu e$$
) < 2.2×10⁻⁸ · $\frac{|(U_L^t)_{31}|^2}{|(U_L^t)_{32}|^2}$

The current BR(B+ \rightarrow K+ μ e) limit yields the weak bound

$$|(U_L^t)_{31}/(U_L^t)_{32}| < 3.7$$

$$\frac{BR(B^{+} \rightarrow K^{+} \mu e)}{BR(B^{+} \rightarrow K^{+} \mu \mu)} = \begin{bmatrix} \beta_{NP}^{2} \\ (\beta_{SM} + \beta_{NP})^{2} \\ = 0.159^{2} \\ \text{according to } R_{K} \end{bmatrix} \frac{|(U_{L}^{\ell})_{31}|^{2}}{|(U_{L}^{\ell})_{32}|^{2}} \cdot \begin{bmatrix} 2 \\ \mu^{+}e^{-} \& \mu^{-} e^{+} \\ \text{modes} \end{bmatrix}$$

$$BR(B^+ \to K^+ \mu e) < 2.2 \times 10^{-8} \cdot \frac{|(U_L^t)_{31}|^2}{|(U_L^t)_{32}|^2}$$

The current BR(B+ \rightarrow K+ μ e) limit yields the weak bound

$$|(U_L^t)_{31}/(U_L^t)_{32}| < 3.7$$

$$lacksquare BR(B^+ o K^+ \mu \, au)$$
 would be even more promising, as it scales with $|(U_L^t)_{33}/(U_L^t)_{32}|^2$

$$\frac{BR(B^{+} \rightarrow K^{+} \mu e)}{BR(B^{+} \rightarrow K^{+} \mu \mu)} = \begin{bmatrix} \frac{\beta_{NP}^{2}}{(\beta_{SM} + \beta_{NP})^{2}} & \frac{|(U_{L}^{t})_{31}|^{2}}{|(U_{L}^{t})_{32}|^{2}} \\ = 0.159^{2} & \text{according to } R_{K} \end{bmatrix}$$

BR(B⁺
$$\rightarrow K^+ \mu e$$
) < 2.2×10⁻⁸ · $\frac{|(U_L^t)_{31}|^2}{|(U_L^t)_{32}|^2}$

The current BR(B+ \rightarrow K+ μ e) limit yields the weak bound

$$|(U_L^t)_{31}/(U_L^t)_{32}| < 3.7$$

 $lacksquare BR(B^+ o K^+ \mu \, au)$ would be even more promising, as it scales with $|(U_L^t)_{33} / (U_L^t)_{32}|^2$

A reliable prediction of the BR requires some more work:

 phase-space factors are substantially different than in the μμ and ee cases (but can easily be accounted for)

Marian Commission of the Commi

Wallian Committee Committ

Again, $B_s \to \mu \tau$ would be even more promising, because it scales as $|(U_L^t)_{33}/(U_L^t)_{32}|^2$ (a potential enhancement factor, actually)

- Again, $B_s \to \mu \tau$ would be even more promising, because it scales as $|(U_L^t)_{33}/(U_L^t)_{32}|^2$ (a potential enhancement factor, actually)
- $\overline{\mathsf{V}}$ An interesting signature outside B physics would be $K \to \pi \ \ell'$

- Again, $B_s \to \mu \tau$ would be even more promising, because it scales as $|(U_L^t)_{33}/(U_L^t)_{32}|^2$ (a potential enhancement factor, actually)
- ightharpoonup An interesting signature outside B physics would be $K \to \pi \ \ell'$

Note, instead, that the "K-physics analogue" of R_K:

$$\frac{BR(K \to \pi \mu \mu)}{BR(K \to \pi e e)}$$
 less interesting as it is long-distance dominated [see D'Ambrosio et al., 1998]

More quantitative LFV predictions More quantitative LFV predictions require knowledge of the U_L^{ℓ}

More quantitative LFV predictions

More quantitative LFV predictions require knowledge of the U_L^e

Reminder:

$$(U_L^\ell)^\dagger Y_\ell U_R^\ell = \hat{Y}_\ell$$

Z.,....

- More quantitative LFV predictions require knowledge of the U_L^{ℓ}
- One approach:

DG, Lane, 1507.01412

Reminder:

$$(U_L^\ell)^\dagger Y_\ell U_R^\ell = \hat{Y}_\ell$$

More quantitative LFV predictions require knowledge of the U_L^e

 $(U_L^\ell)^\dagger Y_\ell U_R^\ell = \hat{Y}_\ell$

Reminder:

One approach:

DG, Lane, 1507.01412

Appelquist-Bai-Piai ansatz:
 the flavor-SU(3) rotations are not all independent. Choosing 3 to be the independent ones
 allows to predict one SM Yukawa in terms of the other two.

