

# First light...

at HAWC Observatory's high altitude TeV gamma ray detector in Mexico

Daniel Fiorino for the HAWC Observatory 4 March 2012 Rencontres de Moriond



# Aerial View of HAWC Site





# The HAWC Observatory



High altitude 4100m

Next generation water **Cherenkov detector** (Milagro was previous generation)

High energy air shower array ~100 GeV - 100 TeV

**Cosmic Ray Origins,** Gamma Ray Bursts, many more

Ititude

ater

Cherenk

HAWC 300 **Upon Completion** 

### HAWC Equipment





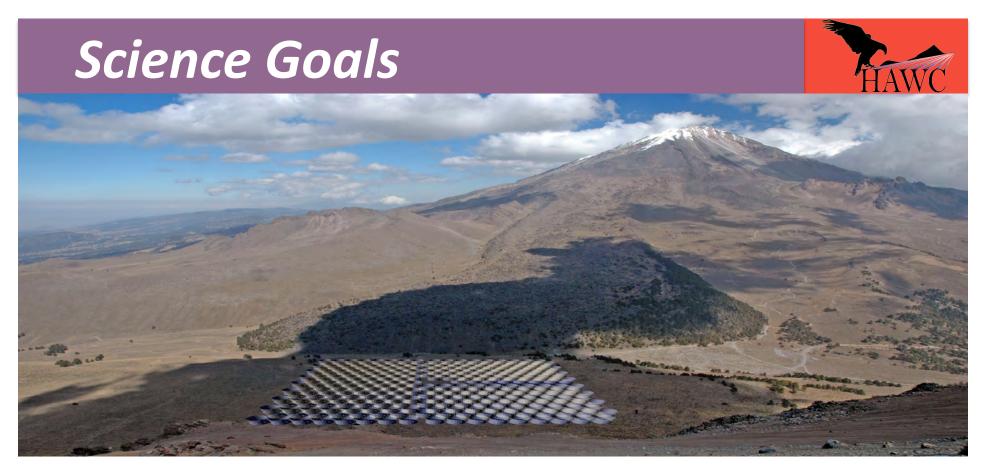
#### **Cherenkov Tanks**

Corrugated Steel
4.0m tall, 7.3m wide
Target for Cherenkov light
4 photomultiplier tubes



#### **Photomultiplier Tubes**

900 Hamamatsu 8" bulb
Reused from Milagro
Sensitive to UV-optical blue
300 High quantum efficiency (~30%)
Lower energy threshold



#### <u>Unbiased Skymap</u> <u>Galactic Sources of Gamma Rays</u>

Supernovae Remnants
Crab Nebula (SN 1054)
Standard Candle... Recent Flares
Extended Objects (e.g. Molecular Clouds)
Galactic Plane

#### Extragalactic (z<0.1) Sources of Gamma Rays

Active Galactic Nuclei
flaring
multiwavelength campaign (Fermi-LAT)
Gamma Ray Bursts

counterpart (Fermi-LAT/optical telescope alerts)
assumptions of spectra, constrain EBL

Nearby Galaxies starburst galaxies (many SNRs)

# VAMOS – Test Array





Verification and Monitoring of Systems

7 Cherenkov tanks

3 months of data

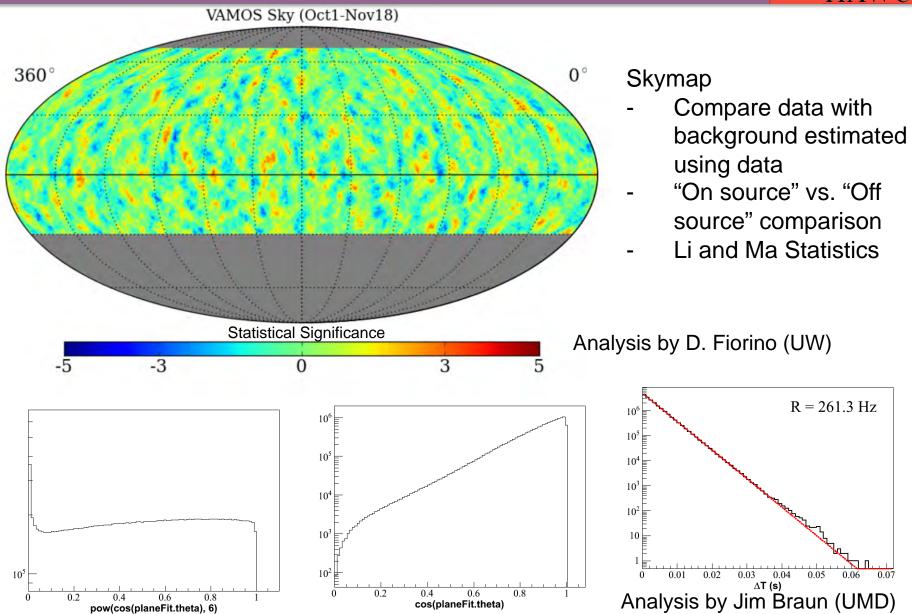
7 photomultipliers per tank

Exceeded set milestones
Improved deployment technique
Online systems running
Analysis Chain -> Skymap



