

Low mass WIMP searches with EDELWEISS III: first results

Thibault de Boissière

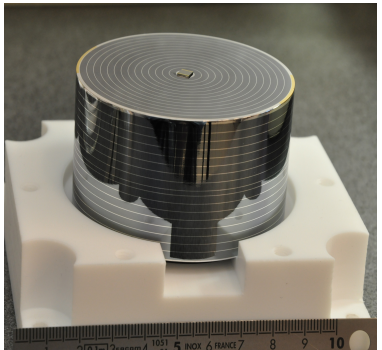
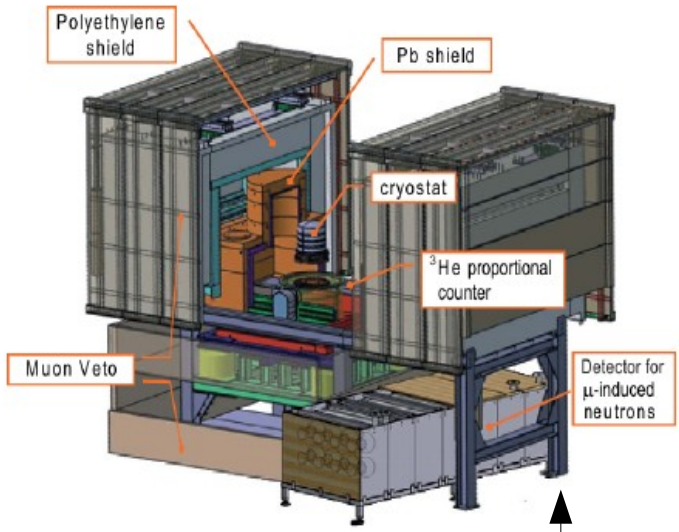


Part I

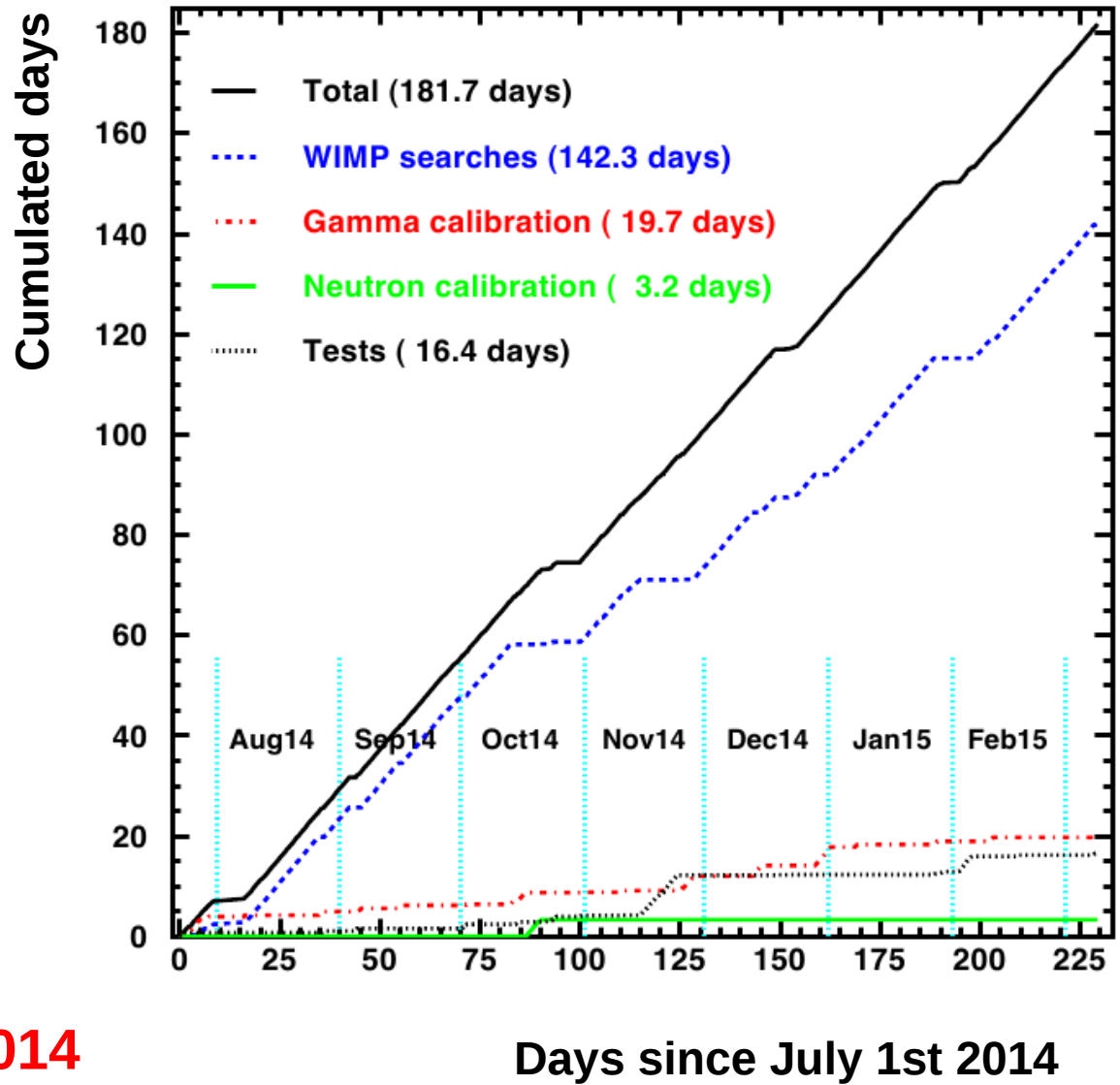
The EDELWEISS-III experiment

The EDELWEISS-III experiment

Setup and status



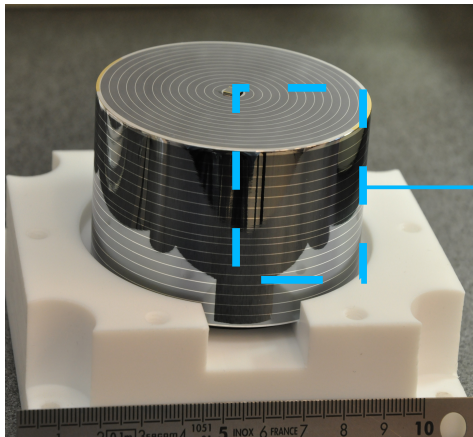
Setup
 Germanium bolometer



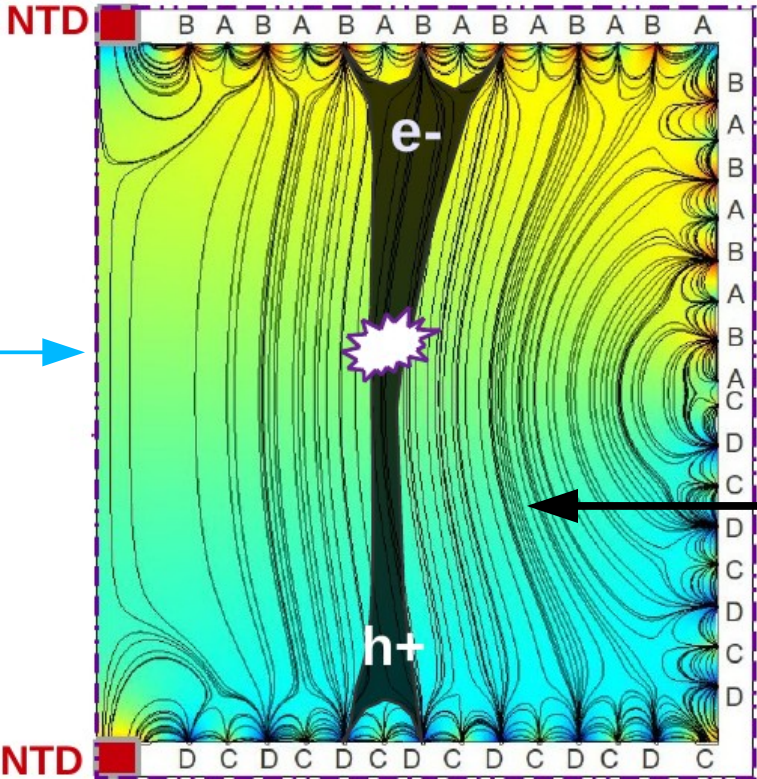
24 detectors running
Physics data since summer 2014
Expect 3000 kg.d for Summer 2015



The EDELWEISS-III experiment: Detector scheme



800 g detector
600 g fiducial mass



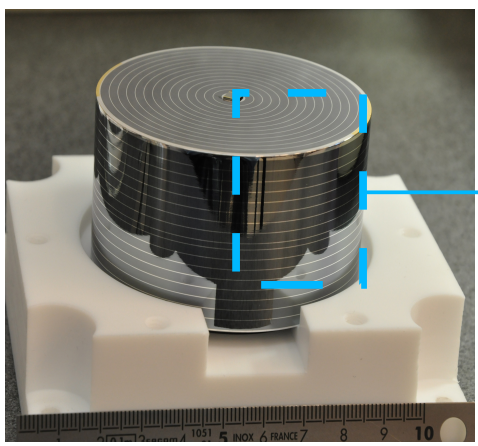
4 ionisation channels:
A,B,C,D (interleaved electrodes)

