

# Search of WIMPs with Liquid Argon: The DarkSide experiment

Claudio Giganti





# Outline

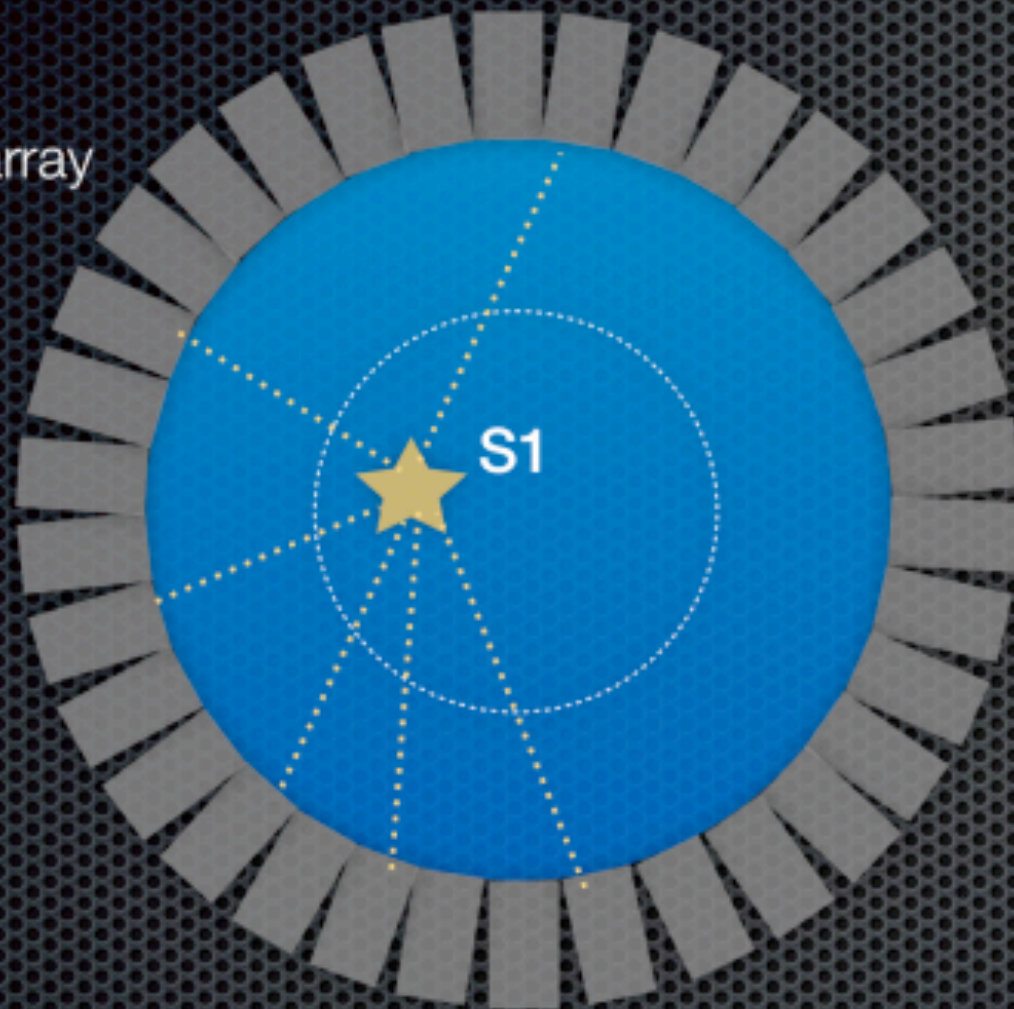
- At our first meeting I introduced the Darkside physics program
  - Search for WIMPs with Liquid Argon detectors
  - In the first phase we are taking data with a 50 kg detector
- Since then there has been few updates that I'll discuss today
  - Darkside-50 released the first paper with dark matter results
  - Definition a plan for the next generation of Darkside detectors
  - French contributions to WIMPs search with Liquid Argon
    - Analysis effort for Darkside-50
    - R&D efforts for Darkside-G2
    - ANR proposal for the study of directionality of Nuclear Recoils with Liquid Argon



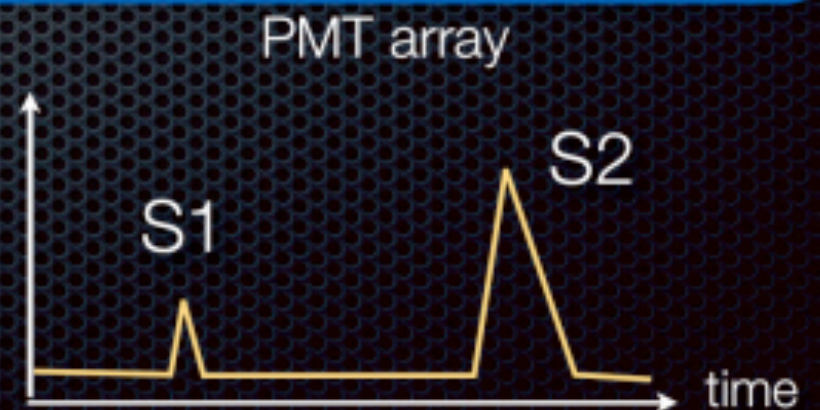
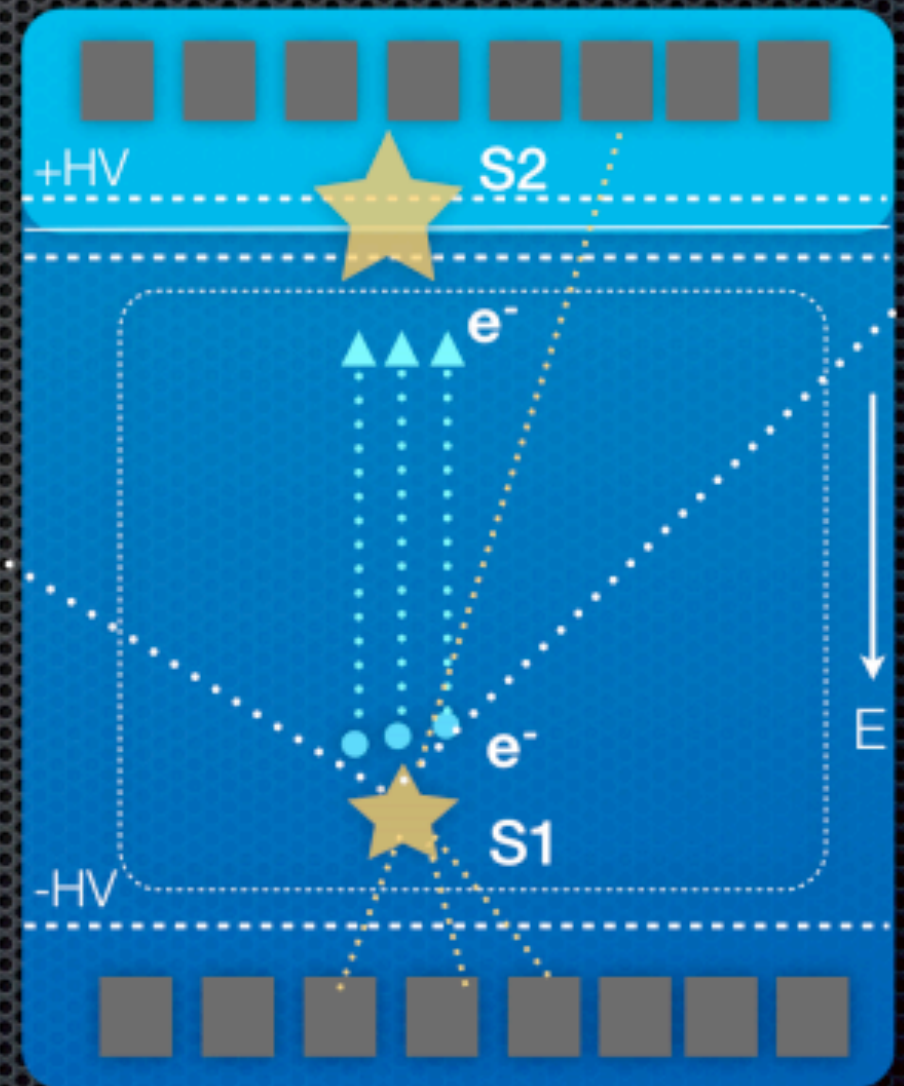
# Noble liquids TPC

Single phase

PMT array



Double phase (TPC)





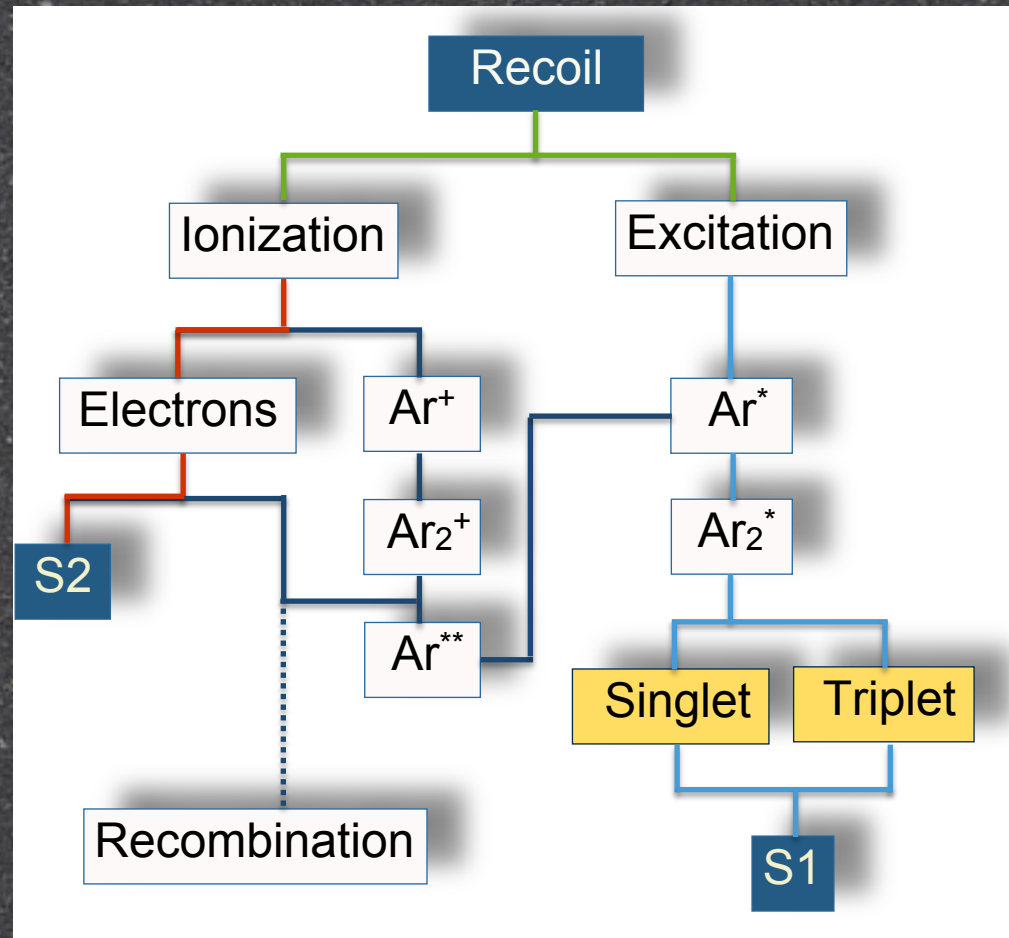
# LAr or LXe

		<i>LAr</i>	<i>LKr</i>	<i>LXe</i>
Physical properties	Atomic number	18	36	54
	Boiling point at 1 bar, $T_b$ (K)	87.3	119.8	165.0
	Density at $T_b$ (g/cm <sup>3</sup> )	1.40	2.41	2.94
Ionisation	$W$ (eV) <sup>1</sup>	23.6	20.5	15.6
	Fano factor	0.11	~0.06	0.041
	Drift velocity (cm/ $\mu$ s) at 3 kV/cm	0.30	0.33	0.26
	Transversal diffusion coefficient at 1 kV/cm (cm <sup>2</sup> /s)	~20		~80
Scintillation	Decay time <sup>2</sup> , fast (ns)	5	2.1	2.2
	slow (ns)	1000	80	27/45
	Emission peak (nm)	127	150	175
	Light yield <sup>2</sup> (phot./Mev)	40000	25000	42000
	Radiation length (cm)	14	4.7	2.8
	Moliere radius (cm)	10.0	6.6	5.7
Excellent discrimination power!				

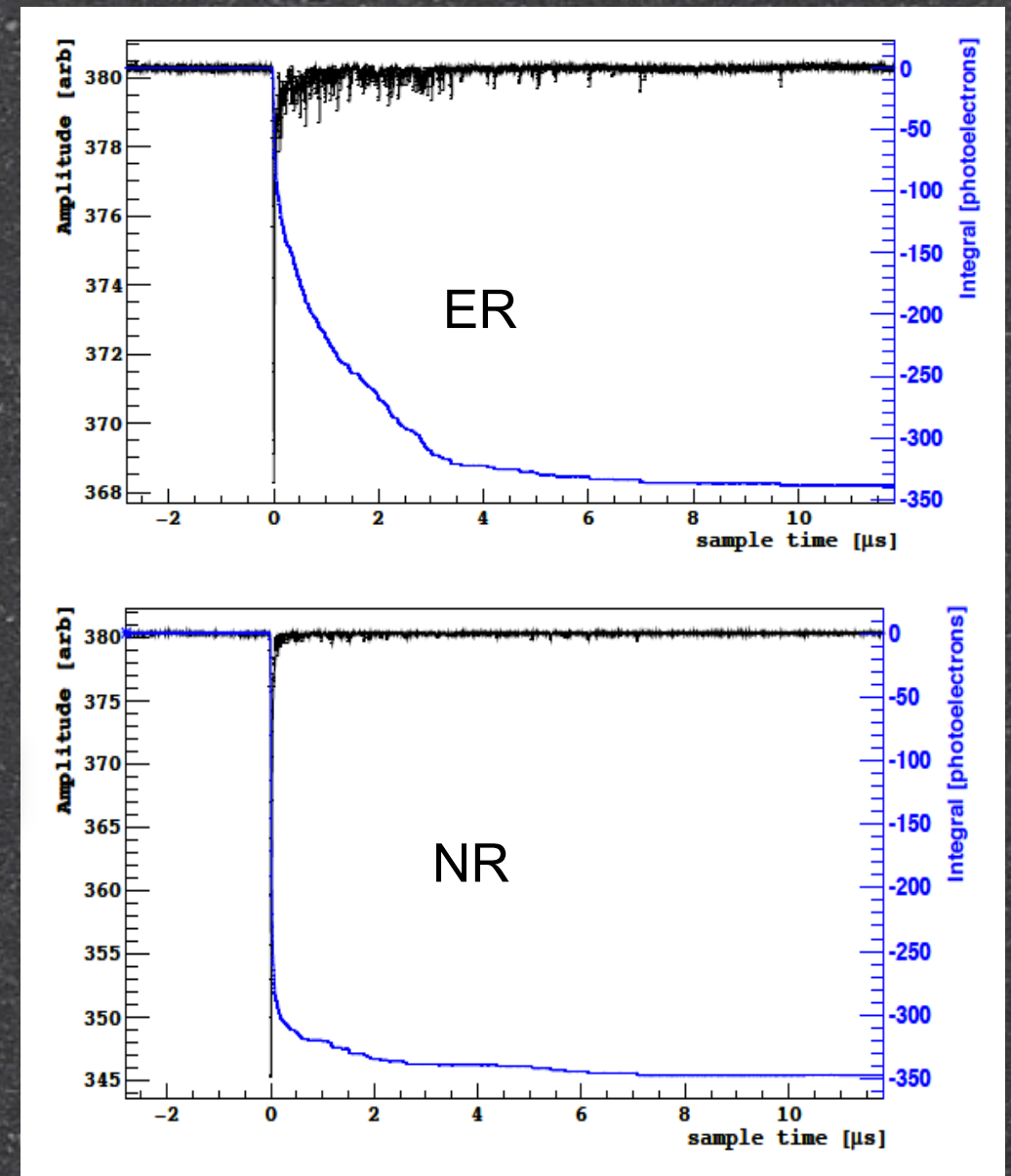
- Liquid Xenon has excellent radio-purity → key ingredient to build large detectors
- Liquid Argon has much better PSD but a serious problem
  - Cosmogenic <sup>39</sup>Ar in atmospheric argon → high rate  $\beta$  emitter → pile-up if you want to build large detectors



