

cLFV search with high intensity muon beams

F. Kapusta
LPNHE Paris



LPC Clermont,
9-10 june 2015

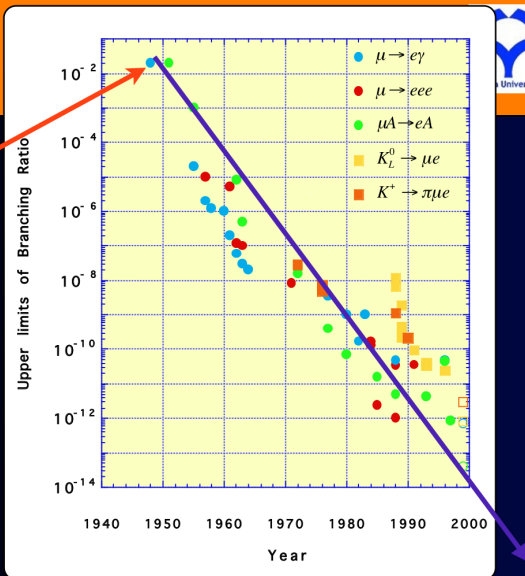
CLFV History

First cLFV search



Бруно Понтекорво

Pontecorvo in 1947

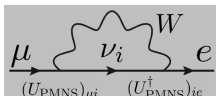


Physics Motivation : Beyond the Standard Model with muons

- ▶ We already know that LF is not conserved from neutrino oscillations.
- ▶ Direct search (Energy Frontier) LHC, ILC : higher energy for heavier new particle(s).

$$|A_{SM} + \varepsilon_{NP}|^2 \simeq |A_{SM}|^2 + 2\text{Re}(A_{SM}\varepsilon_{NP})$$

- ▶ Indirect search (Intensity Frontier): "slight" difference from SM prediction. cLFV in the SM (+ m_ν) is negligibly small.



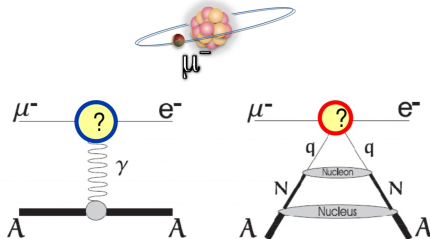
Adding some radiation for energy and momentum conservation.

μ -e transition results from the PMNS mixing.
Tiny neutrino mass suppresses the contribution.
Cheng and Li ('77, '80) Petcov('77).
 $BR(\mu \rightarrow e\gamma) \simeq O(10^{-54})$.

$$|A_{SM} + \varepsilon_{NP}|^2 \simeq |\varepsilon_{NP}|^2 \Rightarrow \text{Rate} \simeq \frac{1}{\Lambda^4}$$

- ▶ Probe the PeV scale with cLFV.

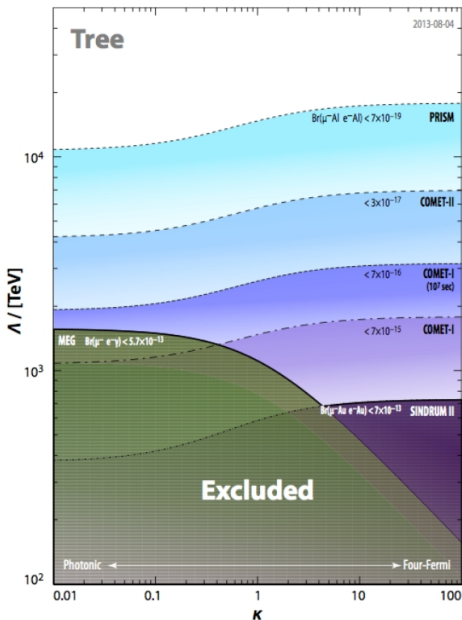
Exclusion diagrams



Dipole and 4-fermion couplings

$\mu^- N \rightarrow e^- N$ low energy effective lagrangian

$$\mathcal{L} = \frac{1}{1+\kappa} \frac{m_\mu}{\Lambda^2} \bar{\mu}_R \sigma^{\mu\nu} e_L \mathcal{F}_{\mu\nu} + \frac{\kappa}{1+\kappa} \frac{1}{\Lambda^2} (\bar{\mu}_L \gamma^\mu e_L) (\bar{q}_L \gamma_\mu q_L) + h.c.$$



