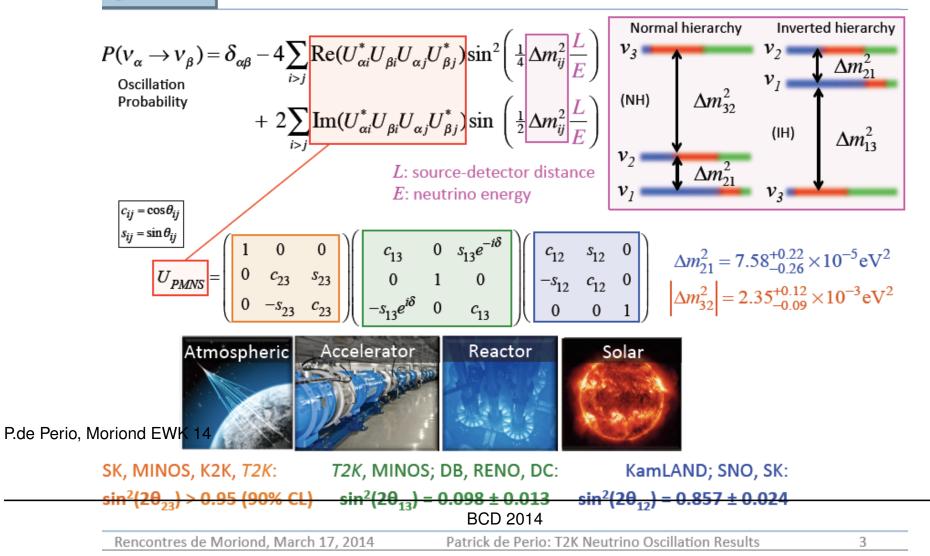
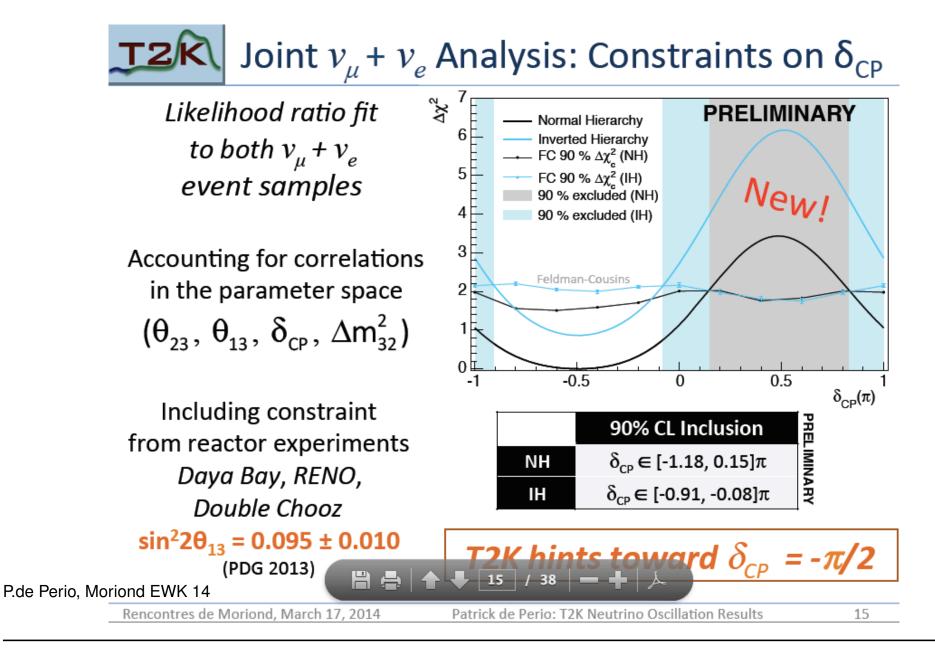
Neutrinos, Cosmology