2 heat channels:
Thermal photon measurement with NTD

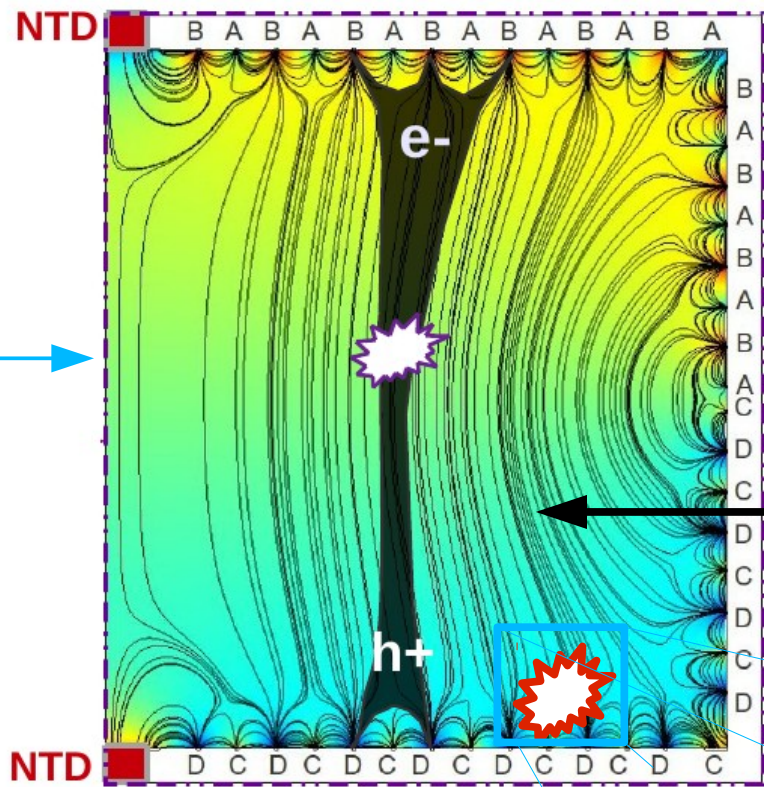
Electric field lines

Bulk (fiducial) event: charge is collected in B and D

The EDELWEISS-III experiment: Detector scheme



800 g detector
 600 g fiducial mass



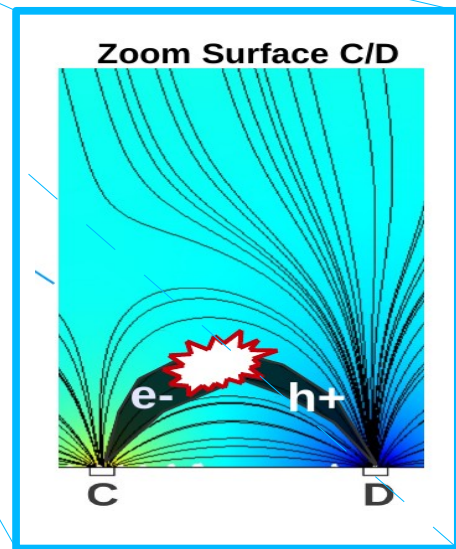
4 ionisation channels:
 A,B,C,D (interleaved electrodes)

2 heat channels:
 Thermal photon measurement with NTD

Electric field lines

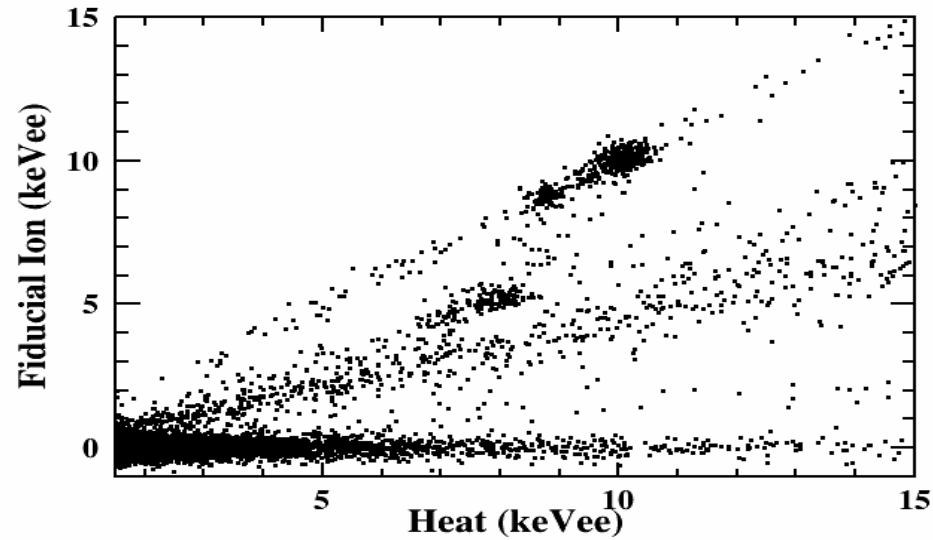
Bulk (fiducial) event: charge is collected in B and D

Surface event: charge is collected in A and B or C and D





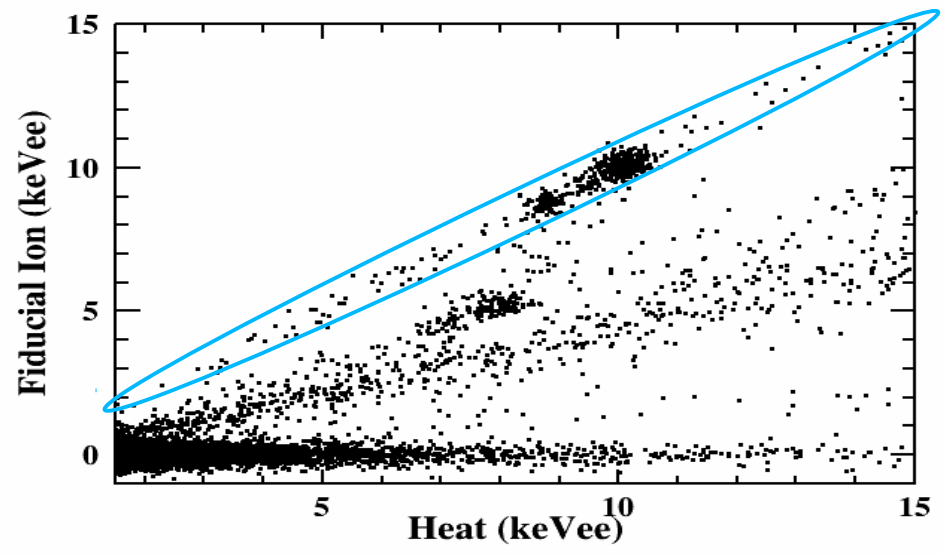
The EDELWEISS-III experiment: Background rejection





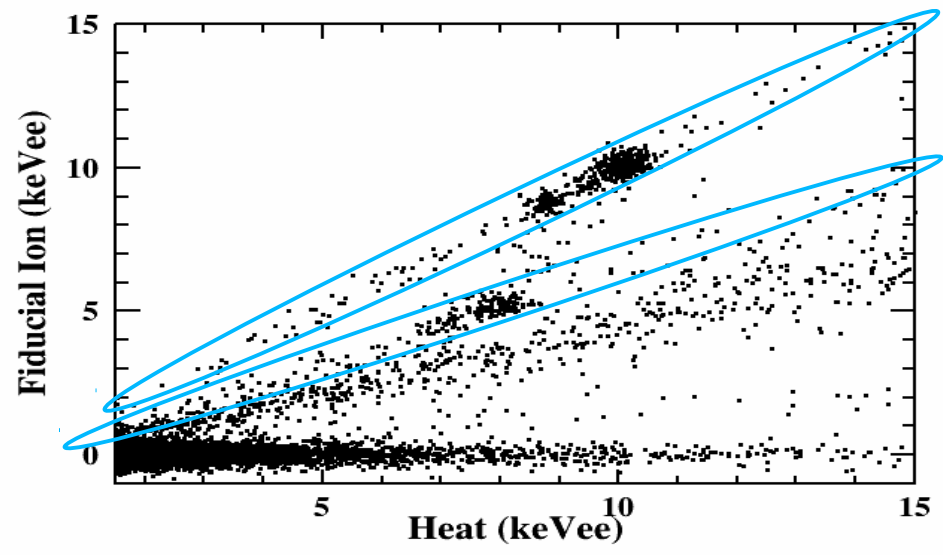
The EDELWEISS-III experiment: Background rejection