# Pulse Shape Discrimination



- Fast decay time (Singlet) ~ 7 ns
- Slow decay time (Triplet) ~ 1600 ns
- NR: ~70% of the energy goes in the singlet → large f<sub>90</sub>
- ER: ~30% of the energy goes in the singlet: small f<sub>90</sub>



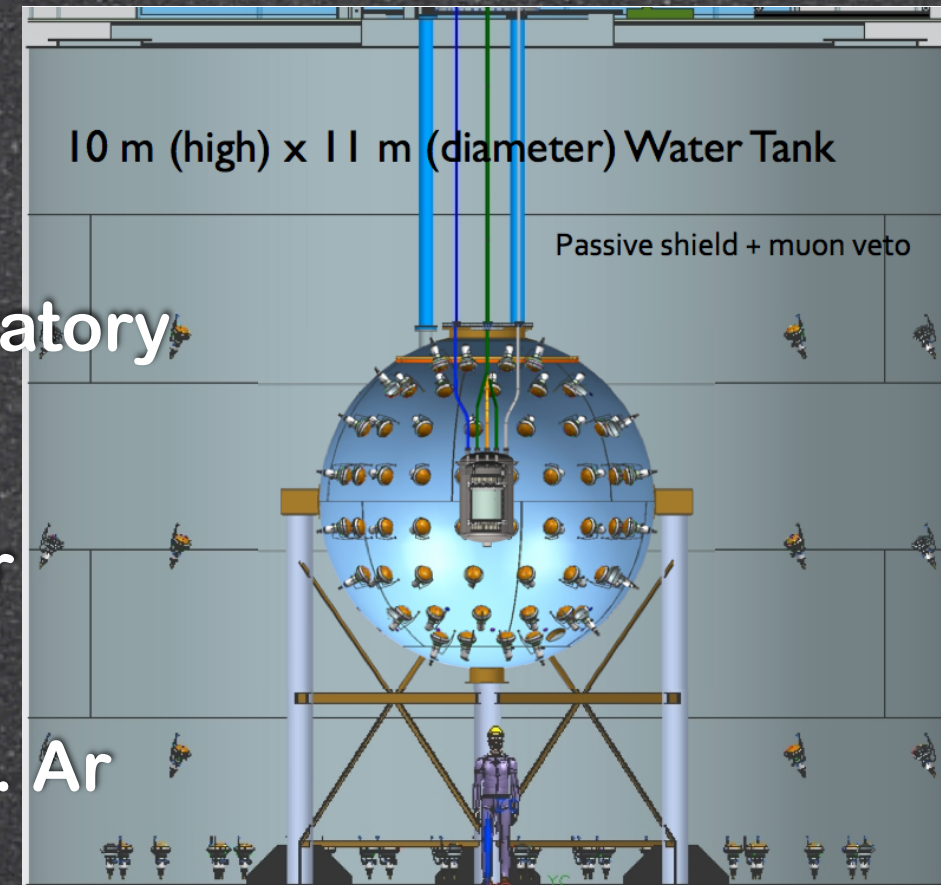
f<sub>90</sub>: Q(0-90 ns)/Q(all)

**Rejection factor > 10<sup>8</sup>**  
 another factor 10<sup>2</sup> from S1/S2



# DarkSide-50

- Experiment installed in the Gran Sasso Laboratory
- Double phase TPC with 50 kg of liquid Argon
- 2 vetoes system: Liquid Scintillator and Water Cherenkov
- Started data taking in January 2014 with Atm. Ar



## Background reduction

Depleted Underground Argon  
Low background materials  
Active Shields against neutrons  
and muons

## Background identification

Pulse Shape Discrimination  
S1/S2 discrimination  
Measure neutron flux in borate  
scintillator  
Position reconstruction

Demonstrate the potential of the  
technology for multi ton  
**background-free** detector



# DarkSide Collaboration

- Collaboration ~ 100 people
- Mainly from US and Italy

## Ukraine

KINR, NAS Ukraine – Kiev

## CHINA

IHEP – Beijing

## POLAND

Jagiellonian University – Krakow

## FRANCE

Université Paris Diderot, CNRS/IN2P3, CEA/IRFU, Observatoire de Paris, Sorbonne Paris Cité – Paris  
IPHC, Université de Strasbourg, CNRS/IN2P3 – Strasbourg

## USA

Augustana College – SD  
Black Hills State University – SD  
Fermilab – IL  
Princeton University – NJ  
SLAC National Accelerator Center – CA  
Temple University – PA  
University of Arkansas – AR  
University of California – Los Angeles, CA  
University of Chicago – IL  
University of Hawaii – HI  
University of Houston – TX  
University of Massachusetts – MA  
Virginia Tech – VA

## ITALY

INFN Laboratori Nazionali del Gran Sasso – Assergi  
Università degli Studi and INFN – Genova  
Università degli Studi and INFN – Milano  
Università degli Studi Federico II and INFN – Napoli  
Università degli Studi and INFN – Perugia  
Università degli Studi Roma Tre and INFN – Roma

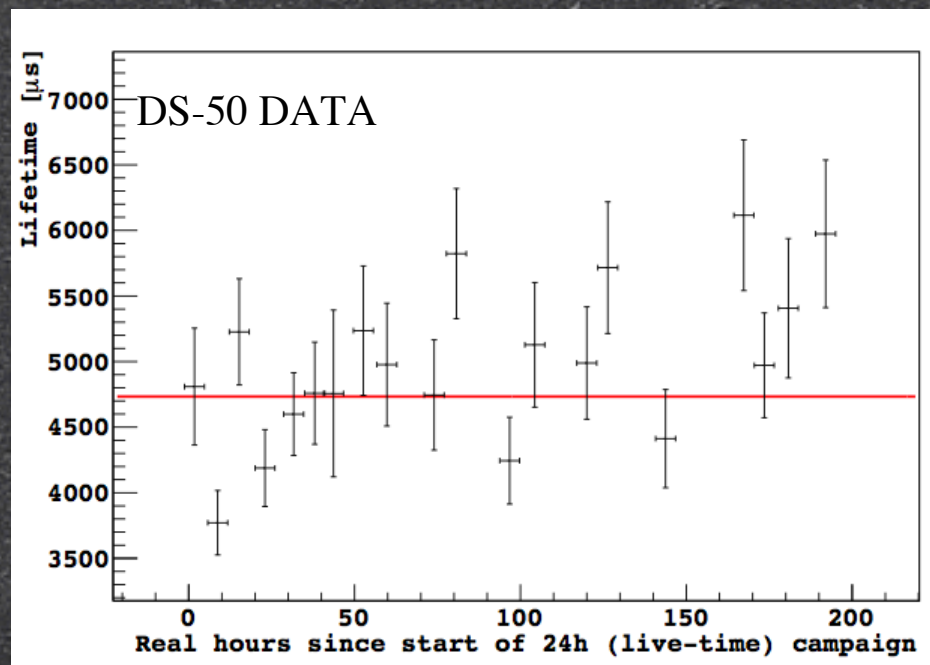
## RUSSIA

Joint Institute for Nuclear Research – Dubna  
Lomonosov Moscow State University – Moscow  
National Research Centre Kurchatov Institute – Moscow  
Saint Petersburg Nuclear Physics Institute – Gatchina

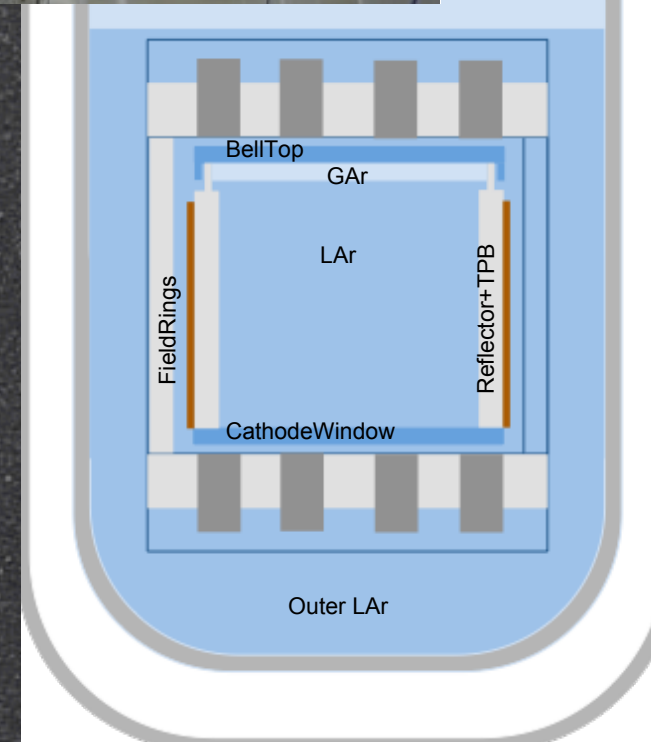


# TPC

- 50 kg active mass of UAr (37 kg FV)
- 19 top + 19 bottom High Quantum Efficiency 3'' PMTs (R11065)
- 36 cm height, 36 cm diameter
- All inner surfaces coated with TPB (used to shift wavelength of Ar scintillation from 128 nm to 420 nm)
- Large electron life-time ( $> 5$  ms)



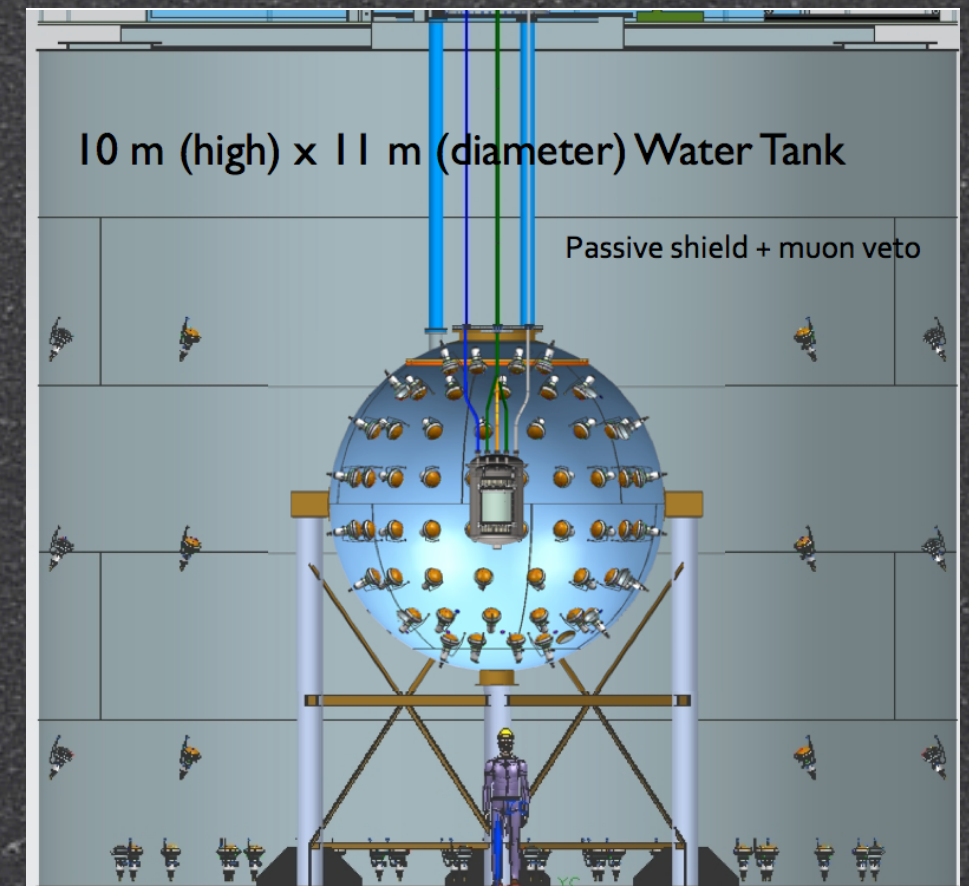
High purity of Argon  
Stable operations of  
electric fields