Slide 36

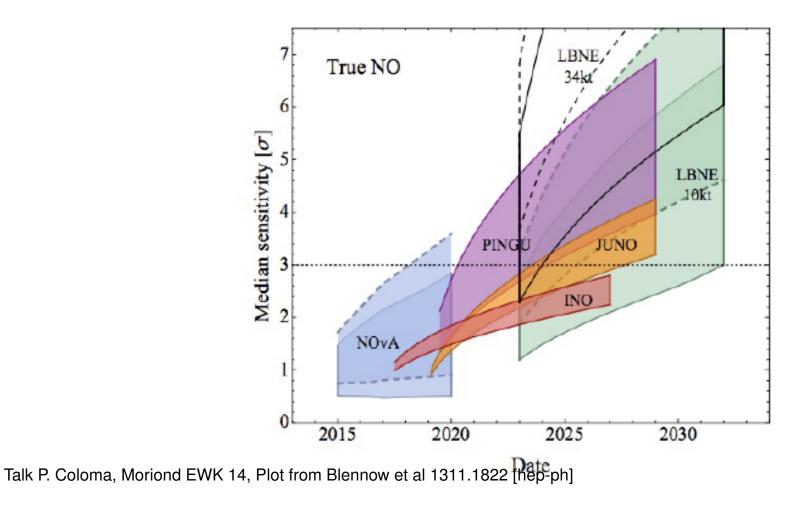
 Δm^2 , Θ_{ij} measured. Q: hierarchy? CP phase δ , More then 3?

Neutrino Oscillations: Current Status





Future sensitivity rejecting IH with oscillation experiments long term perspective; input δ etc varied:



Sterile Neutrinos and Anomalies

$N_{\nu} = 3$ hypothesis succesful re most earth based data.

- however, a few experiments exhibit deviations from the "standard" 3v scenario:
 - LSND & MiniBooNE observe excesses of $\overline{\nu}_e$ events in $\overline{\nu}_\mu$ beams;
 - SBL reactors observe a deficit with respect to the expected $\bar{\nu}$ fission flux;
 - gallium experiments exposed to radioactive sources also find a deficit;
- sterile neutrino models fail to explain satisfactorily all the experimental data:

| Requirement | (3+0) | (2+2) | (3+1) | (3+2) | (1+3+1) |
|--|-------|-------|-------|-------|---------|
| Ordinary neutrino oscillation data | OK | NO | OK | OK | OK |
| $\overline{\nu}_e \rightarrow \overline{\nu}_e$: SBL reactor & gallium data | NO | OK | OK | OK | OK |
| $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$: LSND & MB $\bar{\nu}$ data | NO | OK | OK | OK | OK |
| $\nu_{\mu} \rightarrow \nu_{e}$: MB high-energy ν data | OK | POOR | POOR | OK | OK |
| $\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}$: MB low-energy excess | NO | POOR | POOR | OK | OK |
| $\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{\mu}$: disappearance data | OK | OK | NO | NO | NO |
| Constraints from cosmology | OK | NO | OK | NO | NO |

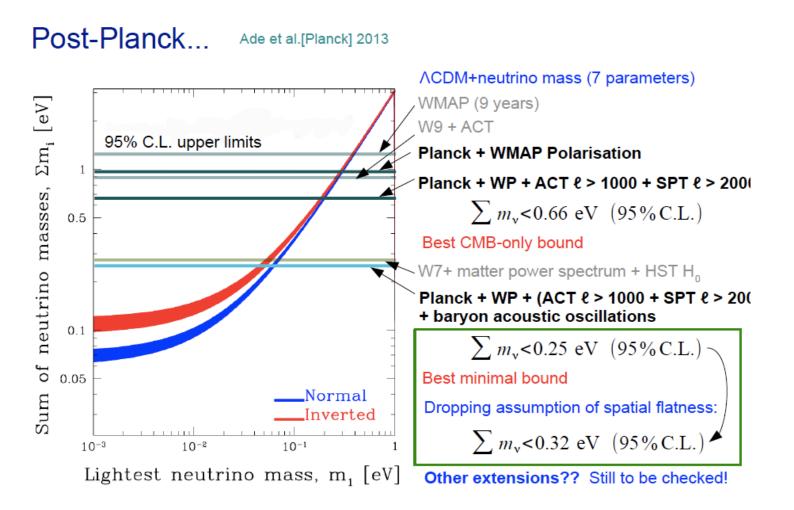
⇒ we are still quite far from the solution of the LSND puzzle!