Fiducial gamma





The EDELWEISS-III experiment: Background rejection

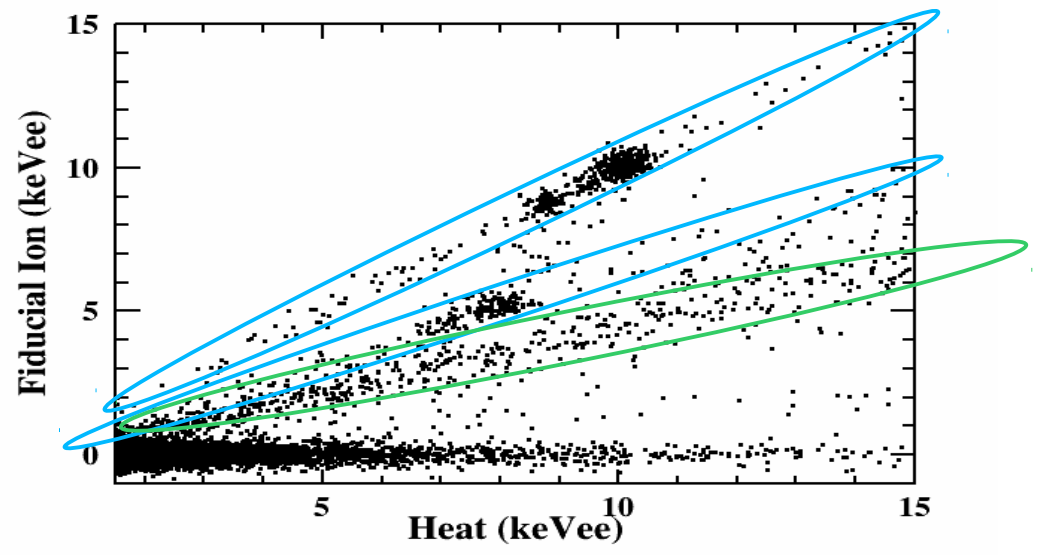


Fiducial gamma

Surface Gamma



The EDELWEISS-III experiment: Background rejection

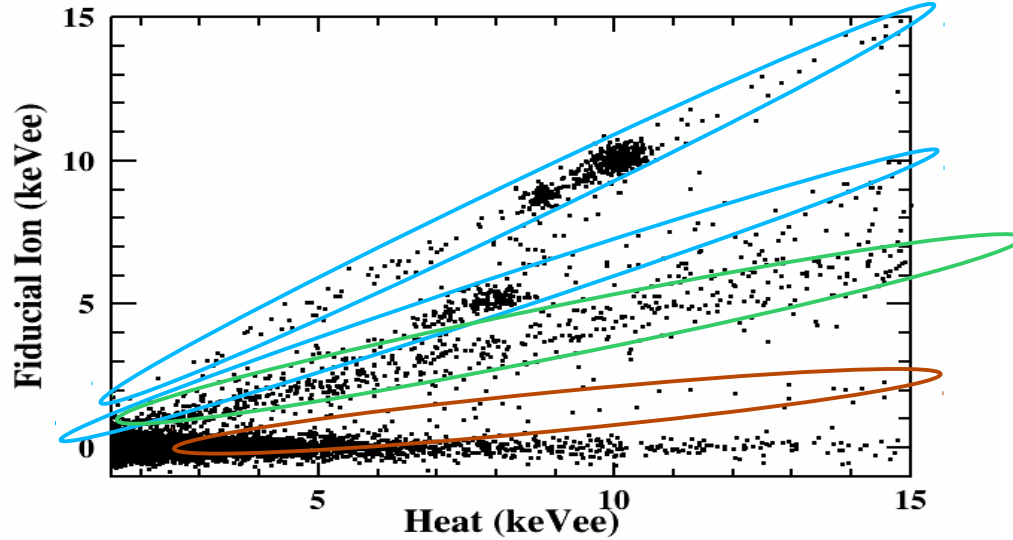


Fiducial gamma

Surface Gamma

Surface Beta

The EDELWEISS-III experiment: Background rejection



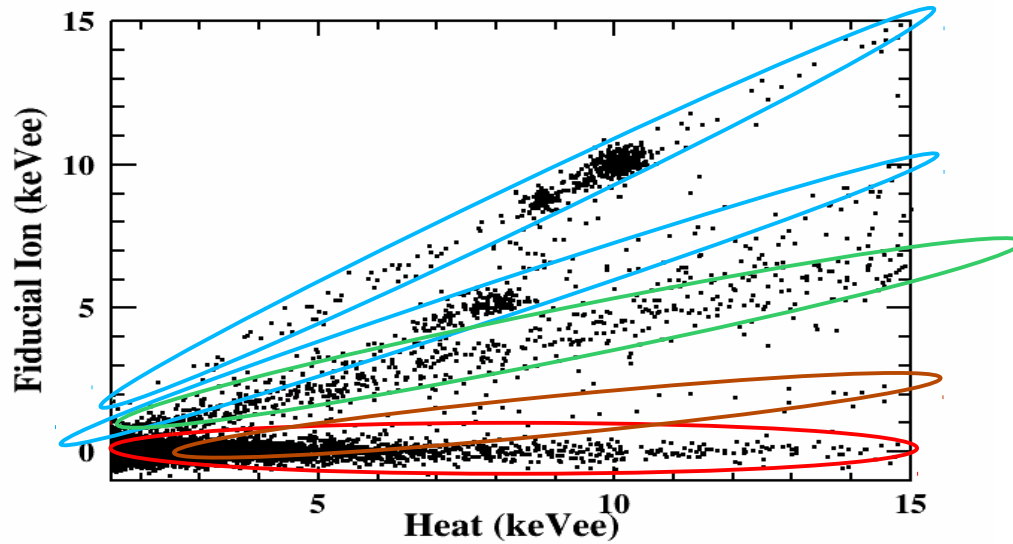
Fiducial gamma

Surface Gamma

Surface Beta

Surface Lead

The EDELWEISS-III experiment: Background rejection



Fiducial gamma

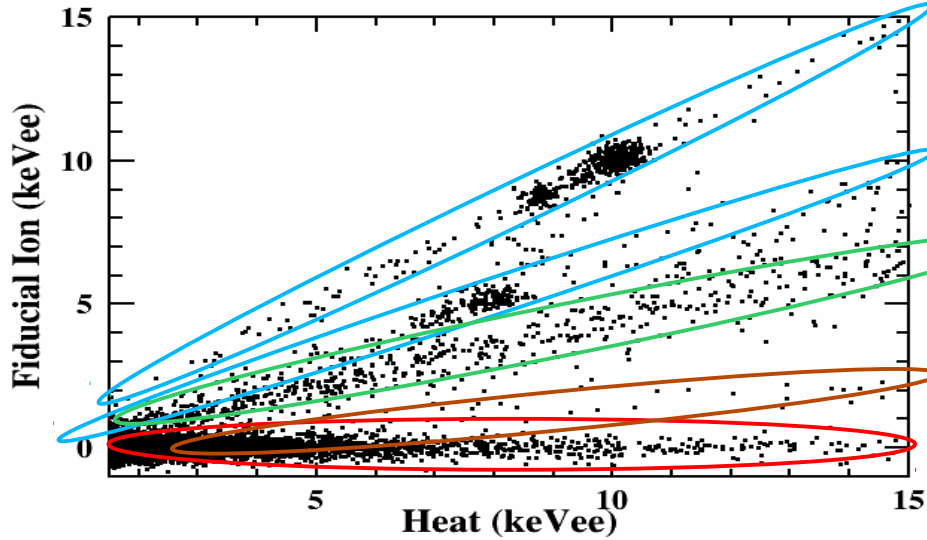
Surface Gamma

Surface Beta

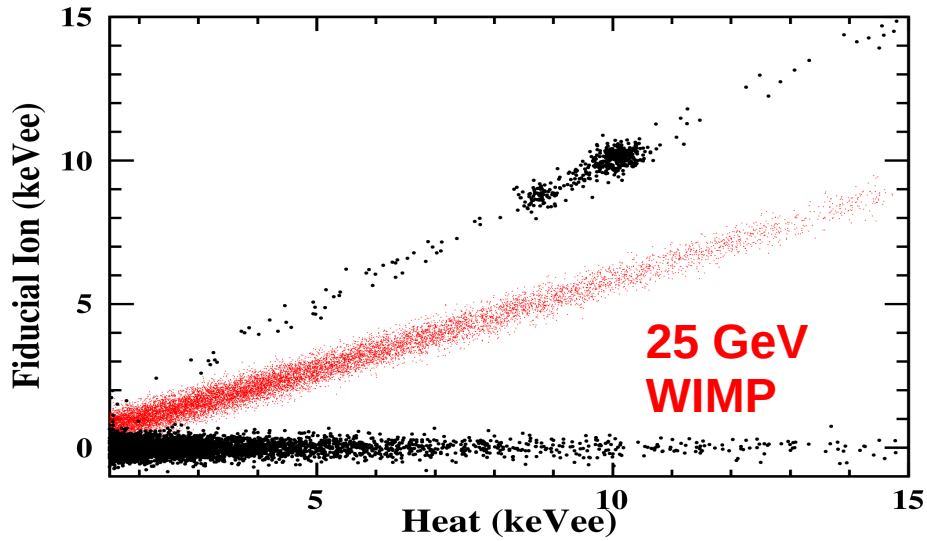
Surface Lead

Heat only

The EDELWEISS-III experiment: Background rejection



- Fiducial gamma
- Surface Gamma
- Surface Beta
- Surface Lead
- Heat only



- We remove surface events using the ionisation channels
- We combine ionisation and heat to discriminate fiducial gamma and heat only versus WIMPs



Part II

First EDWELWEISS-III search for low mass WIMPs



First EDWELWEISS-III search for low mass WIMPs: Foreword

We picked one standard detector:

- Unblinded a small fraction of the data set, Aug – Dec 2014.
- Allows us to prepare the analysis
- Exposure 35 kg.d after cuts.