$\mu \rightarrow e\gamma$ vs. μ -e conversion

André de Gouvêa

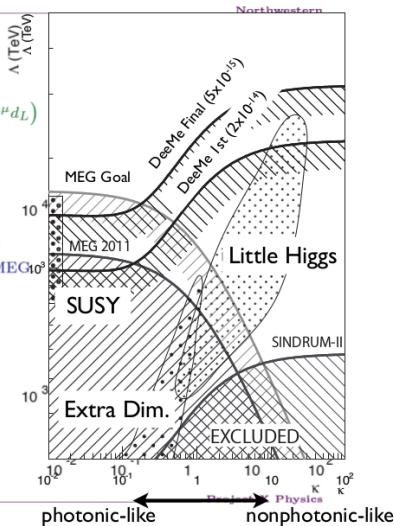
Model Independent Analysis

$$L_{\text{CLFV}} = \frac{m_\mu}{(\kappa+1)\Lambda^2} \bar{\mu} R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(1+\kappa)\Lambda^2} \bar{\mu} L \gamma_\mu e_L (\bar{u}_L \gamma^\mu u_L + \bar{d}_L \gamma^\mu d_L)$$

- $\mu \rightarrow e\text{-conv}$ at 10^{-17} "guaranteed" deeper probe than $\mu \rightarrow e\gamma$ at 10^{-14} .
- We don't think we can do $\mu \rightarrow e\gamma$ better than 10^{-14} . $\mu \rightarrow e\text{-conv}$ "only" way forward after MEG1
- If the LHC does not discover new states $\mu \rightarrow e\text{-conv}$ among very few process that can access 1000+ TeV new physics scale:
tree-level new physics: $\kappa \gg 1$, $\frac{1}{\Lambda^2} \sim \frac{g^2 \theta_{e\mu}}{M_{\text{new}}^2}$.



January 31, 2008



Experimental consequences

Rare decays searches require :

- ▶ Detectors with very good resolution and excellent background rejection.
- ▶ Background includes physical background, beam-related backgrounds, accidentals, cosmic rays and false tracking.
- ▶ As good as possible simulation and tracking are mandatory

Comparison between $\mu \rightarrow e\gamma$ and $\mu - e$ conversion :

	background	challenge	beam intensity
$\mu \rightarrow e\gamma$	accidentals	detector resolution	limited
$\mu - e$ conversion	beam	beam background	no limitation

- ▶ High intensity pulsed muon beams require strict proton beam extinction between pulses.
- ▶ Unlimited discussions on limits on BR or SES between experiments.

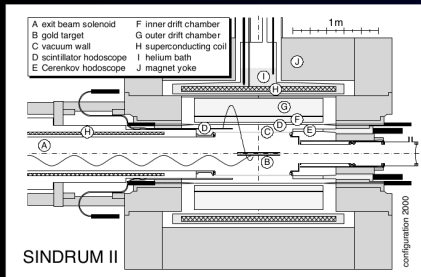
Current bounds and future sensitivities

Process	Experiment	Limit
$\text{BR}(\mu^+ \rightarrow e^+ \gamma)$	MEG('13)	$5.7 \cdot 10^{-13}$
	MEG-II(\geq '16) at PSI	$4 \cdot 10^{-14}$
$\text{BR}(\mu^+ \rightarrow e e e)$	SINDRUM('88)	$1.0 \cdot 10^{-12}$
	Mu3e(\geq '17) at PSI	$\mathcal{O}(10^{-16})$
	MUSIC(\simeq '17)	$\mathcal{O}(10^{-16})$
$\text{R}(\mu \rightarrow e : \text{Au})$	SINDRUM-II('06)	$7 \cdot 10^{-3}$
$\text{R}(\mu \rightarrow e : \text{Al})$	COMET(\simeq '17)	$\mathcal{O}(10^{-17})$
	Mu2e(\simeq '20)	$\mathcal{O}(10^{-17})$
$\text{R}(\mu \rightarrow e : \text{Ti})$	PRISM(\simeq '20)	$\mathcal{O}(10^{-18})$

