

# **Newest Results from MINOS**

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## **The MINOS Collaboration**



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## **The MINOS Experiment**



- **MINOS (Main Injector Neutrino Oscillation Search)** 
  - Long-baseline neutrino oscillation experiment

#### **Basic concept**

- Create a neutrino beam provided by 120 GeV protons from the Fermilab Main Injector
- Measure energy spectrum at the Near Detector, at Fermilab
- Measure energy spectrum at the Far Detector, 735 km away, deep underground in the Soudan Mine.
- Compare Near and Far measurements to study neutrino oscillations



# **MINOS Physics Goals**

- Precise measurements of  $|\Delta m^2_{32}|$  and  $sin^22\theta_{23}$  via  $\nu_{\mu}$  disappearance
- Search for or constrain exotic physics such as sterile  $\boldsymbol{v}$
- Search for sub-dominant  $v_{\mu} \rightarrow v_{e}$  oscillations via  $v_{e}$  appearance
- Compare v,  $\overline{v}$  oscillations
- Atmospheric neutrino and cosmic ray physics
- Study v interactions and cross sections using the high statistics Near Detector data set



## **Event Topologies**

### Monte Carlo

#### $v_u$ CC Event



long µ track & hadronic activity at vertex

#### **NC Event**



short event, often diffuse

#### $v_e$ CC Event



short, with typical EM shower profile

#### **Energy resolution**

- •π<sup>±</sup>: 55%/√E(GeV)
- µ<sup>±</sup>: 6% range, 10% curvature



## Charged Current Analysis of 3.36×10<sup>20</sup> POT of MINOS Data

# - Precision measurement of $|\Delta m^2|$ and $sin^2 2\theta$ -



## **CC Event Selection**

- CC / NC Event classification is performed with a k-nearest neighbor (kNN) based algorithm with four inputs:
  - Track length (planes)
- For hits belonging to the track:
  - 2. Mean pulse height/plane
  - 3. Fluctuation in pulse height
  - 4. Transverse track profile

Near Detector

CC selection eficiency

**MINOS Preliminary** 

NC contamination



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CC efficiency / NC contamination

0.8

0.6

0.4

0.2

0

0

2



## **Systematic Uncertainties**

• The impact of different sources of systematic uncertainty is evaluated by fitting modified MC in place of the data:



The three largest shifts are included as nuisance parameters in the oscillation fit.

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# **CC Energy Spectrum Fit**

 Fit the energy distribution to the oscillation hypothesis:

 $P(\nu_{\mu} \rightarrow \nu_{\tau}) = \sin^2(2\theta) \sin^2\left(\frac{1.27\Delta m^2 L}{E}\right)$ 

- Including the three largest sources of systematic uncertainty as nuisance parameters
  - Absolute hadronic energy scale: 10.3%
  - -Normalization: 4%
  - -NC contamination: 50%





## **Allowed Regions**



#### Most precise measurement of $|\Delta m|^2$ performed to date!

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## **Alternative Hypotheses**





## Neutral Current Analysis of 2.46×10<sup>20</sup> POT of MINOS Data

- Looking for sterile neutrino mixing -



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## **NC/CC Event Separation**

- NC events are typically shorter than CC events
- Expect showers and no tracks or very short tracks reconstructed for NC events
- Main background from inelastic (high-y)  $v_{\mu}$  CC events



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## **NC Analysis Results - Rate**

- Compare the NC energy spectrum with the expectation of standard 3-flavor oscillation physics
  - Depletion of Far Detector NC spectrum may indicate sterile neutrino mixing
- Fix the oscillation parameter values

 $R \equiv \frac{N_{Data} - E}{S_{uc}}$ 

- $\sin^2 2\Theta_{23} = 1$   $\Delta m_{32}^2 = 2.43 \times 10^{-3} \text{ eV}^2$   $\Delta m_{21}^2 = 7.59 \times 10^{-5} \text{ eV}^2$ ,  $\Theta_{12} = 0.61$  from KamLAND+SNO
- $\Theta_{13}$  = 0 or 0.21 (normal MH,  $\delta$ =3 $\pi$ /2) from CHOOZ Limit
  - N.B. CC  $v_e$  are classified as NC by the analysis
- Make comparisons in terms of the *R* statistic:
- For different energy ranges •
  - 0-3 GeV
  - 3-120 GeV
  - All events (0-120 GeV)

Predicted CC background from all flavors

Predicted NC

interaction signal



## **NC Analysis Results - Rate**

- Plot shows the selected FD NC energy spectrum for Data and oscillated MC predictions
- Expect largest NC disappearance for E < 3 GeV if sterile mixing is driven by Δm<sup>2</sup><sub>32</sub>
- Depletion of total NC event rate (1-*R*) < 17% at 90% C.L. for the 0-120 GeV range



Data is consistent with					
no NC deficit at FD					
and thus with no					
sterile neutrino mixing					