# Active vetoes

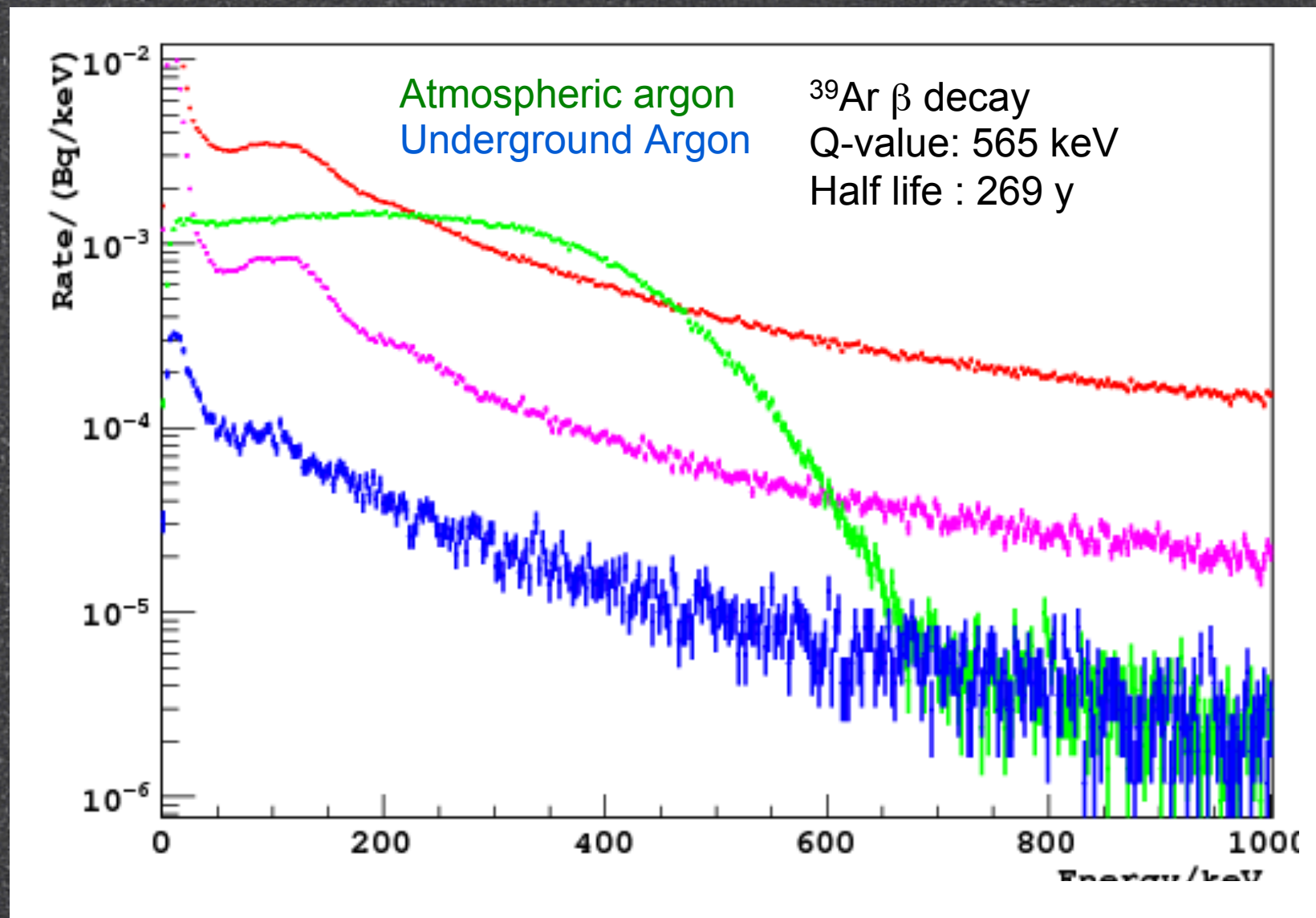
- **Liquid Scintillator Neutron veto**
- 30 tonnes boron-loaded liquid scintillator detector
  - Readout with 110 low-radioactivity PMTs
  - 2 m radius sphere
- Passive shield against neutrons and gamma
- Tag neutrons from TPC through n-capture to measure the neutron flux
- **Muon veto**
  - 1 kton ultra pure water
  - 10 m height, 11 m diameter
  - 80 PMTs
  - Tag cosmogenic neutron events



**Both vetoes are designed to host DarkSide-G2 (5 ton TPC)!**



# Underground Argon



- $^{39}\text{Ar}$   $\beta$ -decay with a rate of  $\sim 1$  Hz/Kg in Atm. Ar
  - Even if you can distinguish ER from NR with PSD it's impossible to build large detectors with AAr due to pile-up
- Solution: use Underground Argon  $\rightarrow$  factor of  $>150$  of depletion



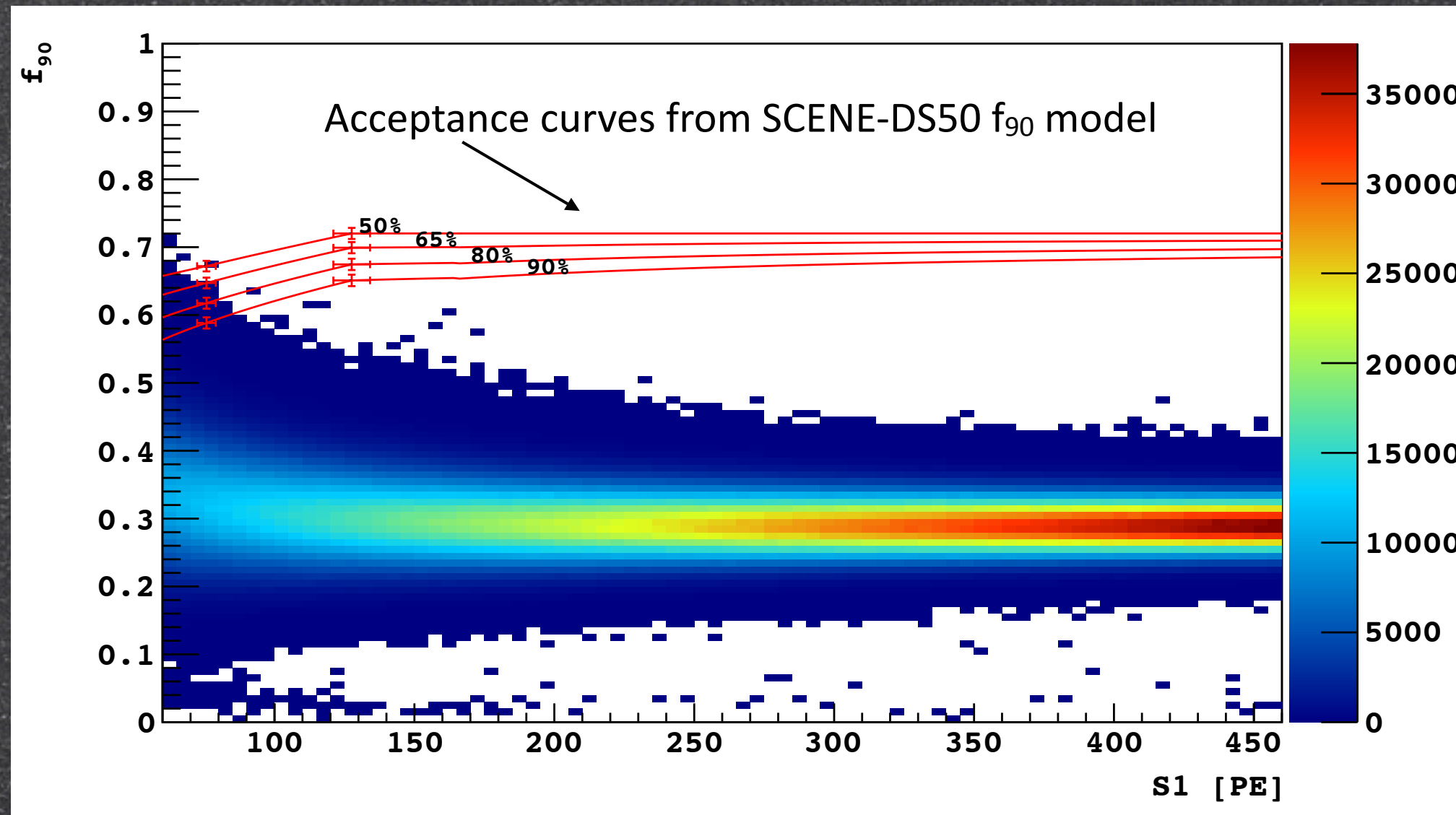
# DarkSide status

- ~2 months of data taking with Atmospheric Argon
- Huge statistics of ER to demonstrate background free operations
  - Number of  $^{39}\text{Ar}$  corresponding to two decades of DarkSide-50 with UAr!
- Release first physics results with this data set  
<http://arxiv.org/pdf/1410.0653v1.pdf>
- Replace Atmospheric Argon with Underground Argon in December
- Start 3 years run with UAR
- Move to Darkside-G2 (ton-scale LAr) in the future



# Background free

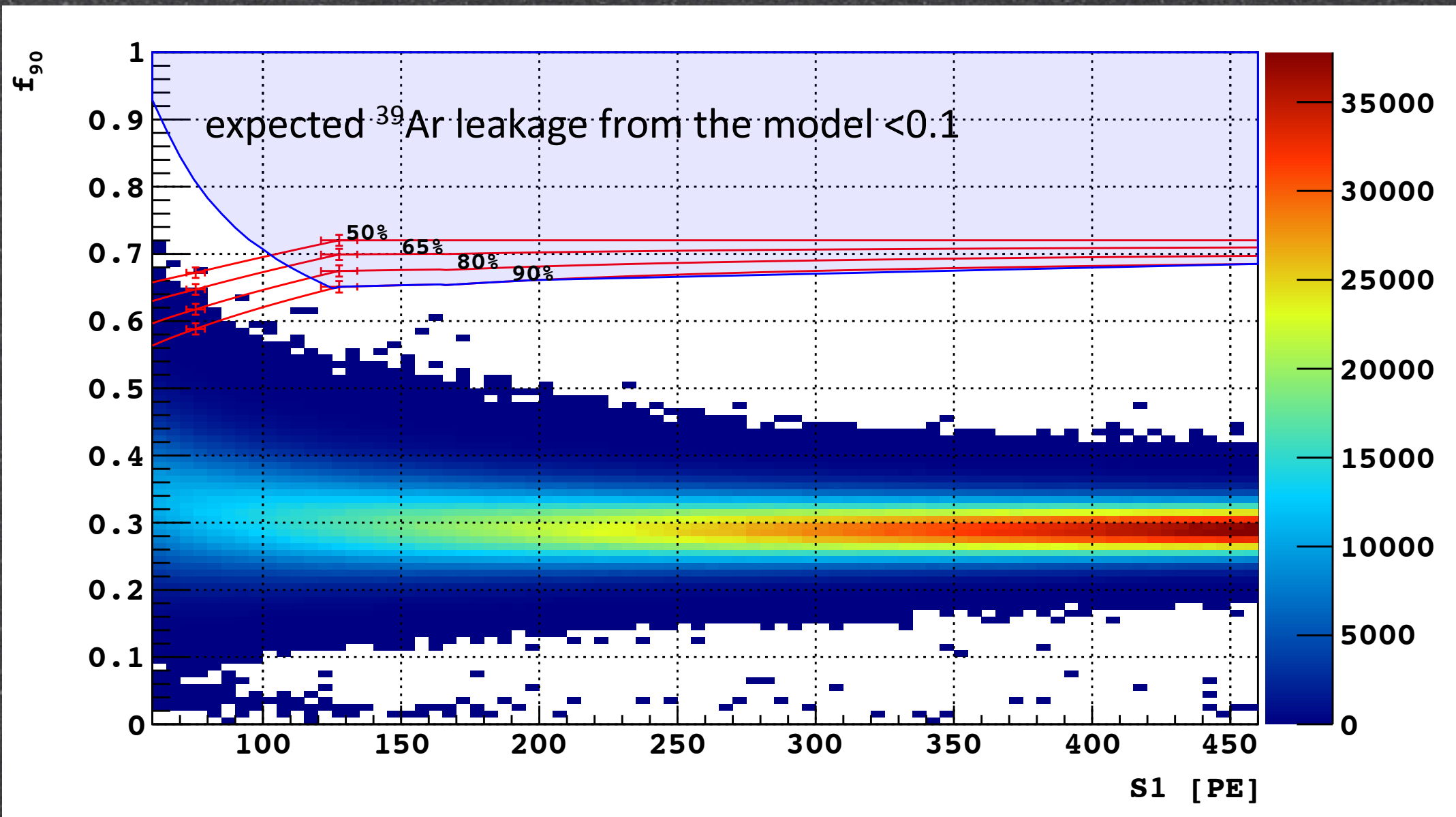
- 2 months of data with AAr  $\rightarrow$   $> 20$  years of UAr
- $15 \times 10^6$  ER selected and none of them has a pulse shape compatible with a NR above





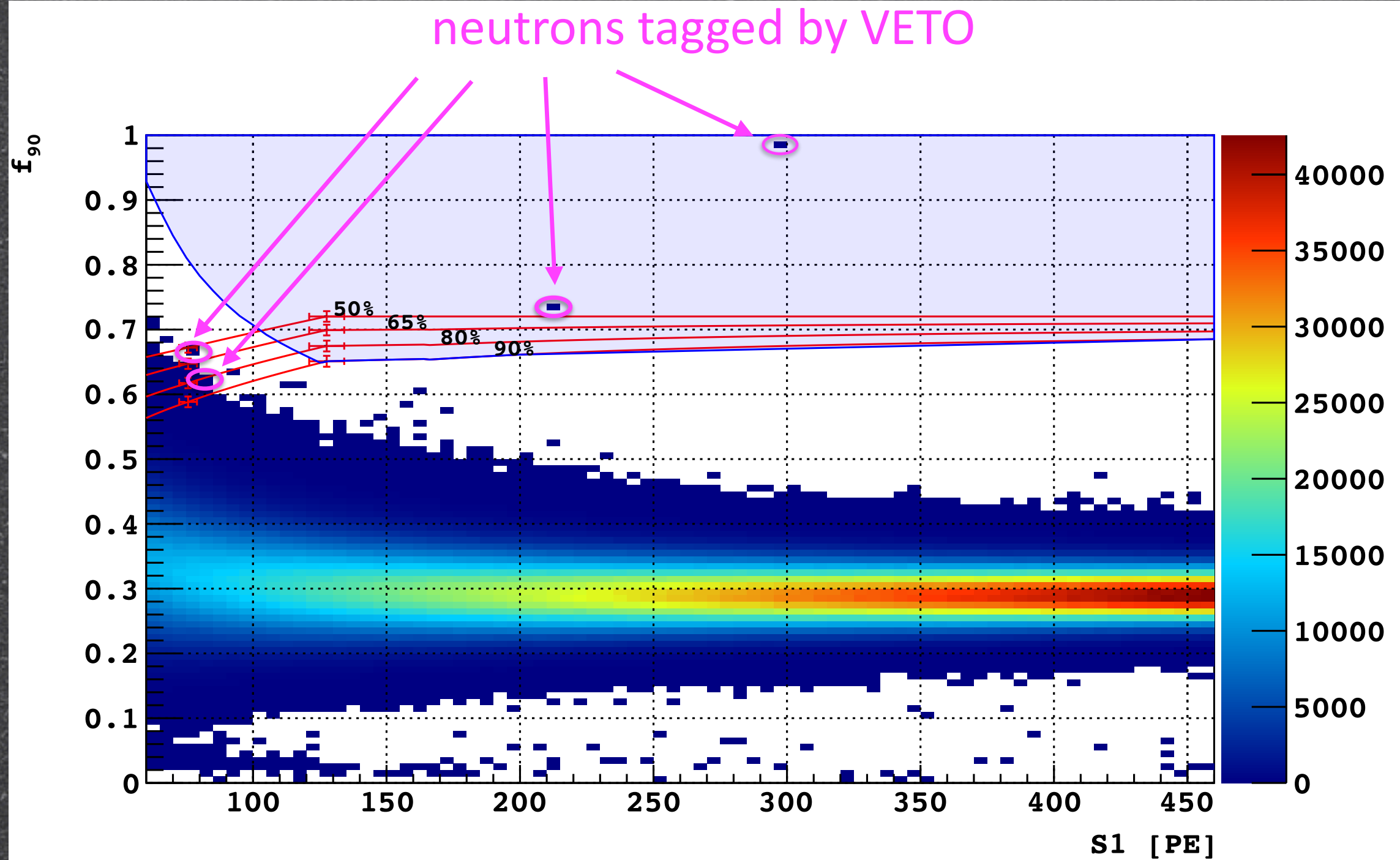
# WIMP search

- Define the WIMP search region as the region in which there are less than 0.1 ER expected into the search region with the current data-set