Michele Maltoni <michele.maltoni@csic.es>

MORIOND EW 2014, 17/03/2014

Talk Maltoni, Moriond EWK 14

 $m_{\nu} < ?$



Effective number of light "sterile" species $N_{eff} =$? Contribution to energy density $\sum_{i=1}^{3} \rho_{i\nu} + \rho_X \propto (3.046 + \Delta N_{eff}))$

Placnk-inferred N_{eff} compatible with 3.046 at better than 2σ .

| 2σ error bars | Planck+WP | | Planck+WP+BAO | | Planck+WP+highL | | Planck+WP+highL+BAO | |
|-----------------------------------|-----------|----------------------------|---------------|------------------------------|-----------------|----------------------------|---------------------|-------------------------------|
| Parameter | Best fit | 95% limits | Best fit | 95% limits | Best fit | 95% limits | Best fit | 95% limits |
| Ω _κ | -0.0105 | $-0.037^{+0.043}_{-0.049}$ | 0.0000 | $0.0000^{+0.0066}_{-0.0067}$ | -0.0111 | $-0.042^{+0.043}_{-0.048}$ | 0.0009 | $-0.0005^{+0.0065}_{-0.0066}$ |
| Σm_{ν} [eV] | 0.022 | < 0.933 | 0.002 | < 0.247 | 0.023 | < 0.663 | 0.000 | < 0.230 |
| <i>N</i> _{eff} | 3.08 | $3.51_{-0.74}^{+0.80}$ | 3.08 | $3.40^{+0.59}_{-0.57}$ | 3.23 | $3.36^{+0.68}_{-0.64}$ | 3.22 | $3.30_{-0.51}^{+0.54}$ |
| $Y_{\rm P}$ | 0.2583 | $0.283^{+0.045}_{-0.048}$ | 0.2736 | $0.283^{+0.043}_{-0.045}$ | 0.2612 | $0.266^{+0.040}_{-0.042}$ | 0.2615 | $0.267^{+0.038}_{-0.040}$ |
| $dn_{\rm s}/d\ln k$ | -0.0090 | $-0.013^{+0.018}_{-0.018}$ | -0.0102 | $-0.013^{+0.018}_{-0.018}$ | -0.0106 | $-0.015^{+0.017}_{-0.017}$ | -0.0103 | $-0.014^{+0.016}_{-0.017}$ |
| <i>r</i> _{0.002} | 0.000 | < 0.120 | 0.000 | < 0.122 | 0.000 | < 0.108 | 0.000 | < 0.111 |
| w | -1.20 | $-1.49^{+0.65}_{-0.57}$ | -1.076 | $-1.13^{+0.24}_{-0.25}$ | -1.20 | $-1.51^{+0.62}_{-0.53}$ | -1.109 | $-1.13^{+0.23}_{-0.25}$ |
| Talk Y.Wong, Moriond EWK 14 , DUT | | | | | | | | |

Wong points out that a discrepancy in the determination of the Hubble parameter H_0 from Hubble space telescope and others

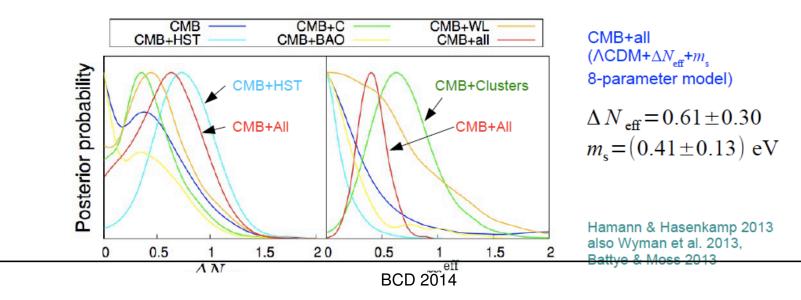
| Parameter | Planck | | Pla | nck+lensing | Planck+WP | |
|-----------------|----------|----------------------|----------|----------------------|-----------|-------------------|
| | Best fit | 68% limits | Best fit | 68% limits | Best fit | 68% limits |
| σ_8 | 0.8344 | 0.834 ± 0.027 | 0.8285 | 0.823 ± 0.018 | 0.8347 | 0.829 ± 0.012 |
| Z _{re} | 11.35 | $11.4^{+4.0}_{-2.8}$ | 11.45 | $10.8^{+3.1}_{-2.5}$ | 11.37 | 11.1 ± 1.1 |
| H_0 | 67.11 | 67.4 ± 1.4 | 68.14 | 67.9 ± 1.5 | 67.04 | 67.3 ± 1.2 |