PSI vs J-PARC : $10^8 \mu/s$ vs $10^{11} \mu/s$

Previous Measurements

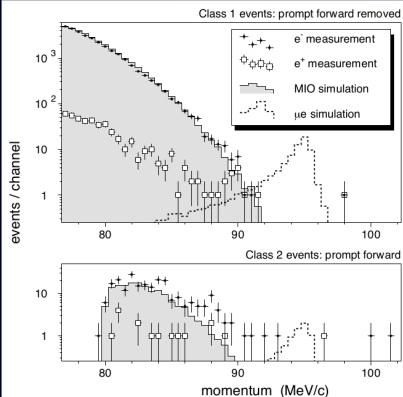
SINDRUM-II (PSI)



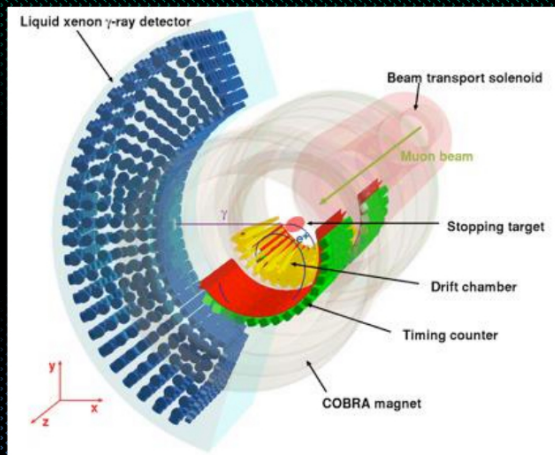
PSI muon beam intensity $\sim 10^{7-8}/\text{sec}$ beam from the PSI cyclotron. To eliminate beam related background from a beam, a beam veto counter was placed. But, it could not work at a high rate.

Published Results (2004)

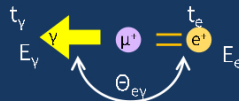
$$B(\mu^- + Au \rightarrow e^- + Au) < 7 \times 10^{-13}$$



MEG experiment



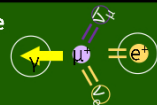
Signal Event



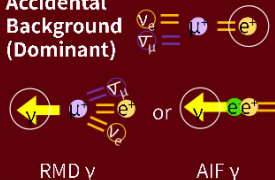
$$E_\gamma = E_e = m_\mu / 2$$

$$t_\gamma = t_e, \Theta_{e\nu} = \pi$$

Radiative Muon Decay



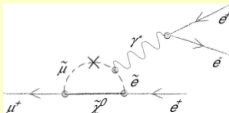
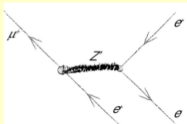
Accidental Background (Dominant)



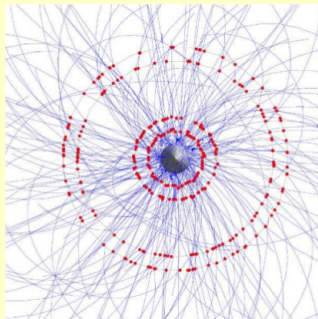
The Mu3e Experiment



- Search for LFV decay: $\mu \rightarrow eee$
- Single event sensitivity of
 10^{-15} Phase I
 $<10^{-16}$ Phase II
- Muon rate 10^8 ($>10^9$) per second
- $O(10)$ ($O(100)$) tracks within 50ns
- Sensitive to New Physics:

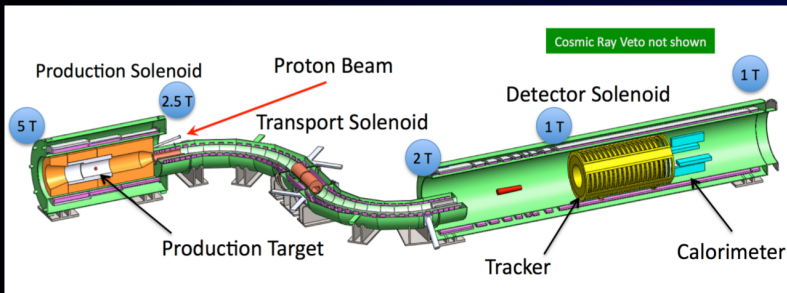
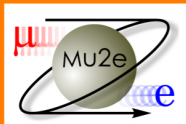


Discussed in Research Proposal:
 → arXiv:1301.6113



All silicon tracker
 based on
 HV-MAPS technology

μ -e conversion : Mu2e at Fermilab



$$B(\mu^- + Al \rightarrow e^- + Al) = 5 \times 10^{-17} \quad (\text{S.E.})$$

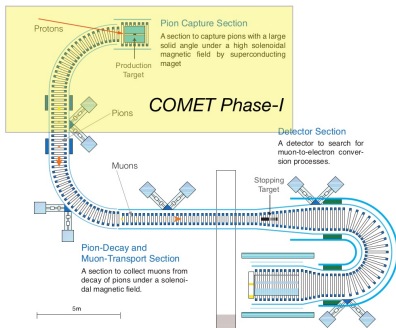
$$B(\mu^- + Al \rightarrow e^- + Al) < 10^{-16} \quad (90\% \text{C.L.})$$

- Reincarnation of MECO at BNL.
- Antiproton buncher ring is used to produce a pulsed proton beam.
- Approved in 2009, and CD0 in 2009, and CD1 in 2011.
- Data taking starts in about 2019.

COMET

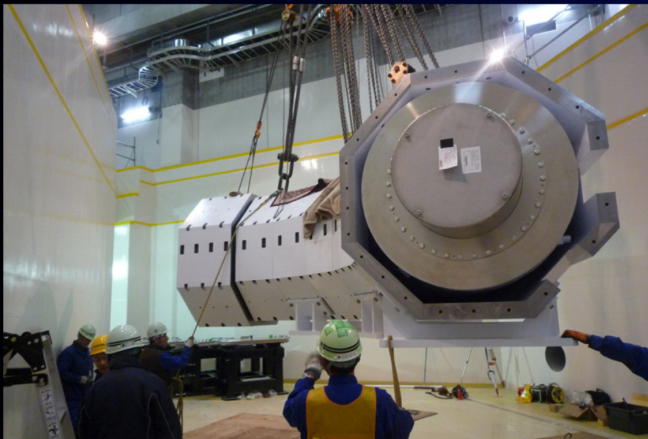
$\mu \rightarrow e$ conversion

- ▶ Staging approach
- ▶ Phase I to achieve 10^{-14} sensitivity and then Phase II



- ▶ Funding approved in JFY 2012 supplementary budget
- ▶ Annex of the current existing hall
- ▶ 8 GeV, pulsed proton beam to produce high-intensity muon beam
- ▶ COMET building finished and Muon Transport Solenoid installed

