# Precision Reactor $\bar{\nu}_e$ Spectrum Predictions and Measurements

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#### Overview



- The reactor antineutrino spectrum is a valuable handle for:
  - Neutrino oscillation physics
  - Safeguards
  - Nuclear applications
- New precision spectrum measurements now available from θ<sub>13</sub> experiments: What are the implications?
- Talk outline
  - Introduction
  - Detailed look at recent results
  - Discussion and implications





#### Introduction

## **Reactor Antineutrino Production**



#### • Reactor $\overline{v}e$ : produced in decay of product beta branches



# Predicting $S_i(E)$ , Neutrinos Per Fission



- Two main methods:
- Ab Initio approach:
  - Calculate spectrum branch-by-branch using beta branch databases: endpoints, decay schemes
  - Problem: Some rare beta branches with little/ incomplete information; infer these additions
- Conversion approach
  - Measure beta spectra directly
  - Convert to  $\overline{V}_e$  using 'virtual beta branches'
  - **Problem:** 'Virtual' spectra not well-defined: what forbiddenness, charge, etc. should they have?
- Devised in 50's, each method has lost and gained favor over the years

Carter, *et al*, Phys. Rev. 113 (1959) King and Perkins, Phys. Rev. 113 (1958)









## Predicting $S_i(E)$ , Neutrinos Per Fission

• Early 80s: ILL  $\overline{v}_e$  data fits newest *ab initio* spectra well

i.e.: Davis, Vogel, *et al.*, **PRC** 24 (1979) ILL: Kwon, et al., **PRD** 24 (1981)

 I980s: New reactor beta spectra: measurements conversion now provides lower systematics

> Schreckenbach, et al., Phys Lett B160 (1985) Schreckenbach, et al., Phys Lett B218 (1989)

I 990s: Bugey measurements fit converted spectrum well

B.Achkar, et al., Phys Lett B374 (1996)

 I980s-2000s: Predicted, measured fluxes agree in Russian, EU, US exps.



## Recent History: Problems Emerge





0.15

0.10

0.05

0.00

-0.05

φ-φITT)/φITT

- Start with ab initio approach
- Subtract this from ILL beta spectra
- Use conversion procedure on remaining beta spectrum: ~10%
- OR Huber: virtual branches only
- Change in flux/spectrum
  - Predicted and measured fluxes no longer agree.
  - Spectrum shifted to higher energy



## Even More Recent History: More Problems



- Spectra from  $\theta_{13}$  experiments disagree with predictions
  - So now, not only the flux, but also the spectra disagree with predictions.
- Let's go over these in a little more detail.



![](_page_8_Picture_0.jpeg)

## **Recent Spectral Measurements**

## $\theta_{13}$ Experiments

![](_page_9_Picture_1.jpeg)

- Large detectors: 10s-ton single-volume LS target
- Long baselines from conventional cores: 0.1 2.0 km
- 'Large' overburdens: 100+ MWE
- Qualities allow low-background, high-resolution measurements

![](_page_9_Figure_6.jpeg)

## Daya Bay

![](_page_10_Picture_1.jpeg)

- Spend a little more time on Daya Bay, since it's my specialty
  - >2σ deviation from Huber/Mueller(U238) over entire spectrum
  - Zoom in on particular region from 4-6 MeV: >4 $\sigma$  deviation from prediction
  - Hints at deviation in other regimes: perhaps not just a 'bump'
  - Also, don't forget the 5.3% flux deficit reported at Neutrino2014...

![](_page_10_Figure_7.jpeg)

## **RENO** and **Double** Chooz

![](_page_11_Picture_1.jpeg)

- RENO:  $3.5\sigma$  deviation from prediction in vicinity of 5MeV
- Double Chooz:  $I.5\sigma$  deviation over full energy range
  - Coming: more stats of largely unoscillated neutrinos with new near detector
- Note that 'bump has negligible effect on  $\theta_{13}$  rate+shape fits.

![](_page_11_Figure_6.jpeg)

## Skeptical Questions

![](_page_12_Picture_1.jpeg)

- These results indicate that measured nuebar spectra do not match predictions based on beta spectrum conversion
- Before we go there:
  - 'Maybe it's just a background that hasn't been properly accounted for...'
  - 'Maybe this is just an absolute energy scale issue'
  - 'Is there any other strange behavior in the way this excess pops up in the data?'