We define a simple region of interest (ROI):

- In particular Heat >1.5 keVee (3.6 keVNR)

We use Boosted Decision Trees within this ROI :

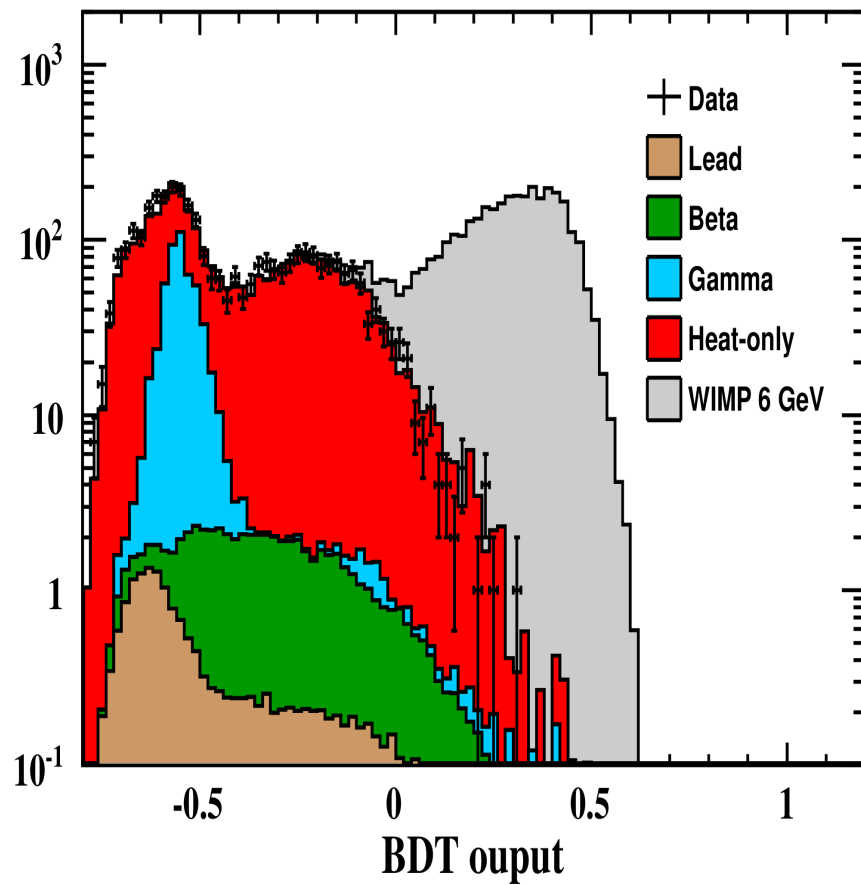
- Combine the 6 variables (4 ionisation and 2 heat) for Signal/Background discrimination

Background models are data driven:

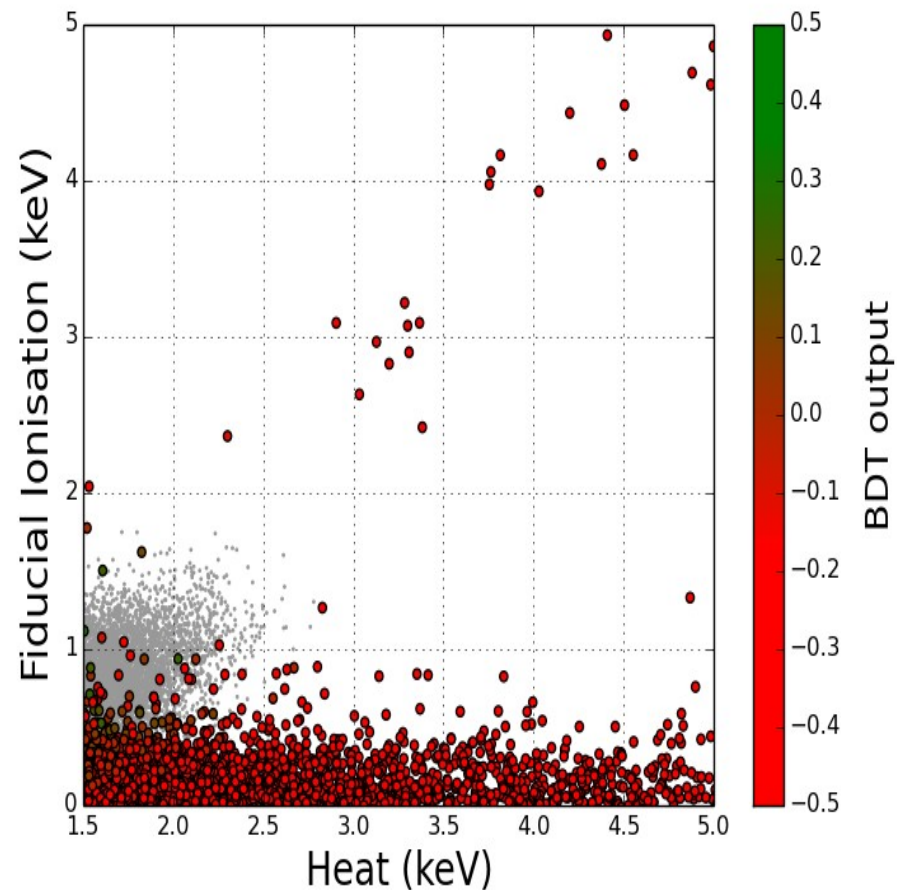
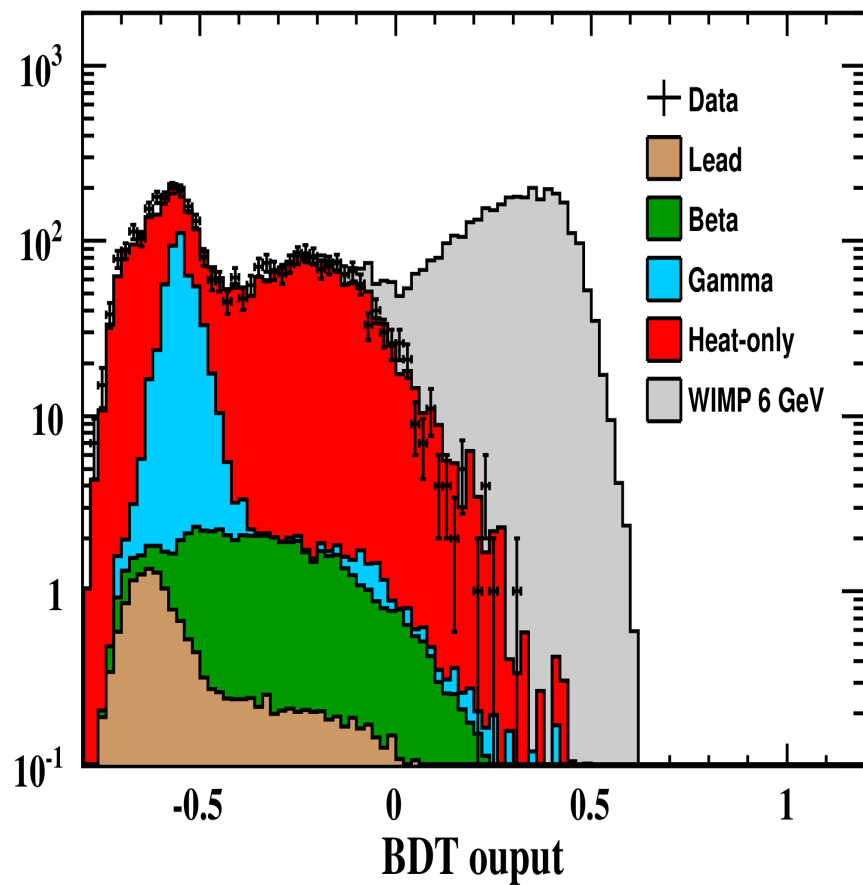
- Use regions without signal (sideband) to build the model
- Use calibrations as crosscheck.