Muon Transport Solenoid at J-PARC



COMET Collaboration



164 collaborators
37 institutes, 13 countries

The COMET Collaboration

R. Akhmetshin^{6,28}, V. Anishchik⁴, M. Aoki²⁹, R. B. Appleby^{8,22}, Y. Arimoto¹⁵, Y. Bagaturia³³, Y. Ban⁴, W. Bertche²², A. Bondar^{6,28}, S. Canfer³⁰, S. Chen²⁵, Y. E. Cheung²⁵, B. Chladze³², D. Clarke³⁰, M. Danilov^{13,23}, P. D. Damcey¹¹, J. David²⁰, W. Da Silva²⁰, C. Densham³⁰, G. Devidze³², P. Dornan¹¹, A. Drutskoy^{13,23}, V. Dugimov¹⁴, A. Edmonds³⁰, L. Epshteyn^{6,27}, P. Evtoukhovich¹⁴, G. Fedotov^{6,28}, M. Finger⁷, M. Finger Jr⁷, Y. Fujii², Y. Fukao¹⁵, J.-F. Genat²⁰, M. Gensabeck²², E. Gillies¹¹, D. Grigoriev^{6,27,28}, K. Gritsay¹⁴, R. Han¹, K. Hasegawa¹⁵, I. H. Hasin²⁹, O. Hayashi²⁹, M. I. Hossain¹⁶, Z. A. Ibrahim²¹, Y. Igarashi¹⁵, F. Ignatov^{6,28}, M. Iio¹⁵, M. Ikono¹⁵, K. Ishibashi¹⁹, S. Ishimoto¹⁵, T. Itahashi²⁹, S. Ito²⁹, T. Iwami²⁹, Y. Iwashita¹⁷, X. S. Jiang², P. Jonsson¹¹, V. Kalinnikov¹⁴, F. Kapusta²⁰, H. Katayama²⁹, K. Kawagoe¹⁹, V. Kazanin^{6,28}, B. Khazin^{36,28}, A. Khvedelidze¹⁴, M. Koike³⁶, G. A. Kozlov¹⁴, B. Krikler¹¹, A. Kulikov¹⁴, E. Kulish¹⁴, Y. Kuno²⁹, Y. Kuriyama¹⁸, Y. Kurochkin⁵, A. Kurup¹¹, B. Lagrange^{11,18}, M. Lancaster³⁵, H. B. Li², W. G. Li², A. Liparteliani³², R. P. Litchfield³⁵, P. Loveridge³⁰, G. Macharashvili¹⁴, Y. Makida¹⁵, Y. Mao³, O. Markin¹³, Y. Matsumoto²⁹, T. Mibe¹⁵, S. Mihara¹⁵, F. Mohamad Idris²¹, K. A. Mohamed Kamal Azmi²¹, A. Moiseenko¹⁴, Y. Mori¹⁸, N. Mosulishvili³², E. Motuk³⁵, Y. Nakai¹⁹, T. Nakamoto¹⁵, Y. Nakazawa²⁹, J. Nash¹¹, M. Nioradze³², H. Nishiguchi¹⁵, T. Numao³⁴, J. O'Dell³⁰, T. Ogitsu¹⁵, K. Oishi¹⁹, K. Okamoto²⁹, C. Omori¹⁵, T. Ota³³, H. Owen²², C. Parkes²², J. Pasternak¹¹, C. Postinari³⁰, V. Pomaradev⁴, A. Popov^{6,28}, V. Rusinov^{13,23}, A. Ryzhenenkov^{6,28}, B. Sabirov¹⁴, N. Saito¹⁵, H. Sakamoto²⁹, P. Sarin¹⁰, K. Sasaki¹⁵, A. Sato²⁹, J. Sato³¹, D. Shemyakin^{6,28}, N. Shiguo¹⁹, D. Shoukavy⁵, M. Shmecka³, M. Sugano³⁰, Y. Takubo¹⁵, M. Tanaka¹⁵, C. V. Tao³⁶, E. Tarkovsky^{13,23}, Y. Tevzadze³², N. D. Thong²⁹, V. Thuan¹², J. Tojo¹⁹, M. Tomasek³, M. Tomizawa¹³, N. H. Tran²⁹, I. Trek³², N. M. Trung²⁹, Z. Tsamalaidze¹⁴, N. Tsvetkov¹⁴, S. Tygier²², T. Uchida¹⁵, Y. Uchida¹¹, K. Ueno¹⁵, S. Umasankar¹⁴, E. Velicheva¹⁴, A. Volkov¹⁴, V. Vrba⁹, W. A. T. Wan Abdullah²¹, M. Warren³⁹, M. Wing³⁵, T. S. Wong²⁹, C. Wu²⁵, G. Xia²², H. Yamaguchi¹⁹, A. Yamamoto¹⁵, M. Yamamaka²⁴, Y. Yang¹⁹, H. Yoshida²⁹, M. Yoshida¹⁵, Y. Yoshii¹⁵, T. Yoshioka¹⁹, Y. Yuan², Y. Yudin^{6,28}, J. Zhang², Y. Zhang²

