



Looking for geo-neutrinos with Borexino

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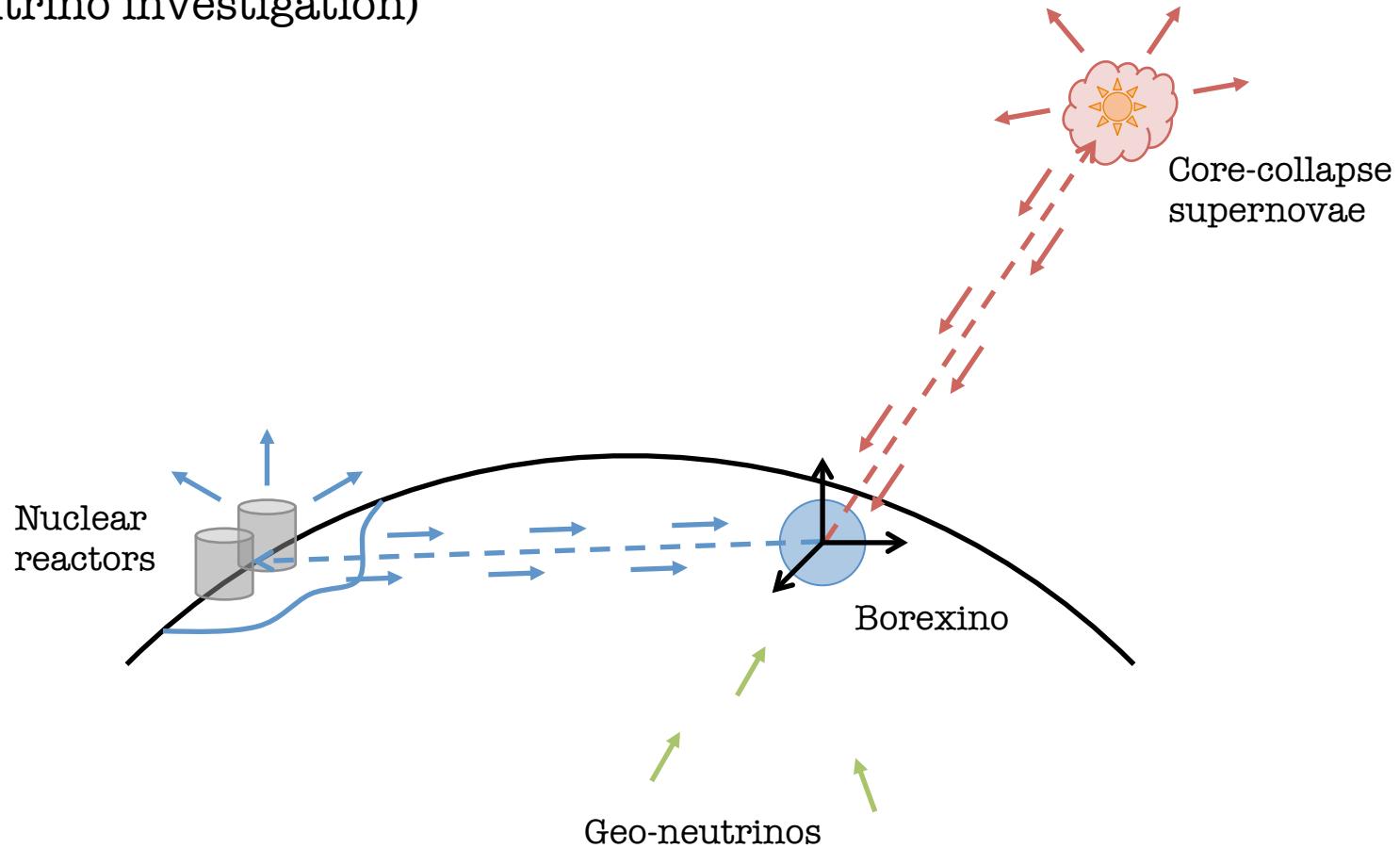
Applied Antineutrino Physics 2014

December 15, 2014

-- Anti- ν_e in Borexino --

-- Borexino is currently exposed to geo-neutrinos and nuclear reactors neutrinos (core-collapse supernovae?)

-- **COMING SOON:** anti- ν_e from ^{144}Ce - ^{144}Pr source
(sterile neutrino investigation)



-- Why geo-neutrinos? --

-- Geo-neutrinos are messengers from the Earth interior

--> Especially of interest for the mantle knowledge

-- Radioactive decays inside the Earth

--> **^{238}U** , **^{235}U** , **^{232}Th** decay series as well as ^{40}K decay are involved and produced ν_e and anti- ν_e called geo-neutrinos

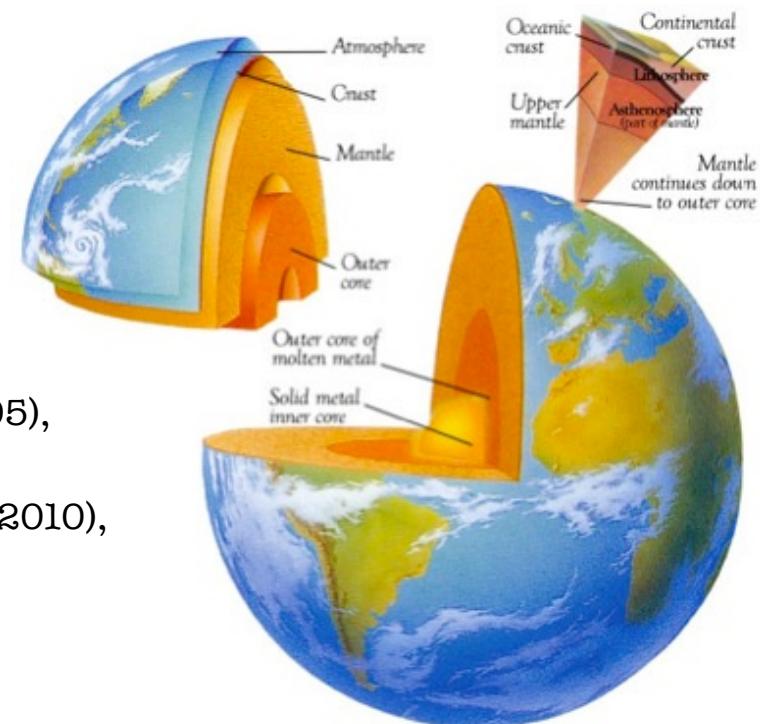
-- Differents Earth models exist (cosmochemical, geochemical, geodynamical etc...) and do not agree between themselves