WIMP search data: Boosted Decision Trees

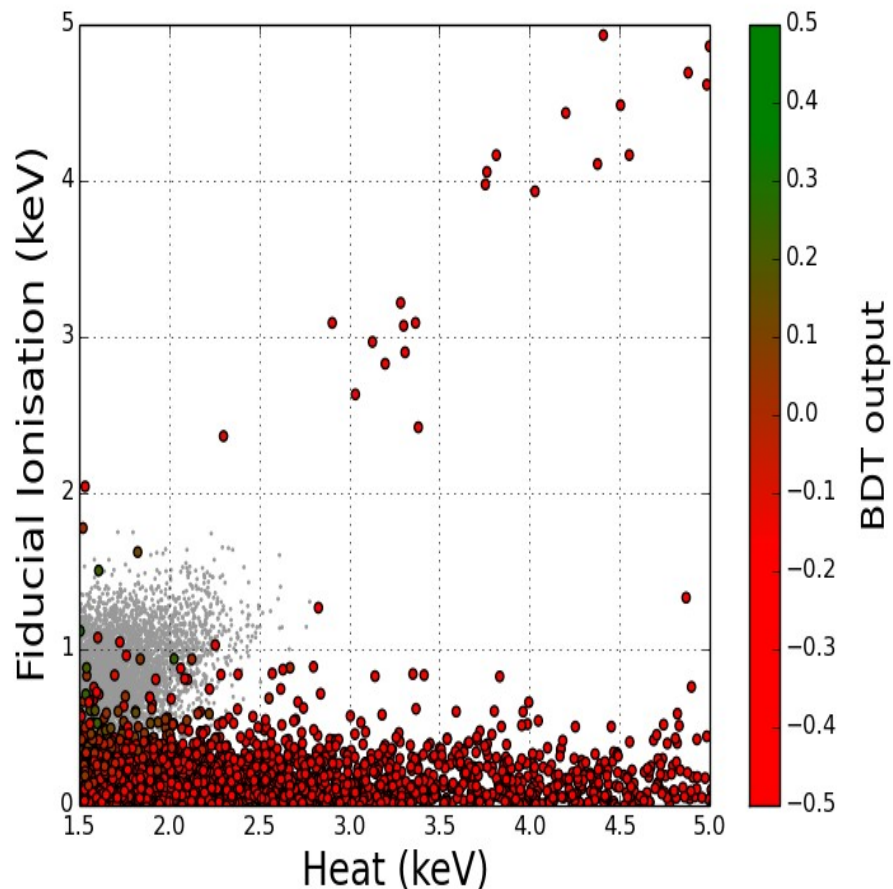
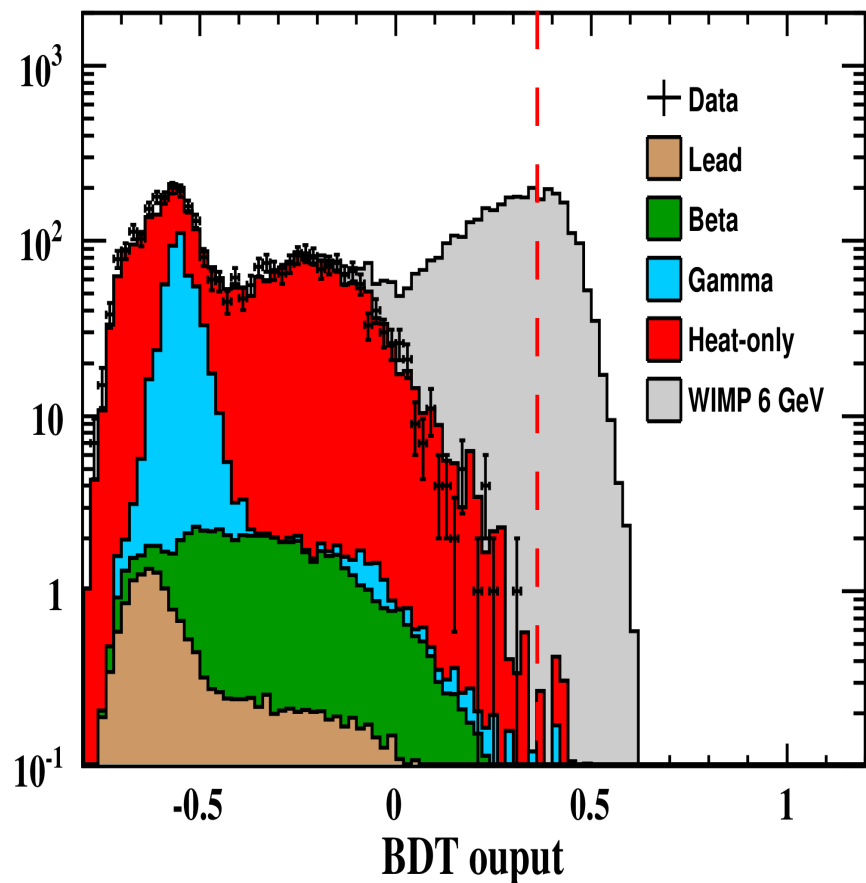