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³Peking University, Beijing, People's Republic of China

⁴Belarusian State University (BSU), Minsk, Belarus

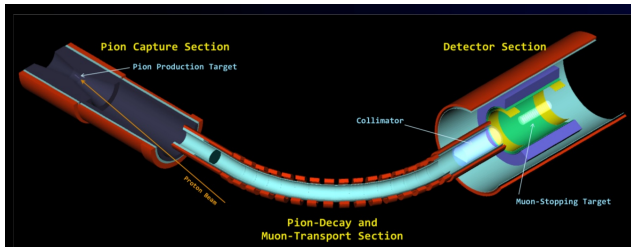
⁵B.I. Stepanov Institute of Physics, National Academy of Sciences of Belarus, Minsk, Belarus



COMET (E21)

COMET Phase I (2016)

- Beam background study and achieve $S.E.S. \simeq 3.10^{-15}$ with 8 GeV - 3.2 kW proton beam, ~ 3 months DAQ

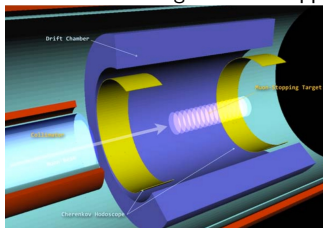


COMET Phase II (2020)

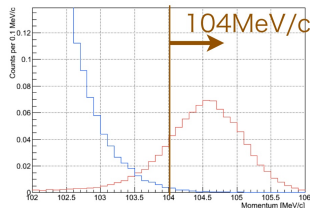
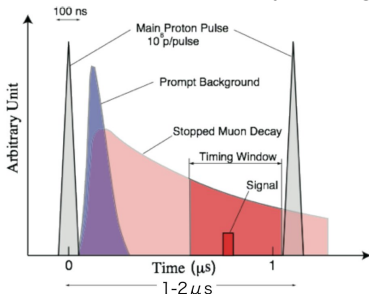
- 8 GeV - 56 kW proton beam, ~ 1 year DAQ to achieve the COMET final goal of $S.E.S \simeq 3.10^{-17}$

France-Japan collaboration in COMET

- CDC and Triggering counter surrounding a muon stopping target



- $\mu \rightarrow e$ conversion signal identified with an energetic electron of 105MeV emitted from a muonic atom with delayed timing.

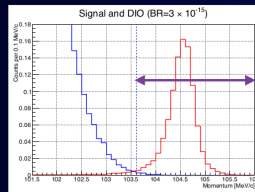


Signal Sensitivity for COMET Phase-I with CyDet

Signal Acceptance

Table 28: Breakdown of the $\mu^- N \rightarrow e^- N$ conversion signal acceptance.

Event selection	Value	Comments
Geometrical acceptance	0.37	
Track quality cuts	0.66	
Momentum selection	0.93	$103.6 \text{ MeV}/c < P_e < 106.0 \text{ MeV}/c$
Timing window	0.3	$700 \text{ ns} < t < 1100 \text{ ns}$
Trigger efficiency	0.8	
DAQ efficiency	0.8	
Track reconstruction efficiency	0.8	
Total	0.043	



Signal Sensitivity

- $f_{\text{cap}} = 0.6$
- $A_e = 0.043$
- $N_\mu = 1.23 \times 10^{16}$ muons

$$B(\mu^- + Al \rightarrow e^- + Al) \sim \frac{1}{N_\mu \cdot f_{\text{cap}} \cdot A_e}$$

$$B(\mu^- + Al \rightarrow e^- + Al) = 3.1 \times 10^{-15}$$

$$B(\mu^- + Al \rightarrow e^- + Al) < 7 \times 10^{-15} \quad (90\% C.L.)$$

Muon intensity

about 0.00052 muons stopped/proton

With 0.4 μA , a running time of about 110 days is needed.