![](_page_12_Picture_7.jpeg)

## Skeptical Question I

'Maybe it's just a background that hasn't been accounted for.'

![](_page_13_Figure_2.jpeg)

Reconstructed Energy (MeV)

## Skeptical Question 2

![](_page_14_Picture_1.jpeg)

#### 'Maybe this is just an absolute energy scale issue'

![](_page_14_Figure_3.jpeg)

![](_page_14_Figure_4.jpeg)

No bump or other strange behavior in B-12 spectrum WRT prediction

Experiments' electronics (and attendant non-linearities) differ greatly, but all see the same structure.

## Skeptical Question 3

RENO Preliminary

![](_page_15_Picture_1.jpeg)

RENO Preliminary

Events/MeV ≥ ₩30000 3500 Far detector Near detector 'Is there any other strange 3000 Data - Data a25000 MCosc MCosc behavior in the way this 2500  $sin^2 2\theta_{13} = 0.100$  $sin^2 2\theta_{11} = 0.100$ 20000 2000  $\Delta m_{f_1}^2 = 2.32 \times 10^{-3} \text{ eV}$  $\Delta m_{1.}^2 = 2.32 \times 10^{-3} \text{ eV}^2$ excess pops up in the data?' 15000 1500 10000 1000E 5000 500 MC) / MC 0.2 (Data - MC) / MC **RENO** Fa **RENO** Near 0.15 0.05 No time-dependent -0.05 -0.05 spectral changes observed Prompt Energy [MeV] Prompt Energy [MeV] 0.22 0.2 Detectors see the same general feature vents (Unity Normalization) 0.18 0.16 eks 17-24 0.14 AD4+AD5+AD6 eks 25-32 AD1+AD2+AD3 0.12 0.1 0.08 0.06 0.04 0.02 1.1 Ratio to Average 1.05E Daya Bay Daya Bay **Preliminary Preliminary** 0.95 0.85 10 12 6 Prompt energy (MeV) Prompt energy (MeV) Prompt Positron Energy (MeV) 10

## Piling On

![](_page_16_Picture_1.jpeg)

Events

300

250

200

150

100

50

- Not just one faulty experiment broad agreement.
- Different electronics and scintillator (to some degree)
- Overburdens, backgrounds vary widely between experiments
- Other notable results:
  - CHOOZ: A hint present, low CL
  - Bugey3: Seems like no feature is present?
    - Large non-scintillating volume in target? Binning?
    - Something else?

![](_page_16_Figure_10.jpeg)

![](_page_16_Figure_11.jpeg)

![](_page_16_Picture_12.jpeg)

![](_page_17_Picture_0.jpeg)

## **Discussion and Implications**

## Discussion

![](_page_18_Picture_1.jpeg)

- Visible discrepancy between measured and predicted fluxes
- Root cause could be systematics in  $\theta_{13}$  experiments, but good evidence exists to doubt this
- Could predictions be the root cause? How exactly?
- What else can we do to clarify the picture?
- How does this relate to non-proliferation? Applications?

## Forbidden Decay Handling in Conversion

![](_page_19_Picture_1.jpeg)

 Conversions make simple assumptions about forbidden-ness of involved beta branches

![](_page_19_Figure_3.jpeg)

- Capable of shifting predicted flux downward by 5%
- Has not been shown what forbidden decay treatment would reproduce both reactor beta and nuebar spectra but it might be possible to do so

FIG. 3: Different treatments of the forbidden GT transitions contributing to the antineutrino spectrum summed over all actinides in the fission burn in mid-cycle [21] of a typical reactor. The left panel shows the ratio of these antineutrino spectra relative to that using the assumptions of Ref. [4]. The right panel shows the spectra weighted by the detection cross section, where the additional curve in black uses the assumptions of Ref. [4]. The spectra are strongly distorted by the forbidden operators, being lower below the peak and in some cases more than 20% larger above the peak than Ref. [4]. The corresponding change in the number of detectable antineutrinos relative to [4] is -0.75%, 5.8% and 1.85% for the  $0^-, 1^-$ , and  $2^-$  forbidden operators, respectively.

## **Recent Ab Initio Predictions**

![](_page_20_Picture_1.jpeg)

- What if we just compare measured shape directly to *ab initio*?
  - Much better agreement in spectrum
  - Not so much on the overall flux...
  - Some spectral features also present in *ab initio* calculations from Mueller, *et al*.