--> Geo-neutrinos are a new source of information

-- Geo-neutrino measurements

--> KamLAND (Nature **436**, 499-503 (2005), Phys. Rev. D **88**, 033001 (2013))

--> Borexino (Phys. Lett. B **687**, 299-304 (2010), Phys. Lett. B **722**, 295-300 (2013))



-- Which geo-neutrinos? --

- Anti- ν_e detection through inverse beta decay interactions
 - > Threshold at 1.8 MeV

	^{238}U	^{235}U	^{232}Th	$^{40}\text{K}(\bar{\nu}_e)$	$^{40}\text{K}(\nu_e)$
$\tau_{1/2}$ (year)	4.47×10^9	7.04×10^8	1.40×10^{10}	1.28×10^9	1.28×10^9
Q (MeV)	51.7	46.4	42.7	1.311	1.505
$Q_{\bar{\nu}_e}$ (pJ)	0.634	0.325	0.358	0.103	-
# $\bar{\nu}_e$	6	4	4	1	-
$\mathcal{R}_{\bar{\nu}_e}$ ($\bar{\nu}_e/(g \cdot s)$)	7.46×10^4	3.20×10^5	1.63×10^4	2.31×10^5	-
# ν_e	-	-	-	-	1
\mathcal{R}_{ν_e} ($\nu_e/(g \cdot s)$)	-	-	-	-	2.77×10^4
E_{\max} (MeV)	3.26	1.23	2.25	1.311	0.044

- Only anti- ν_e from ^{238}U and ^{232}Th can be detected

-- Geo-neutrinos oscillation? --

-- Anti- ν_e from **^{238}U** and **^{232}Th** do oscillate

--> Survival probability of the geo-neutrinos:

$$P_{ee} = \cos^4 \theta_{13} \left(1 - \sin^2(2\theta_{12}) \sin^2 \left(1.27 \frac{\Delta m_{21}^2 (\text{eV}^2) L(\text{m})}{E(\text{MeV})} \right) \right) + \sin^4 \theta_{13}$$

-- Oscillation length around 100 km $\ll R_{\text{Earth}}$

--> Reasonable assumption of an averaged survival probability:

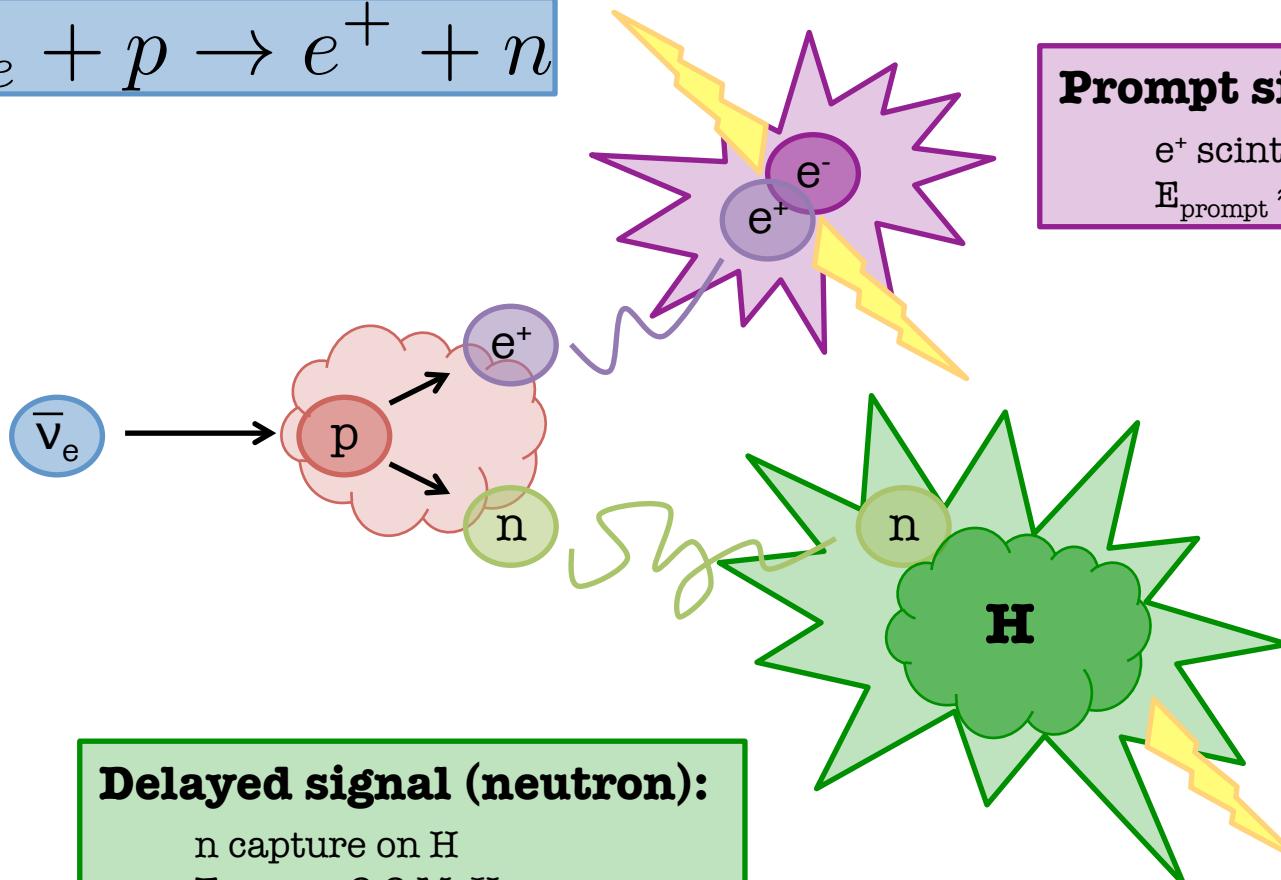
$$\langle P_{ee} \rangle = \cos^4 \theta_{13} \left(1 - \frac{1}{2} \sin^2(2\theta_{12}) \right) + \sin^4 \theta_{13} = 0.55 \pm 0.03$$