WIMP search data: Boosted Decision Trees

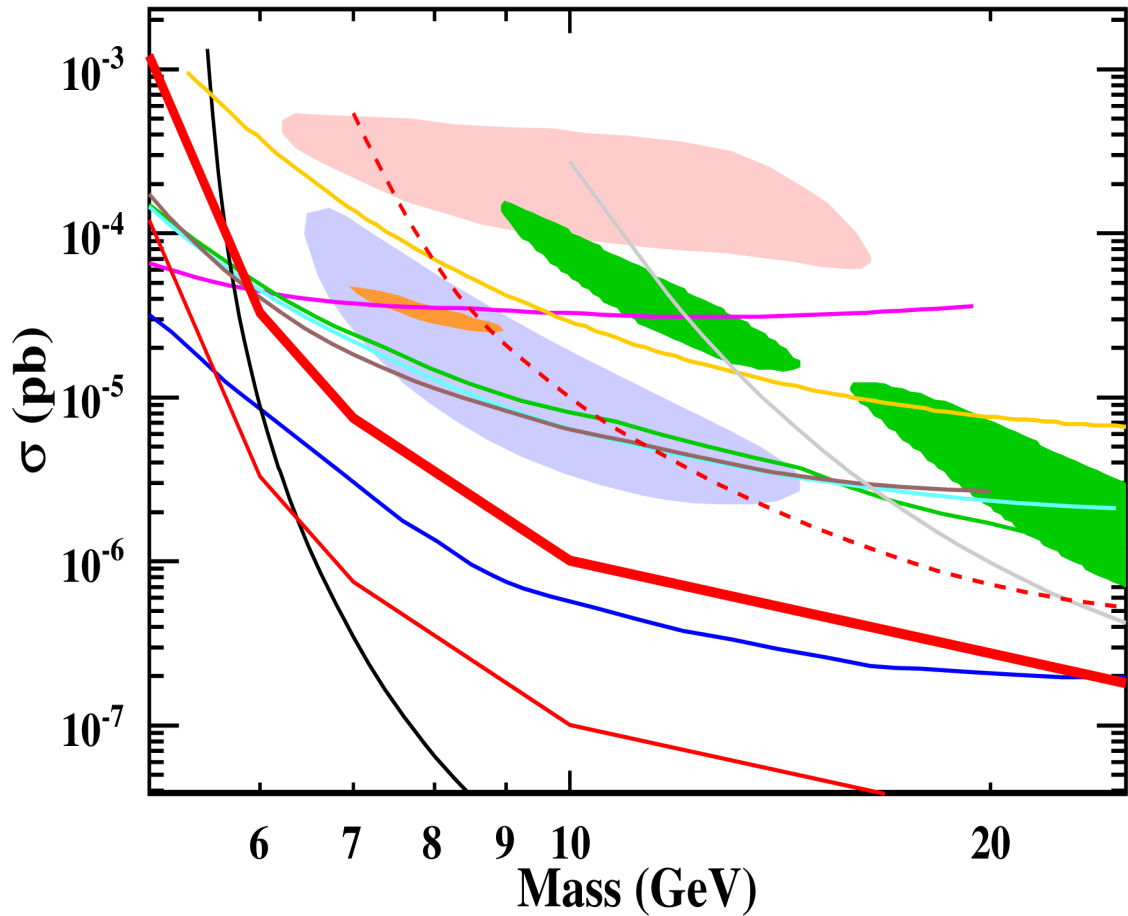


WIMP search data: Boosted Decision Trees

Cut value



First EDELWEISS-III result: low mass exclusion limit



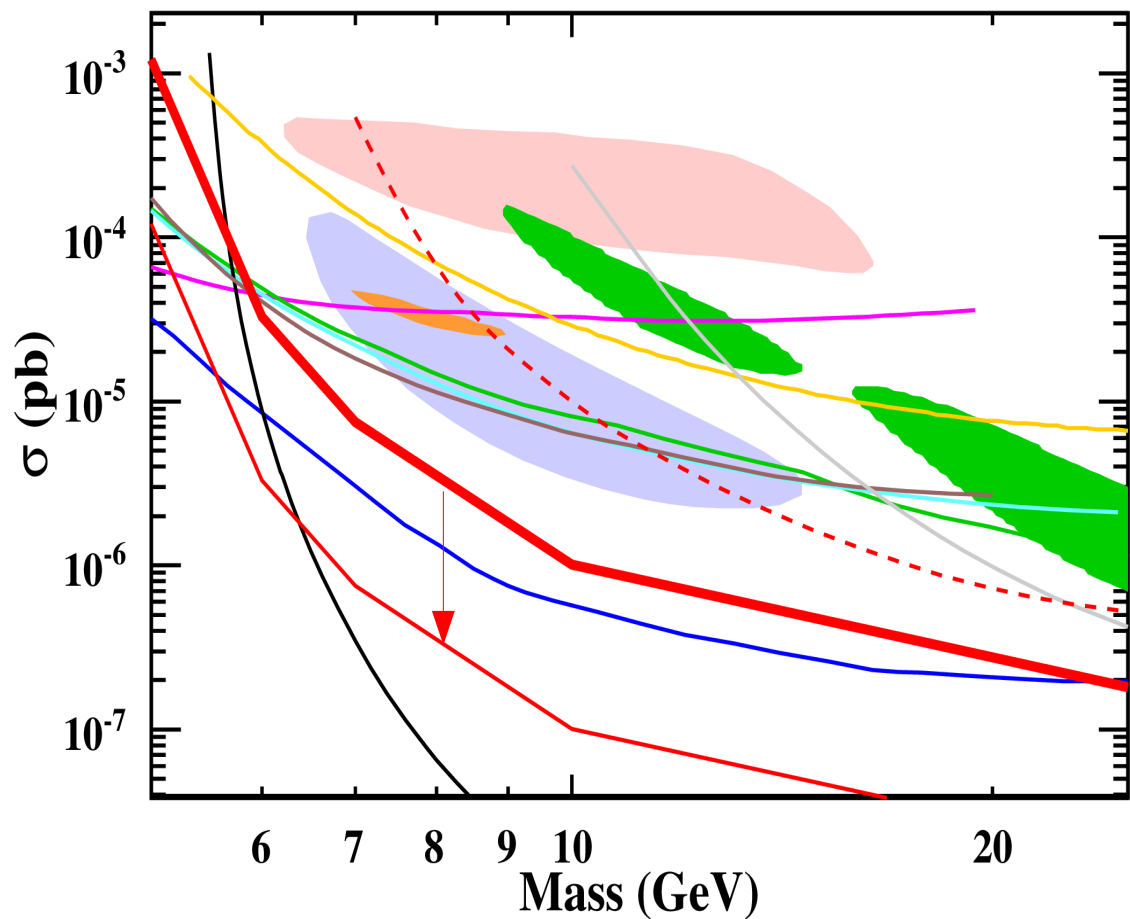
- DAMA
- SCDMS
- CDMSLite
- CDMS Si
- CoGeNT
- LUX
- XENON 10 S2 only
- SIMPLE
- COUPP
- CRESST
- PICO 2L
- EDW II
- EDW III FID837

Limit obtained with a single detector

- Poisson limit after BDT cut
- Analysis threshold: 1.5 keVee (3.6 keVNR)
- 35 kgd exposure after cuts



First EDELWEISS-III result: low mass exclusion limit



- DAMA
- SCDMS
- CDMSLite
- CDMS Si
- CoGeNT
- LUX
- XENON 10 S2 only
- SIMPLE
- COUPP
- CRESST
- PICO 2L
- EDW II
- EDW III FID837

Expect fast improvements in sensitivity:

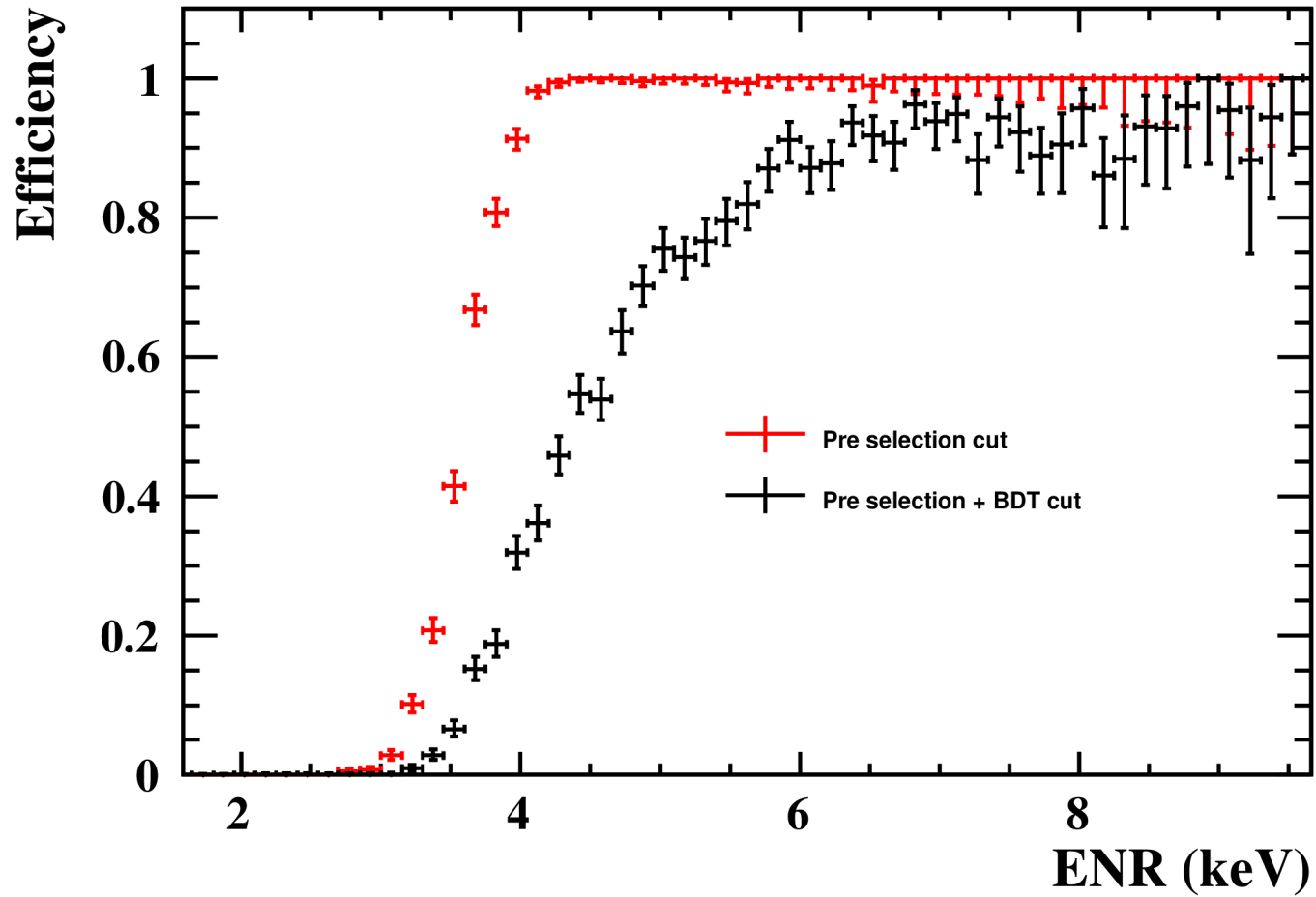
- We already have x10 more data of similar quality
- Run ongoing
- Will decrease the analysis threshold