Hubble space telescope

$$H_0 = 73.8 \pm 2.4 \,\mathrm{km \, s^{-1} \, Mpc^{-1}}$$

Exploit the $N_{\text{eff}} - H_0$ degeneracy and introduce to a large N_{eff} to bring HST and Planck in line with one another.

could be resolved with fourth ν

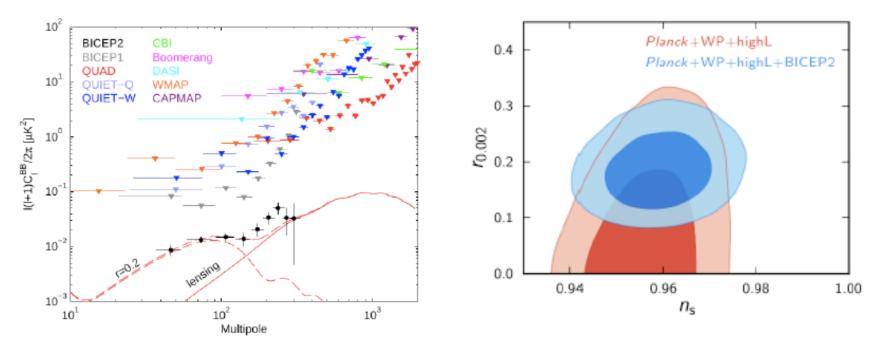


arXiv.org > astro-ph > arXiv:1403.3985

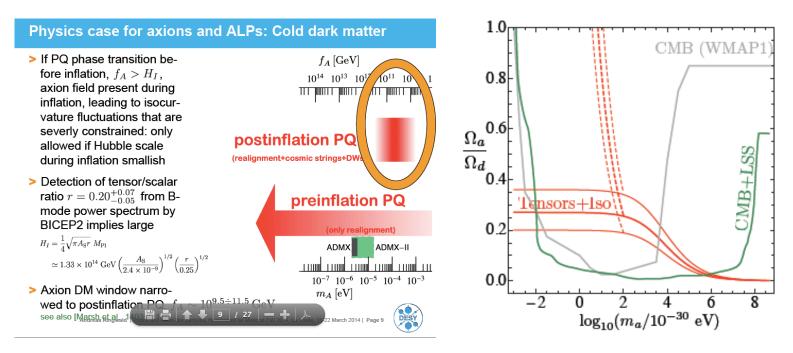
Astrophysics > Cosmology and Nongalactic Astrophysics

BICEP2 I: Detection Of B-mode Polarization at Degree Angular Scales

Comments in talks J.Hamann, A.Ringwald, O.Perdereau, Moriond EWK 14



The Biceps2 measurement (if real) has implications for axions, weakly interactling particles that appear e.q. in dark matter models



Talk by A Ringwald, Moriond EWK14, and plot by Marsh et al 1403.4216

As we discussed the "hot" and vastly different topics, we encountered the established, the unprecise and the unknown.

Many excting topics not covered (dark matter, astroparticles, gravitation, TeV-scale model building ..)

It is clear that fundamental physics is always "research in progress".

It is also clear that many exciting questions remain, that can be tackled in a combined effort of theory and experiment.

We need your ideas.

Part III - Outlook for HEP experiments and machines

- The SM has been built in a continuum of 50 years of experimental and theoretical progresses. It is supported nowadays by two pillars: the EWK and CKM precision global consistency checks.
- We want to shake those pillars w/ improved precision and test the predictions made possible out of the global consistency checks.
- Top quark physics: LHC is a top factory. Its basic properties are measured with increasingly better control of systematics effect.
- Higgs physics: the beginning of a quest. Invaluable window for NP, make a global consistency check of the couplings, determine the shape of the scalar potential. Opening to cosmology, the scalar field brings a no-additional-parameter inflation model.
- Flavour is a central problem of Physics. Flavour Physics might be the solution... Again the beginning of a quest with the precision era. Future facilities to measure neutrinos properties.
- We should not forget in this landscape smaller scale projects and HEP-related dark matter searches.

Hiller, Monteil

BCD 2014



Immediate future (up to ~2022)

- LHC Run II It starts ~now w/ unprecedented energy in the center-of-mass. Search for new phenomena (a lot of hope) and refine in precision the physics outcome of Run I.
- LHCb upgrade and SuperKEKB: the former is disjoint from the LHC upgrade. Meant to cope with the luminosity upgrade of the Run III. SuperKEKB is a *B*-factory equipped w/ a new detector Belle II. Commissioning next year. Both are expected to analyse 50 times the current statistics: the precision era for quark flavour physics. BES (chinese charm factory) is still running.