-- **WARNING:** not used for anti- ν_e from nuclear reactors (individual calculations)

Mixing angles and mass square differences are taken from
Phys. Rev. D **89**, 093018 (2014)

-- Detecting anti- ν_e --

-- Anti- ν_e detection through inverse beta decay interactions



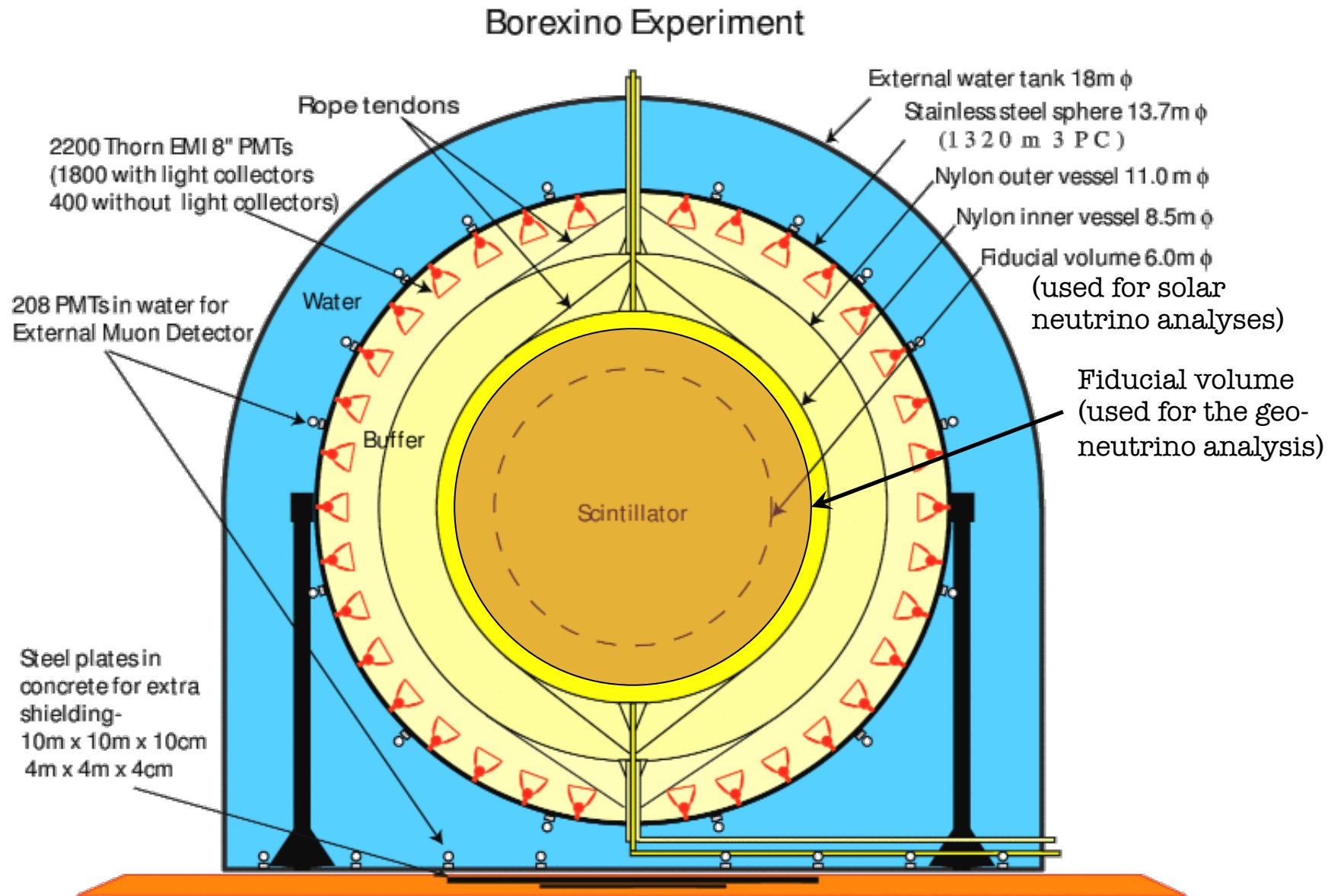
Prompt signal (positron):

$$e^+ \text{ scintillation + annihilation}$$
$$E_{\text{prompt}} \approx E_{\nu} - T_n - 0.8 \text{ MeV}$$

Delayed signal (neutron):

n capture on H
 $E_{\text{delayed}} \approx 2.2 \text{ MeV}$
 $\Delta t = 254.5 \pm 1.8 \mu\text{s}$

-- The Borexino detector --



-- Selecting anti- ν_e --

-- Prompt signal:

-- $Q_{\text{prompt}} > 408 \text{ p.e.}$

-- Fiducial Volume Cut (FVC)

1 MeV ≈ 500 p.e.

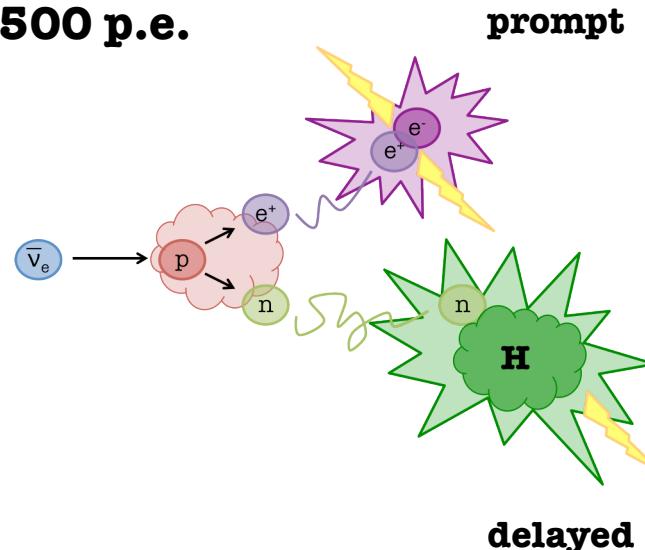
-- Delayed signal:

-- $860 < Q_{\text{delayed}} < 1300 \text{ p.e.}$

-- Coincidence:

-- $20 < \Delta t < 1280 \mu\text{s}$

-- $\Delta R < 100 \text{ cm}$



-- 2 s dead time window applied after an internal muon and 2 ms dead time window applied after an external muon