# First Light





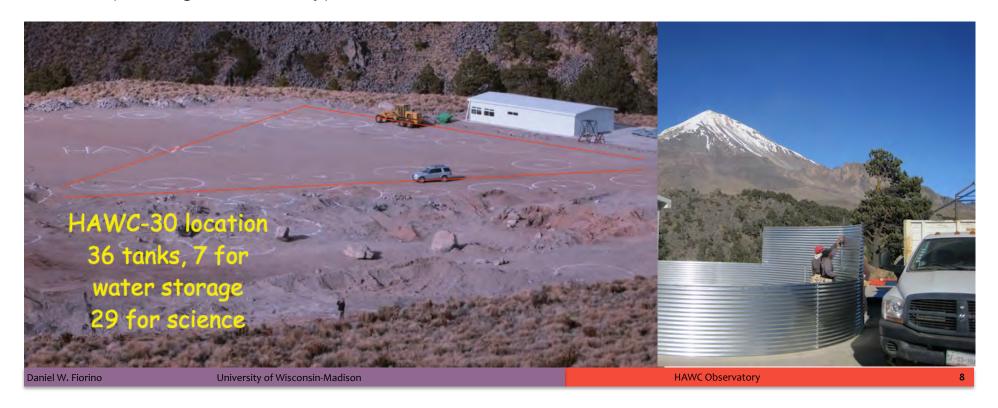
# Stay Tuned...



30 Tanks are scheduled to be operational this Fall (2012)

Site is leveled Tank positions are laid out Electronics ~ May 2012 Verify Cosmic Ray Rates Zenith Alignment Cosmic Ray Anisotropy? Moon Shadow?

HAWC 100 Tanks (> Milagro sensitivity)





### The HAWC Collaboration







Colorado State University George Mason University Georgia Tech University Harvey Mudd University Los Alamos National Lab Michigan State University Michigan Tech University NASA/Goddard Ohio State at Lima
Pennsylvania State University
University of California-Irvine
University of Maryland
University of New Hampshire
University of New Mexico
University of Utah
University of Wisconsin-Madison

Benemerita Universidad Autonomica de Puebla Centro de Investigacion y de Estudios Avanzados Universidad Autonoma de Chiapas Universidad de Guadalajara Universidad de Guanajuato Universidad Michoacana de San Nicolas de Hidalgo Universidad Nacional Autonoma de Mexico

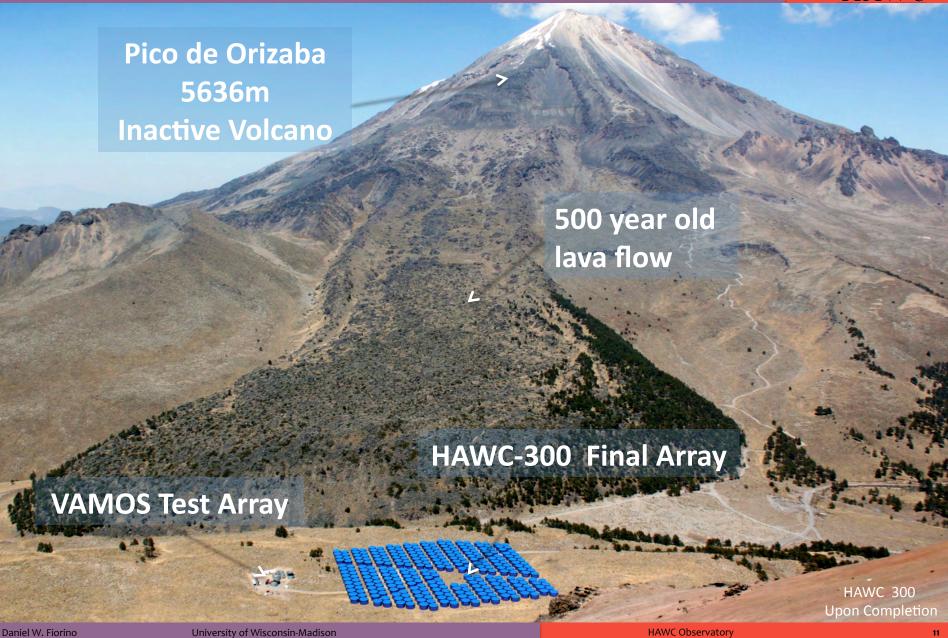






# The HAWC Observatory





# HAWC Improves Upon Milagro

Water Cherenkov experiments provide unbiased information that no others can provide. HAWC is the logical next step.