$E_{\rm reco}  ({\rm GeV})$	$N_{\text{Data}}$	$S_{\rm NC}$	$B_{\rm CC}^{\nu_{\mu}}$	$B_{\rm CC}^{\nu_{\tau}}$	$B_{\rm CC}^{\nu_e}$
0 - 3	100	101.1	11.2	1.0	1.8(9.3)
3 - 120	191	98.0	64.2	3.5	11.8(24.6)
0 - 3	R = 0.	$85 \pm 0.1$	$0 \pm 0.07$	$(0.78 \pm$	$0.10 \pm 0.07$
3 - 120	R = 1.	$14 \pm 0.1$	$4 \pm 0.10$	$(1.02 \pm$	$0.14 \pm 0.10$
0 - 120	R = 0.	$99 \pm 0.0$	$9 \pm 0.07$	$(0.90 \pm$	$0.09 \pm 0.08$



## NC Analysis Results – f<sub>s</sub>Fit

- Assume one sterile neutrino and that mixing between  $\nu_{\mu},\,\nu_{s}$  and  $\nu_{\tau}$  occurs at a single  $\Delta m^{2}$
- Survival and sterile oscillation probabilities become:  $P(\nu_{\mu} - \nu_{\mu}) = 1 - \alpha_{\mu} \sin^{2}(1.27\Delta m^{2}L/E)$  $P(\nu_{\mu} - \nu_{s}) = \alpha_{s} \sin^{2}(1.27\Delta m^{2}L/E)$
- Simultaneous fit to CC and NC energy spectra yields the fraction of  $\nu_{\mu}$  that oscillate to  $\nu_{s}$ :

$$f_{s} = \frac{P(v_{\mu} \to v_{s})}{1 - P(v_{\mu} \to v_{\mu})} = 0.28^{+0.25}_{-0.28} \text{(stat.+syst.)}$$

 $f_s < 0.68$  (90% C.L.)

Submitted to PRL (arXiv:hep-ex/0807.2424)



## v<sub>e</sub> Appearance Analysis

- Constraining  $\theta_{13}$  -



### ve Selection

•NC and short  $v_{\mu}$  CC events are the dominant backgrounds

•Neural Network  $v_e$  selection algorithm based on characteristics of electromagnetic showers

 MC tuned to bubble chamber experiments for hadronization models

•Data/MC comparisons show disagreements due to hadronic model

 Correct the model to match the data using data-driven methods in ND

 Background predictions from two methods agree within statistical uncertainty





Reconstructed Energy (GeV)



# Future $\theta_{13}$ Limits

- Expect 12 signal and 42 bg events at the CHOOZ limit for the current exposure
- Data-driven systematics are hoped to drop to 5% in future years
- Inverted hierarchy shown only for lowest exposure for simplicity





## **Summary and Conclusions**

- The MINOS Experiment is making several contributions to our understanding of Neutrino Physics
- New measurement of atmospheric oscillation parameters from  $v_{\mu}$  disappearance:
  - $-\left( |\Delta m|^2 = (2.43 \pm 0.13) \times 10^{-3} \text{ eV}^2 \right) (68\% \text{ C.L.})$
  - $-(\sin^2(2\theta) > 0.90 (90\% \text{ C.L.})$

- Decay and decoherence models are disfavored at 3.7  $\sigma$  and 5.7  $\sigma$  , respectively

• New results from search for oscillations into sterile neutrinos:

- -(1-R < 17% at 90% C.L., 0 < E < 120 GeV)
- $[f_s < 0.68 \quad (90\% \text{ C.L.})$
- Consistent with no sterile neutrino mixing
- First results on  $v_e$  appearance expected later this year and have sensitivity below the CHOOZ limit.



## **Backup Slides**

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# **The NuMI Neutrino Beam**



Hadron Monitor

• Beam energy spectrum can be modified by varying the relative positions of target and horns.



92.9% 
$$v_{\mu}$$
, 5.8%  $\overline{v}_{\mu}$ , 1.3%  $v_e + \overline{v}_e$ 

- Beam performance:
  - 10µs spill of 120 GeV protons every 2.2s
  - Intensity: 3.0×10<sup>13</sup> POT/spill
  - 0.275 MW beam power
  - 10<sup>18</sup> POT/day





## **Accumulated Beam Data**

Many thanks to Fermilab's Accelerator Division

Total NuMI protons to 00:00 Monday 25 August 2008





## **The MINOS Detectors**

#### **Near Detector**

- Located 1km downstream of the target
- ~1kt (980t) total mass
- Shaped as squashed octagon (4.8×3.8×15m<sup>3</sup>)
- Partially instrumented (282 steel, 153 scintillator planes)
- Fast QIE readout electronics, continuous sampling during beam spill



Muon Spectrometer planes 121 : 281

#### Far Detector

- Located 735km away in Soudan mine, MN
- 5.4kt, 2 supermodules
- Shaped as octagonal prism (8×8×30m<sup>3</sup>)
- 486 steel planes, 484 scintillator planes
- Veto shield (scintillator modules)
- Spill times from Fermilab for beam trigger



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Hadron Shower

planes 61 : 120

Target

planes 0:20 planes 21:60

Veto

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## **Near to Far Extrapolation**

Far detector energy spectrum without oscillations is not the same as the Near detector spectrum