-- No neutron event in the 2 ms time window before the prompt signal and in the 2 ms time window after the delayed signal

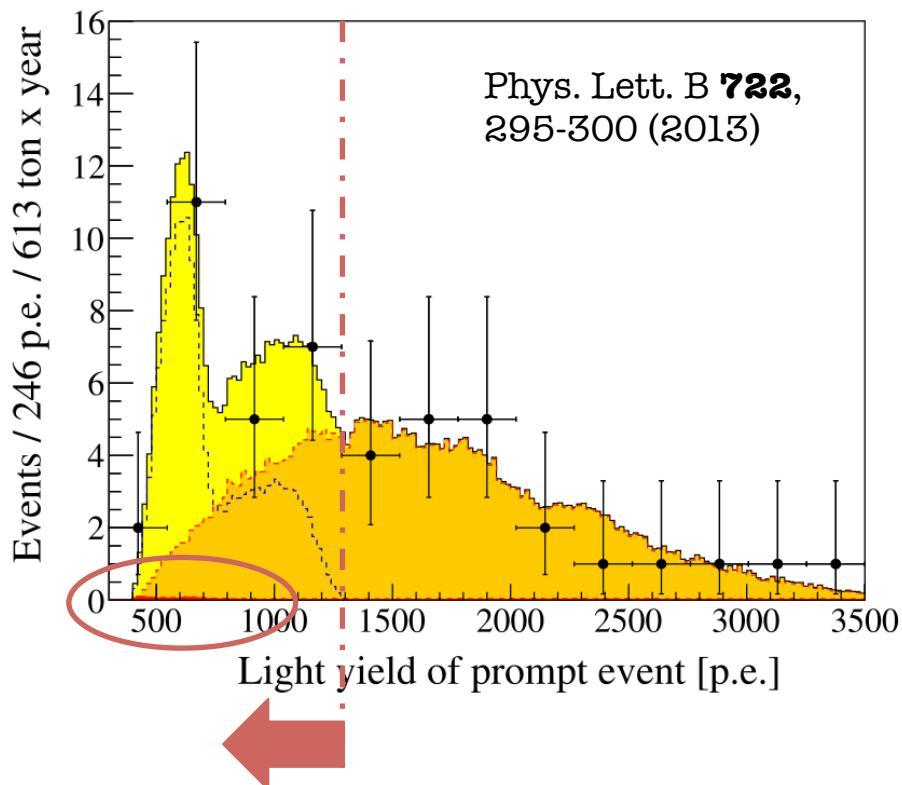
46 candidates

(between December 2007 and August 2012)

-- Anti- ν_e energy spectra --

-- Q_{prompt} spectrum contains both the geo-neutrino and the reactors neutrino components (and the backgrounds)

--> Since $E_{\text{max}}(^{238}\text{U}) = 3.26 \text{ MeV}$ and $E_{\text{max}}(^{232}\text{Th}) = 2.25 \text{ MeV}$, geo-neutrinos stand in the 4 first bins of the Q_{prompt} spectrum



Background source	Events
$^9\text{Li}-^8\text{He}$	0.25 ± 0.18
Fast n 's (μ 's in WT)	<0.07
Fast n 's (μ 's in rock)	<0.28
Untagged muons	0.080 ± 0.007
Accidental coincidences	0.206 ± 0.004
Time corr. background	0.005 ± 0.012
(γ, n)	<0.04
Spontaneous fission in PMTs	0.022 ± 0.002
(α, n) in scintillator	0.13 ± 0.01
(α, n) in the buffer	<0.43
Total	0.70 ± 0.18

Very low background (except for reactor background)!

-- Reactor background --

-- Anti- ν_e from nuclear reactors are the main background

-- Estimation of the expected number of events from the spectral components of ^{235}U , ^{238}U , ^{239}Pu and ^{241}Pu

Number of nuclear reactors considered

Number of months considered

Exposure in month m and includes detector efficiency

$$N_{\text{react}} = \sum_{r=1}^R \sum_{m=1}^M \frac{\eta_m}{4\pi L_r^2} P_{rm} \times \int dE_{\bar{\nu}_e} \sum_{i=1}^4 \frac{f_i}{E_i} \phi_i(E_{\bar{\nu}_e}) \sigma(E_{\bar{\nu}_e}) P_{ee}(E_{\bar{\nu}_e}, L_r)$$

Detector-reactor distance

Power fraction of component i

Effective thermal power of reactor r in month m

Average energy released per fission of component i

-- MC have been developed in order to take into account the 446 nuclear reactors running during the period of interest