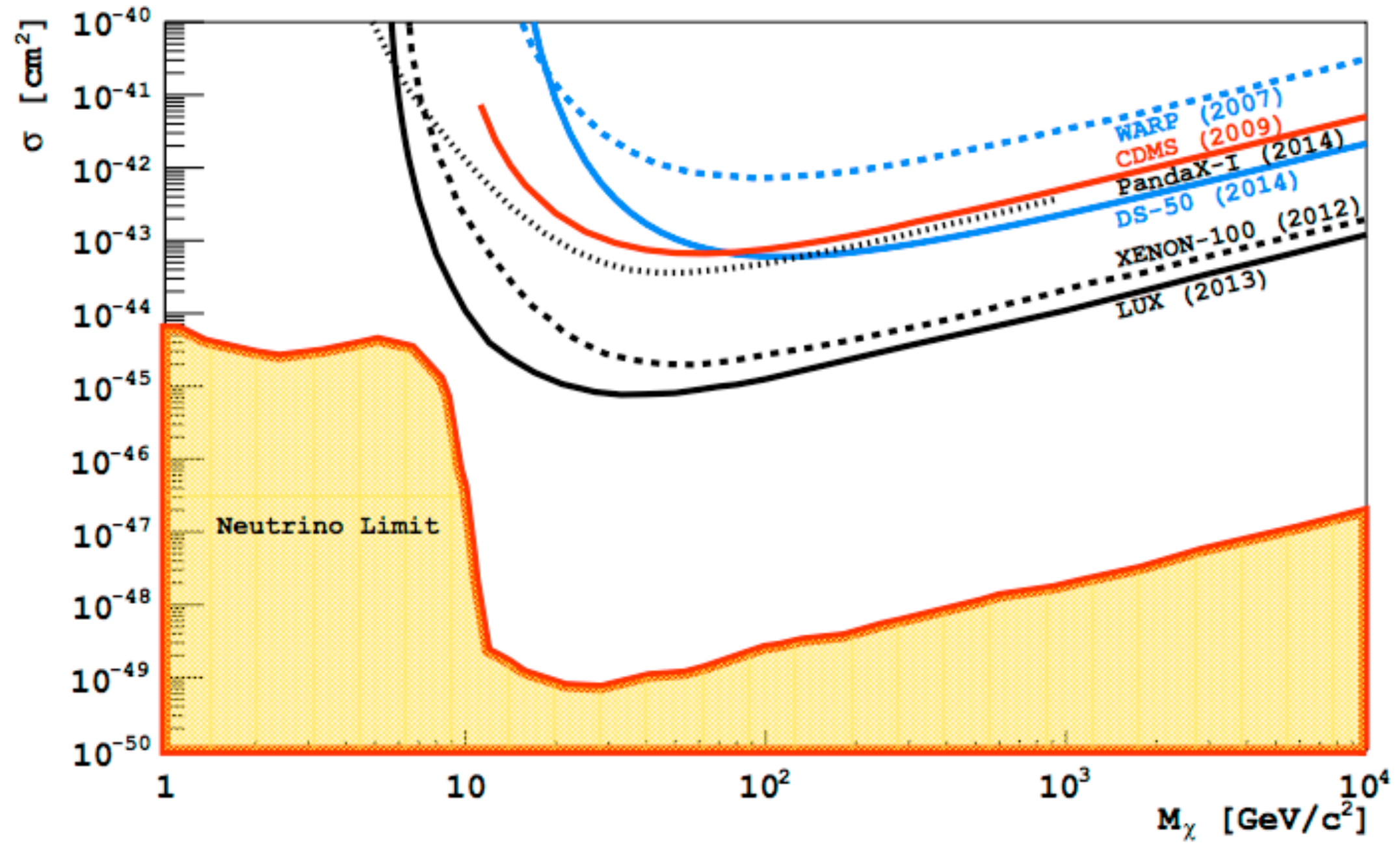


# Nuclear recoils



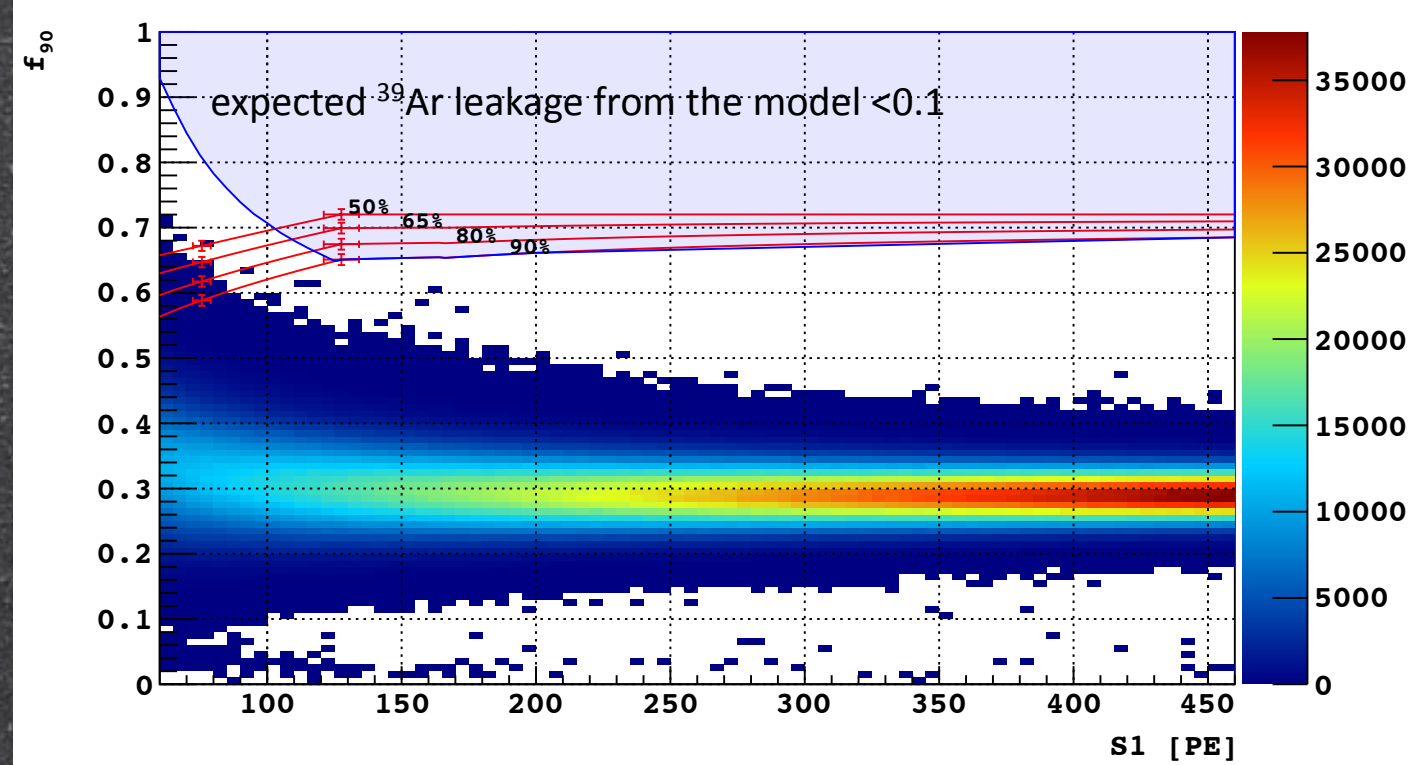
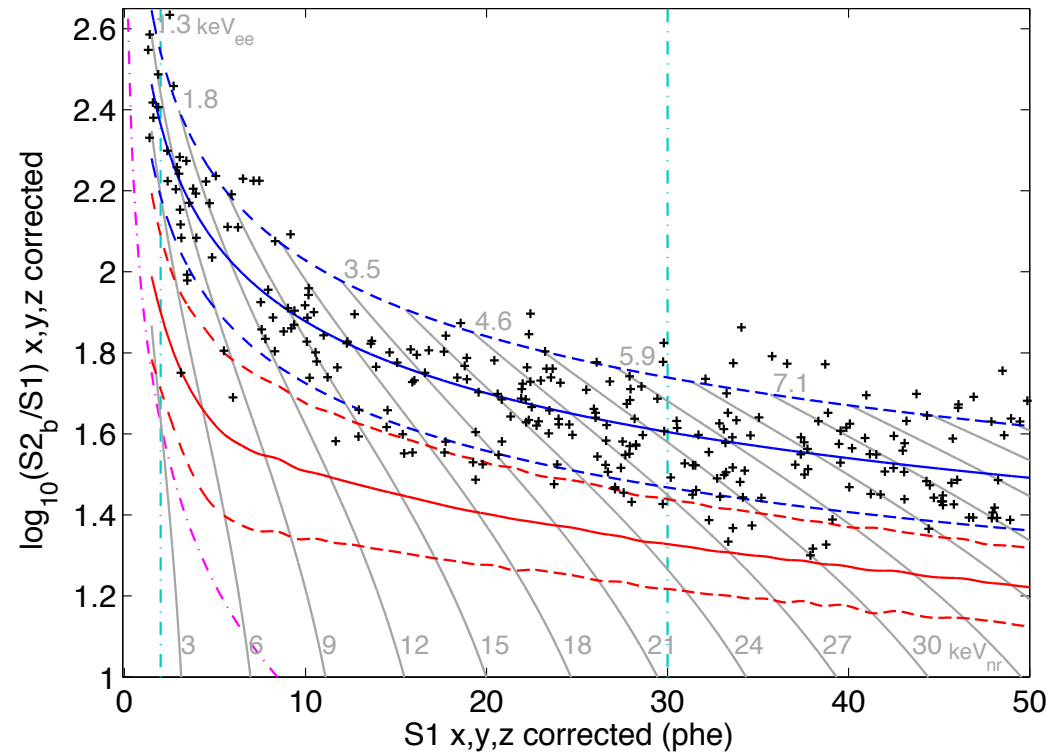


# WIMPs limits





# PSD in LXe and LAr

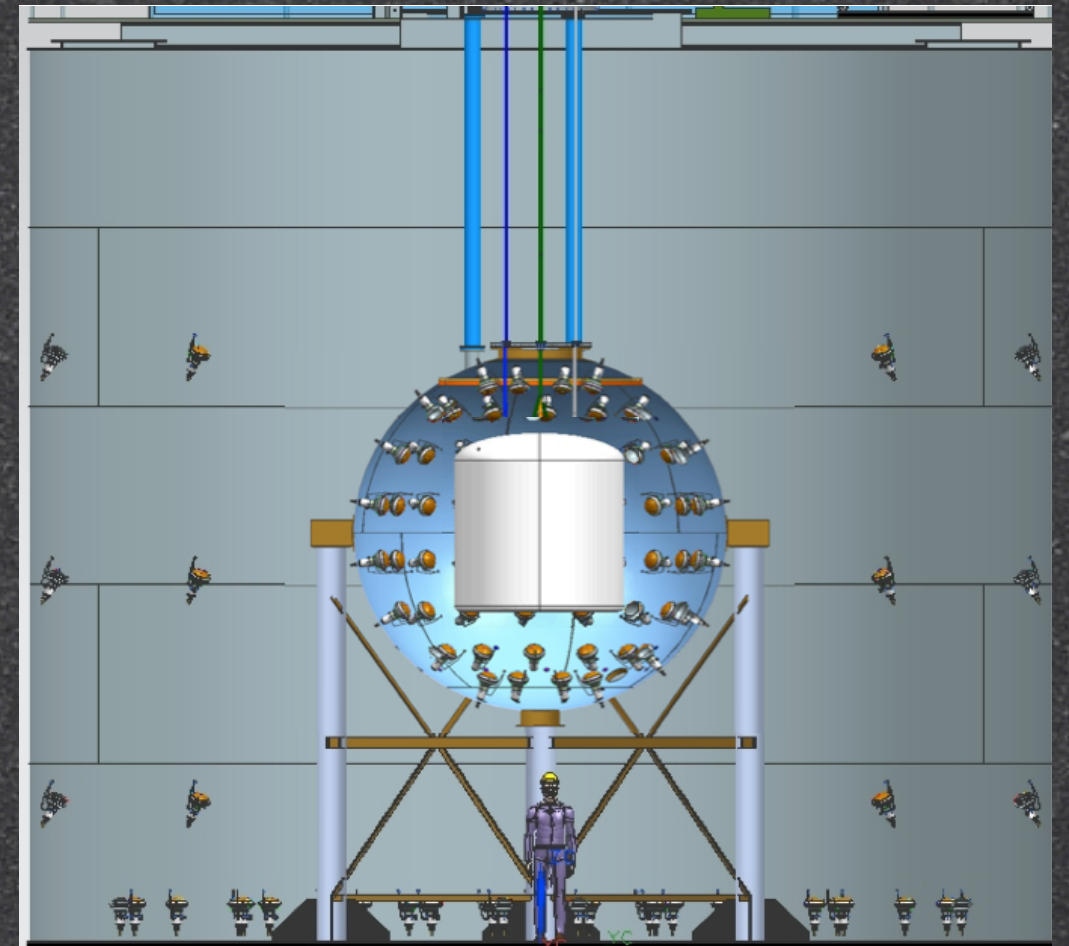


- LXe → ER close to the NR mean
- LAr → no ER leaking in the NR region (not even in the 90% acceptance line for the NR)
  - These results were obtained with 50 days of Atm Argon → >20 years of data with Underground Argon
  - Zero-background detector!