More quantitative LFV predictions require knowledge of the U_L^e

Reminder:

$$(U_L^t)^{\dagger} Y_t U_R^t = \hat{Y}_t$$

One approach:

DG, Lane, 1507.01412

- Appelquist-Bai-Piai ansatz:
 the flavor-SU(3) rotations are not all independent. Choosing 3 to be the independent ones
 allows to predict one SM Yukawa in terms of the other two.
- One can thereby determine Y_{ℓ} in terms of Y_{μ} and Y_{d}

More quantitative LFV predictions require knowledge of the U_L^e

Reminder:

 $(U_L^t)^\dagger Y_t U_R^t = \hat{Y}_t$

One approach:

DG, Lane, 1507.01412

- Appelquist-Bai-Piai ansatz:
 the flavor-SU(3) rotations are not all independent. Choosing 3 to be the independent ones
 allows to predict one SM Yukawa in terms of the other two.
- One can thereby determine Y_e in terms of Y_u and Y_d
- But we don't know Y_u and Y_d entirely, so we take an (independently motivated) model for them, reproducing quark masses and the CKM matrix [Martin-Lane, PRD 2005].

More quantitative LFV predictions require knowledge of the U_L^e

Reminder:

$$(U_L^{\ell})^{\dagger} Y_{\ell} U_R^{\ell} = \hat{Y}_{\ell}$$

• One approach:

DG, Lane, 1507.01412

- Appelquist-Bai-Piai ansatz:
 the flavor-SU(3) rotations are not all independent. Choosing 3 to be the independent ones
 allows to predict one SM Yukawa in terms of the other two.
- One can thereby determine Y_{ℓ} in terms of Y_{μ} and Y_{d}
- But we don't know Y_u and Y_d entirely, so we take an (independently motivated) model for them, reproducing quark masses and the CKM matrix [Martin-Lane, PRD 2005].
- Another approach:

Boucenna, Valle, Vicente, PLB 2015

4.....

More quantitative LFV predictions require knowledge of the U_L^e

Reminder:

$$(U_L^{\ell})^{\dagger} Y_{\ell} U_R^{\ell} = \hat{Y}_{\ell}$$

One approach:

DG, Lane, 1507.01412

- Appelquist-Bai-Piai ansatz:
 the flavor-SU(3) rotations are not all independent. Choosing 3 to be the independent ones
 allows to predict one SM Yukawa in terms of the other two.
- One can thereby determine Y_{ℓ} in terms of Y_{μ} and Y_{d}
- But we don't know Y_u and Y_d entirely, so we take an (independently motivated) model for them, reproducing quark masses and the CKM matrix [Martin-Lane, PRD 2005].
- Another approach:

Boucenna, Valle, Vicente, PLB 2015

- One has $(U_i^{\ell})^{\dagger}U_i^{\nu} = PMNS$ matrix

More quantitative LFV predictions require knowledge of the U_L^e

Reminder:

$$(U_L^{\ell})^{\dagger} Y_{\ell} U_R^{\ell} = \hat{Y}_{\ell}$$

One approach:

DG, Lane, 1507.01412

- Appelquist-Bai-Piai ansatz:
 the flavor-SU(3) rotations are not all independent. Choosing 3 to be the independent ones
 allows to predict one SM Yukawa in terms of the other two.
- One can thereby determine Y_e in terms of Y_u and Y_d
- But we don't know Y_u and Y_d entirely, so we take an (independently motivated) model for them, reproducing quark masses and the CKM matrix [Martin-Lane, PRD 2005].
- Another approach:

Boucenna, Valle, Vicente, PLB 2015

- One has $(U_{i}^{\ell})^{\dagger}U_{i}^{\nu}=PMNS$ matrix
- Taking $U_L^{\ \nu} = 1$, $U_L^{\ \ell}$ can be univocally predicted

LFV predictions in one of the two scenarios of [DG, Lane]