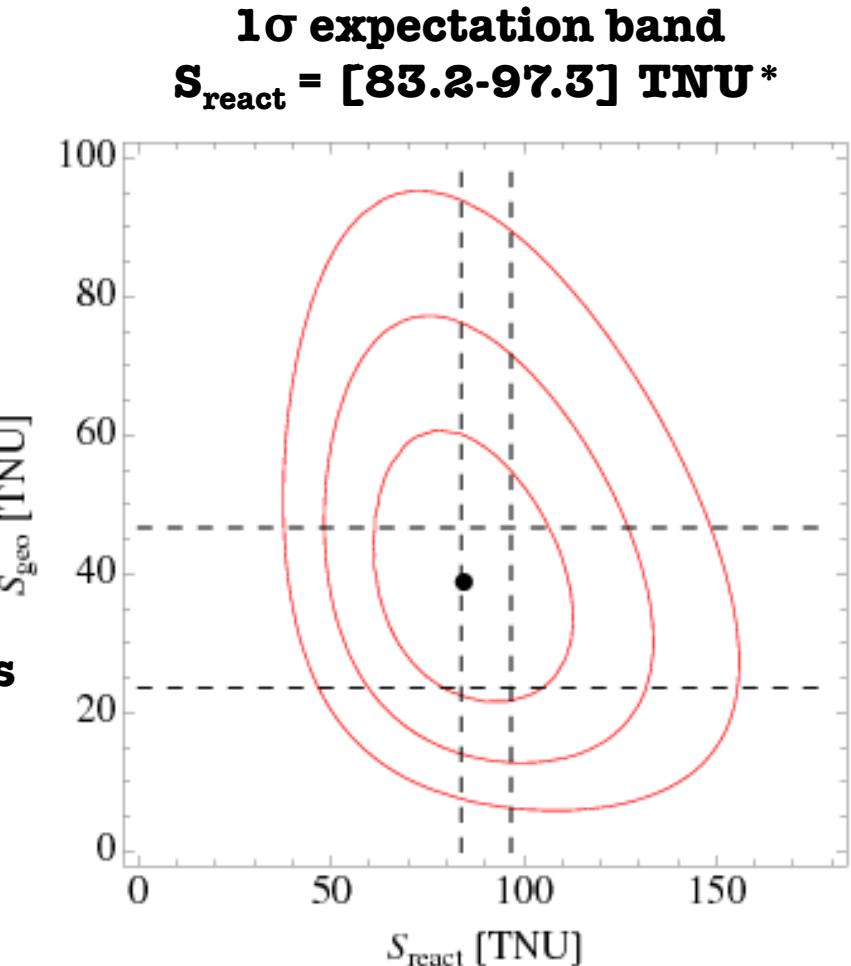
-- Fit analysis --

-- Unbinned maximal likelihood method based on the prompt energy spectra of the candidates

-- The hypothesis that $S_{\text{geo}} = 0$ is rejected at 99.9989% C.L. (4.4σ)

1 σ expectation band
 $S_{\text{geo}} = [26.3-46.6] \text{ TNU}^*$ for
different BSE geological models

$$S_{\text{geo}} = 38.8 \pm 12.0 \text{ TNU}$$



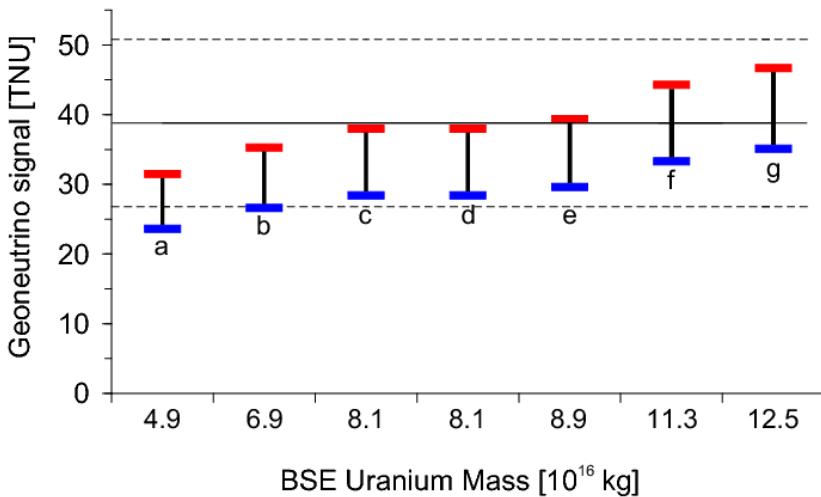
* 1 TNU = 1 event detected over 1 year exposure of 10^{32} target protons at 100 % efficiency

-- BSE geological models --

-- Bulk Silicate Earth (BSE) models describe both the crust and the mantle

- Different BSE models:
 - Cosmochemical
 - Geochemical
 - Geodynamical

BSE S_{geo} [TNU]		Model
- Low -	- High -	
23.6	31.44	Javoy et al. (2010) (a)
26.6	35.24	Lyubetskaya & Korenaga (2007) (b)
28.4	37.94	McDonough & Sun (1995) (c)
28.4	37.94	Allegre et al. (1995) (d)
29.6	39.34	Palme & O'Neil (2004) (e)
33.3	44.24	Anderson (2007) (f)
35.1	46.64	Turcotte & Schubert (2002) (g)



1σ expectation band
 $S_{\text{geo}} = 38.8 \pm 12.0$ TNU from the
Borexino fit analysis

Borexino results in agreement
with BSE models

-- Accessing geo-neutrinos from the mantle --

-- Measured signal = BSE signal = crust signal + mantle signal
where crust = local crust (LOC) + rest of the crust (ROC)

-- Borexino:

- S_{geo} (total) = 38.8 ± 12.0 TNU
- S_{geo} (crust) = 23.4 ± 2.8 TNU

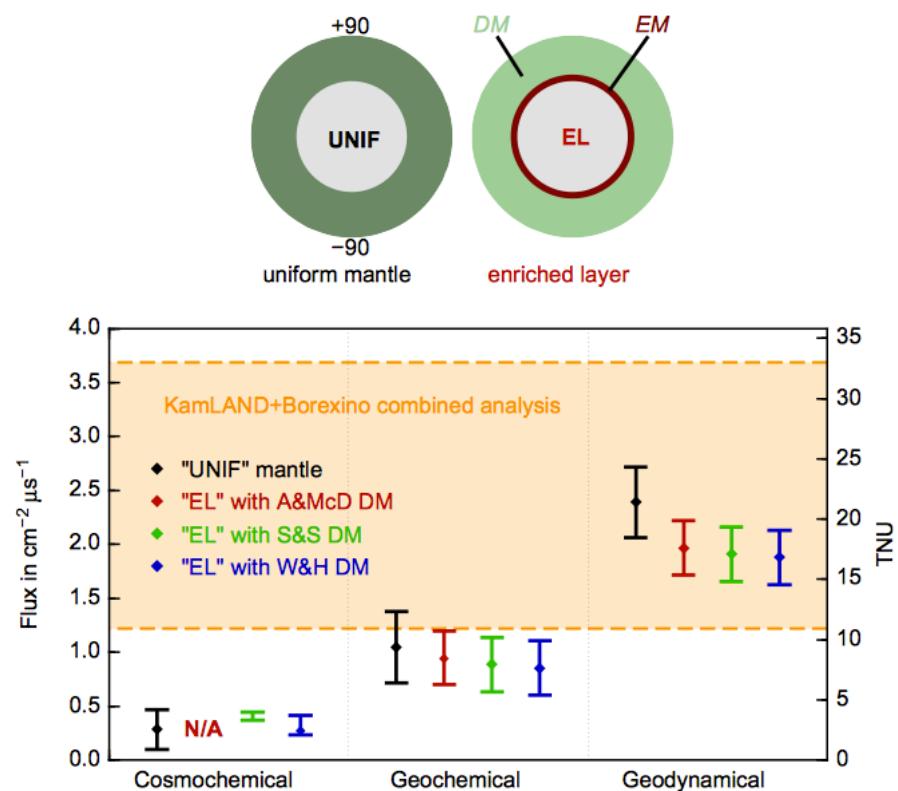
$$\left. \begin{array}{l} - \\ - \end{array} \right] S_{\text{geo}} (\text{mantle}) = 15.4 \pm 12.3 \text{ TNU}$$