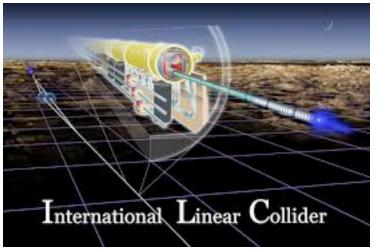
Main (selected) Projects on the Table:

- LHC High Luminosity: likely to happen. 3000 /fb at 13/14 TeV (now 25 /fb at 7/8 TeV).
- e+e- linear collider (90 800 GeV), known as ILC.
- e⁺e⁻ circular collider (90 375 GeV), known as FCC-*ee*.
- proton-proton circular collider (100 TeV), known as FCC-pp
- e⁺e⁻ linear collider (3 TeV), known as CLIC.
- Long Baseline neutrinos.

Part III - Selected projects for HEP.

Few comments in order for e+e- linear collider

- Physics Case established since 15 years: m_W, m_t, H properties and couplings (including trilinear), VV at high energy.
- Cold technology for accelerating cavities (< 800 GeV). They are mature since year ~2000.
- ~31 km baseline. Lumi 10³⁴ cm⁻²s⁻¹
- Japan declared interest.
- http://www.linearcollider.org/ILC/Press/ Images-and-graphics



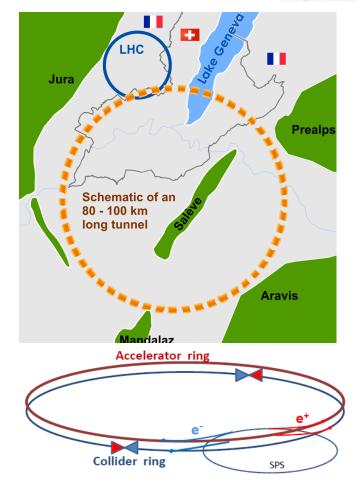




Part III - Selected projects for HEP.

Few comments in order for e⁺e⁻ circular collider

- Emerged in the last couple of years. Embodied in Future Circular Collider CERN project.
- 80-100 km circunference. Lumi at the Z pole 10³⁶ cm⁻²s⁻¹ (!).
- Physics Case: *Z*, *W*, top and *H* physics (in the event nothing new is seen at LHC Run II). Given the expected precision, calculations at three loops are necessary.
- Design study undergoing for the european strategy examination in 2018.
- http://tlep.web.cern.ch/

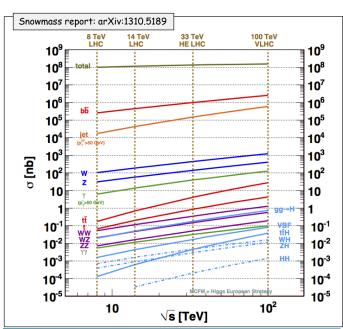


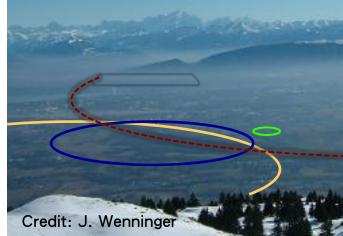
Part III - Selected projects for HEP.



Few comments in order for pp circular collider

- Same tunnel. Same idea as LEP/LHC complex. Embodied in Future Circular Collider CERN project.
- 100 TeV Lumi of 5. 10³⁴ cm⁻²s⁻. Main challenge is as for LHC the magnets (16-20 T required).
- Physics Case: multiply LHC cross-sections by factors 7 to 60 for physics process vs 1.25 for total pp. In particular, the shape of the scalar potential can be accurately measured.
- Design study undergoing for the european strategy examination in 2018.
- <u>http://tlep.web.cern.ch/content/fcc-hh</u>





- You're studying Particle Physics and Cosmology in fascinating times.
- The modern Physics provides a mature and motivated view of the Nature at the scale of the elementary interactions below the TeV scale.
- That makes a very solid ground.
- There's so much more to understand. This is/will happen in facilities opened worldwide, theory and experiment intricated. Research requires a global network. Transmitting the knowledge too. We've tried today! And will hopefully build more in the future.
- Hope you enjoyed your afternoon.