	$B^+ o K^+ \mu^\pm au^\mp$	$B^{\scriptscriptstyle +} ightarrow K^{\scriptscriptstyle +} \ e^{\scriptscriptstyle \pm} \ au^{\scriptscriptstyle \mp}$	$B^{\scriptscriptstyle +} ightarrow K^{\scriptscriptstyle +} \ e^{\scriptscriptstyle \pm} \ \mu^{\scriptscriptstyle \mp}$
	1.14×10^{-8}	3.84×10^{-10}	0.52×10^{-9}
Exp:	$< 4.8 \times 10^{-5}$	$< 3.0 \times 10^{-5}$	$< 9.1 \times 10^{-8}$

LFV predictions in one of the two scenarios of [DG, Lane]

	$B^+ \rightarrow K^+ \mu^{\pm} \tau^{\mp}$	$B^+ \to K^+ e^{\pm} \tau^{\mp}$	$B^+ \rightarrow K^+ e^{\pm} \mu^{\mp}$
	1.14×10^{-8}	3.84×10^{-10}	0.52×10^{-9}
Exp:	$< 4.8 \times 10^{-5}$	$< 3.0 \times 10^{-5}$	$< 9.1 \times 10^{-8}$

	$B_s \rightarrow \mu^{\pm} \tau^{\mp}$	$B_s \rightarrow e^{\pm} \tau^{\mp}$	$B_s \rightarrow e^{\pm} \mu^{\mp}$
	$1.37 imes 10^{-8}$	4.57×10^{-10}	1.73×10^{-12}
Exp:			$< 1.1 \times 10^{-8}$

All predictions are phase-space corrected.

4......

For a recent discussion:
Alonso, Grinstein, Martin-Camalich,

• Being defined above the EWSB scale, our assumed operator $G\ \bar{b}'_L \gamma^{\lambda} b'_L \ \bar{\tau}'_L \gamma_{\lambda} \tau'_L$ must actually be made invariant under $SU(3)_c \times SU(2)_L \times U(1)_{\gamma}$

For a recent discussion: Alonso, Grinstein, Martin-Camalich,

Being defined above the EWSB scale, our assumed operator $G \ \bar{b}'_L \gamma^{\lambda} b'_L \ \bar{\tau}'_L \gamma_{\lambda} \tau'_L$ must actually be made invariant under $SU(3)_c \times SU(2)_t \times U(1)_y$

Bhattacharya, Datta, London, Shivashankara, PLB 15

$$\bar{b}'_L \gamma^{\lambda} b'_L \bar{\tau}'_L \gamma_{\lambda} \tau'_L$$



$$\begin{array}{c} \text{SU(2)}_{\text{\tiny L}} \\ & \stackrel{\frown}{\bar{Q}}{}^{\prime}{}_{L}\,\gamma^{\lambda}Q{}^{\prime}{}_{L}\,\,\bar{L}{}^{\prime}{}_{L}\gamma_{\lambda}L{}^{\prime}{}_{L} \\ & \stackrel{\frown}{\bar{Q}}{}^{\prime}{}_{L}\,\gamma^{\lambda}Q{}^{\prime}{}_{L}\,\,\bar{L}{}^{\prime}{}_{L}\gamma_{\lambda}L{}^{\prime}{}_{L} \end{array}$$

$$ar{Q}^{\prime i}_{L} \gamma^{\lambda} Q^{\prime j}_{L} \, ar{L}^{\prime j}_{L} \gamma_{\lambda} L^{\prime j}_{L}$$

[also charged-current int's]

For a recent discussion: Alonso, Grinstein, Martin-Camalich,

Being defined above the EWSB scale, our assumed operator $G \ \bar{b}'_L \gamma^{\lambda} b'_L \ \bar{\tau}'_L \gamma_{\lambda} \tau'_L$ must actually be made invariant under $SU(3)_c \times SU(2)_t \times U(1)_y$

Bhattacharya, Datta, London, Shivashankara, PLB 15

$$\begin{array}{c} \mathbf{S}^{15} \\ \bar{b}^{\,\prime}_{\,L} \, \gamma^{\lambda} b^{\,\prime}_{\,L} \, \bar{\tau}^{\,\prime}_{\,L} \gamma_{\lambda} \tau^{\,\prime}_{\,L} \end{array} \qquad \begin{array}{c} \mathbf{SU(2)}_{\!L} \\ \\ \bar{b}^{\,\prime}_{\,L} \, \gamma^{\lambda} b^{\,\prime}_{\,L} \, \bar{\tau}^{\,\prime}_{\,L} \gamma_{\lambda} L^{\,\prime}_{\,L} \end{array} \qquad \begin{array}{c} \mathbf{SU(2)}_{\!L} \\ \\ \bar{Q}^{\,\prime}_{\,L} \, \gamma^{\lambda} Q^{\,\prime}_{\,L} \, \bar{L}^{\,\prime}_{\,L} \gamma_{\lambda} L^{\,\prime}_{\,L} \end{array}$$