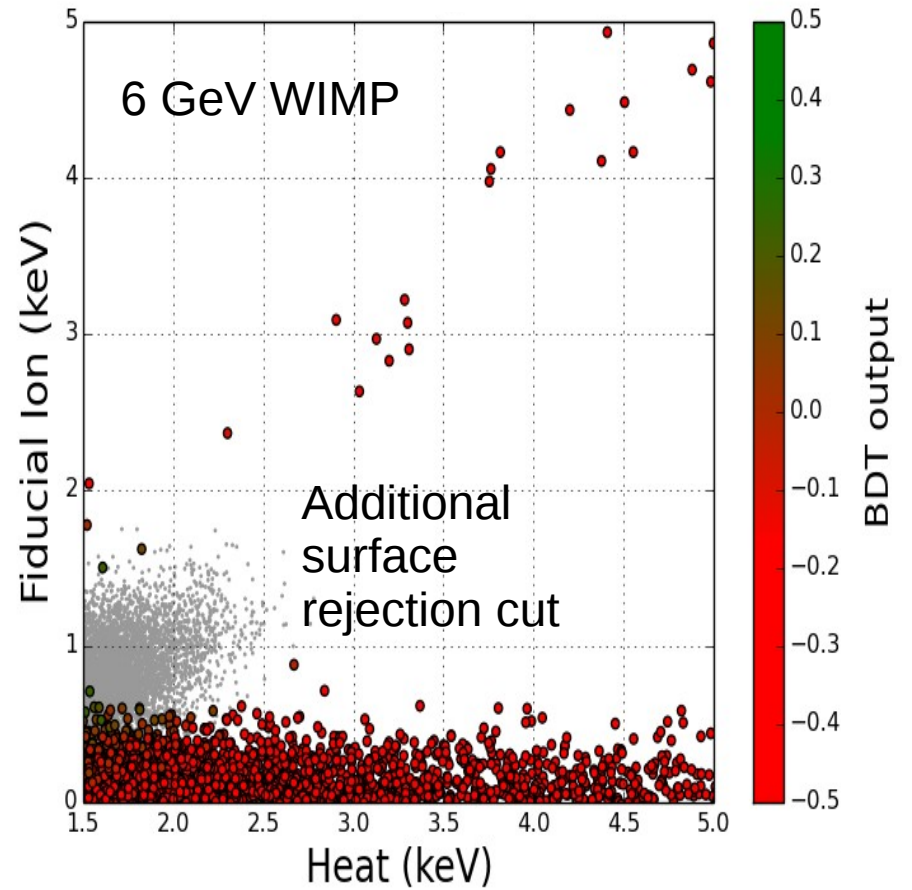
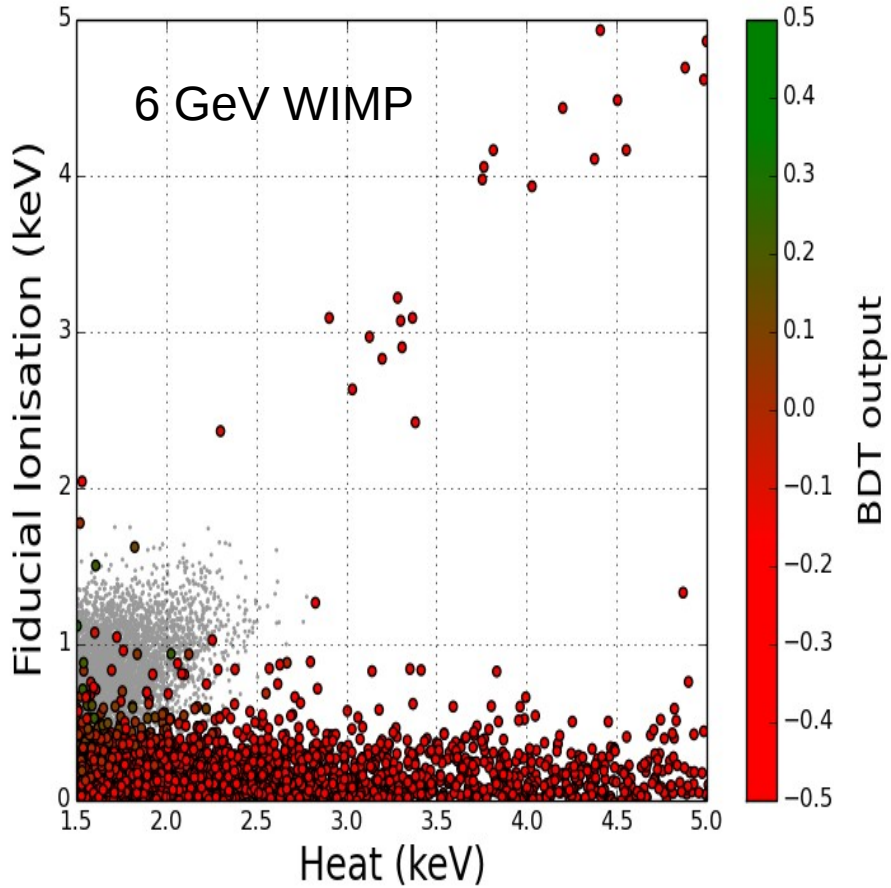


Backup



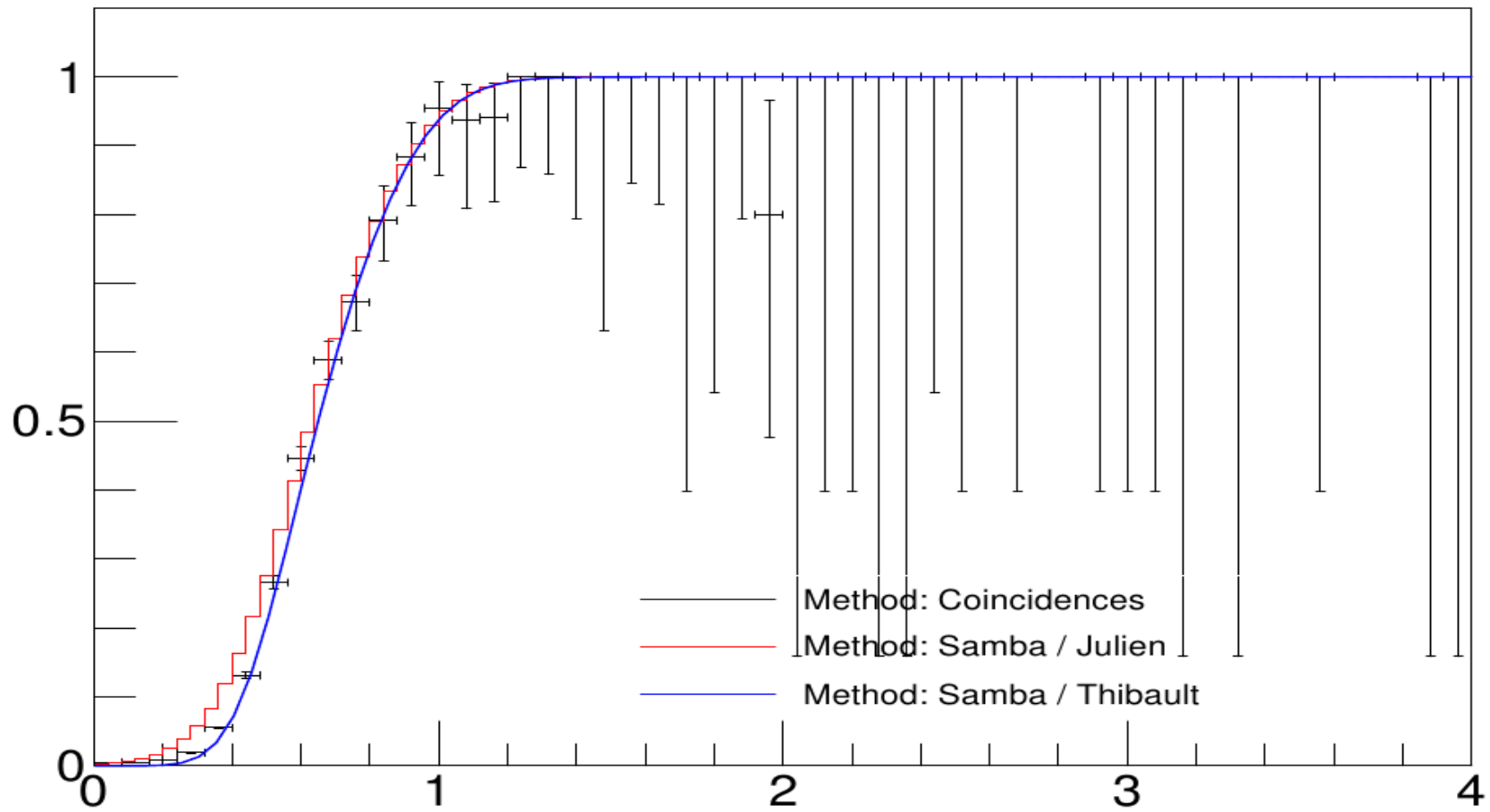
For 10 GeV WIMP

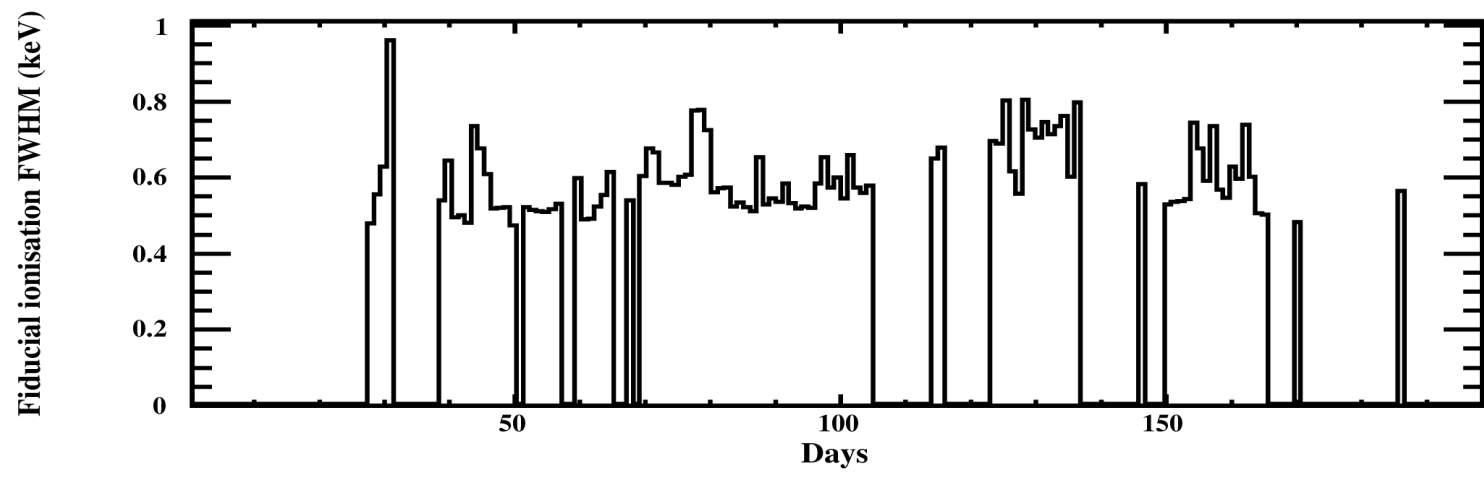
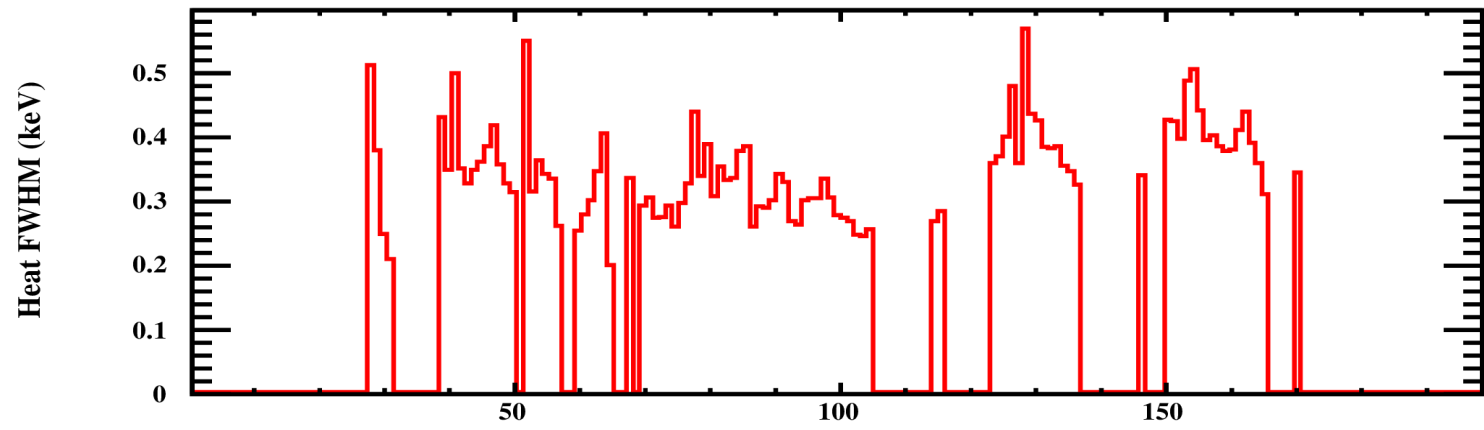