Background Estimate for COMET Phase-I with CyDet

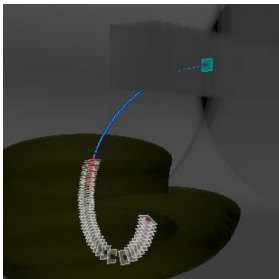
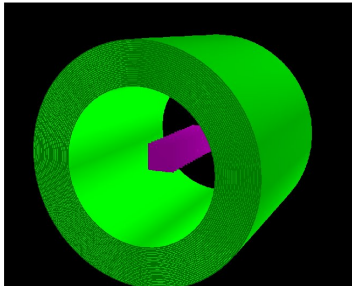


Table 30: Summary of the estimated background events for a single-event sensitivity of 3.1×10^{-15} with a proton extinction factor of 3×10^{-11} .

Type	Background	Estimated events
Physics	Muon decay in orbit	0.01
Physics	Radiative muon capture	5.6×10^{-4}
Physics	Neutron emission after muon capture	< 0.001
Physics	Charged particle emission after muon capture	< 0.001
Prompt Beam	Beam electrons (prompt)	8.3×10^{-4}
Prompt Beam	Muon decay in flight (prompt)	$\leq 2.0 \times 10^{-4}$
Prompt Beam	Pion decay in flight (prompt)	$\leq 2.3 \times 10^{-3}$
Prompt Beam	Other beam particles (prompt)	$\leq 2.8 \times 10^{-6}$
Prompt Beam	Radiative pion capture(prompt)	2.3×10^{-4}
Delayed Beam	Beam electrons (delayed)	~ 0
Delayed Beam	Muon decay in flight (delayed)	~ 0
Delayed Beam	Pion decay in flight (delayed)	~ 0
Delayed Beam	Radiative pion capture (delayed)	~ 0
Delayed Beam	Anti-proton induced backgrounds	0.007
Others	Electrons from cosmic ray muons	< 0.0001
Total		0.019

France-Japan collaboration in COMET

- ▶ LPNHE R&D for an active muon stopping target in order to get an additional point for the electron trajectory (CM11 - 2013)

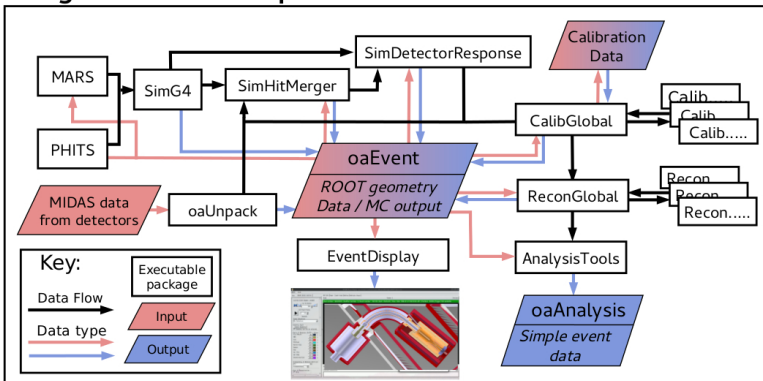


- ▶ Simulation and reconstruction with GENFIT within ICEDUST(Integrated Comet Experiment Data User Software Toolkit), the new COMET Software Framework rooted in T2K ND280.
- ▶ Discussion on the possibility of a beam test of ATLAS pixels at J-PARC with Kyushu University.
- ▶ MARS and ICEDUST installed at CCIN2P3 (thanks to Yonny Cardenas).

ICEDUST

Overview

Integrated Comet Experiment Data User Software Toolkit



Recent history

- ▶ COMET Software Framework: from ND280 to ICEDUST.
Imperial College London lead : Ajit Kurup, Ben Krikler
- ▶ Common COMET g-2/EDM FJPLP Workshop (Paris, 20-21 february 2014)



Workshop on silicon detectors for g-2/EDM/COMET experiments

20-21 février 2014
LPNHE Paris
Dorian Perrin-Benveniste

Overview
Agenda
Timetable
Registration
1. Registration Form
List of registrants
Access to LPNHE

The g-2/EDM and COMET experiments will detect positrons and electrons with silicon microstrips and pixels detectors.
The current status and the future needs will be reviewed and discussed.