# Darkside-G2

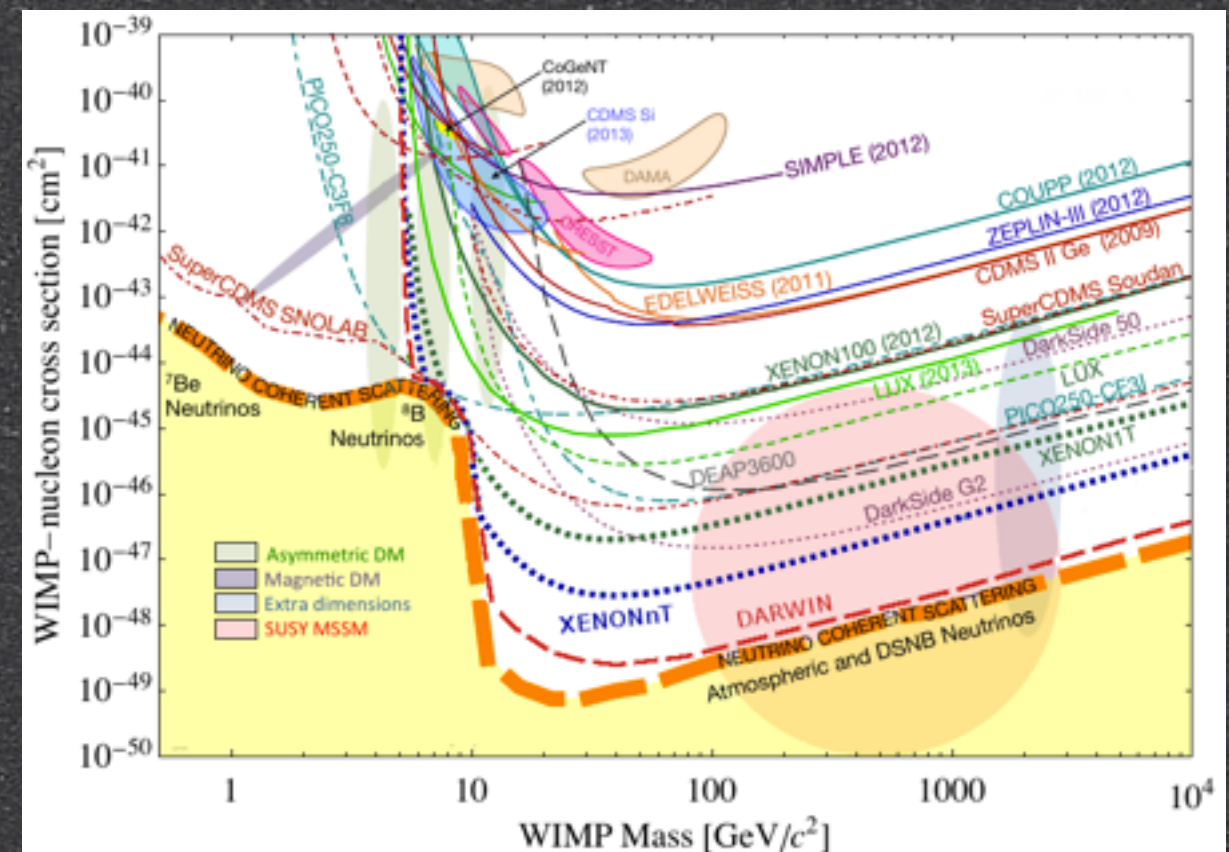
- Results presented by Darkside gave very good boost to the Darkside operations with funding agencies
- Ton-scale TPC to be installed inside the same veto systems currently used by DS-50
- Replace PMTs with SiPM → cheaper, cleaner, smaller
- Ask for funding mainly to INFN + contribution from DOE (depleted Argon)
- Discussions will be done in the next few months, start building in the next couple of years
  - Good time to join





# LAr on long term

- The Xenon technology still has the better limits for WIMPs
- But the Argon technology proved to be free of electron recoils while keeping a 90% acceptance for NR
  - Thanks to the PSD in Liquid Argon
  - LUX have some backgrounds even keeping a 50% acceptance for NR
- Necessary to continue the Argon program in view of the next generations detectors
- LAr is certainly an interesting technology for the ultimate WIMP detector (Darwin?)





# Joining Darkside

- The APC and IPHC groups are already part of Darkside-50
- We have recently joined the collaboration and in general they would be very happy to have additional collaborators
  - Some discussion with Cristian Galbiati (spokesman)
  - We can contribute to the analysis of Darkside-50 data
  - Contributions to R&D for the next phases of this program → SiPM and characterization of the detector
- Meeting in Milan the 11th and 12th of December to strengthen the European part of the collaboration
  - We plan to attend this meeting (me and Sandro)



# French laboratories in DS

- 2 laboratories are part of the DarkSide collaboration (APC, IPHC)
  - APC: Davide Franco, Alessandra Tonazzo, Stefano Perasso (post-doc), Paolo Agnes (PhD)
  - IPHC: Anselmo Meregaglia, Cecile Jollet
- Their main contribution was to write the GEANT4 based MC simulation of the experiment → used for the on-going analyses
- We are now working on the analysis of DS-50
  - Luca Agostino is working on the analysis of DarkSide
  - Analysis of the data from the on-going calibration campaign
  - Analysis of the data with Underground Ar next year
- Participate to the necessary R&D for the next generation
  - Sandro and Stefano are interested in these developments



# French laboratories in DS

- 2 laboratories are involved in the DarkSide collaboration (APC, IPHC)
  - APC: Davide D'Amico (post-doc), Paolo A. Nardulli (post-doc)
  - IPHC: Anselmo Delgado (post-doc)
- Their main contribution is the development of a ROOT based MC simulation and the analysis of the experimental data
- We are now working on the development of a new DarkSide collaboration campaign
  - Luca Agostini (post-doc)
  - Analysis of the data from the first campaign
  - Analysis of the data from the second campaign
- Participate to the next generation of experiments
  - Sandro and Stefano

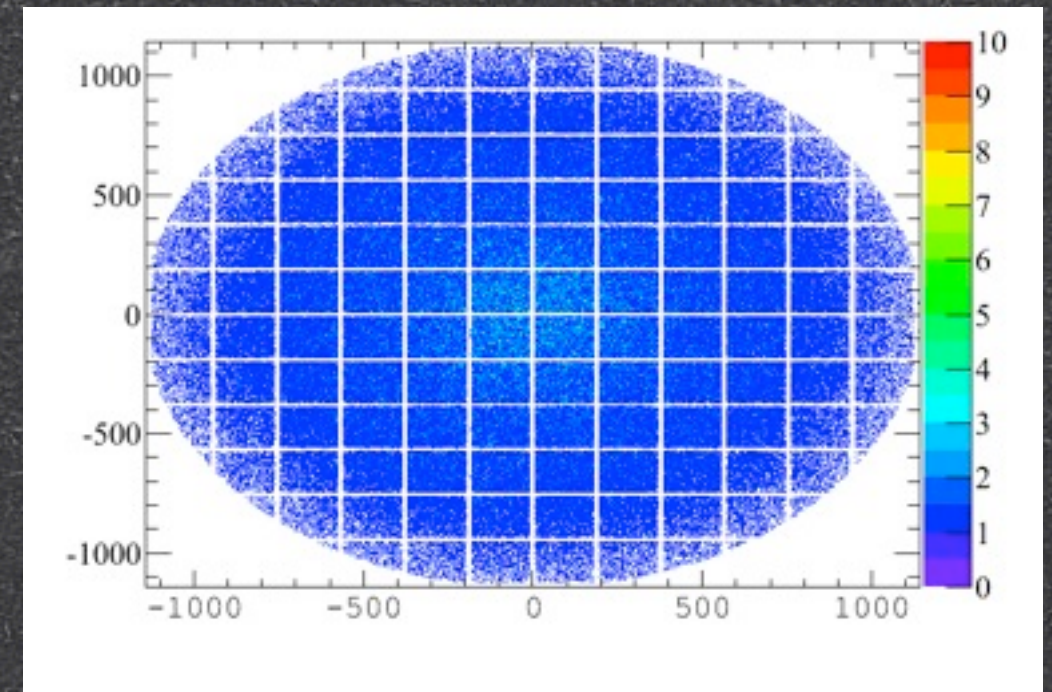




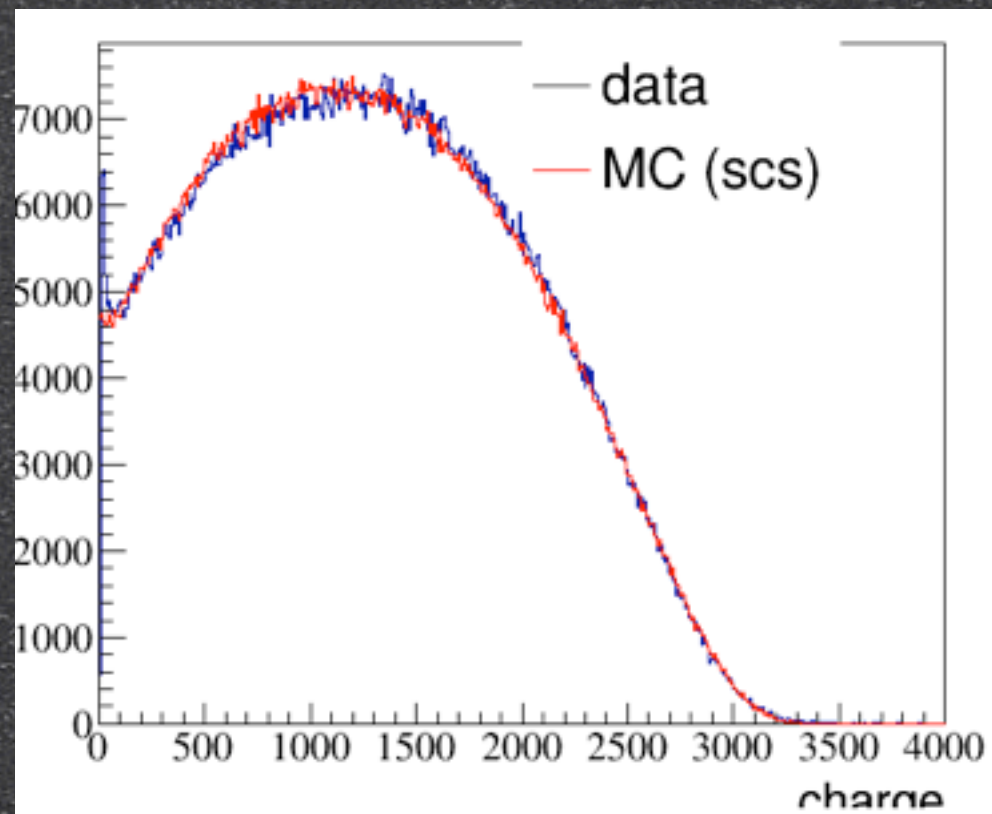
# Some examples

- For DS-50 analysis there's an existing MC framework developed at APC
- Easy to use, but still a lot of work to do to fully characterize the detector

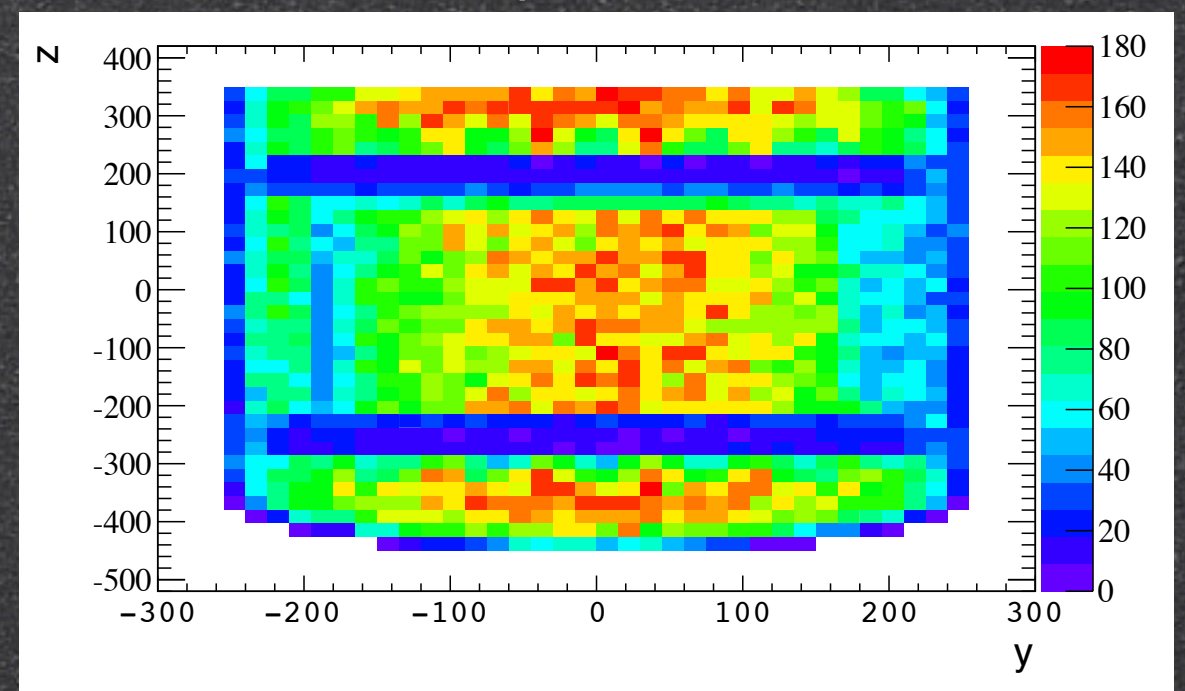
TPC with SiPM on the top



Data/MC comparison for  $^{39}\text{Ar}$



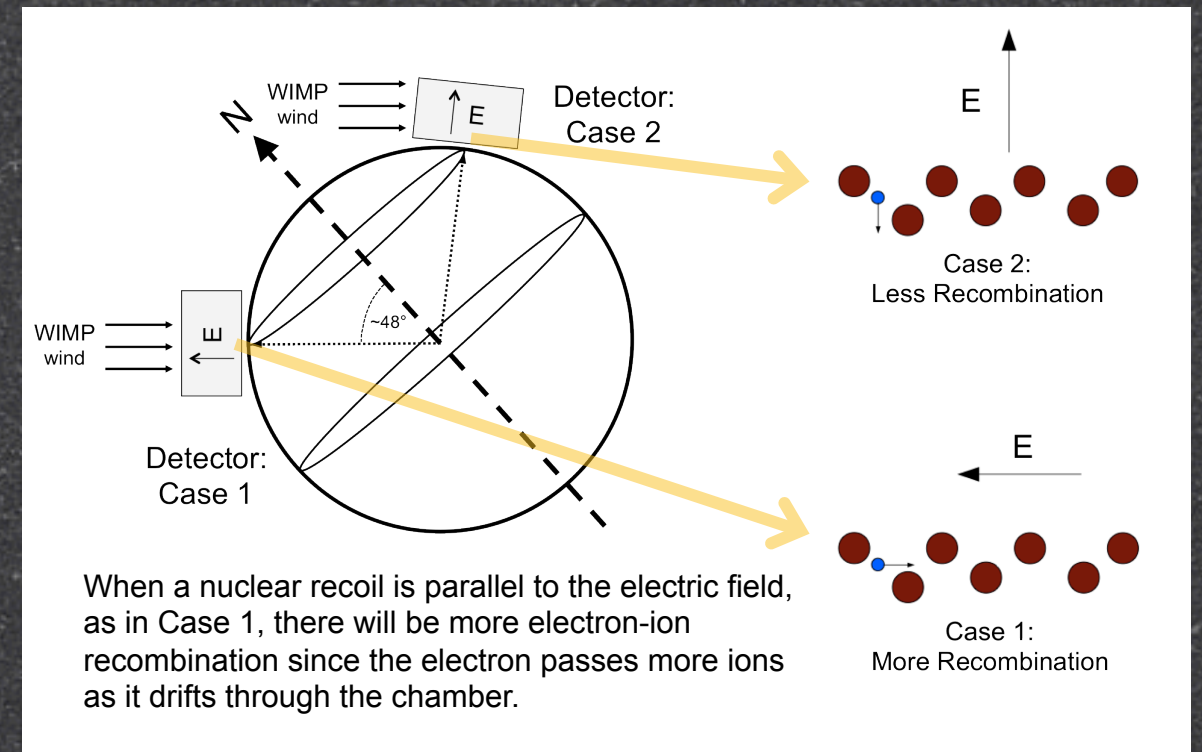
Simulation of  $\gamma$  inside Darkside-50