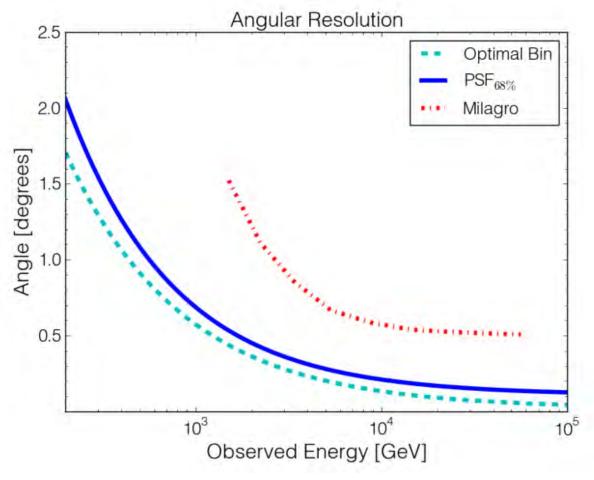
Optically isolated tanks

Better resolution (timing and charge)

Higher altitude
Closer to shower maximum
More particles, lower energy threshold

High quantum efficiency PMTs
More Cherenkov light yield

### Improved Angular Resolution



Error in reconstructing arrival direction

Simulated showers (CORSIKA) incident on simulated HAWC (GEANT4)

Competitive with imaging air Cherenkov telescopes at highest energies

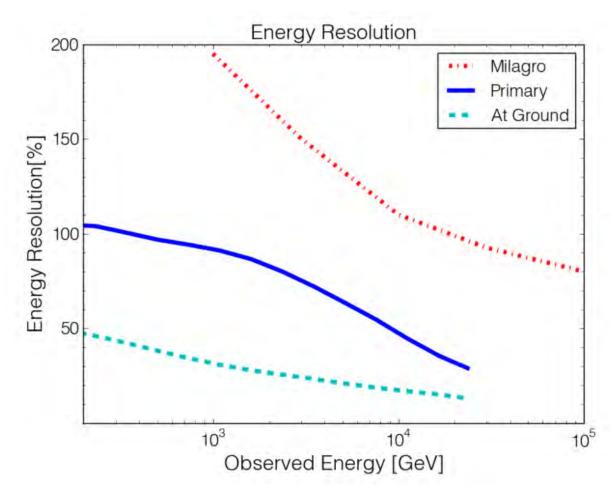
Optical isolation

Dense sampling

**HAWC Observatory** 

PSF = Point spread function

### Improved Energy Resolution



Error in reconstructing energy

Simulated showers (CORSIKA) incident on simulated HAWC (GEANT4)

Improved source spectra over Milagro

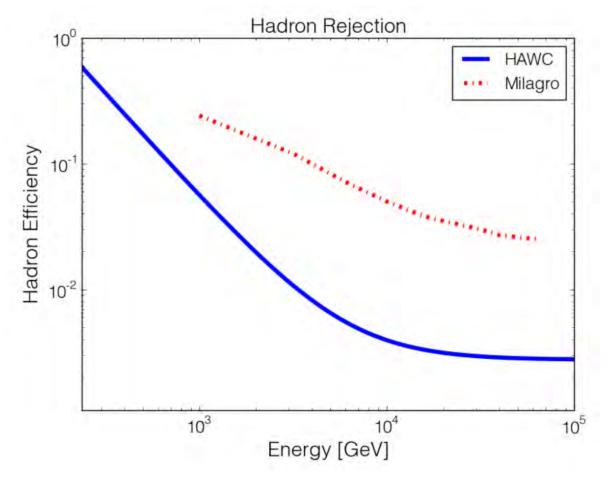
Energy measured at ground is not a perfect indicator of primary energy

Primary Actual error in estimating primary energy

At Ground Energy resolution to energy at ground level

Daniel W. Fiorino

# Improved y/Hadron Separation



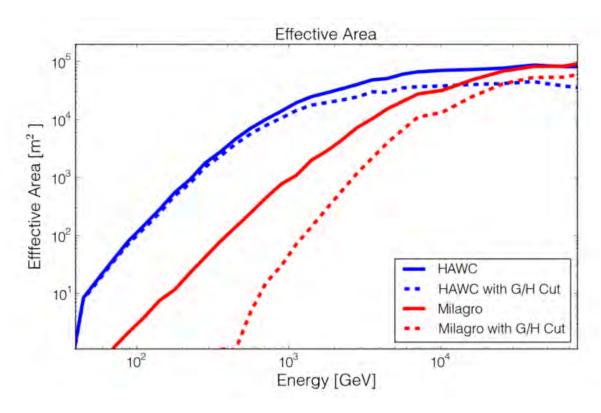
Error in determining species

Simulated showers (CORSIKA) incident on simulated HAWC (GEANT4)

Optical isolation

High quantum efficiency PMTs

# Improved Effective Area



Effective area of detector of shower

Simulated showers (CORSIKA) incident on simulated HAWC (GEANT4)

Trigger on lower energy showers

Milagro had outriggers