Démarre 20 févr. 2014 09:00
Fini 21 févr. 2014 19:00
Europe/Paris

LPNHE Paris
1213-PC08-11

- ▶ 3rd Workshop on Muon g-2, EDM and Flavour Violation in the LHC Era in december 2014



- ▶ It was the right time for a decision from CCIN2P3 director and IN2P3 Particle Scientific Deputy Director to create a comet group to allow "foreign collaborators" to register and use CCIN2P3 machines.

Software Activity

Software group structure, january 2014

- Software group involves 13+ people

Sub-group coordinator: Ajit Kurup

Sam Tygier:

Fluka

Ben Krikler:

SimG4, overall framework

Per Johnsson:

Unit tests, ND280 support

Kazuki Ueno:

Straw tracker

Wilfrid da Silva, Frederic**Kapusta:**

GENFIT, Active Target

Andy Edmonds

MARS, SimG4

Chen Wu

Build system, repository, CyDet

Phill Litchfield

Offline databases, ND280 support

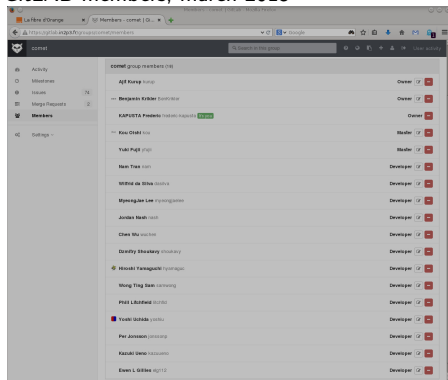
Fedor Ignatov

Reconstruction

Vladimir Kalinnikov, Elena**Velicheva**

ECAL

GitLAB members, march 2015



Recent events

- ▶ Accepted proposal from the french group to use the CCIN2P3 computing power and support in order to prepare a Grid computing at the COMET Collaboration level.
- ▶ Accepted proposal to have gitlab.in2p3.fr hosting the COMET software in order to ease the collaborative work.
- ▶ ICEDUST is running with MARS using a common 1 TB of semi-permanent space on /sps/hep/comet.
- ▶ SimMARS has been tested and optimized at CCIN2P3.
- ▶ A MySQL database is available for parameters storage.
- ▶ muon.in2p3.fr is a french website under construction to unify μ^+ (g-2/EDM) and μ^- (COMET) experiments for BSM physics, the official COMET website being comet.kek.jp
- ▶ AtCM16, CCIN2P3 was accepted as a COMET Member.
- ▶ Full Simulation is currently running at CCIN2P3 with 20 TB of data storage using iRODS for data sharing.

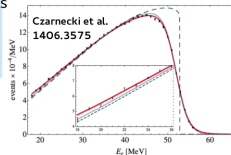
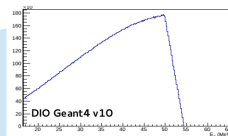
Questions and a proposal

- From the "3rd Workshop on Muon g-2, EDM and Flavour Violation ...":
 - new idea of cLFV search $\mu^- e^- \rightarrow e^- e^-$ in muonic atom.
 - upgrade of the DIO spectrum using Czarnecki last computation...

Simulation: Intrinsic BG

ICEDUST
Simulation

- Switch to Geant4 v10
 - Muon stopping on all Z:
 - Capture rates: Suzuki, Measday, Roalsvig, Phys.Rev. C35 (1987)
 - Decay rates: Mukhopadhyay Phys. Rep. 30 (1977)
- Include muon and pion capture processes as extended physics models
 - Radiative μ and π capture
 - Charged particle emission after μ capture
 - Neutron emission after μ capture
 - μ decay in orbit using most recent calculations
 - Czarnecki et al. Phys. Rev. D 9 (2014)



December 2014

27

Ben Krikler

- Expected contribution from french theorists : model predictions and "advertising plots".
- You are invited to give a talk at CM17 beginning of september in Paris and ... maybe join COMET.
- "4th Workshop on Muon g-2, EDM and Flavour Violation in the LHC Era" in 2016 organized in Marseilles by Marc Knecht.