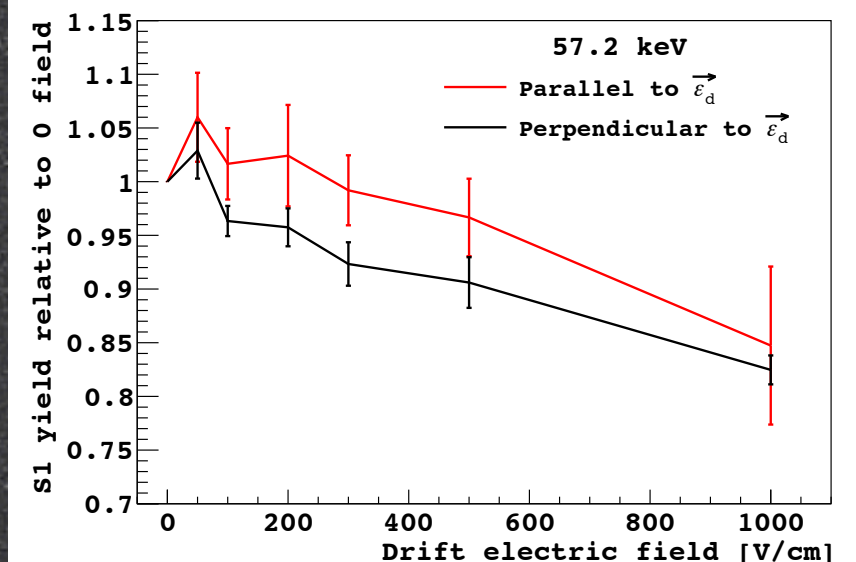


# Directionality with LAr

- S1 is different if the electric field is parallel or perpendicular to the nuclear recoil
- Some weak hints of this effect in LAr have been observed in SCENE
- This effect might strengthen the significance of few WIMPs candidates if they will be observed in DS-G2
- There's a large interest in some of the DS groups to investigate this effect
- We made an ANR proposal to build a small TPC
  - Study directionality
  - Develop SiPM



## Angular dependence from SCENE



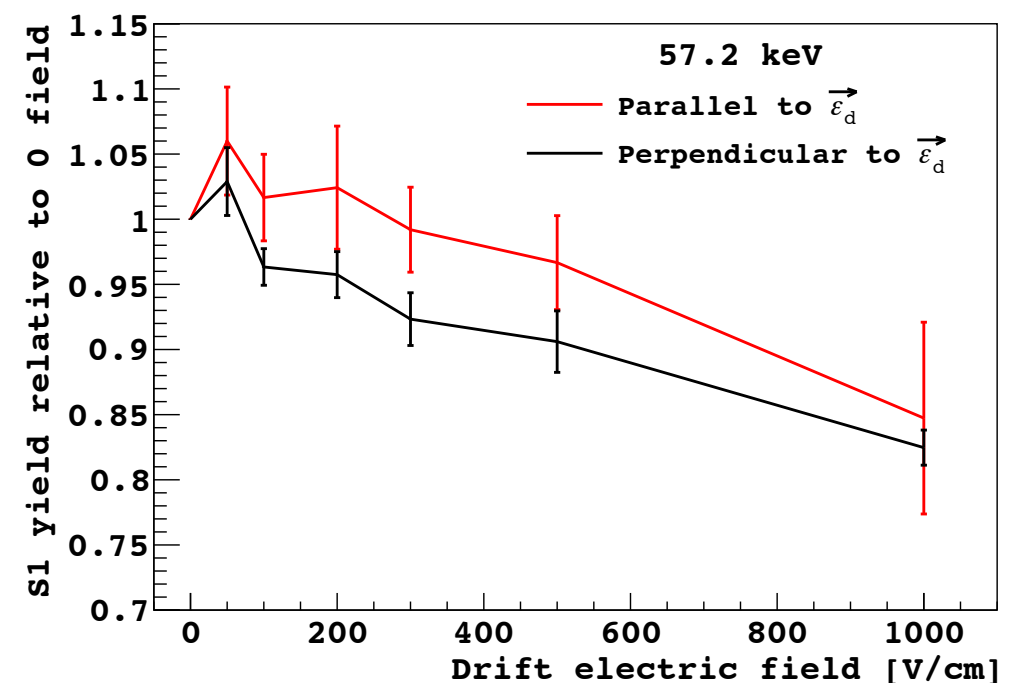
<http://arxiv.org/abs/1406.4825>



# Directionality in France

- The main limitation to study the directionality is to have enough beam time allocate
  - Neutrons with a well defined energy
- Writing the ANR we discussed with the director of IPNO where they have an excellent neutron beam for these kind of studies (Tandem accelerator)
- Faizal Azaiez participate to the ANR proposal
- We will have further discussion with him and Cristian at the begin of 2015 to see if we can have a collaboration also independently on the ANR

## Angular dependence from SCENE



<http://arxiv.org/abs/1406.4825>



# SiPM for DS-G2

- Advantages of SiPM for DS-G2
  - Higher QE → reduce threshold for WIMP search
  - Smaller than PMTs → increase the FV
  - Better PSD thanks to the higher Single Photo-Electron resolution
  - Smaller backgrounds than PMTs that might mimic NR in the TPC
- Our contribution to the ANR would be mainly for the electronics part
  - Development of SiPM
  - Readout electronics for the TPC
- Possibility to contribute independently from the ANR under investigation



# Conclusions

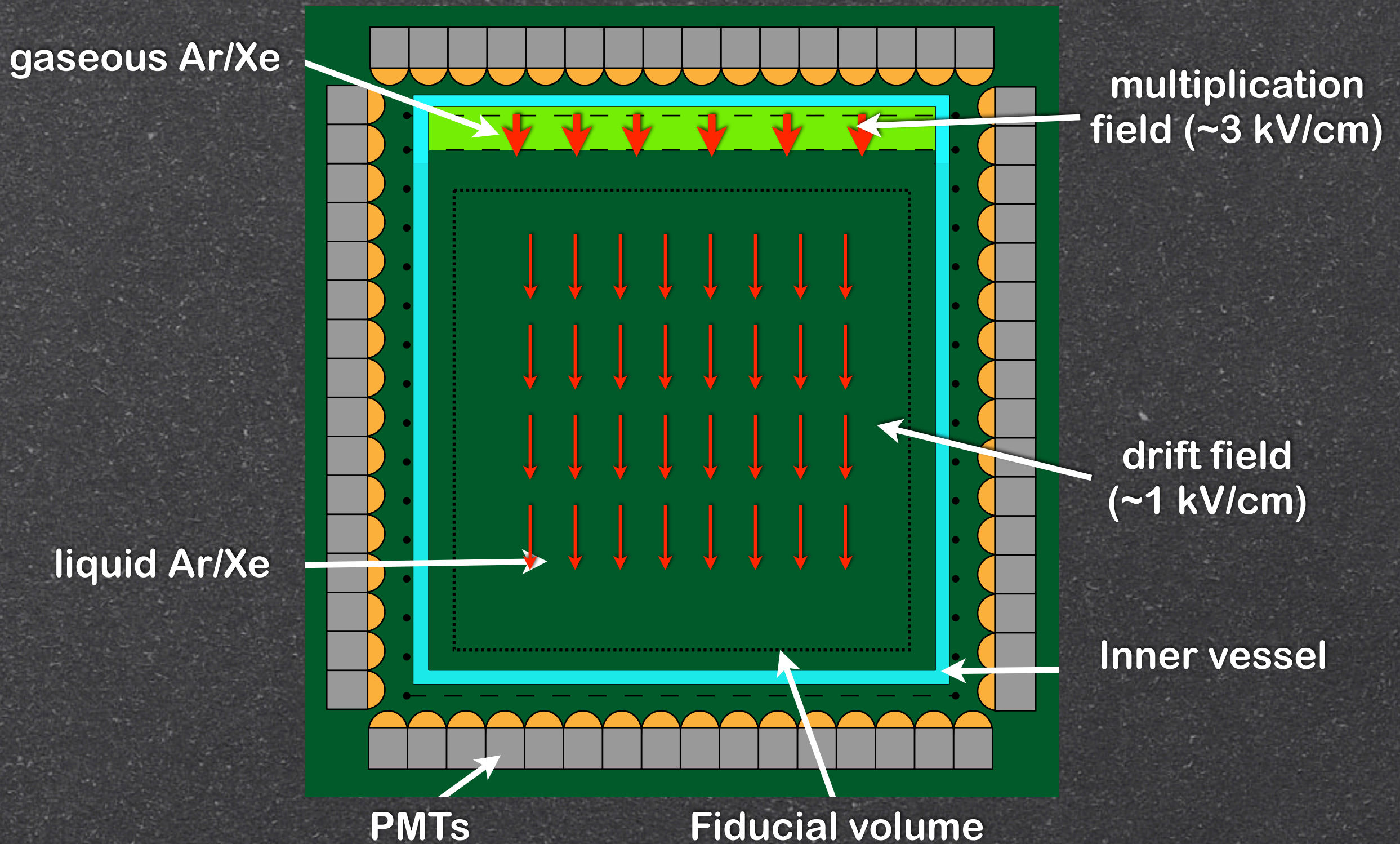
- Liquid Argon TPC are good candidates to build large detector to search for direct dark matter
  - Current limits are behind the ones from Liquid Xenon
  - Disadvantage :  $^{39}\text{Ar}$   $\beta$  emitter  $\rightarrow$  use Underground Argon
  - Advantage: Pulse Shape Discrimination in LAr better than in LXe
- **Darkside-50** is proving that it's possible to build a **background-free** detector to search for WIMPs
- A ton-scale detector can be built in the next years
- French groups are already working on DarkSide
  - The collaboration will be happy to have new collaborators
- I believe it's a good option for us to get involved in an experiment searching for WIMPs
  - Analysis of DarkSide-50 data
  - Necessary R&D towards DarkSide-G2