-- KamLAND:

- S_{geo} (mantle) = 5.0 ± 7.3 TNU

-- Combined (Sramek et al. (2013),
Ludhova and Zavatarelli (2013))

-- Discrimination between different models is not yet possible, even when combining Borexino and KamLAND



-- Fit analysis with Th/U ratio left free --

-- Fit of U and Th spectra independently

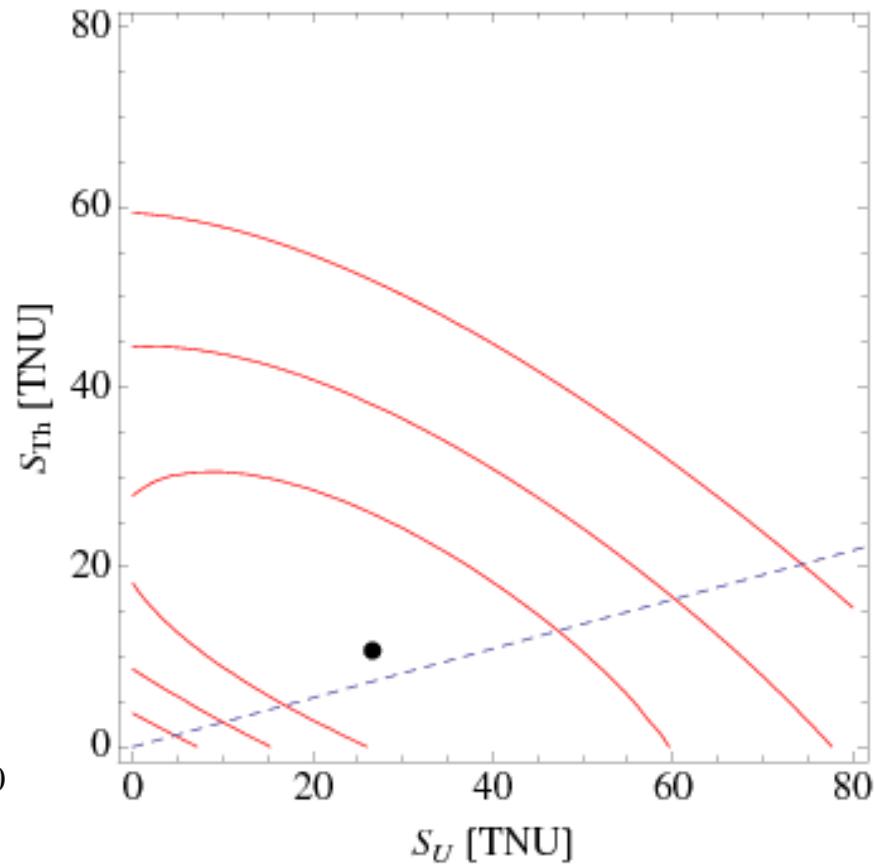
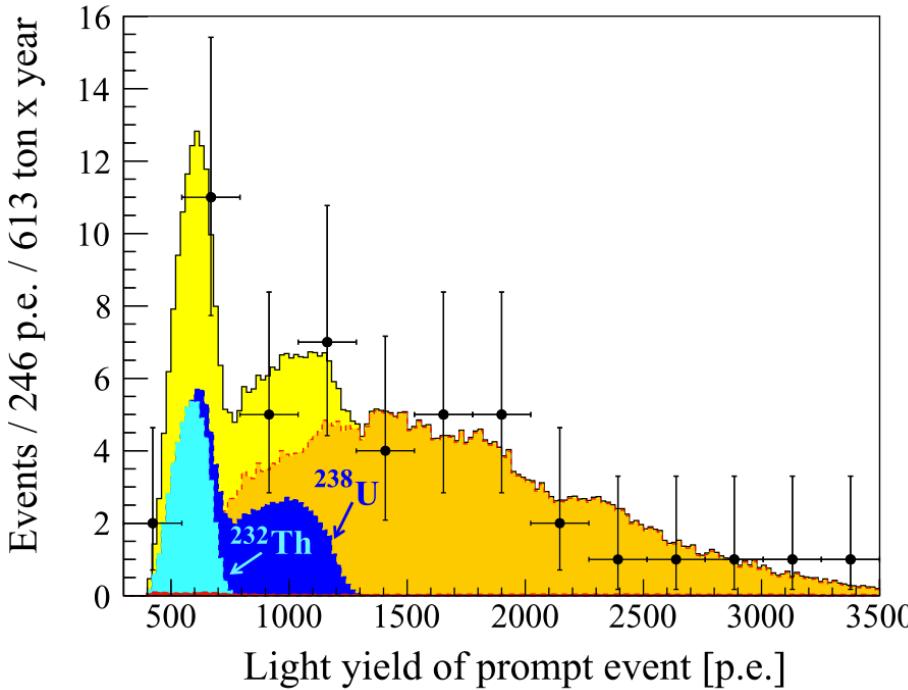
$$S_U = 26.5 \pm 19.5 \text{ TNU}$$

$$S_{\text{Th}} = 10.6 \pm 12.7 \text{ TNU}$$

-- Th/U ratio left free in the fit

--> Best fit value compatible

with the chondritic value of 3.9



-- Conclusion and perspectives --

- Geo-neutrinos are of interest, especially for the mantle knowledge
 - > Borexino helps geoscience to better understand our planet
- 2 years more of statistics
 - > Borexino alone should be able to reach the 5σ signal
- Future large scale detectors coming up soon
 - > SNO+
 - > JUNO
 - > LENA
 - > Hanohano

-- Anti- ν_e from ^{144}Ce - ^{144}Pr decays --> SOX --

-- Several anomalies to be understood

- ① LSND and MiniBooNE
- ② Radiochemical solar experiments (GALLEX and SAGE)
- ③ Short baseline reactor experiments