[neutral-current int's only]

[also charged-current int's]

Thus, the generated structures are all of:

$$t't'v'_{\tau}v'_{\tau}$$

$$t't'\tau'\tau'$$

$$t't'v'_{\tau}v'_{\tau}$$
, $t't'\tau'\tau'$, $b'b'v'_{\tau}v'_{\tau}$, $b'b'\tau'\tau'$, $t'b'\tau'v'_{\tau}$,

$$b'b'\tau'\tau'$$

$$t'b'\tau'\nu'$$

For a recent discussion: Alonso, Grinstein, Martin-Camalich,

Being defined above the EWSB scale, our assumed operator $G \ \bar{b}'_L \gamma^{\lambda} b'_L \ \bar{\tau}'_L \gamma_{\lambda} \tau'_L$ must actually be made invariant under $SU(3)_c \times SU(2)_t \times U(1)_y$

Bhattacharya, Datta, London, Shivashankara, PLB 15

$$\bar{b}'_L \gamma^{\lambda} b'_L \bar{\tau}'_L \gamma_{\lambda} \tau'_L$$

$$\begin{array}{c} \operatorname{SU(2)_L} \\ \bar{b}\,{}'_L\,\gamma^\lambda b\,{}'_L\,\bar{\tau}\,{}'_L\gamma_\lambda \tau\,{}'_L \end{array} \qquad \begin{array}{c} \operatorname{SU(2)_L} \\ \\ \operatorname{inv.} \end{array} \qquad \begin{array}{c} \bullet \quad \bar{Q}\,{}'_L\gamma^\lambda Q\,{}'_L\,\bar{L}\,{}'_L\gamma_\lambda L\,{}'_L \\ \\ \bullet \quad \bar{Q}\,{}'_L\gamma^\lambda Q\,{}'_L\,\bar{L}\,{}'_L\gamma_\lambda L\,{}'_L \end{array}$$

[neutral-current int's only]

[also charged-current int's]

Thus, the generated structures are all of:

$$t't'v'_{\tau}v'_{\tau}$$
,

$$t't'\tau'\tau'$$

$$b'b'v'_{\tau}v'_{\tau}$$
,

$$b'b'\tau'\tau'$$

$$t't'v'_{\tau}v'_{\tau}$$
, $t't'\tau'\tau'$, $b'b'v'_{\tau}v'_{\tau}$, $b'b'\tau'\tau'$, $t'b'\tau'v'_{\tau}$,

·

For a recent discussion: Alonso, Grinstein, Martin-Camalich,

Being defined above the EWSB scale, our assumed operator $G \bar{b}'_L \chi^{\lambda} b'_L \bar{\tau}'_L \gamma_{\lambda} \tau'_L$ must actually be made invariant under $SU(3)_c \times SU(2)_t \times U(1)_y$

Bhattacharya, Datta, London, Shivashankara, PLB 15

$$\bar{b}'_L \gamma^{\lambda} b'_L \bar{\tau}'_L \gamma_{\lambda} \tau'_L$$



$$ar{Q}^{\prime i}_{L} \gamma^{\lambda} Q^{\prime j}_{L} \, ar{L}^{\prime j}_{L} \gamma_{\lambda} L^{\prime i}_{L}$$

Thus, the generated structures are all of:

$$t't'v'_{\tau}v'_{\tau}$$

$$t't'\tau'\tau'$$

$$b'b'v'_{\tau}v'_{\tau}$$
,

$$b'b'\tau'\tau'$$

$$t't'v'_{\tau}v'_{\tau}$$
, $t't'\tau'\tau'$, $b'b'v'_{\tau}v'_{\tau}$, $b'b'\tau'\tau'$, $t'b'\tau'v'_{\tau}$

After rotation to the mass basis (unprimed), the last structure contributes to $\Gamma(b \rightarrow c \, \tau \, \overline{v}_{\scriptscriptstyle i})$



Can explain BaBar deviations on $R(D^{(*)}) = \frac{BR(\bar{B} \rightarrow D^{(*)+} \tau^- \bar{\nu}_{\tau})}{BR(\bar{B} \rightarrow D^{(*)+} \ell^- \bar{\nu}_{\ell})}$ (D* channel confirmed by LHCb)