# Back-up



# Noble liquids TPC

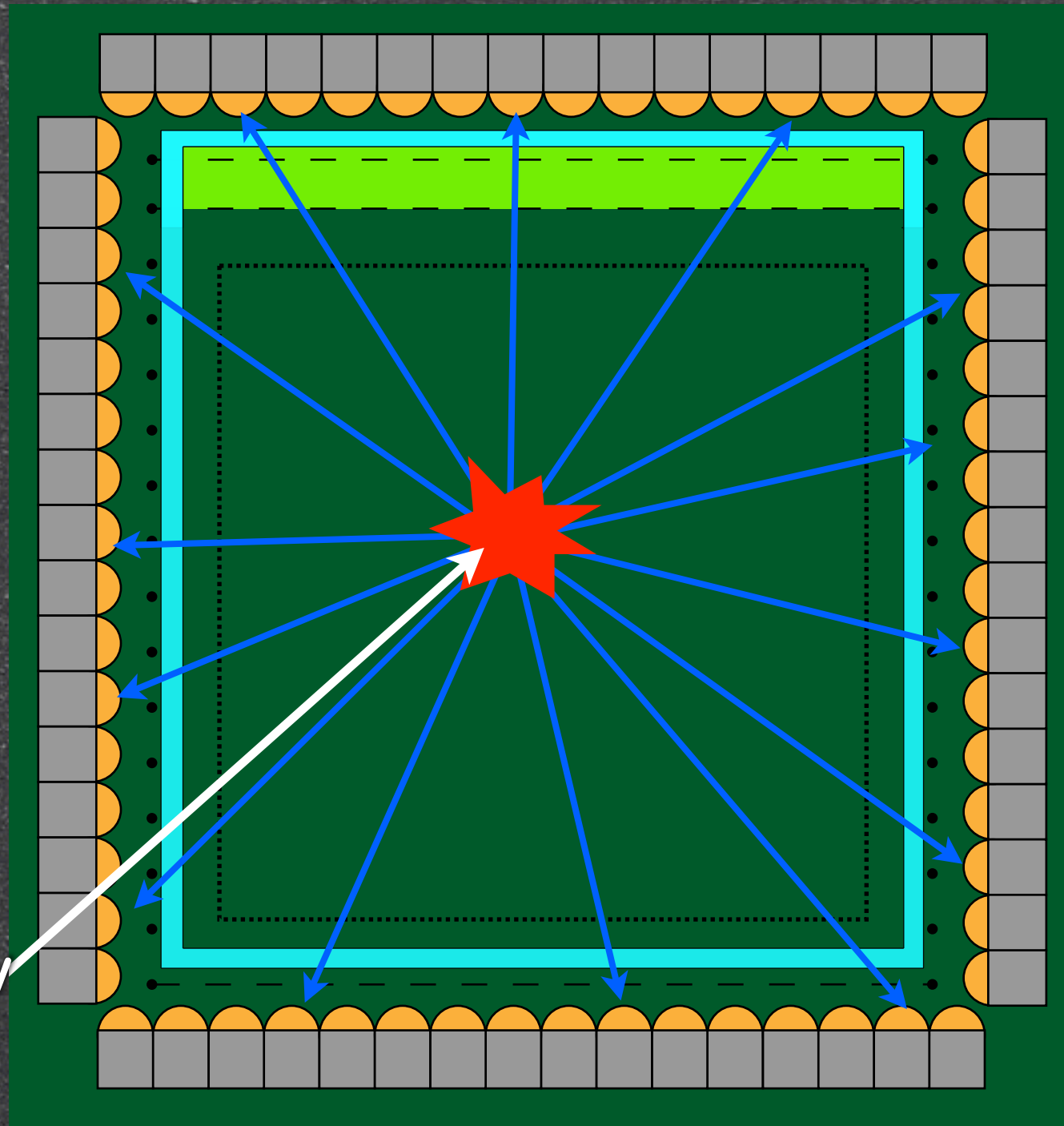




# Noble liquids: scintillation

Primary  
scintillation  
photons emitted  
and detected →  
**S1**

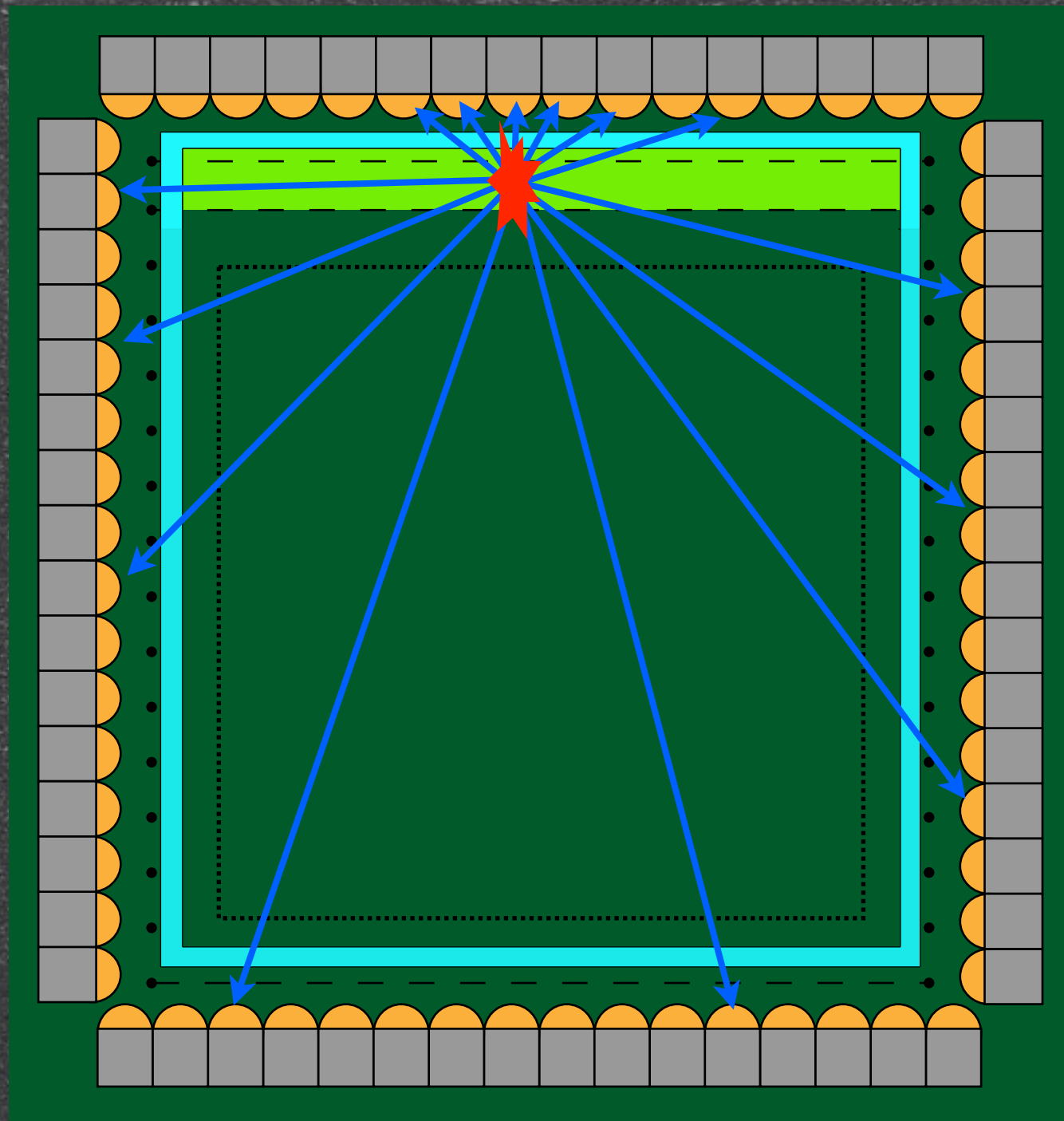
WIMP scatter  
deposits energy  
in the FV





# Noble liquids: ionization

Ionized  
electrons drift to  
the gas region  
where  
secondary  
photons are  
emitted and  
detected → S2





# WIMP direct detection

Low rate ( $\sim 1$  ev/ton/yr  
@  $\sigma = 10^{-47} \text{cm}^2$ )

Large masses

Low energy nuclear  
recoils ( $< 100$  keV)

Low energy thresholds

Background  
suppression

Deep underground  
Passive and active shielding  
Low radioactivity  
Discrimination of ER from NR

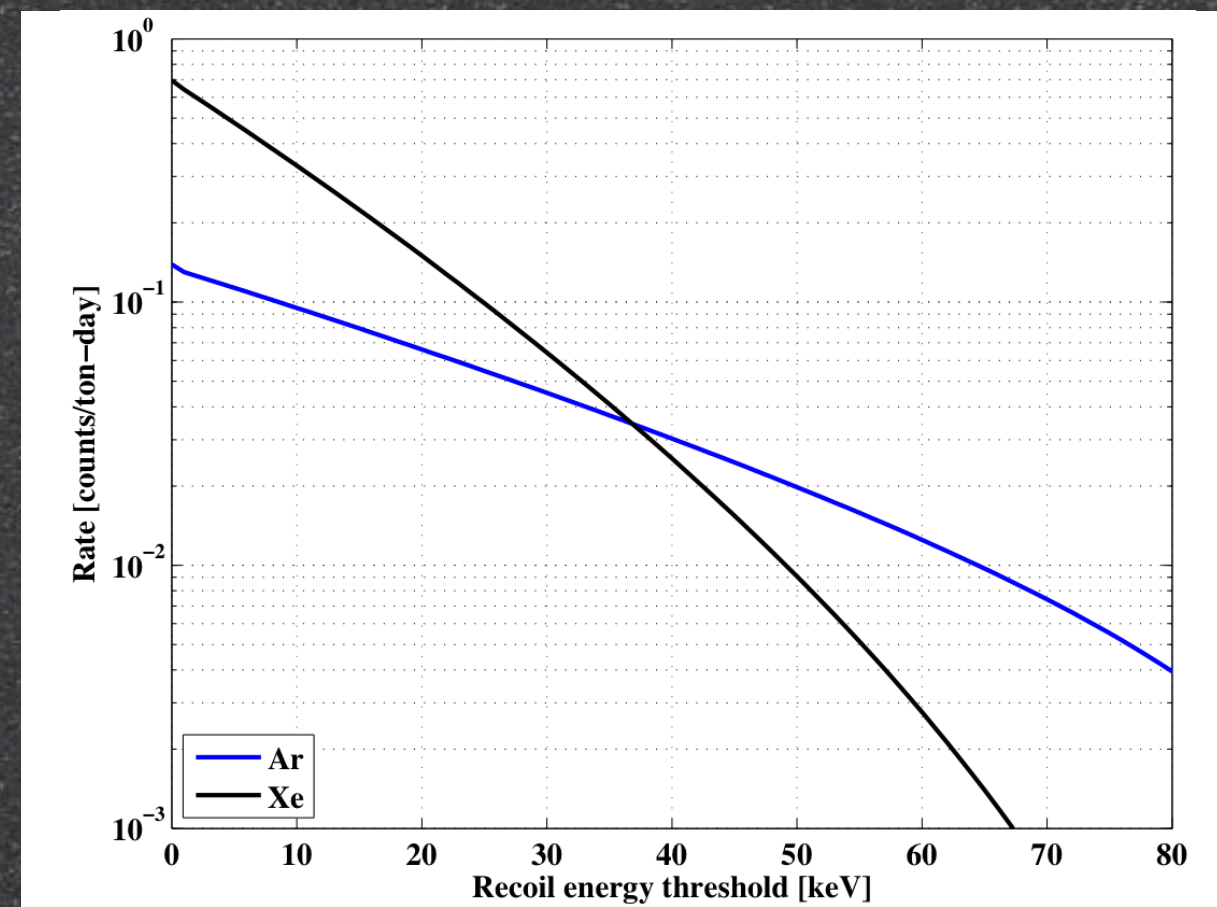
WIMPs and neutrons  $\rightarrow$  Nuclear Recoils  
 $\beta, \gamma \rightarrow$  Electron Recoils



# Noble liquids

- Dense and relatively inexpensive → Large masses
- Easy to purify
- Use electrons ionization and photons scintillation
  - High ionization ( $W \sim 20$  eV)
  - High scintillation yield ( $\sim 40$  photons/MeV)
- Discriminate electron recoils from nuclear recoils

2 liquids used in DM experiments:  
Liquid Xenon  
Liquid Argon

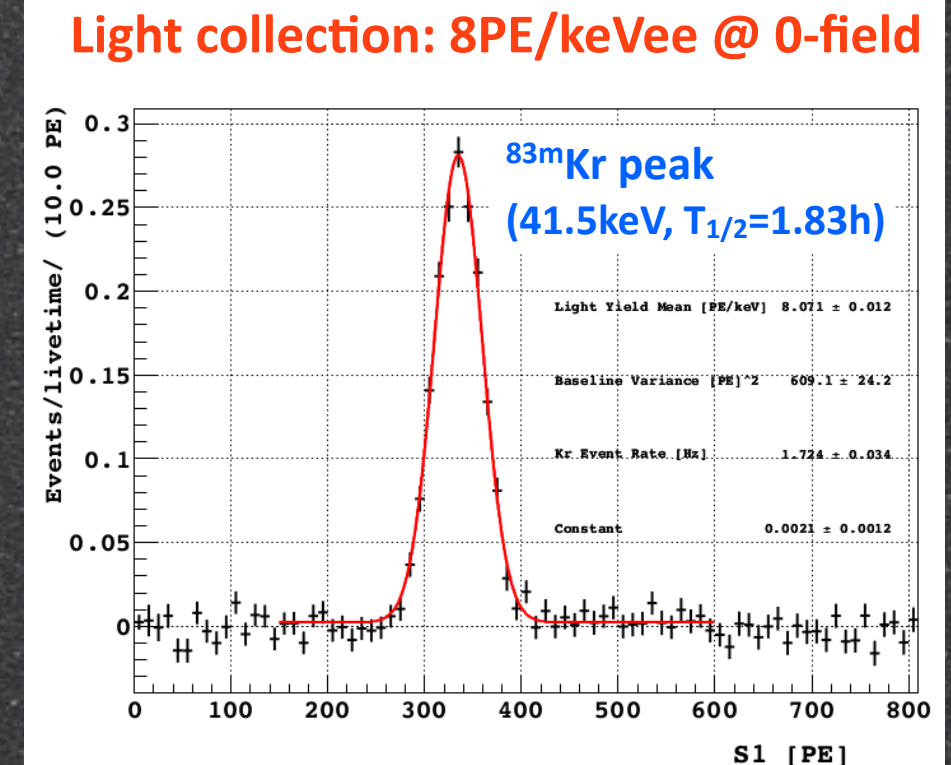
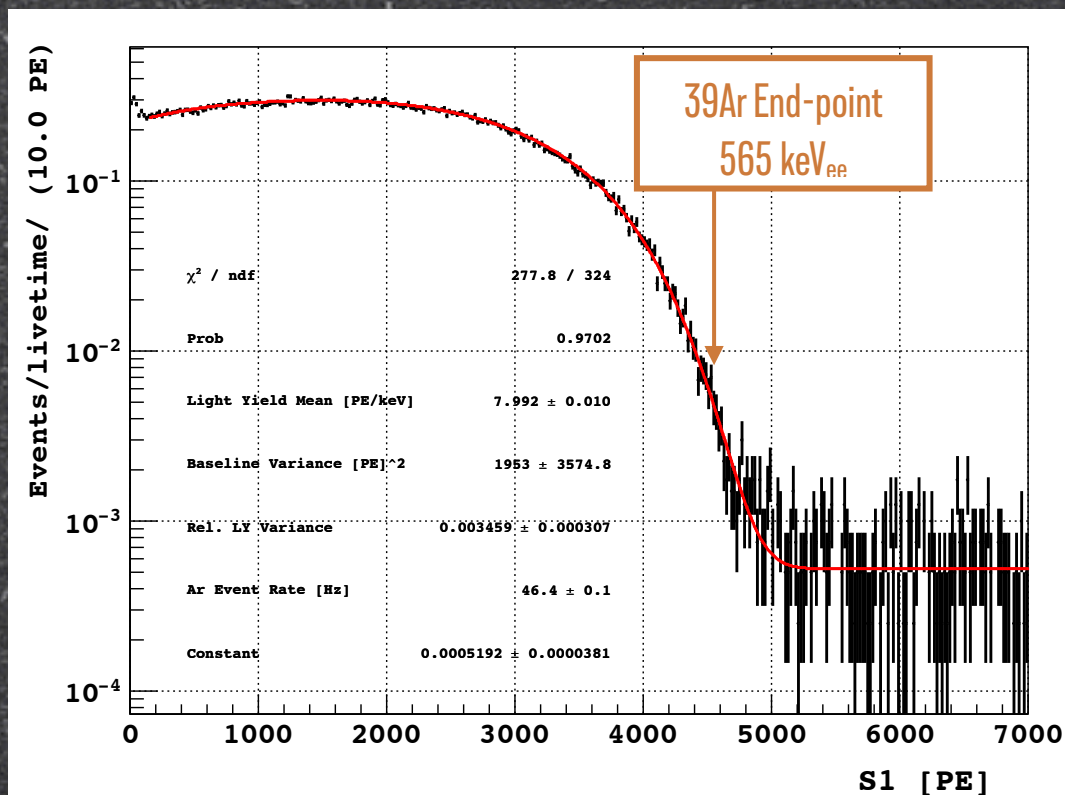




# Electron recoils in DS-50

- Use  $^{39}\text{Ar}$  and a Kr source to estimate the light yield for the scintillation in DS-50
- This is a very important parameter because the pulse shape discrimination critically depends on the number of photons produced in the scintillation process