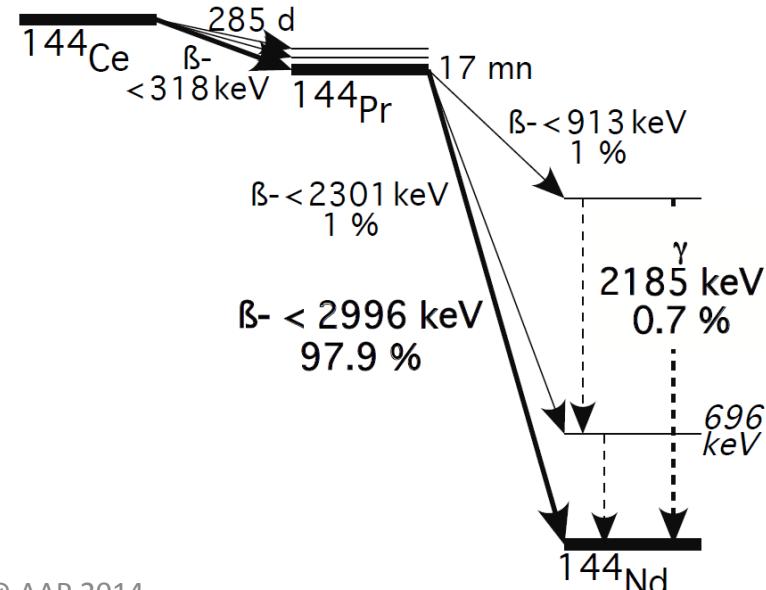
-- Investigation on a possible new neutrino oscillation

--> **SOX** experiment (Short Oscillations with BoreXino)

How? Bringing a 100 kCi ^{144}Ce - ^{144}Pr source (anti- ν_e emitter) below Borexino and looking for the L/E pattern

WHEN? End of 2015

STAY TUNED!

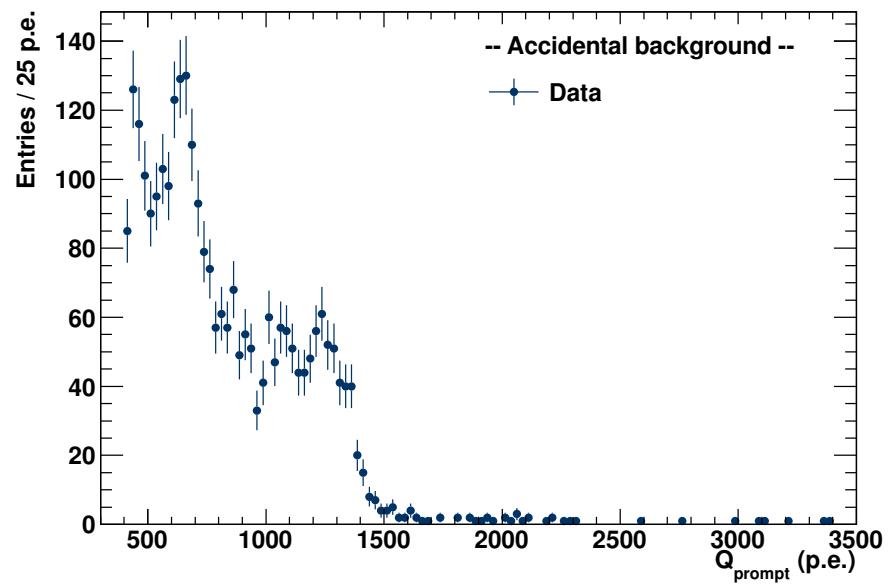
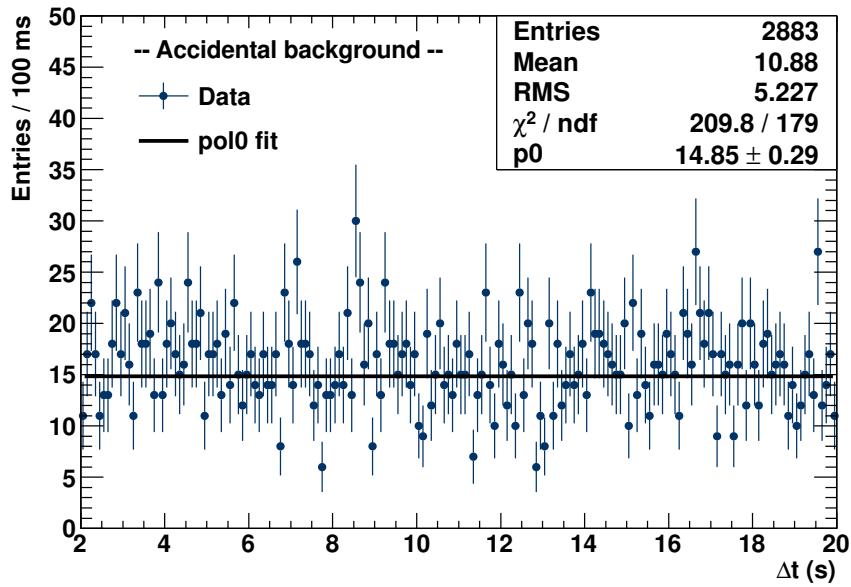