- Start with near detector data and extrapolate to the far detector
  - Use Monte Carlo to provide corrections due to energy smearing and acceptance
  - Encode pion decay kinematics and the geometry of the beamline into a beam transport matrix used to transform the ND spectrum into the FD energy spectrum





## **Data Sensitivity**

For a true value at the Sensitivity l∆m²l (×10<sup>-3</sup>eV²) 5.2 best fit point of:  $|\Delta m^2| = 2.43 \times 10^{-3} \, eV^2$  $sin^{2}(2\theta) = 1.00$ , **MINOS** data 26.5% of unconstrained fits have a fit value of  $sin^2(2q) \ge 1.07$ . 90% c.l. contours 0.8 0.85 0.95 0.9  $sin^2(2\theta)$ 

## **Neutral Current NC Energy Spectrum**

NC selected Data and MC energy spectra for Near Detector



- Good agreement between Data and Monte Carlo
- Discrepancies much smaller than systematic uncertainties
- NC events are selected with 90% efficiency and 60% purity



## **Systematic Errors**

- **Relative Normalization**: ±4%
  - POT counting, Near/Far reconstruction efficiency, fiducial mass
- **Relative Hadronic Calibration:** ±3%
  - Inter-Detector calibration uncertainty
- Absolute Hadronic Calibration: ±11%
  - Hadronic Shower Energy Scale(±6%), Intranuclear rescattering(±10%)
- Muon energy scale: ±2%
  - Uncertainty in dE/dX in MC
- CC Contamination of NC-like sample: ±15%
- NC contamination of CC-like sample: ±25%
- Cross-section uncertainties:
  - $m_A$  (qe) and  $m_A$  (res): ±15%
  - KNO scaling: ±33%
- **Poorly reconstructed events:** ±10%
- Near Detector NC Selection: ±8% in 0-1 GeV bin
- Far Detector NC Selection: ±4% if E < 1 GeV,</li>
   <1.6% if E > 1 GeV
- Beam uncertainty: 1σ error band around beam fit results

	$0-3~{ m GeV}$	$3 - 120 \mathrm{GeV}$
Absolute $E_{had}$	$\pm < 0.01$	$\pm 0.05$
Relative $E_{had}$	$\pm 0.03$	$\pm 0.04$
Normalization	$\pm 0.04$	$\pm 0.08$
Near detector selection	$\pm 0.02$	_
$\nu_{\mu}$ -CC background	$\pm 0.03$	$\pm 0.01$
Total:	$\pm 0.07$	$\pm 0.10$

## Effect of the most relevant systematic uncertainties on **R**



# $v_{\mu}$ to $v_{sterile}$ in SuperK



- High energy v experience matter effects which suppress oscillations to sterile v
  - Matter effects not seen in upμ or high-energy PC data
  - Reduction in neutral current interactions also not seen
  - constrains  $v_s$  component of  $v_u$  disappearance oscillations
- Pure  $v_{\mu}$ -> $v_s$  disfavored
  - $v_s$  fraction < 20% at 90% c.l.
- Result published only in conference proceedings



## ve Selection

- •Neural Network  $v_e$  selection algorithm based on characteristics of electromagnetic showers
- •MC tuned to bubble chamber experiments for hadronization models
- •Data/MC comparisons show disagreements due to hadronic model
- •Developed two data-driven methods to correct the model to match the data
- Muon Removed CC Events (MRCC)
  - Use well understood  $\nu_{\mu}$  CC data sample with removed track hits to correct NC event number
  - Beam  $\nu_e$  known from MC, subtract from NC component to obtain  $\nu_\mu$  CC
- •Horns on/off
  - pions are not focused with horns off and energy spectrum peak disappears
  - Estimate NC and  $\nu_{\mu}$  CC from differences between horns on/off data samples and MC
- Extrapolate each background to FD to obtain data-driven sensitivities







## **Muon Removal**

>20% Data/MC discrepancy in both the standard v<sub>e</sub> and the muon removed CC samples

- Comparisons of standard Data and MC shower topological distributions disagree in the same way as does MRCC data with MRCC MC
  - So MC hadronic shower production/modeling is a major contribution to the disagreement.
  - Kinematic phase space of MRCC and selected NC events matches well, but MRCC and selected CC events do not.
- The MRCC sample is thus used to make ad-hoc correction to the model to NC events per bin
  - Beam  $v_e$  from MC, CC events are the remainder





## Horn on/off Method

• After applying  $v_e$  selection cuts to Near Detector data, the composition of the selected events is quite different with the NuMI focusing horns on or off.

$$N_{on} = N_{NC} + N_{CC} + N_{e}$$
(1)  
$$N_{off} = r_{NC} * N_{NC} + r_{CC} * N_{CC} + r_{e} * N_{e}$$
(2)

from MC:  

$$\mathbf{r}_{NC(CC,e)} = \mathbf{v}_{NC(CC,e)}^{off}/\mathbf{N}_{NC(CC,e)}^{on}$$
  
• Get horn on/off ratios from MC,  
then solve for NC and CC  
backgrounds in bins of energy,  
get beam  $v_e$  from the beam  
MC (a well understood number)  
– Independent of hadronic modeling

**Reconstructed Energy (GeV)**