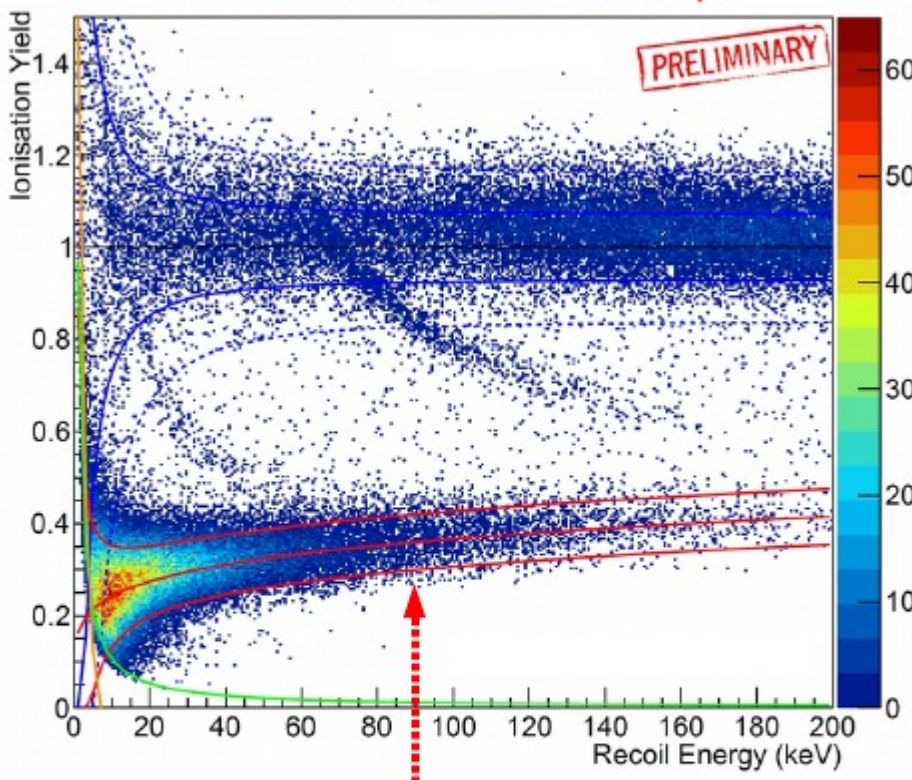


Online trigger efficiency





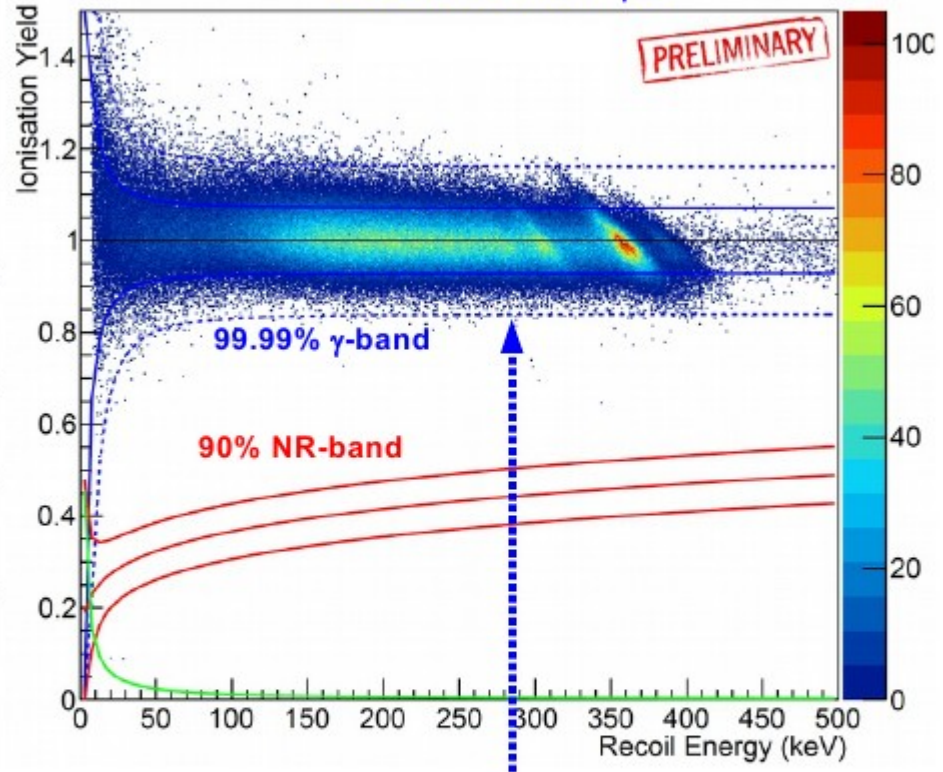
AmBe calibration: n + 60% γ 's



Nuclear Recoils
 $Q \approx 1/3$
 ("quenching")

Ionization yield
 $Q = E_{\text{ion}} / E_{\text{recoil}}$

^{133}Ba calibration: γ 's

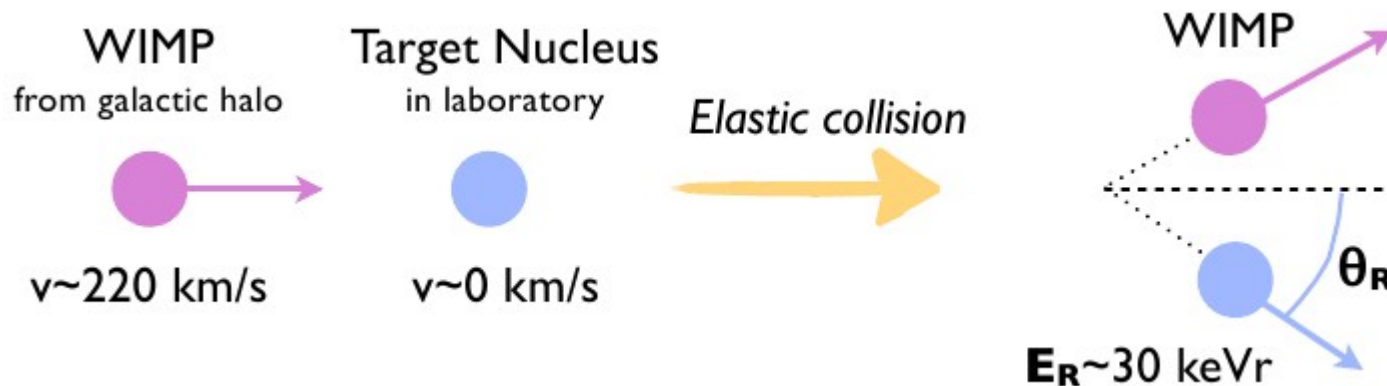


Electron Recoils
 $Q = 1$
 (by normalization)

WIMP direct detection

Brief review

- We have **evidence** for Dark Matter: CMB, Lensing, Galaxy Rotation curves...
- Well motivated **candidate**: WIMP (massive, stable, charge neutral, weakly interacting)
 - It provides the missing matter density in a natural way
- Detection** scheme:



Challenges:

- Many backgrounds
- Low energy scale for the recoil ($\sim \text{keV}$)
- Low event rate ($< \text{evt/kg/year}$)