![](_page_20_Figure_6.jpeg)

![](_page_21_Picture_0.jpeg)

## New Data, New Constraints

- Q: How do we clarify this picture?
  - A: Make new measurements, get more handles!
- Upcoming short-baseline experiments have opportunity to measure absolute spectrum while searching for oscillations
  - High statistics: certainly on par with  $\theta_{13}$  measurements
  - Better resolution = better discrimination power between models
  - HEU spectrum measurement = additional handle to test models
- Further clarity can be valuable to neutrino, nuclear, nonproliferation, and applications communities

## Implications for Non-Proliferation

![](_page_22_Picture_1.jpeg)

- What is spectral shape difference between U-235 and Pu-239?
  - Huber, Mueller predict spectral difference, but don't predict the right spectrum
  - Ab initio calculations also suggest a spectral difference, but not identically
- Without this knowledge, more uncertainty in modelling/ demonstrating Pu239 production monitoring with antineutrinos
  - Measuring this difference directly could resolve this uncertainty, provide fodder for model-fitting

![](_page_22_Figure_7.jpeg)

23

![](_page_23_Picture_0.jpeg)

## Implications for Nuclear Applications

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- Why is there more decay heat than predicted 3-3000s after a reactor is turned off???
- Means we need higher cooling safety factors during reactor-off periods: this costs \$\$\$!!!
- 5 MeV nuebar 'bump' produced by many isotopes of great concern to this decay heat measurement
- High-res measurement may constrain individual isotopes
  - Direct check on concerning **ENSDF** nuclear data
    - TOTALLY different systematics!
  - Isotopes: Rb-92, Sr-97, Cs-142

Figure 3. Electromagnetic decay heat following thermal fission burst of <sup>239</sup>Pu - data from JENDL, JEF-2.2, JEFF-3.1 and ENDF/B-VI are shown together with experimental data from Yayoi, Lowell and Oak Ridge National Laboratory

![](_page_23_Figure_10.jpeg)

## Implications for Oscillation Physics

![](_page_24_Picture_1.jpeg)

- Pointed out yesterday: can we really use existing flux uncertainty estimates in our SBL sterile searches?
  - $\theta_{13}$  experiments appear to show these error bands are too small

![](_page_24_Figure_4.jpeg)

## Implications for Oscillation Physics

![](_page_25_Picture_1.jpeg)

- Will fine structure affect measurement of mass hierarchy?
  - Magnitude of spectral features in flux comparable to that of mass hierarchy
  - However, hierarchy gives very distinct energy-dependent signature
- Knowledge of underlying structure will improve confidence in a hierarchy-related spectral distortion measurement

![](_page_25_Figure_6.jpeg)

## Summary

![](_page_26_Picture_1.jpeg)

- State-of-the-art reactor spectrum predictions are not matched by recent direct nuebar spectrum measurements
- These are the same predictions used to:
  - Produce reactor flux estimates for the 'reactor antineutrino anomaly'
  - Benchmark neutrino oscillation results
  - Demonstrate Pu-239 production monitoring using antineutrinos
- New high-resolution measurements of HEU and LEU fuel will be essential to clarifying this picture

![](_page_27_Picture_0.jpeg)

## END

## Historical Context

![](_page_28_Picture_1.jpeg)

- A similar experimental setup in the past: Bugey-3
  - Segmented short-baseline LiLS detector
- PROSPECT Pros:
  - Smaller reactor core, closer to core: better for SBL oscillation search
  - Stable scintillator: Bugey's degraded after a few months in near detector!
  - Smaller target dead volume: ~2% versus >15% for Bugey
  - Aim for better light yield, PSD
- PROSPECT Con: No Overburden
  - 14+ mwe (Bugey-3), <10 mwe (PROSPECT)
  - Bugey had 25:1 S:B

![](_page_28_Figure_12.jpeg)

## Beta Decay Recap

![](_page_29_Picture_1.jpeg)

![](_page_29_Figure_2.jpeg)

30

E, [MeV]

## Reactor Spectroscopy: Example

![](_page_30_Picture_1.jpeg)

- TAGS: Total absorption gamma spectroscopy
- Measure total gamma energy, not individual gamma energies
- Allows ID of levels, BRs much easier

![](_page_30_Figure_5.jpeg)

- If branching ratios are known better, decay released in those decays will be modelled better
- Better model = smaller safety factor = \$\$\$ saved.