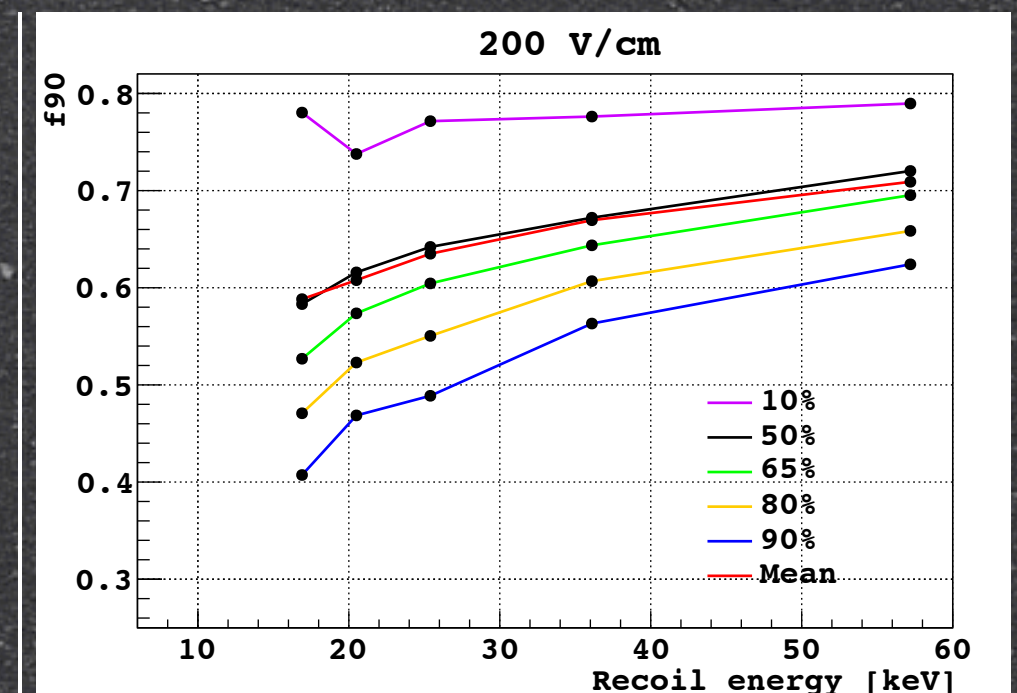
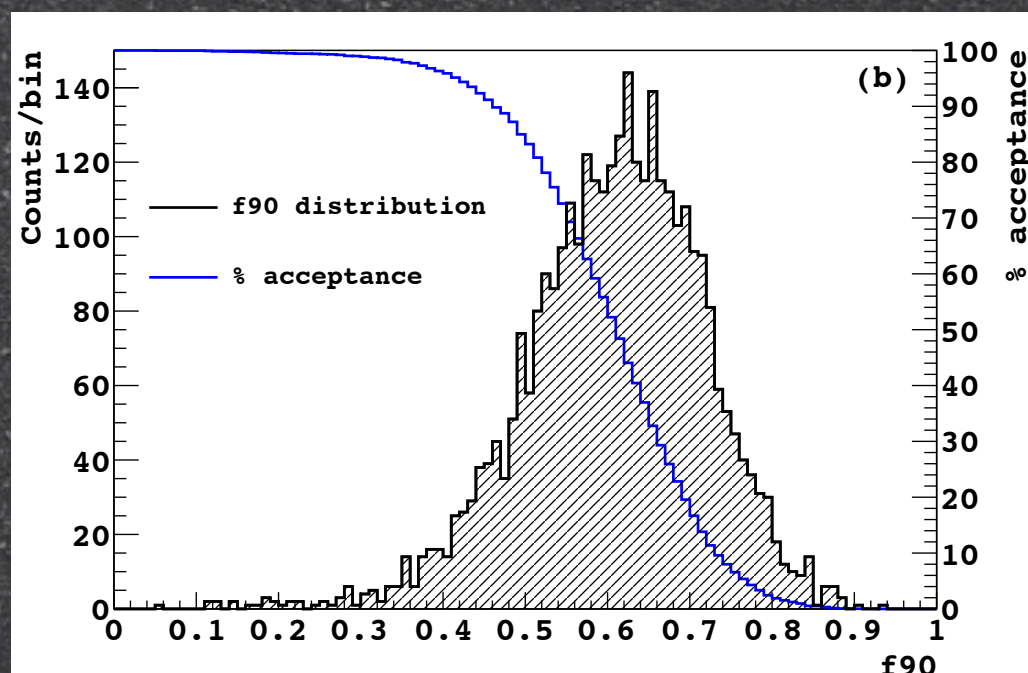
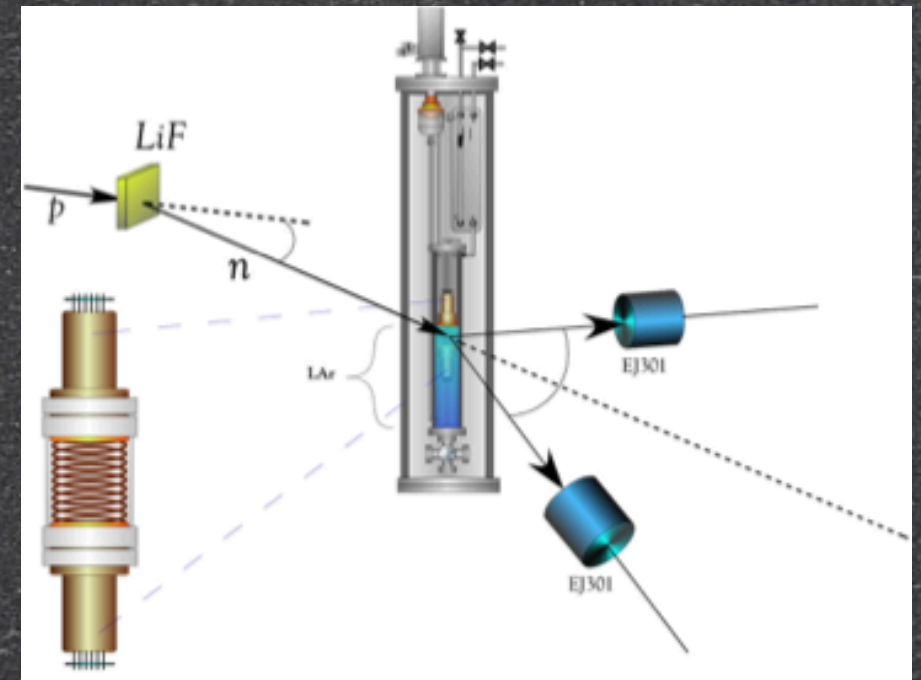
8 PE/keVee @ 0 field  
(better than design  $\rightarrow$  6 PE/keVee)





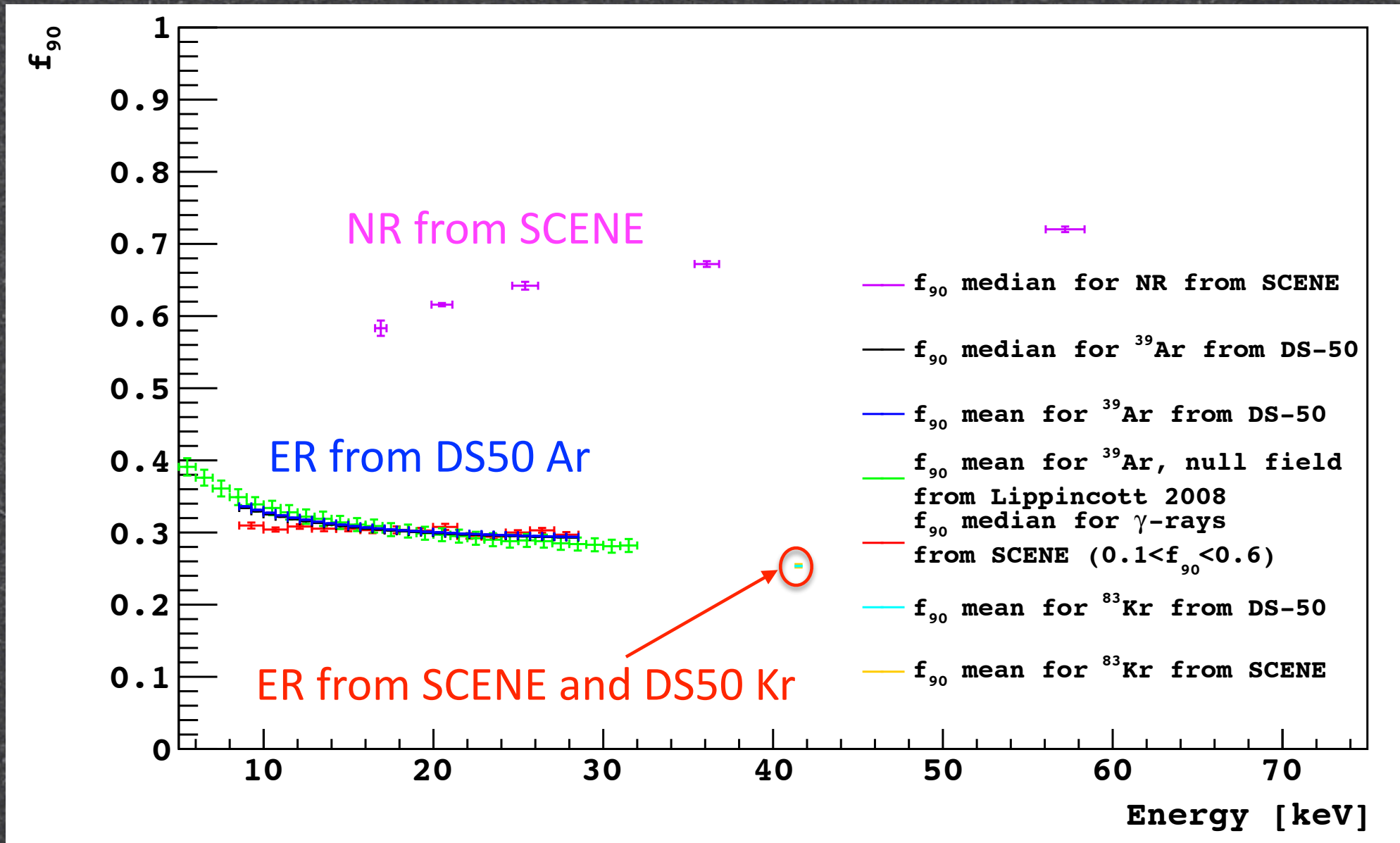
# NR calibration: SCENE

- Calibrating the response for the nuclear recoils is more difficult because you need single scatter in the TPC
- A calibration experiment has been performed putting a small TPC onto a neutron beam and measuring the scintillation produced by single scatter nuclear recoils of known energy





# ER vs NR



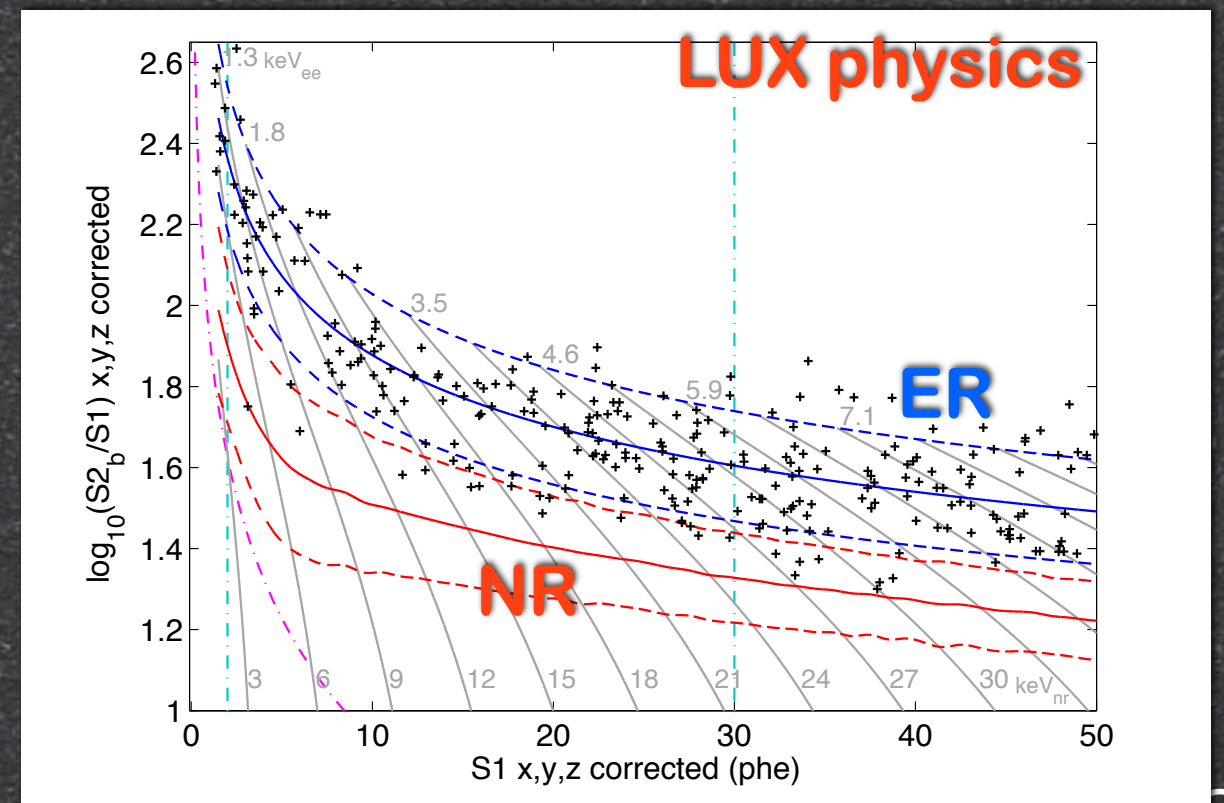
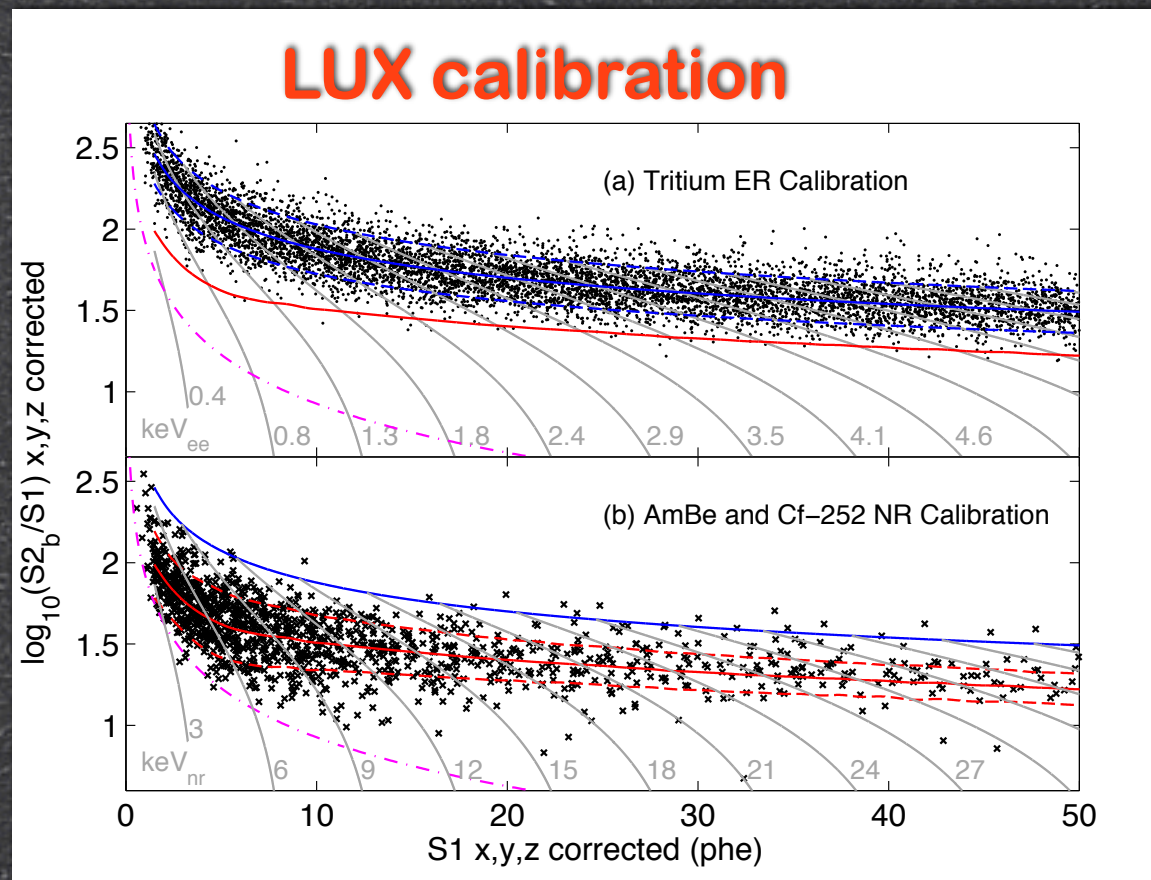
- $f_{90}$  for ER  $\sim 0.3$
- $f_{90}$  for NR  $\sim 0.7$
- This is the variable used to perform the PSD



# Argon vs Xenon

- Xenon experiments are currently putting best limits on DM thanks to their large volume and the lowest activity of the Xenon
- On the other side the PSD of the Xenon is worst than the one obtainable with LAr → a background-free experiment is only possible with LAr

LUX results → 0.6 ER events expected below NR mean, some more observed close to the NR mean





# DarkSide-50 sensitivity

- The goal of DS-50 is to demonstrate that DM search with zero background can be done
- Due to the small size of the detector even after 3 years of running with UAr the limit will be just slightly better than the current limits from LUX
- Once the technology is established plans to build a larger detector → DarkSide-G2