Thank you for your attention

-- Accidental background --

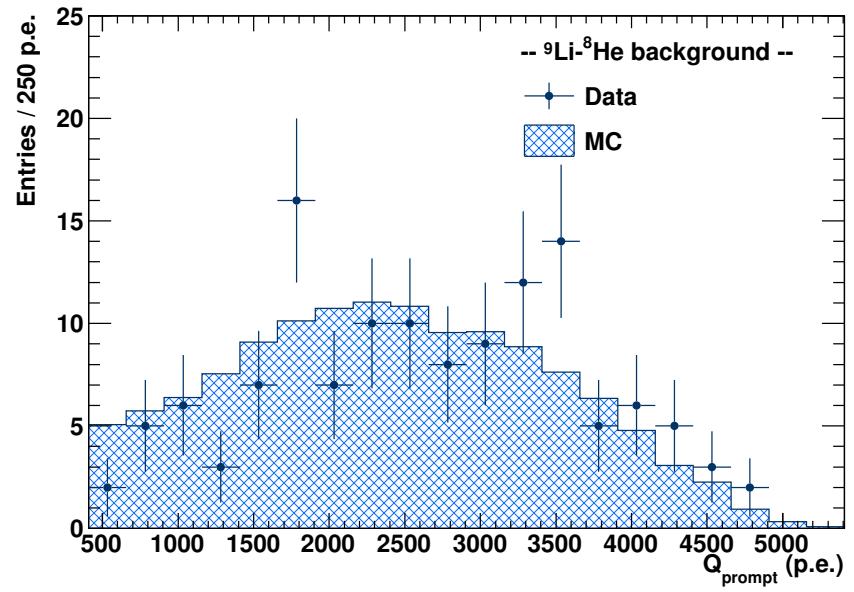
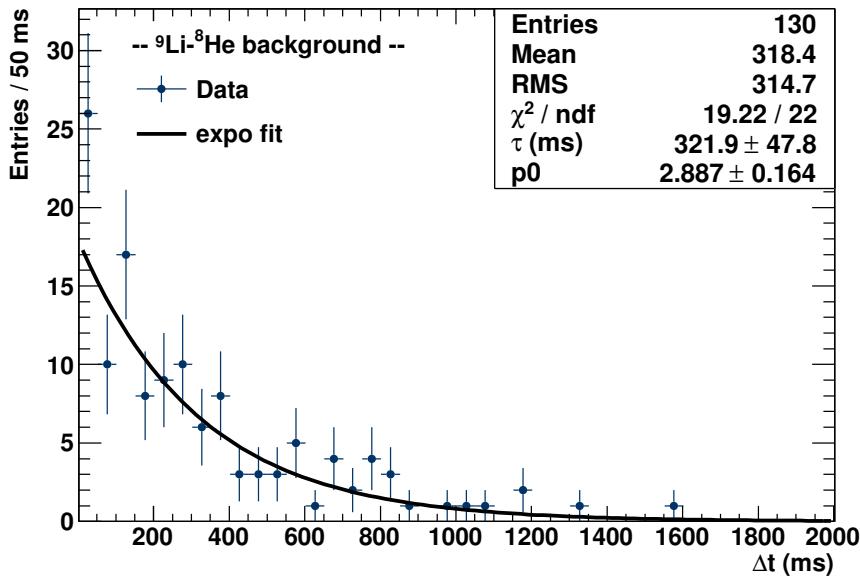
-- Looking for the off time coincidence in the [2 s, 20 s] range after the prompt signal



-- Δt distribution between prompt signal and delayed signal

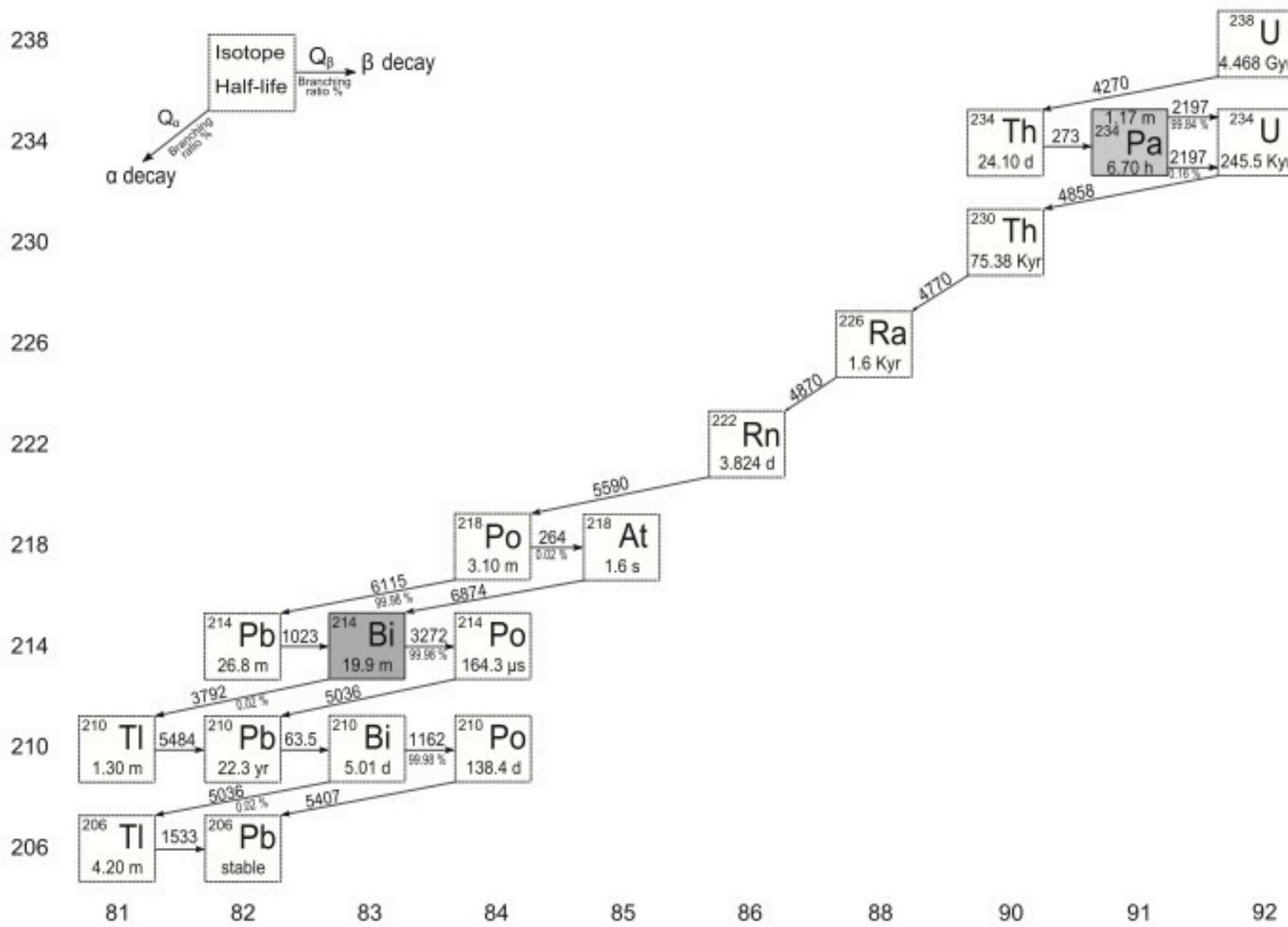
-- ${}^9\text{Li}-{}^8\text{He}$ background --

-- Same anti- ν_e cuts except that we are now looking in the [2 ms, 2 s] range after an internal muon



-- Δt distribution between internal muon and prompt signal

-- ^{238}U decay chain --



-- ^{232}Th decay chain --

